

The gravitino mass and extra dimensions

Leonardo Bersigotti

IMPRS recruiting workshop, MPP

25 November 2024



We need to find a theory of Quantum Gravity (QG).

To unify **Quantum Field Theory (QFT)** and **General Relativity (GR),** and to address difficult fundamental questions such as **Black Hole interior** and **"big bang" singularity**.



We need to find a theory of Quantum Gravity (QG).

• The most developed candidate is **String Theory**.



We need to find a theory of Quantum Gravity (QG).

- The most developed candidate is **String Theory**.
 - Explored through **String Phenomenology**, with two approaches:
 - **Top-down:** String model-building.
 - Bottom-up: Effective Field Theory (EFT) analysis.



We need to find a theory of Quantum Gravity (QG).

- The most developed candidate is **String Theory**.
 - Explored through **String Phenomenology**, with two approaches:
 - **Top-down:** String model-building.
 - Bottom-up: Effective Field Theory (EFT) analysis.



We need to find a theory of Quantum Gravity (QG)

- The most developed candidate is String Theory.
 - Explored through **String Phenomenology**, with two approaches:
 - **Top-down:** String model-building.
 - Bottom-up: Effective Field Theory (EFT) analysis.

The structure of QG imposes strong constraints at lower energies.

Swampland Program

Vafa 2005

- e.g. No global symmetries. Misner, Wheeler 1957
- e.g. Quantum break-time of de Sitter. Dvali, Gómes, Zell 2017
- e.g. Supersymmetry from gravitational theta vacua. Dvali, Kobakhide, Sakhelashvili 2024



We need to find a theory of Quantum Gravity (QG).

- The most developed candidate is String Theory.
 - Explored through **String Phenomenology**, with two approaches:





• Super String Theory in its simplest formulation comes with two key characteristics:

Supersymmetry (SUSY)

Extra dimensions (EDs)



• Super String Theory in its simplest formulation comes with two key characteristics:





• Super String Theory in its simplest formulation comes with two key characteristics:



- Max-Planck-Institut für Physik
- Super String Theory in its simplest formulation comes with two key characteristics:



- Max-Planck-Institut für Physik Δ₄-Δ₉≥±±
- Super String Theory in its simplest formulation comes with two key characteristics:



• The mass of the gravitino $m_{3/2}$, the supersymmetric partner of the graviton, sets the scale of SUSY breaking in a quasi-flat spacetime, as observed today.

We are interested in exploring phenomenological connections between $m_{3/2}$ and EDs.

2. Exploring connections





Interesting observation:

The experimental **lower bound** on the **gravitino mass** is of the same order of the **upper bound** on the size of **extra dimensions**.

 $m_{3/2} > 0.1 \, eV$

Derived from LHC bounds on m_{soft} .

 $l < 38.6 \, \mu m$

Derived from constraints on deviations from gravitational Newton's law.







Interesting observation:

The experimental **lower bound** on the **gravitino mass** is of the same order of the **upper bound** on the size of **extra dimensions**.



Derived from LHC bounds on m_{soft} .



Derived from constraints on deviations from gravitational Newton's law.

This relationship aligns naturally with a special constraint on the gravitino mass:

Cribiori, Lüst, Scalisi 2021 Castellano, Font, Herraez, Ibañez 2021

The Gravitino Conjecture (GC): In Planck units, the limit of small gravitino mass m_{3/2} → 0 always corresponds to the massless limit of an infinite tower of states, yielding the breakdown of the EFT.





• The gravitino mass $m_{3/2}$ is linked to the compactification scale l:



2.1 Exploring connections



Purposes:



We listed those predictions concerning the **SUSY breaking scale** and **EDs**.



1. Let's assume that the volume of the internal manifold is dominated by a large *p*-cycle, \mathcal{V}_p .

 $p \; {\rm equal} \; {\rm radii} \; {\rm means} \; p \; {\rm large} \; {\rm EDs}$

• The **Kähler potential** is modified:

• Type IIA:
$$K \simeq -\frac{6}{p} \ln \mathcal{V}_p + \dots$$

• Type IIB: $K \simeq -\frac{12}{p} \ln \mathcal{V}_p + \dots$



1. Let's assume that the volume of the internal manifold is dominated by a large *p*-cycle, \mathcal{V}_p .

• The Kähler potential is modified:

• Type IIA:
$$K \simeq -\frac{6}{p} \ln \mathcal{V}_p + \dots$$

• Type IIB: $K \simeq -\frac{12}{p} \ln \mathcal{V}_p + \dots$

 $p \; {\rm equal} \; {\rm radii} \; {\rm means} \; p \; {\rm large} \; {\rm EDs}$

Note: While not a primary goal, this result holds in the context of string compactification.



1. Let's assume that the volume of the internal manifold is dominated by a large p-cycle, \mathcal{V}_p .

• The **Kähler potential** is modified:

• Type IIA:
$$K \simeq -\frac{6}{p} \ln \mathcal{V}_p + \dots$$

• Type IIB: $K \simeq -\frac{12}{p} \ln \mathcal{V}_p + \dots$

 $p \; {\rm equal} \; {\rm radii} \; {\rm means} \; p \; {\rm large} \; {\rm EDs}$

Note: While not a primary goal, this result holds in the context of string compactification.

2. Using $l, m_{3/2} \propto \mathcal{V}_p$, we refined the relation $l^{-1} \sim (m_{3/2})^n$ to determine n = n(p).

• New **bounds** on *n*:

• Type IIA:
$$\frac{1}{3} < n \le \frac{2+p}{2p}$$

• Type IIB: $\frac{2+p}{12} < n \le \frac{2+p}{2p}$



1. Let's assume that the volume of the internal manifold is dominated by a large p-cycle, \mathcal{V}_p .

• The **Kähler potential** is modified:

• Type IIA:
$$K \simeq -\frac{6}{p} \ln \mathcal{V}_p + \dots$$

• Type IIB: $K \simeq -\frac{12}{p} \ln \mathcal{V}_p + \dots$

p equal radii means p large EDs

Note: While not a primary goal, this result holds in the context of string compactification.

2. Using $l, m_{3/2} \propto \mathcal{V}_p$, we extend the relation $l^{-1} \sim (m_{3/2})^n$ to determine n = n(p).

• New **bounds** on *n*:

• Type IIA:
$$\frac{1}{3} < n \le \frac{2+p}{2p}$$

• Type IIB: $\frac{2+p}{12} < n \le \frac{2+p}{2p}$

- Lower bounds: perturbative string regime.
- Upper bounds: $M_{SUSY} \leq \Lambda_{sp}$ (UV cutoff).

Dvali 2007



3. Drawing scenarios concerning PHENO implications:

Easily detectable scenarios within the next generation of experiments

Note: torsion-balance experiments for extra dimensional gravitational signatures like the one in (Lee, Adelberger, Cook, Fleisher, Heckel 2020).



3. Drawing scenarios concerning PHENO implications:

• Easily detectable scenarios within the next generation of experiments



• p = 1 is the only possibility at this scale. Hannestad, Raffelt 2004 • Possible range: $1 < n \le \frac{3}{2}$ • Predictions: $0.1 eV \le m_{3/2} \le 1 GeV$ $\Lambda_{sp} \sim 10^9 GeV$ $10^4 GeV \le M_{SUSY} \le 10^9 GeV$





3. Drawing scenarios concerning PHENO implications:

- Easily detectable scenarios within the next generation of experiments
 - Recall the experimental constraint $l < 38.6 \,\mu m$, therefore consider $l \sim \mu m$.

• p = 1 is the only possibility at this scale. Hannestad, Raffelt 2004 • Possible range: $1 < n \le \frac{3}{2}$ • Predictions: $0.1 eV \le m_{3/2} \le 1 GeV$ $\Lambda_{sp} \sim 10^9 GeV$ $10^4 GeV \le M_{SUSY} \le 10^9 GeV$



Scenarios with larger gravitino mass

 $m_{3/2}\sim 10^5\,GeV$

- Wide range of possible scenarios **across all values** of *p*.
- All of them require small EDs.



• Scenarios with larger gravitino mass - Tables

	IIA: $m_{3/2} = 10^5 GeV$ and $M_{SUSY} = 10^{11} GeV$				$\underline{ ext{IIB}}: m_{3/2} = 10^5 GeV \hspace{0.1 cm} ext{and} \hspace{0.1 cm} M_{SUSY} = 10^{11} GeV$			
p values	n values	$l\left(m ight)$	species scale (GeV)	p values	n values	$l\left(m ight)$	species scale (GeV)	
1	$\tfrac{1}{3} < n \leq \tfrac{3}{2}$	$2.32\cdot 10^{-30} - 9.84\cdot 10^{-15}$	$8.06 \cdot 10^{16} - 4.98 \cdot 10^{11} \sim M_{SUSY}$	1	$\tfrac{1}{4} < n \leq \tfrac{3}{2}$	$1.78\cdot 10^{-31} - 9.84\cdot 10^{-15}$	$1.90 \cdot 10^{17} - 4.98 \cdot 10^{11} \sim M_{SUSY}$	
2	$\tfrac{1}{3} < n \leq 1$	$2.32\cdot 10^{-30} - 1.98\cdot 10^{-21}$	$1.45\cdot 10^{16} - 4.98\cdot 10^{11}$	2	$rac{1}{3} < n \leq 1$	$2.32\cdot 10^{-30} - 1.98\cdot 10^{-21}$	$1.90\cdot 10^{17} - 4.98\cdot 10^{11}$	
3	$\tfrac{1}{3} < n \leq \tfrac{5}{6}$	$2.32\cdot 10^{-30} - 1.16\cdot 10^{-23}$	$5.19\cdot 10^{15} - 4.98\cdot 10^{11}$	3	$\tfrac{5}{12} < n \leq \tfrac{5}{6}$	$3.04\cdot 10^{-29} - 1.16\cdot 10^{-23}$	$1.11\cdot 10^{15} - 4.98\cdot 10^{11}$	
4	$\tfrac{1}{3} < n \leq \tfrac{3}{4}$	$2.32\cdot 10^{-30} - 8.86\cdot 10^{-25}$	$2.62\cdot 10^{15} - 4.98\cdot 10^{11}$	4	$\frac{1}{2} < n \le \frac{3}{4}$	$3.97\cdot 10^{-28} - 8.86\cdot 10^{-25}$	$8.5\cdot 10^{13} - 4.98\cdot 10^{11}$	
5	$\tfrac{1}{3} < n \leq \tfrac{7}{10}$	$2.32\cdot 10^{-30} - 1.89\cdot 10^{-25}$	$1.60\cdot 10^{15} - 4.98\cdot 10^{11}$	5	$\tfrac{7}{12} < n \leq \tfrac{7}{10}$	$5.19\cdot 10^{-27} - 1.89\cdot 10^{-25}$	$6.51\cdot 10^{12} - 4.98\cdot 10^{11}$	
6	$\tfrac{1}{3} < n \leq \tfrac{2}{3}$	$2.32\cdot 10^{-30} - 6.78\cdot 10^{-26}$	$1.11\cdot 10^{15} - 4.98\cdot 10^{11}$	6	$n=rac{2}{3}$	$6.78\cdot10^{-26}$	$4.98\cdot 10^{11}$	

Summary





- Assuming the Gravitino Conjecture:
 - We can make valuable predictions regarding SUSY by observing EDs and viceversa.
- The connection is exploited by using String Compactification results:
 - The internal manifold's volume impacts Kähler prefactors.
 - Including $l, m_{3/2} \propto \mathcal{V}_p$ in $l^{-1} \sim (m_{3/2})^n$ leads to **bounds on** n.
- We collected **Phenomenological implications** in scenarios.
- We performed calculations without considering any proportionality parameter in $l^{-1} \sim (m_{3/2})^n$.

Outlook



• We found phenomenologically interesting scenarios that are similar to the **ADD model**.

Arkani-Hamed, Dimopoulos, Dvali 1998

- In the prediction extrapolated for p = 1 and $l \sim \mu m$ we found a phenomenologically viable scenario similar to the **Dark Dimension scenario**. Montero, Vafa, Valenzuela 2022
- It would be interesting to study how our findings in SUSY-breaking scale may change cosmological predictions (e.g. **Dark Matter candidates**, conditions for **Primordial Black Holes** formation, number of light species).

Obied, Dvorkin, Gonzalo Vafa 2023 Anchordoqui, Antoniadis, Lüst 2022

 It would be interesting to understand Black Hole instabilities, such as Gregory-Laflamme, in relation to scenarios with large extra dimensions for possible additional phenomenological constraints.



Thank you for your attention!

This work and other aspects are unpublished yet, [Scalisi, LB, Masias] to appear.

Leonardo Bersigotti

IMPRS recruiting workshop, MPP

25 November 2024