Moduli, Fluxes & the Hierarchy Problem

Sebastian Halter

Max-Planck-Institute for Physics, Munich

IMPRS/GK Young Scientists Workshop Wildbad Kreuth

July 28th, 2011

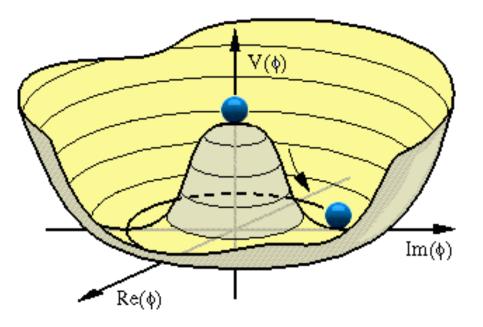
Outline

- 1. What are moduli?
- 2. Why are they important?
- 3. Why are they annoying?
- 4. What are fluxes?
- 5. Why are they useful?
- 6. What else do we need?
- 7. How to solve the hierarchy problem?

1. What are moduli?

Moduli $\leftarrow \leftarrow (4d \text{ scalar fields with } m \approx 0 \& M_P^{-1} \text{ couplings})$

←→ related to "degeneracy" of vacuum state



Geometric moduli ←→ fluctuations of metric along extradimensions

 \leftarrow (determine size and shape of compact space

Simplest example: compactify on circle S^1

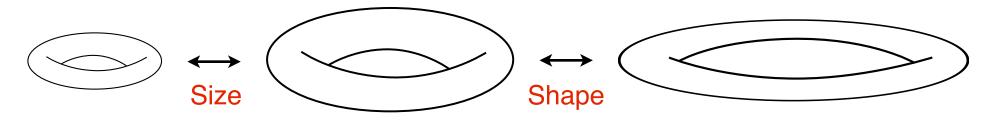
 \longleftrightarrow 1 modulus: fluctuation of radius $\iff G_{55} \sim R$

1. What are moduli?

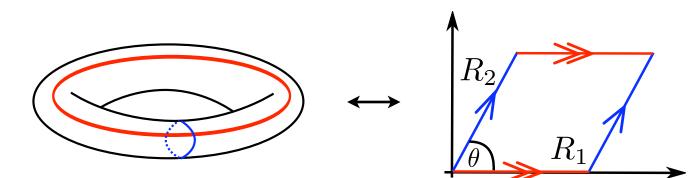
Next simplest example: compactify on 2 circles

$$\longleftrightarrow$$
 compactify on torus T^2

Size & shape is difference between



How to quantify these differences? Consider 2 "base" circles



Define 2 complex moduli fields:

$$T \sim R_1 R_2 \sin \theta + i B_{12} \& U \sim \frac{R_2}{R_1} \sin \theta + i \frac{R_2}{R_1} \cos \theta$$

2. Why are they important?

ANY parameter of 4d low-energy effective field theory, e.g.

a) gauge couplings,

b) or Yukawa couplings,

in string theory compactifications is of the generic form

 $g = #(\text{computable number}) \cdot f(\text{moduli})$

Example: 4d effective gauge coupling of 5d gauge boson on a circle

$$\frac{1}{g_{4d}^2} = \frac{R}{g_{5d}^2}$$

Main challenge for today: "Moduli stabilisation"

Values of all couplings in the vacuum? \longleftrightarrow What is $\langle R \rangle$ etc.?

What is effective 4d scalar potential for moduli? What minima exist?

Minima \longleftrightarrow Moduli acquire mass & expectation values

3. Why are they annoying?

Just compactification manifold: moduli either have

a) (approx.) no potential at all, i.e. moduli unfixed

b) or a "runaway" potential towards decompactification!

Well, that sucks... But why does it happen all the time?

Intuitive reason: think of compact (sub-)manifolds as "bubbles"

Bubbles have surface tension, i.e. curvature

Surface tension or curvature cost potential energy to maintain

VTheory wants to live in 10d & must be "forced" to 4d!Image: Typical potential energy:Typical potential energy:R $V \sim \pm \frac{\#}{R^n}$ R $\to \infty \text{ or } 0$

4. What are fluxes?

Fluxes are "generalized" electromagnetic fields, just with more indices Basic reason why they appear in string theory:

a) point particle: worldline couples to electromagnetic field ("1-form")

$$X_{\mu}(\tau) \rightarrow \oint d\Sigma_{\mu} A^{\mu}$$

b) string: worldsheet couples to "2-form" (2 antisymmetric indices)

$$X_{\mu}(\tau,\sigma) \rightarrow \oint d\Sigma_{\mu\nu} B^{\mu\nu}$$

c) Dp-brane: (p+1)-dim. worldvolume couples to "(p+1)-form"

$$X_{\mu}(\xi_0,\ldots,\xi_p) \to \oint d\Sigma_{\mu_0\ldots\mu_p} C^{\mu_0\ldots\mu_p}$$

Field Strength: $A_{\mu} \to F_{\mu\nu} = \partial_{[\mu}A_{\nu]}, B_{\mu\nu} \to H_{\mu\nu\rho} = \partial_{[\mu}B_{\nu\rho]}$ etc.

5. Why are they useful?

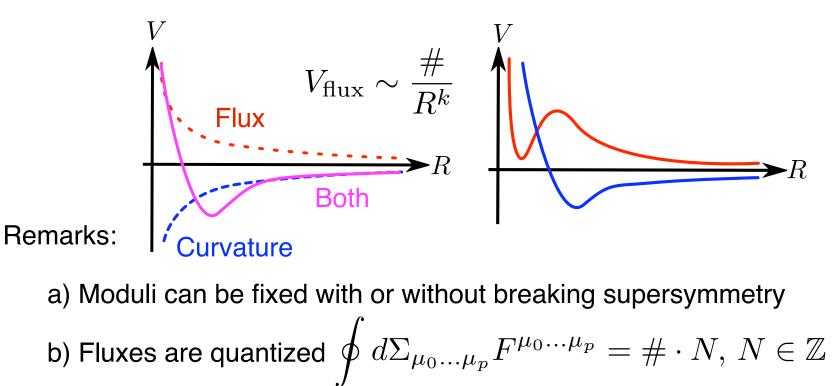
Need: additional energy source to balance surface tension

Fluxes: turn on "magnetic fields" in internal compact space

a) (p+1)-dim. manifold can support flux of p-form

e.g. magnetic field $F_{\mu\nu}$ on sphere S^2

b) "Magnetic" field energy density depends on $R \; !$



6. What else do we need?

Typically: fluxes stabilize most but not all moduli...

In particular: overall volume \mathcal{V}_6 can be still free... What else can we do?

Use "non-perturbative" effect: D7-brane wrapping a 4-dim. submanifold

a) 4d & 8d gauge coupling related by $g_{4d}^{-2} = g_{8d}^{-2} \cdot R^4$

b) Non-perturbative contributions to potential $\propto e^{-\#/g_{4d}^2}$

crucial: not visible in perturbative limit $g_{4d} \rightarrow 0$

c) Strong coupling phenomenon: "gaugino" condensation $\langle \lambda \overline{\lambda} \rangle \neq 0$ Similar to other fermion condensates, e.g.

a) Chiral symmetry breaking in QCD $\langle q \bar{q} \rangle
eq 0$

b) Technicolor EWSB $\langle \psi \bar{\psi} \rangle \neq 0$

All geometric moduli can be stabilised in principle...

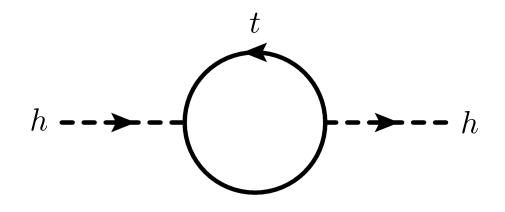
7. How to solve the hierarchy problem?

What is the hierarchy problem?

It is not about why $m_h^2 \ll M_P^2$

It is about the "radiative instability" of this hierarchy!

Quantum corrections induce large mass: $m_h^2 \sim \Lambda^2 + m_{h,0}^2$



Rough classification of solutions:

a) Eliminate quantum corrections above $\Lambda \sim {
m TeV}$ (SUSY)

b) Explain why $\Lambda \sim {
m TeV}$ instead of $\Lambda \sim M_P$ (everything else)

7. How to solve the hierarchy problem?

Moduli can solve it in both ways:

a) Act as hidden sector for spontaneous SUSY breaking

b) Make gravity strongly coupled at $\sim TeV$

How to make gravity strongly coupled?

a) Experimentally only M_P is fixed

b) String theory relates M_P and string scale $M_s \sim \ell_s^{-1}$ by

$$M_P^2 \sim M_s^2 \cdot (\mathcal{V}_6 \ell_s^{-6})$$

c) Gravity becomes strongly coupled at $\sim TeV$ if

 $M_s \sim \text{TeV} \& \mathcal{V}_6 \sim 10^{30} \ell_s^6$

d) LARGE Volume Scenario (or ADD scenario)

 \leftrightarrow gravity diluted in extradimensions



Summary

Moduli are generic to string theory compactifications:

Geometric moduli characterize size & shape of extradimensions

Moduli determine couplings of 4d effective field theory

Decompactification problem:

Surface tension: moduli have "runaway" potentials $V \sim \pm \frac{\#}{R^n}$ Fluxes are also generic to string theory compactifications Internal magnetic fields: balance surface tension $V_{\rm flux} \sim \frac{\#}{R^k}$

Non-perturbative effects: corrections to potential $\propto e^{-\#/g_{4d}^2}$

$$g_{4d}^{-2} = g_{8d}^{-2} \cdot R^4$$

Hierarchy problem: make gravity strongly coupled at $\sim {
m TeV}$

LARGE Volume / ADD scenario $M_s \sim {\rm TeV} \ \& \ \mathcal{V}_6 \sim 10^{30} \ell_s^6$