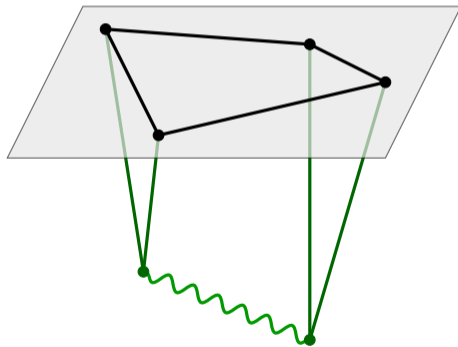


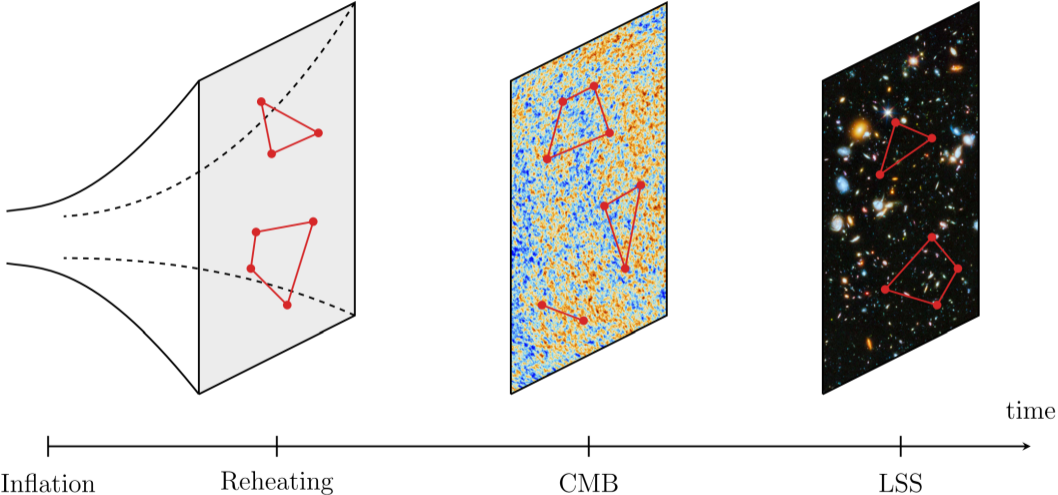
Cosmological Correlators at Loop Order

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25.11.2024



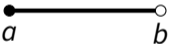
Cosmological Correlators

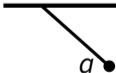


equal-time in-in correlators in the Schwinger-Keldysh path integral formalism:

$$\langle \phi^{a_1}(\eta, x_1) \dots \phi^{a_N}(\eta, x_N) \rangle = \int_{\phi_+(\eta_f, x) = \phi_-(\eta_f, x)} \mathcal{D}[\phi_+, \phi_-] e^{iS[\phi_+] - iS[\phi_-]} \phi_+^{a_1}(\eta, x_1) \dots \phi_+^{a_N}(\eta, x_N)$$

Feynman rules (example: $\mathcal{L}_{\text{int}} \supset \frac{\lambda}{4!} a(\eta)^4 \phi^4$):


 $= \mathcal{D}_{ab}^\nu(k; \eta_1, \eta_2)$


 $= \mathcal{G}_a^\nu(k; \eta)$



$= ib\lambda \int_{-\infty}^0 d\eta a(\eta)^4 \dots$

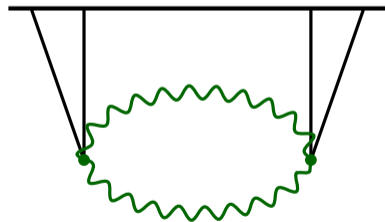
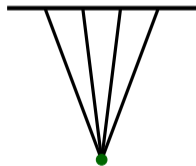
Why is this so much harder than 'usual' QFTs in Minkowski space?

- time-dependent de Sitter background metric: $ds^2 = \frac{1}{(-\eta)^2}(-d\eta^2 + dx^2)$ where $\eta \in (-\infty, 0)$ is the conformal time
- mode functions:

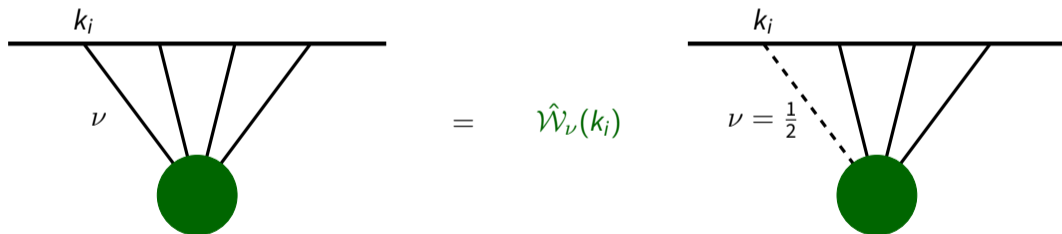
$$u_k(\eta) = -i \frac{\sqrt{\pi}}{2} e^{i\pi(\frac{\nu}{2} + \frac{1}{4})} (-\eta)^{d/2} H_\nu^{(1)}(-k\eta)$$

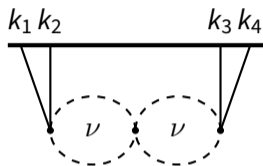
- $\nu = \sqrt{d^2/4 - m^2}$ related to mass
- propagators up to time ordering: $\mathcal{D}_{ab}(k; \eta_1, \eta_2) \sim u_k(\eta_1) u_k^*(\eta_2)$

- **conformally coupled scalars** ($\nu = 1/2$): a lot of progress \rightarrow results > 2 loop order, cosmological polytopes, differential equations, ...
- **massless, minimally-coupled scalars** ($\nu = 3/2$): also several loop results \rightarrow study of IR divergences, ...
- arbitrary massive exchanges ($\nu = \sqrt{d^2/4 - m^2}$): some tree-level results + 1 loop bubble



- External legs are not the problem!
- relevant for cosmology: massless scalars
- also: massive bulk-to-boundary propagators related to conformally coupled case by **weight-shifting operators**





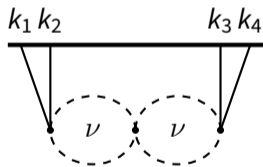
$$\sim \int_{-\infty}^0 \left[\prod_{j=1}^3 d\eta_j (-\eta_j)^{p_j} \right] e^{iak_{12}\eta_1 + ick_{34}\eta_3} \mathcal{Q}_{ab}^\nu(k_s; \eta_1, \eta_2) \mathcal{Q}_{bc}^\nu(k_s; \eta_2, \eta_3)$$

Bubble loop integral:

$$\mathcal{Q}_{ab}^\nu(k; \eta_1, \eta_2) \equiv \int \frac{d^d q}{(2\pi)^d} \mathcal{D}_{ab}^\nu(q; \eta_1, \eta_2) \mathcal{D}_{ab}^\nu(|k - q|; \eta_2, \eta_3)$$

Loop integral can be simplified using

$$\mathcal{Q}_{ab}^{\nu}(k; \eta_1, \eta_2) = \int_{-i\infty}^{+i\infty} d\nu' \frac{\nu'}{\pi i} \rho_{\nu}^{dS}(\nu') \mathcal{D}_{ab}^{\nu'}(k; \eta_1, \eta_2)$$



$$= \int_{-i\infty}^{+i\infty} d\nu' d\nu'' \frac{\nu'}{\pi i} \frac{\nu''}{\pi i} \rho_{\nu}^{dS}(\nu') \rho_{\nu}^{dS}(\nu'')$$

