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2nd Workshop on Silicon Sensors for Radiation Detection and Quantum Applications 13 May 2025

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X-ray satellite missions with PN-CCD Detectors

- XMM-Newton: ESA satellite motivated PN-CCD detector development: *launch Dec. 1999*
- eROSITA: Optimization of PN-CCD detector: *launch July 2019*
- Einstein Probe: launch January 2024
- **SVOM:** *launch June 2024*
- **eXTP:** *launch January 2030 (planned)*
- **SPICE:** PN-CCD detector development for **future X-ray missions**

XMM-Newton



EPIC-PN camera of XMM-Newton

• PN-CCD: single X-ray photon counting detector with high energy, time and spatial resolution





XMM-Newton



PN-CCD



- transfer gates: pn-junctions
- fully depleted = 300 μm
 - → back-illuminated: high QE
 - + radiation hard against soft protons
- parallel transfer & readout architecture
- 6 cm x 6 cm PN-CCD

pixel size

CAMEX & TIMEX

150 x 150 μm²

to EPEA

to EPEA

Quadrant #3

Quadrant #2

- = 12 CCDs (64 channels x 200 rows)
- readout time: 4.6 ms / 3 cm²CCD unit

XMM-Newton



EPIC-PN camera operating modes

Strüder et al., A&A (2001)

mode	field of view	time	out of time	life time	brightest point source
	(FOV)	resolution	(OOT) events	with OOT events	for XMM
	in pixel format	in ms	in $\%$	in $\%$	in counts s^{-1}
	in arc min.				in erg cm ^{-2} s ^{-1} *
\mathbf{FF}	398×384	73.3	6.2	100	6
	27.2×26.2				8.1×10^{-12}
eFF	398×384	199.2	2.3	100	for extended sources
	27.2×26.2				only
LW	198×384	47.7	0.15	94.9	9
	13.5×26.2				1.2×10^{-11}
SW	63×64	5.7	1.1	71.0	100
	4.3×4.4				1.4×10^{-10}
TI	199×64	0.03	100	100	4000
	13.6×4.4				5.9×10^{-9}
BU	20×64	0.007	depends on	3.0	60 000
	1.4×4.4		\mathbf{PSF}		8.1×10^{-8}

XMM-Newton

EPIC-PN camera operating modes





Full Frame & Extended Full Frame







Large Window

0

0

 \bullet

0

Timing Mode in Quadrant 1

XMM-Newton EPIC-PN camera spectra + imaging



Calibration spectrum with the internal radioactive source



Strüder et al., NIM A 512 (2003)

XMM-Newton

EPIC-PN

Status of PN-CCD detector aboard XMM-Newton till today:

• all 12 CCDs operational

• operating parameters unchanged

Energy resolution: Δ FWHM/(FWHM \cdot t) ~ 0.2% /y @1.5keV

Very minor degradation as expected







EPIC-PN camera summary

- EPIC-PN camera onboard XMM-Newton operates successfully since 1999 till today
- Performance still excellent + stable
- All instruments (EPIC-PN, EPIC-MOS1+2, RGS1+2) are fully functional
- XMM-Newton observation proposals over-subscribed by factor 7 (most EPIC-PN)
- Spacecraft fully healthy and fuel until 2034+





eROSITA PN-CCD Detector





- ► frame store
- Excellent energy resolution in energy band [0.2 keV; 10 keV]
- Image area: 28.8 x 28.8 mm²
- Pixel size: 75 x 75 µm² → 384 x 384 Pixel ⇒ 384 parallel signal processing channels









Image readout: 9.2 ms

- eROSITA: cycle time: 50 ms
 - minimize heat dissipation ((40 ms CAMEX off)
 - \rightarrow T = -95°C (best wrt radiation damage)
 - on-board event processing



→ option 1: external filter option 2: new: on-chip filter





MPE

eROSITA Camera (Assembly)





PN-CCD Detectors in Space



All PN-CCD Detector Modules tested (GEPARD chamber) at MPE with 55 Fe \rightarrow performance test + voltage optimization

 $FWHM(5.9keV) \le 140eV$ Noise ≈ 2.5 el. ENC, # bad pixels ~ 0

Calibration + E2E test: PUMA + Panter test facilities (MPE)



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eROSITA telescope array assembled at MPE



Size: 1.9 m Ø x 3.5 m Mass: 808 kg Power: 522 W (max.) Data rate: 400 MB/day (average) 600 MB/day (max.)









Launch on 13 July 2019 (12:31 UTC)





in space: eROSITA fully functional - similar performance as measured on ground





eROSITA (extended ROentgen Survey with an Imaging Telescope Array)

eROSITA wide-field (1°) X-ray telescope developed under MPE's leadership:

- 7 Wolter-1 mirror modules (54 nested mirror shells) +
- 7 cameras with PN-CCD detectors incl. electronics + filter wheel

eROSITA onboard SRG → halo orbit around L2 Since Dec. 2019: SRG/eROSITA performed **all-sky survey**:

whole celestial sphere is mapped once every 6 months Eight all-sky charts were planned until Dec. 2023.

26 February 2022: eROSITA in **safe mode** for political reasons (cooperation with Russia suspended)





eROSITA data release by German eROSTIA consortium (31 Jan. 2024): eRASS1 catalogue: 170 million X-ray photons

• observations 12 Dec. 2019 - 11 June 2020

 >900 000 sources, including ~710.000 AGNs 180 000 X-ray emitting stars in our Milky Way, 12 000 clusters of galaxies, plus binary stars, supernova remnants, pulsars



Photons colour-coded according to energy (red 0.3-0.6 keV, green 0.6-1 keV, blue 1-2.3 keV).



Einstein Probe

cnes



 PN-CCD detectors of eROSITA type in operation on Chinese-European
 X-ray mission Einstein Probe "A New Horizon in Detecting Cosmic

> X-ray Transients" launched on January 9, 2024



Einstein Probe

Exploring the dynamic X-ray Universe

MP

WXT: lobster-eye-based wide-field telescope:
12 MPO (3800 square degrees) + 48 CMOS detectors
FXT: 2 telescopes (eROSITA optics (FoV=1°) + PN-CCD detector)

eesa

Dedicated PN-CCD wafers for EP project by HLL (R. Richter et al.)





MPE



cnes





eesa

FXT "First Light" observation: Puppis A SNR (Credit: Weimin Yuan)

Status end of 2024:

- >70 transient events detected
- thousands of candidates
- > 500 stellar flares
- > 100 alerts sent to international community, guiding follow-up observations from ground based and space instruments





SVOM MXT



SVOM: Chinese-French mission dedicated to **gamma-ray bursts**

French MXT (Microchannel X-ray Telescope) instrument **prototype of eROSITA PN-CCD:** 256x256 pixels (75x75µm²)

• MPE provided PN-CCDs, CAMEX readout ASICs, know-how for assembly and operation of detector.

Launch on **22 June 2024** from the Xichang MXT successfully commissioned + in operation SVOM has detected **113** gamma-ray bursts until May 1st 2025





eXTP satellite mission

enhanced X-ray Timing and Polarimetry

- eXTP SFA planned to be equipped with six 19-cell SDD detectors
- Request in Dec. 2024 for PN-CCD detector of EP-FXT type: imaging + spectroscopy
- eXTP satellite launch: January 2030





SPICE (Small Pixel CCD Experiment)

Paul Nandra motivated CCD technology development for

Future X-ray missions with high-angular resolution X-ray mirrors (~arcsec)

Detector requirements :

- smaller pixels
- larger pixel arrays
- high time resolution
- low heat dissipation \rightarrow sensor temperature (~ -80°C) on S/C

SPICE focal plane with **4 Mpixel**

Architecture: 4 quadrants with 1024x1024 pixels for IM

→ 4096 readout channels (heat dissipation!)

pixel size: 36 x $36\mu m^2 \rightarrow signal charge$ spread over up to 3x3 pixelscentroiding $\rightarrow subpixel resolution$

Time resolution ≥ 30ms

higher time resolution by window or timing mode

Readout ASIC: CAMEX

SPICE detector development in progress:

- HLL: CCD sensor layout & fabrication
- MPE: Requirements (science)

Detector board incl. mech.-thermal design suitable for flight; Thermal vacuum chamber Lab electronics (→ flight electronics)

Data analysis 2nd Workshop on Silicon Sensors for Radiation Detection and Quantum Applications, N. Meidinger, 13 May 2025







SPICE PN-CCD Detector





FoV(4MPixel) ≈ 7.6 x 7.6 cm² → Very innovative design developed for SPICE



Thank you

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