MPP Project Review 2024 - COSINUS Group

Updates from COSINUS

December 9, 2024



M.Stahlberg Postdoctoral Researcher



















Founded in 2018; Construction of experiment approved in 2020

Construction of experiment approved in 2020

~40 Scientists from 8 institutes in 5 countries





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G S S HEPHY INSTITUTE OF PHYSICS OF HELSINKI INSTITUTE OF PHYSICS



Inauguration on April 18, 2024!





Why COSINUS? - The "puzzle" of DAMA/LIBRA



- Direct search for dark matter (DM): halo of particles causes modulating rate of interactions in earth-bound detector
- Observed by DAMA/LIBRA since ~25 years (period + **phase**!)
- Incompatible with other direct detection experiments in "standard" scenario
- DAMA/LIBRA measures scintillation light from NaI crystals **
- Recent improvement of threshold to ~1keV, experiment switched off









Amaré, Julio, et al. "Dark Matter Searches Using Nal (TI) at the Canfranc Underground Laboratory: Past, Present and Future." Universe 8.2 (2022): 75.



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Direct Search for Dark Matter with Nal

...does the choice of Nal as target material have an influence?



Adhikari, G., et al. "Three-year annual modulation search with COSINE-100." Physical Review D 106.5 (2022): 052005.

Caveats:

- Influence of threshold/efficiency
- Limited by exposure (annual modulation)
- Quenching Factor; amount of light observed for NR events
- * Location?



COSINUS Detectors

- Read out phonon signal in addition to scintillation light ** -> NaI as cryogenic bolometer/calorimeter
- Use existing sensor technology & expertise (TES) from CRESST *
- Modified for "difficult" materials as NaI -> remoTES *



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Angloher, G., et al. "First measurements of remoTES cryogenic calorimeters: easy-to-fabricate particle detectors for a wide choice of target materials." Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment 1045 (2023): 167532.





COSINUS Detectors

* -> Identification of individual nuclear recoil events (instead of modulation search)







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Nuclear recoils from DM interactions produce less scintillation light than electron interactions ("quenching")



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Going Underground

* 3600m w.e. overburden to shield cosmic radiation

* Water Cherenkov muon veto

* Dry dilution refrigerator (precooling via pulse tube)



Experimental Setup at LNGS















Experimental Setup at LNGS









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Tank closed!





Cryogenic Facility at LNGS







- * Dry dilution refrigerator by Cryoconcept
- * Reached base temperature of 9.4mK in March 2024
- * Final commissioning step ongoing
- * Currently at 9.6mK !



Cherenkov Muon Veto

- * Setup modelled & tested at MPP
- * PMTs tested & alive after closing of tank
- Installed water purification system
- * Will start filling end of this year / early next year

Rate of cosmogenic neutrons: No veto: (3.5 ± 0.7) cts per (kg yr) With veto: $<(0.30 \pm 0.02)$ cts per (kg yr)













COSINUS Cryogenic Facility at MPP





"GG" cryostat in the old lab

- * Cryostat delivered in July 2023
- * Installed and operated before the move!
- * Base temperature below 7mK













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Move to Garching - February 2024





- COSINUS Cryostat at MPP fully commissioned! *
- Move of the lab completed quickly *
- Base temperature below 6mK *
- Further optimizing noise conditions * (vibration decoupling)



-> Successful measurements with new prototypes!

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Operation in Garching









Prototype Results

- Previous measurements with Si and TeO2 absorbers to optimize detector design
- * Series of measurements with Si beaker light detectors
- * Published results from two NaI prototypes measured in 2022
- * Resolution improved from 2keV to 400eV with 3.7g crystal
- Recent prototype: resolution ~100eV with 3.7g





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<u>Angloher, G., et al. "Particle discrimination in a NaI crystal using the COSINUS remote TES design."</u> *Physical Review D* 109.8 (2024): 082003.

<u>Angloher, G., et al. "Deep-underground dark matter search with a COSINUS detector prototype."</u> <u>*Physical Review D* 110.4 (2024): 043010.</u>







Nal + Evaporated Au

- * Reproducible process
- * Fabrication of thin films possible
- Recent prototype: resolution ~100eV with 3.7g *



In cooperation with R. Götz, Chair of A. Bandarenka

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Tectra evaporation machine @TUM











-> Will be featured in Nature collection about direct DM search

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Physics Reach of COSINUS

- Advantage through low energy threshold and event-by-event background discrimination
- * First phase: 10 modules with 30g of NaI each



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- * Search for Dark Matter-Electron scattering with cryogenic light detectors Zema, V., et al. "Dark matter-electron scattering search using cryogenic light detectors." Physical Review D 110.12 (2024): 123012.
- * Support of the COSINUS & CRESST DAQ infrastructure by the "Beratende Ausschuss für EDV-Anlagen in der Max-Planck-Gesellschaft" (BAR)
- * Planned collaboration with the HLL (HalbLeiterLabor) for TES production
- * Measurement of COSINUS NaI crystals in a neutron beam facility at Triangle Universities National Laboratory (TUNL)
- CRYOCLUSTER group of experiments (CRESST, COSINUS, NUCLEUS) **
- Collaboration with groups at TUM for measurements, DAQ testing & prototype preparation *
- * Increasing importance of cryogenic facilities & detectors with respect to quantum computing applications

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R&D Projects & Synergies

















COSINUS Group at MPI Physics

Lead: K. Schäffner (supported by MPRG program, recently appointed by Lise-Meitner-Exzellenzprogramm) Postdocs: Vanessa Zema*, Maximilian Hughes, Martin Stahlberg PhD students: Moritz Kellermann, Mukund Bharadwaj, Clemens Dittmar, Kumrie Shera Master students: Maximilian Gapp, Kilian Heim, Sarah Braun, Lutz Ziegele Working students: Elisa Gaido*, Pablo Figueroa Falla*, Henrik Ansorge Part time: K. Ackermann, G. Angloher, T. Frank

*funded by Klaus Tschira foundation

+ A lot of support from everyone at the institute (Administration, workshops, PR team, IT, Haustechnik ...)

-> THANK YOU! <-

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Support from CRESST





Senior scientis

(Senior scientist





COSINUS

Summary & Outlook



- * COSINUS experimental facility at LNGS completed Final commissioning steps ongoing
- Promising results from detector prototypes
- Preparation of final detectors ongoing
- Datataking early-mid 2025 (!) *
- * Many R&D projects and interconnections in COSINUS group



Thank you for your attention!







Backup: Physics Reach, MI comparison



Kahlhoefer, Felix, et al. "Model-independent comparison of annual modulation and total rate with direct detection experiments." Journal of Cosmology and Astroparticle Physics 2018.05 (2018): 074.

\rightarrow total rate R \geq S modulation amplitude

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Run 1: 2025-2026

- 8 detector modules
- 30 g Nal

Increasing exposure & lowering threshold allows **model-independent (i.e. halo-independent)** comparison to DAMA/LIBRA

-> exclusion of arbitrary recoil spectrum



Backup: Quenching Factor

Modified from: Joo, H. W., et al. "Quenching factor measurement for NaI (TI) scintillation crystal." Astroparticle Physics 108 (2019): 50-56.



- Interpretation of energy scale necessary for scintillation light-only experiments such as DAMA/LIBRA
- Quenching factor : conversion between electron equivalent and nuclear recoil energy scale
- * Energy-dependent!
- * Discrepancy between literature values



Backup: COSINUS Glovebox



- * Dedicated, custom-built glovebox for detector handling & assembly
- * Includes space for bonding machine * NaI is hygroscopic -> N2 atmosphere
- * Dockable transport box for finished detectors

