# Building a minimal dark matter model with kinetic mixing

- a small cookbook -

Sophia Borowka

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Extending the Standard Model gauge group

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \xrightarrow{\text{sp. symmetry}} SU(3)_C \otimes U(1)_{\text{em}}$$

by an extra  $U(1)_X$  via kinetic mixing

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_X \xrightarrow{\text{sp. symmetry}} SU(3)_C \otimes U(1)_{\text{em}}$$

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- few additional parameters
  - tight constraints possible
  - lesser assumptions can be made by model builder
- operators are of dimensions < 4</li>
  - renormalizability garanteed

The first minimal dark matter models proposed by C.Boehm and P. Fayet '03.

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### How to start

Let us first examine the SM and the dark gauge group independently

 $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \qquad \qquad U(1)_X$ 

and build the Lagrangian for the dark  $U(1)_X$ 

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- $\mathcal{L}_{\mathcal{X}} = -$  kinetic field term
  - + particle field term
  - + source term

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# The field Lagrangians

Each field has its own Lagrangian

• Dirac field  $\rightarrow$  Dirac particles (spin 1/2)

$$\mathcal{L}_{\mathcal{D}} = ar{\psi}(i\partial\!\!\!/_{\mu} - m_\psi)\psi$$

• Klein-Gordon field  $\rightarrow$  scalar particles (spin 0)

$$\mathcal{L}_{\mathcal{KG}} = (\partial_\mu \phi)^\dagger (\partial^\mu \phi) - m_\phi^2 \phi^\dagger \phi$$

• Vector field  $\rightarrow$  vector particles (spin 1)

$$\mathcal{L}_{\mathcal{V}} = -rac{1}{4}(\partial_{\mu}\mathcal{A}_{
u} - \partial_{
u}\mathcal{A}_{\mu})(\partial^{\mu}\mathcal{A}^{
u} - \partial^{
u}\mathcal{A}^{\mu}) + rac{1}{2}m_{\mathcal{A}}^{2}\mathcal{A}_{\mu}\mathcal{A}^{\mu}$$

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- $\mathcal{L}_{\mathcal{X}} = -$  kinetic field term
  - + particle field term
  - + source term

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## The kinetic field term

The field term consists of two contracted field strength tensors and a mass term in the massive case

$$-rac{1}{4} C_{\mu
u} C^{\mu
u} + rac{1}{2} m_X^2 C_\mu C^\mu$$

where the field strength tensor in the Abelian case is

$$C_{\mu\nu} = \partial_{\mu}C_{\nu} - \partial_{\nu}C_{\mu}$$
.

In a non-Abelian case

$$C^{a}_{\mu\nu} = \partial_{\mu}C^{a}_{\nu} - \partial_{\nu}C^{a}_{\mu} + gf^{abc}C^{b}_{\mu}C^{c}_{\nu}$$

$$\begin{aligned} \mathsf{Fun fact}: \ C^a_{\mu\nu}C^{a\mu\nu} &= (\partial_\mu C^a_\nu - \partial_\nu C^a_\mu)^2 + 2(\partial_\mu C^a_\nu - \partial_\nu C^a_\mu)gf^{abc}C^{b\mu}C^{c\nu} \\ &+ g^2 f^{eab}f^{ecd}C^a_\mu C^b_\nu C^{c\mu}C^{d\nu} \end{aligned}$$

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$$\mathcal{L}_{\mathcal{X}} = -\frac{1}{4}C_{\mu
u}C^{\mu
u} + \frac{1}{2}m_{X}^{2}C_{\mu}C^{\mu}$$
  
+ particle field term

+ source term

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What type of particles are interacting with the new vector boson  $C_{\mu}$ ?

• Dirac field  $\rightarrow$  Dirac particles (spin 1/2)

 $\bar{\chi}(i\partial_{\mu}-m_{\chi})\chi$ 

• Klein-Gordon field  $\rightarrow$  scalar particles (spin 0)

 $(\partial_{\mu}\chi)^{\dagger}(\partial^{\mu}\chi) - m_{\chi}^{2}\chi^{\dagger}\chi$ 

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$$\mathcal{L}_{\mathcal{X}} = -\frac{1}{4}C_{\mu\nu}C^{\mu\nu} + \frac{1}{2}m_{X}^{2}C_{\mu}C^{\mu} + \bar{\chi}(i\partial_{\mu} - m_{\chi})\chi + \text{ source term}$$

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The source is a current resulting from local gauge invariance

 $g_X J_X^\mu C_\mu$  .

• Dirac particles  $\rightarrow$  Dirac current

$$J_X^\mu = \epsilon^{\chi}_{Z'} \bar{\chi} \gamma^\mu \chi$$

• Scalar particles  $\rightarrow$  Klein-Gordon current

$$J_{X}^{\mu} = \epsilon_{Z'}^{\chi} \chi^{\dagger} (i\overleftarrow{\partial_{\mu}} - i\overrightarrow{\partial^{\mu}}) \chi$$

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# Where are the couplings?

The couplings are "hidden" in the source term

• Dirac current:  $\bar{\chi}\gamma^{\mu}g_{X}\epsilon^{\chi}_{Z'}\chi C_{\mu}$ 

 $g_{\chi}\epsilon^{\chi}_{\tau'}\gamma^{\mu}$ 



$$g_X \epsilon_{Z'}^{\chi} (p_2 - p_1)^{\mu}$$

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$$\mathcal{L}_{\mathcal{X}} = -\frac{1}{4}C_{\mu\nu}C^{\mu\nu} + \frac{1}{2}m_{X}^{2}C_{\mu}C^{\mu} + \bar{\chi}(i\partial_{\mu} - m_{\chi})\chi + \bar{\chi}\gamma^{\mu}g_{X}\epsilon_{Z'}^{\chi}\chi C_{\mu}$$

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# Kinetic mixing - introducing a link



B. Holdom '86

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Kinetic mixing constitutes a possibility to link the Standard Model with the dark sector:

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_X - \frac{\epsilon}{2} C_{\mu\nu} B^{\mu\nu} \tag{1}$$

bring the field terms into bilinear form

$$\begin{split} \mathcal{L}_{\text{field}} &= -\frac{1}{4} B_{\mu\nu} B^{\mu\nu} - \frac{1}{4} F^{a}_{\mu\nu} F^{a\mu\nu} - \frac{1}{4} C_{\mu\nu} C^{\mu\nu} - \frac{\epsilon}{2} C_{\mu\nu} B^{\mu\nu} + \dots \\ &= -\frac{1}{4} \vec{F}^{T} \begin{pmatrix} 1 & \epsilon & 0 \\ \epsilon & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \vec{F} + \dots \end{split}$$

with  $\vec{F} = (C_{\mu\nu}, B_{\mu\nu}, F^3_{\mu\nu})^T$ 

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## Mass matrix

bring the mass terms into bilinear form

$$\mathcal{L}_{\text{mass}} = \frac{1}{2}(0, v) \left[ \frac{g_1^2}{4} (B_{\mu})^2 + \frac{g_2^2}{4} (A_{\mu}^3)^2 + \frac{g_1 g_2}{4} A_{\mu}^3 B^{\mu} \right] \begin{pmatrix} 0 \\ v \end{pmatrix} + \frac{1}{2} m_X^2 (C_{\mu})^2 + \dots \\ = \frac{1}{2} \vec{V}^T \begin{pmatrix} m_X^2 & 0 & 0 \\ 0 & \frac{1}{4} v^2 g_1^2 & -\frac{1}{4} v^2 g_1 g_2 \\ 0 & -\frac{1}{4} v^2 g_1 g_2 & \frac{1}{4} v^2 g_2^2 \end{pmatrix} \vec{V} + \dots \\ = \vec{V} = (C - B - A^3) \vec{L}$$

with  $V = (C_{\mu}, B_{\mu}, A_{\mu}^{s})$ 

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Eigenvalues of the rotated mass matrix M

$$m_{\gamma}^2 = 0, \ m_Z^2 = \frac{(q-p)}{2}, \ m_{Z'}^2 = \frac{(q+p)}{2}$$

with

$$q = m_X^2 eta + rac{g_1^2 eta + g_2^2}{4} v^2$$
,  $p = \sqrt{q^2 - 4m_X^2 (m_Z^{SM})^2 eta}$ 

and  $\beta = \frac{1}{1-\epsilon^2}$ .

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# **Kinetic Mixing - Neutral Current Couplings**



In canonical normalization

$$\mathcal{L}_{\mathcal{NC}} = \frac{g_2}{2\cos\theta_W} \bar{f} \gamma^\mu [(v_f' - \gamma_5 a_f') Z_\mu' + (v_f - \gamma_5 a_f) Z_\mu] f + e\bar{f} \gamma^\mu Q_f A_\mu f + \bar{\chi} \gamma^\mu [\epsilon_{Z'}^{\chi} Z_\mu' + \epsilon_Z^{\chi} Z_\mu + \epsilon_{\gamma}^{\chi} A_\mu^{\gamma}] \chi$$

For small kinetic mixing and arbitrary m<sub>Z'</sub> the coupling factors to SM fermions f and dark fermions χ are

$$\begin{aligned} v_f, a_f &= v_f^{SM}, a_f^{SM} + \mathcal{O}(\epsilon^2) & v_f' \propto \epsilon + \mathcal{O}(\epsilon^3) \\ \epsilon_Z^{\chi} \propto \epsilon + \mathcal{O}(\epsilon^3) & a_f' \propto \epsilon + \mathcal{O}(\epsilon^3) \\ \epsilon_{\gamma}^{\chi} &= 0 \text{ (no millicharge).} & \epsilon_{Z'}^{\chi} \propto g_X Q_X + \mathcal{O}(\epsilon^2) \end{aligned}$$

## A dark matter cross section prediction



#### preliminary

S. Borowka (MPI for Physics )

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#### Summary

- learned how to construct a dark matter minimal Lagrangian
- used kinetic mixing to link the Standard Model with the dark sector
- discovery potential in collider and direct dark matter detection experiments

Outlook

update limits

#### Literature

- Scalar Dark Matter candidates (C. Boehm, P. Fayet '03): arxiv:hep-ph/0305261
- ► Two U(1)'s and e charge shifts (B. Holdom '86): Phys. Lett. B, 166, 196
- First Dark Matter Results from the XENON100 Experiment (XENON100 Collaboration '10): arxiv:1005.0380 [astro-ph]
- Status and Sensitivity Projections for the XENON100 Dark Matter Experiment (E. Aprile, L. Baudis '09): arxiv:0902.4253 [astro-ph.IM]