

Bundesministerium für Bildung und Forschung





Searches for Dark Matter at the LHC



IMPRS Young Scientist Workshop at Ringberg Castle Makoto Teshima | Max-Planck-Institut für Physik, München

Dark Matter (DM)



- Much evidence for the presence of DM
 - ► Galaxy rotation curves, gravitational lensing etc. (But only gravitational interaction)
- ▶ 23% of the universe is DM
- Explaining DM is one of the main motivations for physics beyond the standard model
- One candidate for DM is WIMP

Weakly Interacting Massive Particle (WIMP)

- Most favourite DM candidate
- Stable, non-baryonic, electrically neutral
- Mostly WIMP taken as Dirac Fermion
- Expected mass range: 1 GeV-10 TeV



Ways to search for Dark Matter



Mono-X searches

- ► At colliders: WIMP pair production
- However, without additional particles, WIMPs do not create any signal
- WIMP production in association with visible SM particle(s) X results in signatures with E_T^{miss}
- ▶ Search for Mono-X with X = jet, γ , W, Z, H
- Mono-jet is the most powerful search for DM at LHC



Simulation of WIMP production

- EFT
 - Effective Field Theory (EFT): 4-point interaction
 - Used in LHC Run 1
 - But EFT invalid at large momentum transfer

- Simplified model -

- ▶ Used in LHC Run 2
- Expanding the contact interaction of DM with SM particles, including a new massiv mediator
- Spin 1 and 0 mediator
- ► 4 free parameters (m_{\chi}, m_{med}, g_{SM}, g_{DM})



Simulation of WIMP production

Mediator Type	g_{DM}	\boldsymbol{g}_q
Vector	1	0.25
Axial-Vector	1	0.25
Scalar	1	1
Pseudo Scalar	1	1

 ATLAS-CMS Dark Matter Forum (arXiv:1507.00966) define benchmark models Simplified model -

- ▶ Used in LHC Run 2
- Expanding the contact interaction of DM with the SM, include a mediator
- Spin 1 and 0 mediator
- ► 4 free parameters (m_{\chi}, m_{med}, g_{SM}, g_{DM})



Mono-jet search at ATLAS





Mono-jet search at ATLAS

- Signal region:
 - At least 1 jet with $p_T > 250 \text{ GeV}$
 - $E_{\rm T}^{\rm miss} > 250 {\rm ~GeV}$
 - Lepton veto
 - And up to 3 jets with p_T > 30 GeV arrowed
- Dominant backgrounds: Z+jets, W+jets
- Control regions are used to study backgrounds
 - ► Determine normalization factor and test in signal free validation regions



Mono-jet exclusion limits

 $\begin{array}{l} \mbox{Perturbativity limits: } \frac{m_{DM}^2 g_{DM}^2}{(\pi M_{med}^2)} < \frac{1}{2} \\ \mbox{For } g_{DM} {=} 1 {: } m_{DM} < \sqrt{\frac{\pi}{2}} M_{med} \end{array}$



Exclusion limits for all Mono/Di-jet searches



Exclusion limits for all Mono/Di-jet searches



Comparison to direct detection experiments

Axial-vector mediator: $\sigma_{\rm SD} \propto \frac{g^2_{\rm DM} \mu^2_{\rm n\chi}}{M^4_{\rm med}}$, $\mu_{\rm n_{\chi}} = \frac{m_{\rm n} m_{\rm DM}}{(m_{\rm n} + m_{\rm DM})}$



Comparison to direct detection experiments



Summary

- Most favoured DM candidate: WIMP
- Searches for DM at colliders
- Example: ATLAS Mono-jet analysis
- LHC searches for DM are complementary to (in)direct DM searches
- Next talk: Dark Matter and the Higgs Boson at ATLAS Speaker: Rainer Röhrig

Backup

Weakly Interacting Massive Particle (WIMP)

- In early universe WIMPs are present in large number
- ► They are in the remain equilibrium $\chi\chi \longleftrightarrow ff$
- Universe expanded and cooled
 - Density are reduced through pair annihilation
- Annihilation rate drops below the expansion rate of the Universe and the DM interactions become insufficient to maintain thermal equilibrium (freeze out)



DM models

invisible DM search (Mono-X)



DM Mediator search (DI-X)



▶ spin 1 and 0 mediator, 4 free parameters $(m_{\chi}, m_{med}, g_{SM}, g_{DM})$

Fitting Strategy Mono-Jet

- Background only fit
- Fit applied using 10 MET bins for model dependent interpretation
- Fit applied using inclusive MET cuts for model-indipendent interpretation

Exclusive bins	Inclusive bins	
(used for interpretation)	(used for model-independent results)	
$250 \text{GeV} < p_{\text{T}}^{\text{recoil}} < 300 \text{GeV}$	$p_{\rm T}^{\rm recoil} > 250 {\rm GeV}$	
$300 \text{GeV} < p_{\text{T}}^{\text{recoil}} < 350 \text{GeV}$	$p_{\rm T}^{\rm recoil} > 300 {\rm GeV}$	
$350 \mathrm{GeV} < p_{\mathrm{T}}^{\mathrm{recoil}} < 400 \mathrm{GeV}$	$p_{\rm T}^{\rm recoil} > 350 {\rm GeV}$	
$400 \text{GeV} < p_{\text{T}}^{\text{recoil}} < 500 \text{GeV}$	$p_{\rm T}^{\rm recoil} > 400 {\rm GeV}$	
$500 \text{GeV} < p_{\text{T}}^{\text{recoil}} < 600 \text{GeV}$	$p_{\rm T}^{\rm recoil} > 500 {\rm GeV}$	
$600 \mathrm{GeV} < p_{\mathrm{T}}^{\mathrm{recoil}} < 700 \mathrm{GeV}$	$p_{\rm T}^{\rm recoil} > 600 {\rm GeV}$	
$700 \mathrm{GeV} < p_{\mathrm{T}}^{\mathrm{recoil}} < 800 \mathrm{GeV}$	$p_{\rm T}^{\rm recoil} > 700 {\rm GeV}$	
$800 \text{GeV} < p_{\text{T}}^{\text{recoil}} < 900 \text{GeV}$	$p_{\rm T}^{\rm recoil} > 800 {\rm GeV}$	
$900 \mathrm{GeV} < p_{\mathrm{T}}^{\mathrm{recoil}} < 1000 \mathrm{GeV}$	$p_{\rm T}^{\rm recoil} > 900 {\rm GeV}$	
$p_{\rm T}^{\rm recoil} > 1000{\rm GeV}$	$p_{\rm T}^{\rm recoil} > 1000{\rm GeV}$	