

# Recent MAGIC results on galactic sources

Julian Krause  
on behalf of the MAGIC collaboration

Max-Planck-Institut für Physik

April 1st 2011

# Outline

- 1 Mono-data
  - Region of SNR G65.1+0.6
  - Cygnus X-3

- 2 Stereo
  - Performance
  - LSI+61 303
  - HESS J0632+057

- 3 Summary & Outlook

# MAGIC phase I (2006-2009)



# Region of SNR G65.1+0.6

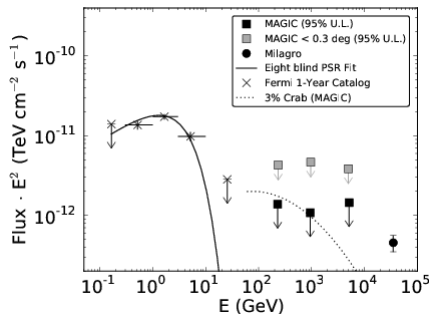
## Motivation

- The region contains 2 Milagro-detected Fermi bright sources (Abdo et al. 2009b)
  - ▶ 1FGL J1954.3+2836 ( $4.3 \sigma$  by Milagro between 10-50 TeV)
  - ▶ 1FGL J1958.6+2845 ( $4.0 \sigma$  by Milagro between 10-50 TeV)
- As of now both sources are identified as gamma-ray pulsars (Abdo et al. 2010)
  - ▶ Periods: 290ms, 92.7ms
  - ▶ Ages: 69.5kyrs, 21kyrs
  - ▶ Cutoffs: 2.9GeV, 1.2GeV

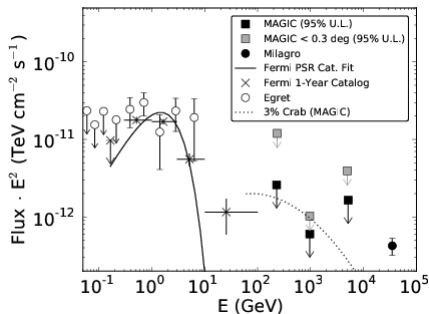
# MAGIC observations in 2009

- 25.5h of good quality data
- differential flux upper limits

J1954



J1958

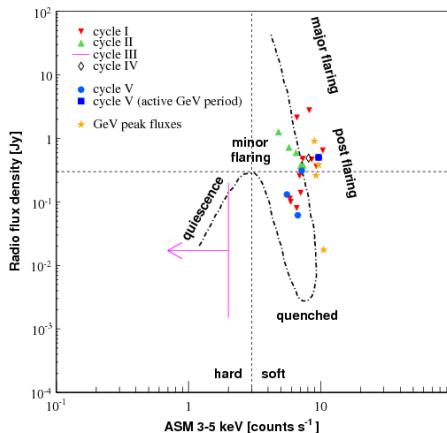


arXiv:1007.3359v1; Abdo et al. 2009b, 2009c, 2010; Albert et al. 2008a; Hartman et al. 1999; Saz Parkinson et al. 2010

## Conclusion

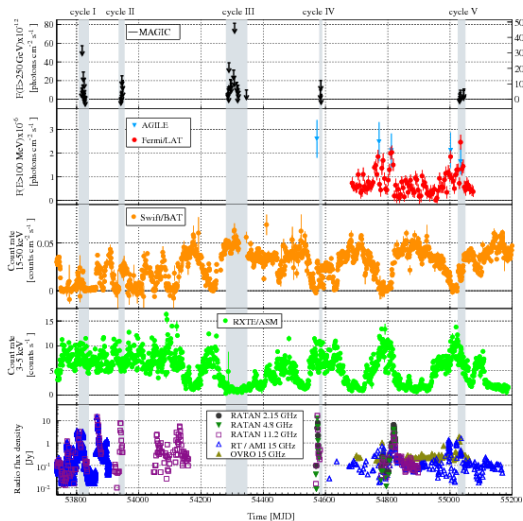
Probably pulsar wind nebula with an inverse Compton peak  $> 1$  TeV

# Cygnus X-3



- X-ray binary of a Wolf-Rayet star and unknown compact object
- Orbital period: 4.8h

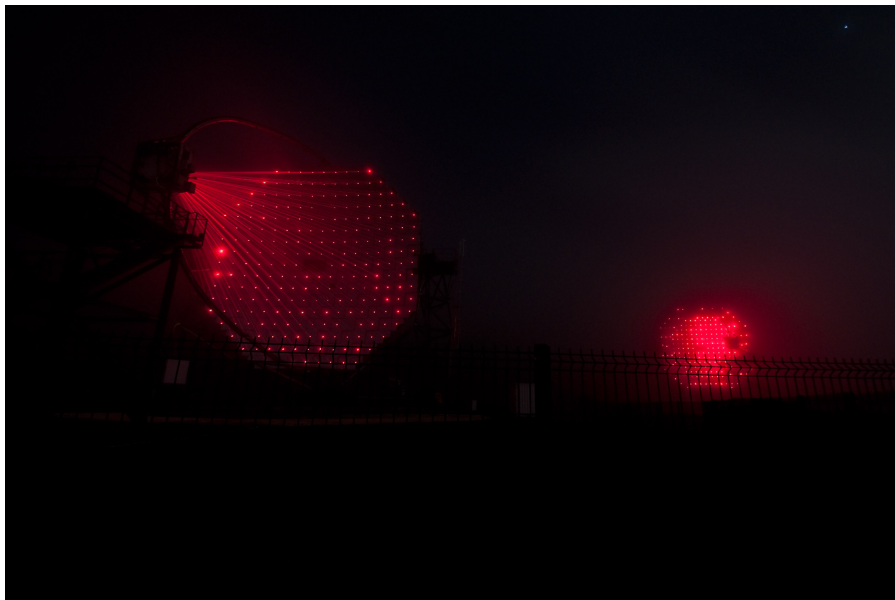
# MAGIC campaign over 4 years (2006-2009) for 56.7h



95% CI upper limits  
Energy > 250 GeV

- All data: 1.3 % crab
- Soft state: 2.5 % crab
- Hard state: 1.1% crab
- GeV Flare: < 6 % crab

# MAGIC phase II (since fall 2009)





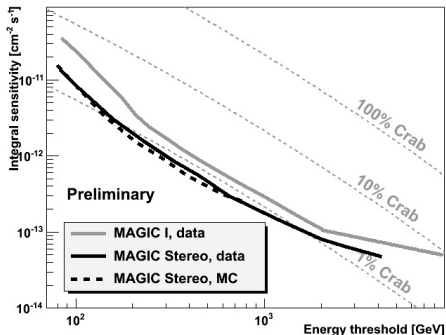
# Improvements due to the stereo system

- Stronger Background suppression
  - ▶ less muons
  - ▶ less accidentals
- 3D information of the shower
  - ▶ Precise estimation of the shower height maximum
  - ▶ Precise estimation of the impact parameter
- Higher energy resolution
- Higher angular resolution
- Better Off-axis performance

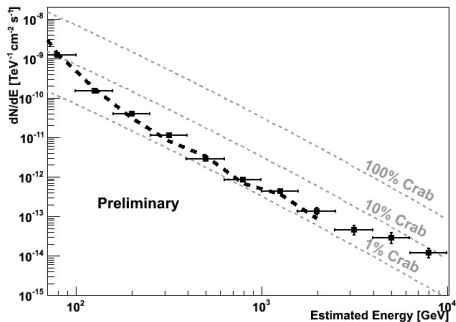
⇒ **Sensitivity improves by a factor  $\approx 2$**

# Sensitivity curves

## Integral



## Differential

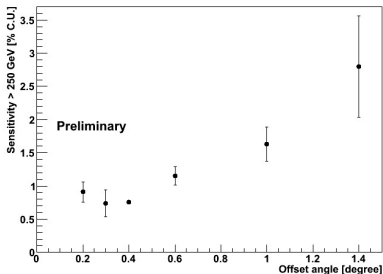


- 0.8 % Crab above 250 GeV
- Sensitivity is calculated as  $N_{\text{excess}} / \sqrt{N_{\text{bgd}}} = 5$  after 50 hours
- Differential sensitivity additionally requires a
  - ▶ Minimum of 10 Excess events
  - ▶ Signal  $\geq 5\%$  background

# Benefit for galactic sources

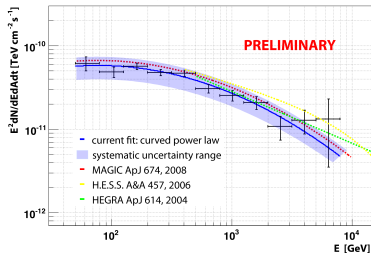
- Improved performance for
  - ▶ Off-axis sources
  - ▶ Extended sources
  - ▶ Variable sources
  
- Allows for
  - ▶ Morphological studies
  - ▶ High resolution spectra

## Off-axis performance



## Crab Nebula SED MAGIC Stereo

November 13-15th 2009, 190min effective observation time



# LSI+61 303

- System of a Be star and a compact object of unknown nature
- Orbital period in radio of  $26.4960 \pm 0.028$  days (Gregory 2002)
- Soft x-ray outburst modulated by orbital period (Paredes et al. 1997)
- Indication that X-ray outbursts change phases over the years (Torres et al. 2010)

## VHE gamma-rays

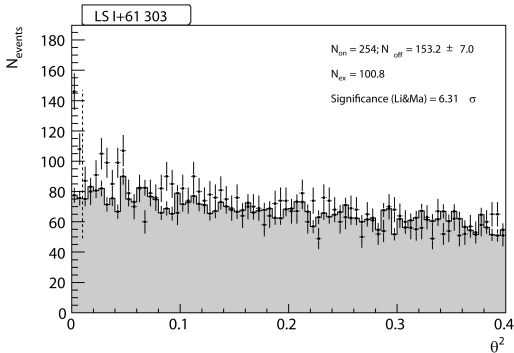
- System discovered by MAGIC in 2006 (Albert et al. 2006, Anderhub et al. 2009)
  - ▶ Orbital period:  $26.6 \pm 0.2$  days
  - ▶ Outburst in orbital phase interval 0.6-0.7 (periastron 0.275)
  - ▶ strong evidence for X-ray/gamma-ray correlation
- Emission confirmed by VERITAS in 2007 (Acciari et al. 2008)
  - ▶ Fall 2008 to early 2009 no detection
  - ▶ Fall 2009 no detection

# MAGIC Observations from 2009/10/15 to 2010/01/22

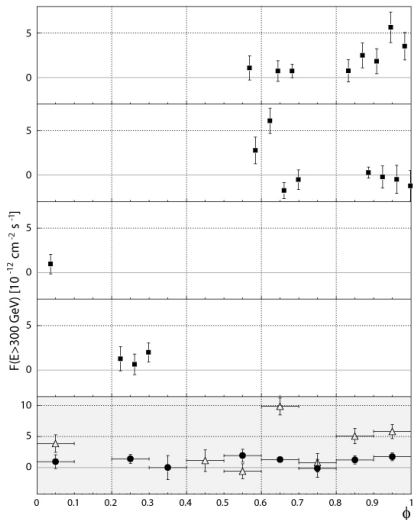
- Observations of consecutive nights with at least 3h observations
- 4 orbital cycles covered
- 48.4h of good quality data

## LSI is there

- 6.31  $\sigma$  detection
- Flux = 1.3 % Crab ( $> 300\text{GeV}$ )



# Periodicity and spectral study



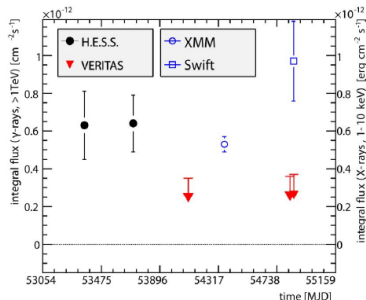
- Flux in phase 0.6-0.7 reduced by a factor 10
- Maximum flux at phase 0.62 with  $\approx 5.4\%$  crab
- Spectrum comparable to previous measurement

## Possible Explanation

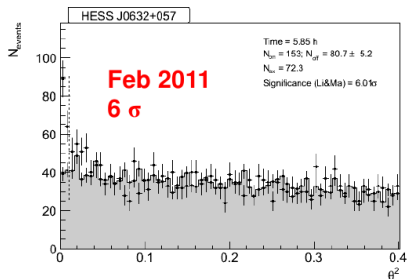
Increase of the stellar wind  
 $\Rightarrow$  cascading effects

# HESS J0632+057

- Discovered by H.E.S.S.
- Variable due to non detection by VERITAS
- Variable X-ray source at same position
- Possible binary, Be star MW148 position is comparable with VHE and X-ray sources
- no clear binray signs in radial velocity measurments



# Outburst in January&February 2011



- Swift-XRT (Atel #3152)
  - ▶ flux increases by a factor 3 from Jan 23th 2011 on till at least Feb 6th
  - ▶ probably binary with orbit of 310-320 days
- VERITAS (Atel #3153)
  - ▶ 7&8 February 2011 detected with  $8 \sigma$  in 5.6h
  - ▶ Flux  $\approx 4\%$  crab  $> 400$  GeV
- MAGIC (Atel #3161)
  - ▶ 7-9 February 2011 detected with  $> 5 \sigma$
  - ▶ Flux  $\approx 3.4\%$  crab  $> 200$  GeV



# Summary & Outlook

## Summary

- End of MAGIC phase I
  - ▶ Cygnus X-3 95 %UL ( $>250\text{GeV}$ ) of 1.3% crab
  - ▶ J1954 95 %UL ( $\approx 1\text{TeV}$ ) of 3% crab
  - ▶ J1958 95 %UL ( $\approx 1\text{TeV}$ ) of 2% crab
- Start of MAGIC phase II
  - ▶ Detected LSI+61 303 ( $<50\text{h}$ ); Flux ( $>300\text{GeV}$ ) 1.3% crab
  - ▶ Detected HESS J0632+057 ( $<6\text{h}$ ); Flux ( $>200\text{GeV}$ ) 3.4% crab
  - ▶ ...

## Outlook

- New galactic sources will probably be weak and/or extended
- More news to come on the conferences this summer