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# Hamiltonian Cobordism: Einstein's Equations meet the Cobordism Conjecture

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IMPRS Recruitment Workshop, MPP

July 17, 2025

# Motivation

# Do we truly need a theory of Quantum Gravity (QG)?

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A consistent theory of QG promises not only conceptual completeness but also the potential to resolve these deep empirical mysteries.

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- \* The network of string **dualities** (S-duality, T-duality, mirror symmetry) reveals deep non-perturbative equivalences between apparently different theories.
- \* **Rich mathematical structure**: the study of string theory has driven profound advances in mathematics.

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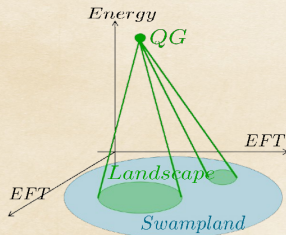
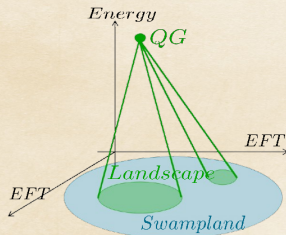


Figure: Swampland and Landscape

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- \* Top-Down & Bottom-Up Approaches
- \* Predictive Power
- \* Guiding Model Building

Figure: Swampland and Landscape

[Vafa 2005; Palti 2019; van Beest, Calderon-Infante, Mirfendereski, Valenzuela 2022]

Background

# No Global Symmetries Conjecture

*There are no global symmetries in QG, i.e. any symmetry is either broken or gauged*

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Figure: Gedankenexperiment - Black Hole entropy

# Topology and QG

In QG, spacetime can have **non-trivial topology**!



Figure: Different topologies

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... BUT non-trivial topologies lead to **(topological) global symmetry**, which are detected by **bordism groups**.

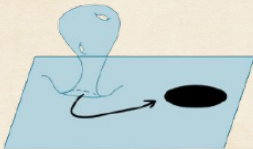


Figure: Gedankenexperiment - Non-trivial topology

# Cobordism Conjecture

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*All cobordism classes must vanish*  $\Omega_d^{QG} = 0$ .

[McNamara, Vafa 2019]

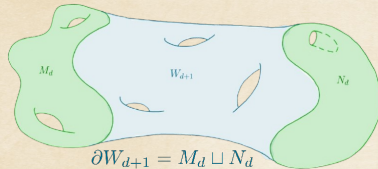


Figure: Cobordism

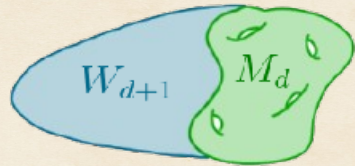


Figure: Cobordism to nothing

Predictive power: **existence of additional defects!**

# Hamiltonian Cobordism

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Topologically, the answer is yes, but once gravity is in the picture, one needs to define a metric that satisfies Einstein's equations.

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Dynamical refinement of the Cobordism Conjecture.

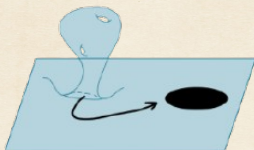


Figure: Non-trivial topology

# Hamiltonian Cobordism: ADM decomposition

Hamiltonian description of GR. → Spatial hypersurfaces ("slices") evolve in time.

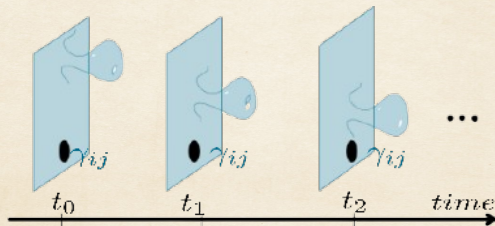


Figure: Foliation of spacetime

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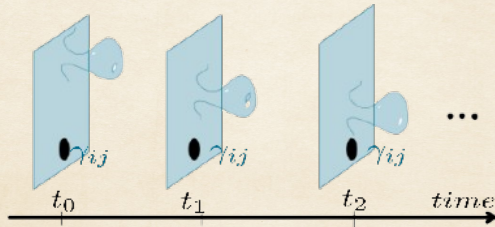


Figure: Foliation of spacetime

$$g_{00} = -\alpha^2 + \gamma^{ij}\beta_i\beta_j, \quad g_{0i} = g_{i0} = \beta_i, \quad g_{ij} = \gamma_{ij} \quad (1)$$

$$ds^2 = g_{\mu\nu}dx^\mu dx^\nu = (-\alpha^2 + \gamma_{ij}\beta^i\beta^j)dt^2 + 2\beta_i dt dx^i + \gamma_{ij}dx^i dx^j, \quad (2)$$

[Arnowitt, Deser, Misner 1959; Bertschinger 2005; Lao 2021]

# Hamiltonian Cobordism: Initial & Energy Conditions

**Hamiltonian and momentum constraints** for GR as initial conditions:

$$-\mathcal{H} = \alpha^2 T^{00} \quad \text{with} \quad \mathcal{H} = K_{ij}K^{ij} - K^2 - {}^{(d)}R, \quad (3)$$

$$\alpha\mathcal{H}^i + \beta^i\mathcal{H} = \alpha^2 T^{0i} \quad \text{with} \quad \mathcal{H}^i = 2\nabla_j(K^{ij} - K\gamma^{ij}). \quad (4)$$

For asymptotically flat manifold  $|x| \rightarrow \infty$ :  $\alpha(x_i, t) \rightarrow 1$ ,  $\beta_i(x_i, t) \rightarrow 0$ , and  $\gamma_{ij}(x_i, t) \rightarrow \delta_{ij}$ .

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**Weak Energy Condition (WEC):**

*Given any future-directed timelike vector field  $X^\mu$ ,  $T_{\mu\nu}X^\mu X^\nu \geq 0$ .*

$$\implies T^{00} \geq 0 \quad (5)$$

# Hamiltonian Cobordism: Assumptions & Results

Assume there exists a metric for the gluing such that:

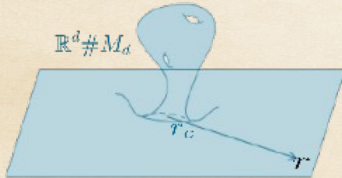


Figure: Initial Configuration

- \*  $\gamma_{ij}(x_i, t) = \begin{cases} \gamma_{ij}^{M_d} & \text{for } r < r_c \\ \delta_{ij} + \mathcal{O}(r^{-(d-2)}) & \text{for } r \geq r_c \end{cases}$
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- \* Smoothness
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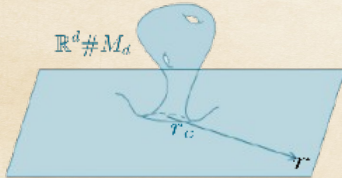


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$M_d$  must admit a metric of positive scalar curvature (PSC).

# Hamiltonian Cobordism: Importance of PSC

**STEP 1:** Conformal transformation of  $\gamma$  to obtain another metric  $\hat{\gamma}$  which is also asymptotically flat and whose scalar curvature is everywhere non-negative.

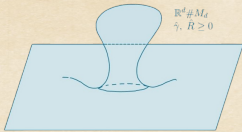


Figure: Step 1

$$\hat{R} = e^{-2\varphi} \left[ R - \frac{4(d-1)}{(d-2)} e^{-(d-2)\varphi/2} \Delta \left( e^{(d-2)\varphi/2} \right) \right] \quad (6)$$

$$e^{2\varphi} \hat{R} + 6e^{-\varphi} \Delta(e^\varphi) + K^2 = \alpha^2 T^{00} + K_{ij} K^{ij} \quad (7)$$

$$\Delta f + \frac{K^2}{6} f \leq 0. \quad (8)$$

# Hamiltonian Cobordism: Importance of PSC

**STEP 2:** Conformal transformation on  $\hat{\gamma}$  to a manifold  $S^d \# M_d$  is *compact* with metric  $\hat{\gamma}_c$  and  $\hat{R}_c \geq 0$ .

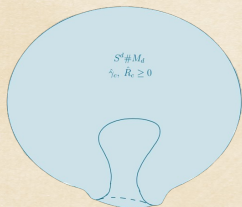


Figure: Step 2

$$u(r) = \begin{cases} 1 & \text{for } r < r_0, \\ u_s(r) & \text{for } r \rightarrow \infty, \end{cases} \quad (9)$$

with smooth boundary conditions at  $r = r_0$ :

$$u(r_0) = 1, \quad u'(r_0) = u''(r_0) = 0 \quad \text{and} \quad \hat{R}_c|_{\substack{r \rightarrow r_0 \\ r > r_0}} = 0, \quad (10)$$

# Hamiltonian Cobordism: Importance of PSC

**STEP 3:** Conformal transformation on  $\hat{\gamma}_p$  to a manifold with positive Yamabe invariant, which implies that with another conformal rescaling one obtains a PSC metric.

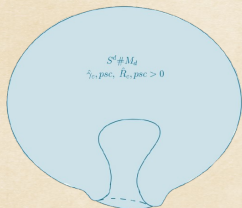


Figure: Step 3

$$\text{Yamabe functional: } \mathcal{E}(\hat{\gamma}_c, \nu) = \frac{\int_{M_d \# R^d} \hat{R}_{c,psc} dV_{\hat{\gamma}_{c,psc}}}{\left( \int_{M_d \# R^d} dV_{\hat{\gamma}_{c,psc}} \right)^{\frac{d-2}{d}}}, \quad (11)$$

$$\text{Yamabe invariant: } Y(M_d \# R^d, \hat{\gamma}_c) = \inf_{\nu \neq 0} \mathcal{E}(\hat{\gamma}_c, \nu). \quad (12)$$

# Hamiltonian Cobordism: Results

**If a compact manifold can be smoothly glued to flat spacetime in a way consistent with general relativity, then it must admit a metric of PSC.**

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**Cobordism Conjecture requires refinement  $\Omega \rightarrow \Omega_{PSC}$ .**

Possible implications:

- \* New bordism classes might appear, some known ones might be trivialized;
- \* Defects/singularities beyond classical GR (e.g., Casimir potentials, orbifolds, ... )

# Outlook

# Hamiltonian Cobordism: Outlook

- \* How do these ideas interact with other known Swampland conjectures?
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**Important first and innovative step in connecting the Cobordism Conjecture to dynamics.**

**Thank you!**

Backup slides

# AdS/CFT Proof

Assumption: splittability (you can "split" the boundary in subsets)

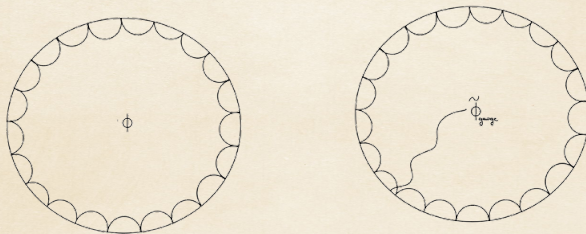


Figure: AdS/CFT proof

# No Global Symmetries Conjecture

The No Global Symmetries conjecture is believed to hold when the notion of symmetry is extended, i.e., not only for ordinary 0-form symmetries but also for generalized symmetries.

## **Definition:** $p$ -form global symmetry

A  $p$ -form global symmetry acts on  $p$ -dimensional 'defects' or operators. Its conserved charge lives on a  $(d - p - 1)$ -dimensional sphere surrounding the defect in a  $d$ -dimensional spacetime.

# Cobordism and String Theory

Cobordism classes form abelian groups depending on the dimension of the manifold and the conserved structure.

## **Definition:** Cobordism Group of Quantum Gravity $\Omega^{QG}$

- \*  $\Omega^{QG,D}$ : This group classifies all  $D$ -dimensional theories of quantum gravity, where the  $D$ -dimensions can be non-compact.
- \*  $\Omega_k^{QG}$ : This group classifies all theories compactified on a  $k$ -dimensional manifold.

# PSC Bordism?

*Should we consider PSC bordism?*

A priori, the above result just means that we should drop all manifolds that do not admit a PSC metric; it doesn't necessarily mean that we should consider the bordism of PSC manifolds.

To do so, we should be able to argue that the interpolating  $(d + 1)$ -dimensional manifold should also have PSC.

# Yamabe Invariant and PSC

$\hat{\gamma}_{c,psc} = \nu^{\frac{4}{d-2}} \hat{\gamma}_c$  with a function  $\nu$  that satisfies  $\nu(r) > 0 \forall r$ .

**Yamabe equation:**  $L_{\hat{\gamma}_c} \nu = \hat{R}_{c,psc} \nu^{\frac{d+2}{d-2}}$  where  $L_{\hat{\gamma}_c} = -\frac{4(d-1)}{d-2} \Delta_{\hat{\gamma}_c} + \hat{R}_c$  (13)

**Yamabe functional:** 
$$\mathcal{E}(\hat{\gamma}_c, \nu) = \frac{\int_M \hat{R}_{c,psc} dV_{\hat{\gamma}_{c,psc}}}{\left(\int_M dV_{\hat{\gamma}_{c,psc}}\right)^{\frac{d-2}{d}}} = \frac{\int_M \left(\frac{4(d-1)}{d-2} |\nabla_{\hat{\gamma}_c} \nu|^2 + \hat{R}_c \nu^2\right) dV_{\hat{\gamma}_c}}{\left(\int_M \nu^{\frac{2d}{d-2}} dV_{\hat{\gamma}_c}\right)^{\frac{d-2}{d}}},$$
 (14)

where we have used the fact that  $dV_{\hat{\gamma}_{c,psc}} = \nu^{\frac{2d}{d-2}} dV_{\hat{\gamma}_c}$  and since  $M$  has no boundary, we have

$$\int_M \nu \Delta_{\hat{\gamma}_c} \nu dV_{\hat{\gamma}_c} = - \int_M |\nabla_{\hat{\gamma}_c} \nu|^2 dV_{\hat{\gamma}_c}. \quad (15)$$

$$\text{Yamabe invariant: } Y(M, \hat{\gamma}_c) = \inf_{\nu \neq 0} \mathcal{E}(\hat{\gamma}_c, \nu), \quad (16)$$

where the infimum is taken over the smooth real-valued positive functions  $\nu$ .

By construction, the metric  $\hat{\gamma}_c$  is such that  $\hat{R}_c \geq 0$  everywhere and in particular  $\hat{R}_c > 0$  in the sphere. We can calculate that  $Y(M_d \# R^d, \hat{\gamma}_c) > 0$ .

Therefore, it follows that:  $\sigma(M_d \# R^d) = \sup_{[g] \in \mathcal{C}_M} Y_{[g]} > 0$  and one can conclude that the manifold admits a metric of positive scalar curvature (**Yamabe problem**).