## A quantum picture of de Sitter spacetime

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#### Work with Gia Dvali and César Gomez

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- Idea: The world is fundamentally quantum
- $\Rightarrow$  Classical solution = collective effect of appropriate quanta (corpuscules)<sup>1</sup>

<sup>1</sup> G. Dvali and C. Gomez, *Quantum Compositeness of Gravity: Black Holes, AdS and Inflation*, arXiv:1312.4795.

- Idea: The world is fundamentally quantum
- $\Rightarrow$  Classical solution = collective effect of appropriate quanta (corpuscules)<sup>1</sup>
  - Tehseen's talk: Solitons as corpuscular bound states<sup>2</sup>
- $\Rightarrow$  Topological properties determined by number of corpuscules

- <sup>1</sup> G. Dvali and C. Gomez, *Quantum Compositeness of Gravity: Black Holes, AdS and Inflation,* arXiv:1312.4795.
- <sup>2</sup> G. Dvali, C. Gomez, L. Gründing and T. Rug, *Towards a Quantum Theory of Solitons*, arXiv:1508.03074.



#### 2 Application to Particle production



The quantum state of de Sitter  $_{\odot OO}$ 

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# De Sitter metric

• Cosmological constant  $\Lambda (\propto H^2)$ 

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## De Sitter metric

- Cosmological constant  $\Lambda \ (\propto H^2)$
- Metric for small times:

$$\mathrm{d}s^2 = (1 + \Lambda t^2)(\mathrm{d}t^2 - \mathrm{d}\vec{x}^2) + \dots$$

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• Canonically normalized Newtonian potential

$$\Phi = \frac{M_p}{2} \Lambda t^2$$

The quantum state of de Sitter  $\bullet \circ \circ$ 

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• Goal: Obtain  $\Phi$  as classical limit of a graviton bound state



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### Bound-state gravitons

- Two different Fock spaces:
  - $\hat{a}_{\vec{L}}^{\dagger}$  creates **free** gravitons.
  - $\hat{b}_{\vec{k}}^{\dagger}$  creates **bound-state** gravitons.

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#### Claim

Bound-state graviton 
$$(m = 0) =$$
 Free graviton  $(m = \sqrt{\Lambda})$ 

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- Conditions on the quantum state  $|N_{\Lambda}\rangle$ :
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- Conditions on the quantum state  $|N_{\Lambda}\rangle$ :
  - Spatially homogeneous  $\Rightarrow$  0 momentum
  - Maximally classical  $\Rightarrow$  Coherent state
- Only free parameter left:  $N \propto \langle N_{\Lambda} | b_{\vec{0}} b_{\vec{0}}^{\dagger} | N_{\Lambda} \rangle$

## Classical limit

- Expectation value in Hubble patch:
  - $\langle \textit{N}_{\Lambda}|\hat{\Phi}|\textit{N}_{\Lambda}\rangle$

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# Classical limit

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$$\langle N_{\Lambda} | \hat{\Phi} | N_{\Lambda} \rangle = \langle N_{\Lambda} | \int_{\vec{k}} \left( \hat{b}_{\vec{k}} e^{-i\omega_{\vec{k}} t} e^{i\vec{k}\cdot\vec{x}} + \text{h.c.} \right) | N_{\Lambda} \rangle$$

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• Representation of  $\Phi$  independent of source

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#### Decay constant



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$$E_1 + E_2 = \sqrt{\Lambda}$$

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• Quantum correction because of back-reaction  $(N' \neq N)$ 

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## Final state of the metric

• Metric changes because of back-reaction (Inaccessible in semi-classical limit  $N \to \infty$ )

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## Final state of the metric

- Metric changes because of back-reaction (Inaccessible in semi-classical limit  $N \to \infty$ )
- Initial de Sitter metric only valid as long as  $N N' \ll N$  $\Rightarrow$  Quantum break time<sup>3</sup>:

$$\Delta t pprox N\Gamma^{-1} = rac{M_p^2}{\Lambda^{1.5}}$$

<sup>3</sup> G. Dvali and C. Gomez, *Quantum Exclusion of Positive Cosmological Constant?*, arXiv:1412.8077.

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⇒ Final state without classical metric description?

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# Outlook

#### Summary

- De Sitter metric as classical limit of graviton state
- Particle production because of graviton decay
- 1/N-correction of the rate caused by back-reaction
- Quantum evolution of the metric

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#### Future research

- Minkowski as graviton state
- Model final de Sitter state
- Inflationary scenarios
- Other metrics such as AdS