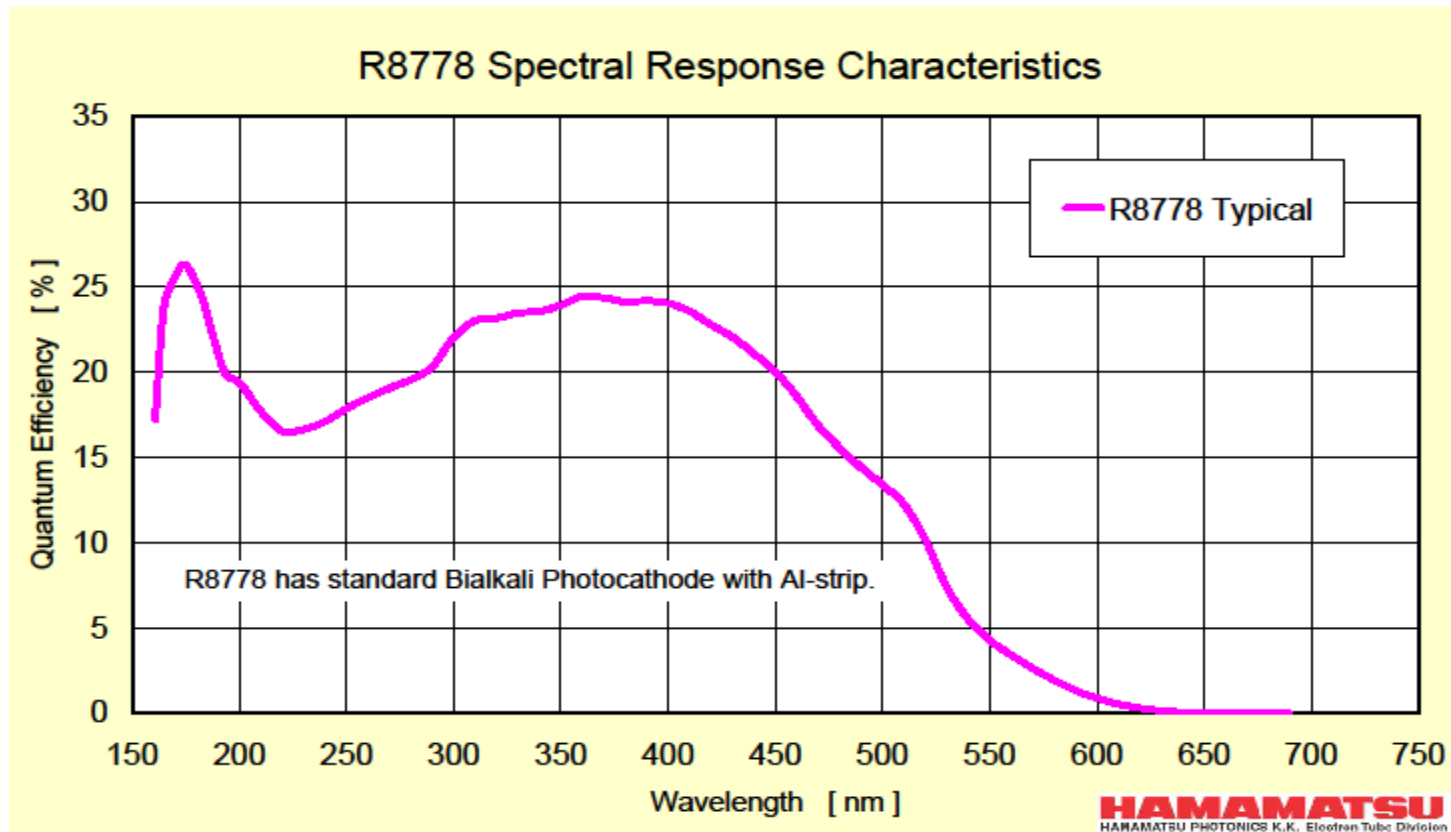


CAN A PHOTON GENERATE MORE THAN ONE ELECTRON?

Jon Howorth

Ralph Powell



Some Ultraviolet Spectral Sensitivity
Characteristics of Broadband
Multialkali Photocathodes

C B Johnson, L Bonney, R F Floryan

ITT Electro-Optical Products Division
3700 E. Pontiac Street
Fort Wayne, IN 46081

SPIE Vol. 932 Ultraviolet Technology II (1998)

Electron Energies Distribution

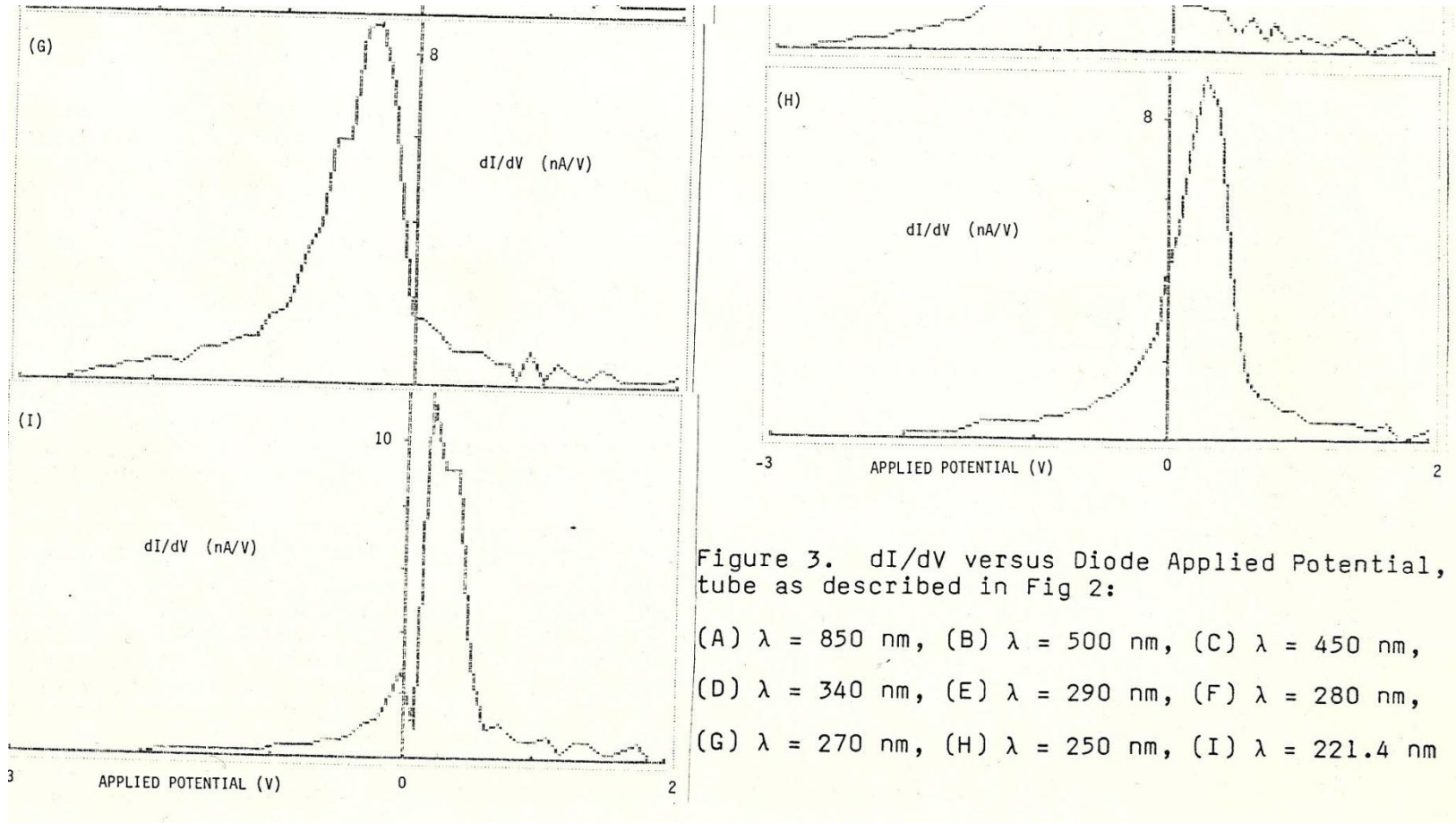
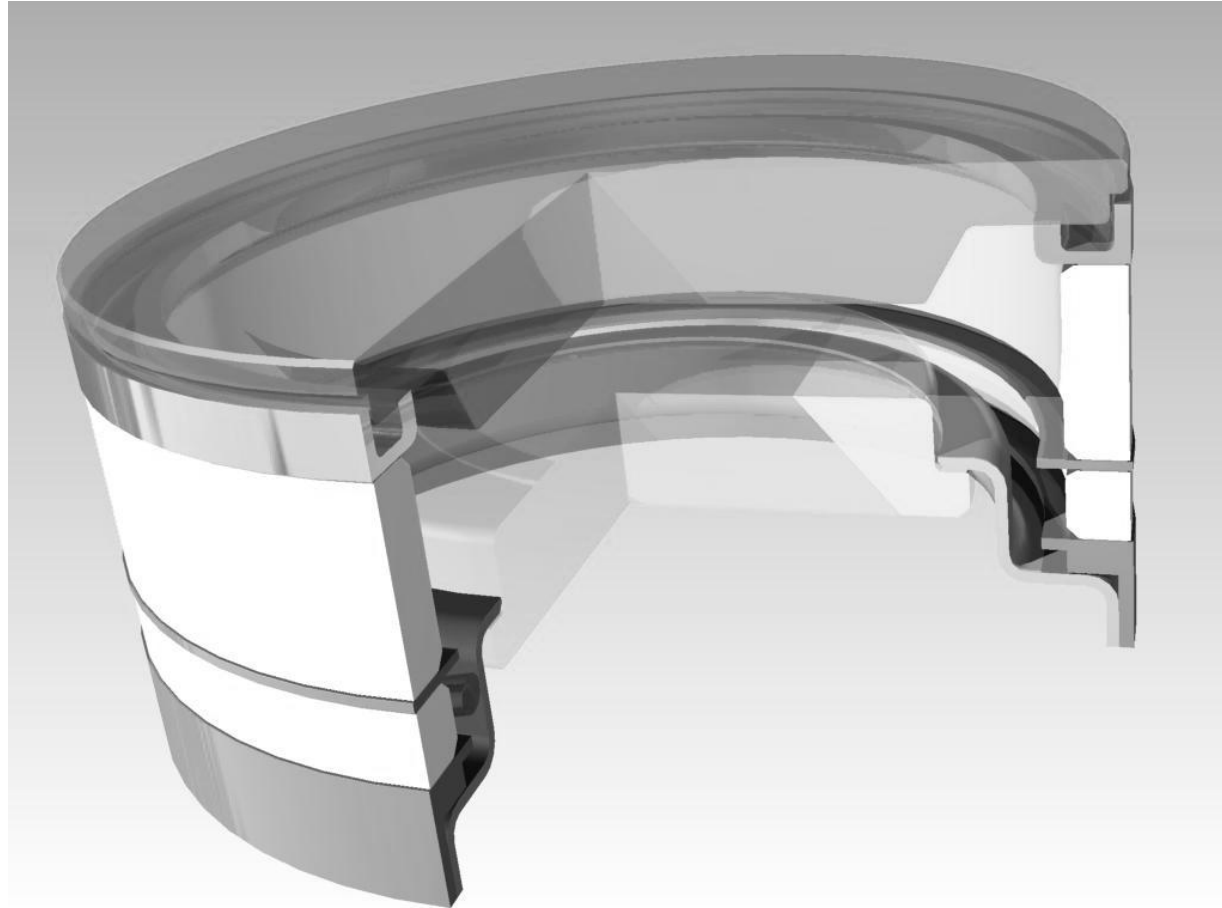


Figure 3. dI/dV versus Diode Applied Potential, tube as described in Fig 2:

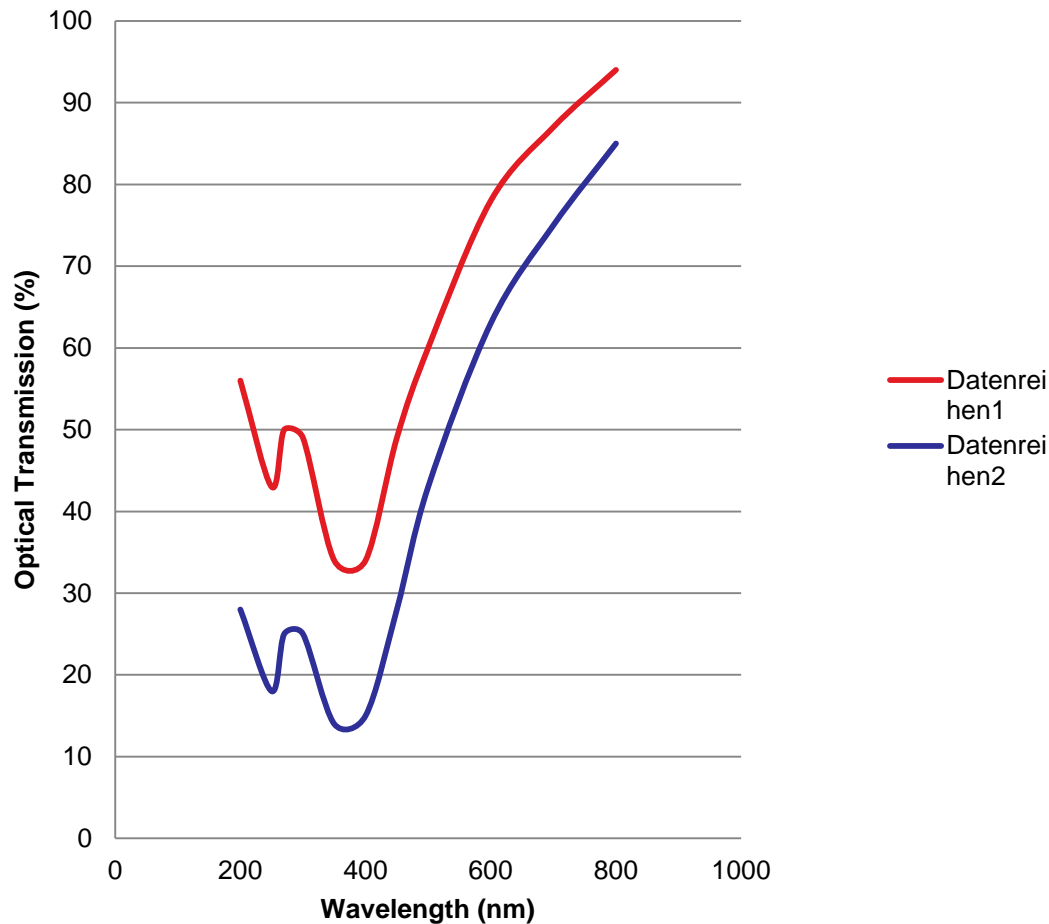
- (A) $\lambda = 850$ nm, (B) $\lambda = 500$ nm, (C) $\lambda = 450$ nm,
- (D) $\lambda = 340$ nm, (E) $\lambda = 290$ nm, (F) $\lambda = 280$ nm,
- (G) $\lambda = 270$ nm, (H) $\lambda = 250$ nm, (I) $\lambda = 221.4$ nm

We first formulated the following simple physical model to explain this behavior. Low energy photons, near the threshold of photoemission, impact with enough energy to excite valence band electrons into the conduction band with sufficient excess energy to overcome the electron affinity at the photocathode-to-vacuum interface and escape into the vacuum. As the photon energy increases more electrons are excited into the conduction band in greater numbers and with greater energies, and the photocurrent increases. This trend continues with increasing photon energy until electrons arrive at the conduction band with sufficient energy to generate additional hole-electron pairs at collisions with the lattice. These inelastic collisions produce two slow electrons that have a lower combined probability to escape into the vacuum, and the photocurrent begins to fall with increased photon energy. As the photon energy continues to increase, a point is reached for which the primary and secondary electrons have sufficient combined energies to escape from the photocathode in sufficient numbers to cause the photocurrent to increase to a second maximum, ie $n = 3$. At the next maximum, two hole-electron pairs are generated by the primary high-energy electron, and the resulting three slower electrons have a combined lower probability to escape, and the photocurrent drops as the photon energy increases.

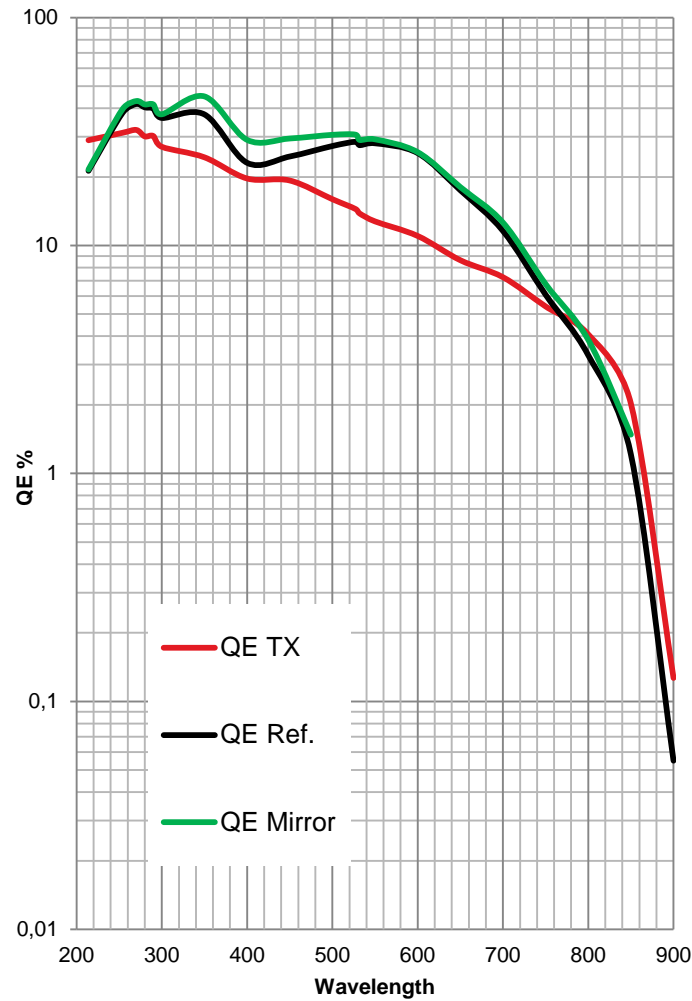
Test Cell



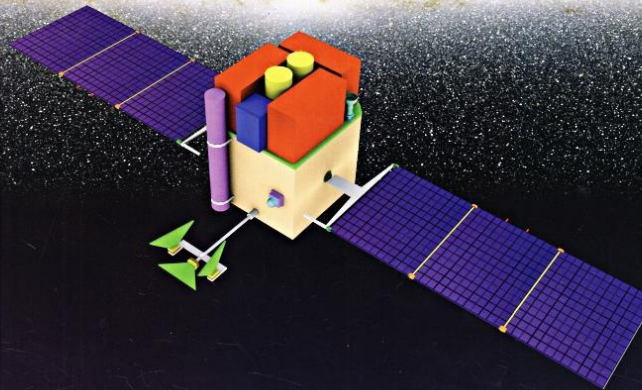
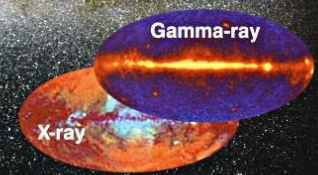
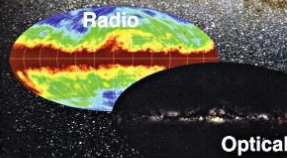
OPTICAL TRANSMISSION S-20 PHOTOCATHODES



Reflection/Transmission Test Cell



ASTROSAT



Science Objectives

- Multiwavelength spectral and timing studies of X-ray binaries, pulsars, SNRs, AGNs, galaxy clusters
- High resolution studies of galaxy morphology in UV
- Sky survey in hard X-ray and UV bands
- Detection of new X-ray transients and routine monitoring of bright X-ray sources
- UV background studies

Multiwavelength studies

Most astronomical objects in the Universe emit radiation spanning the complete electromagnetic spectrum from long wavelength radio waves to very short wavelength gamma rays. Thus for a detailed understanding of the physical processes that give rise to wavelength dependent, time variable phenomena, it is essential to carry out simultaneous multiwavelength observations.

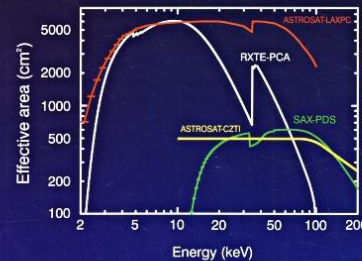
Ultra-Violet Imaging Telescope (UVIT)



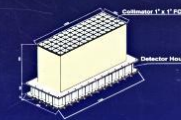
Soft X-Ray Telescope (SXT)



Astrosat is India's first satellite dedicated to multiwavelength astronomy. It will be launched into a 650 km, low inclination ($<10^\circ$) orbit by the PSLV in 2007. The 5 payloads permit simultaneous observations of cosmic objects at UV, soft X-ray and hard X-ray regions of the electromagnetic spectrum.



Large Area Xenon Proportional Counter (LAXPC)



Scanning Sky Monitor (SSM)



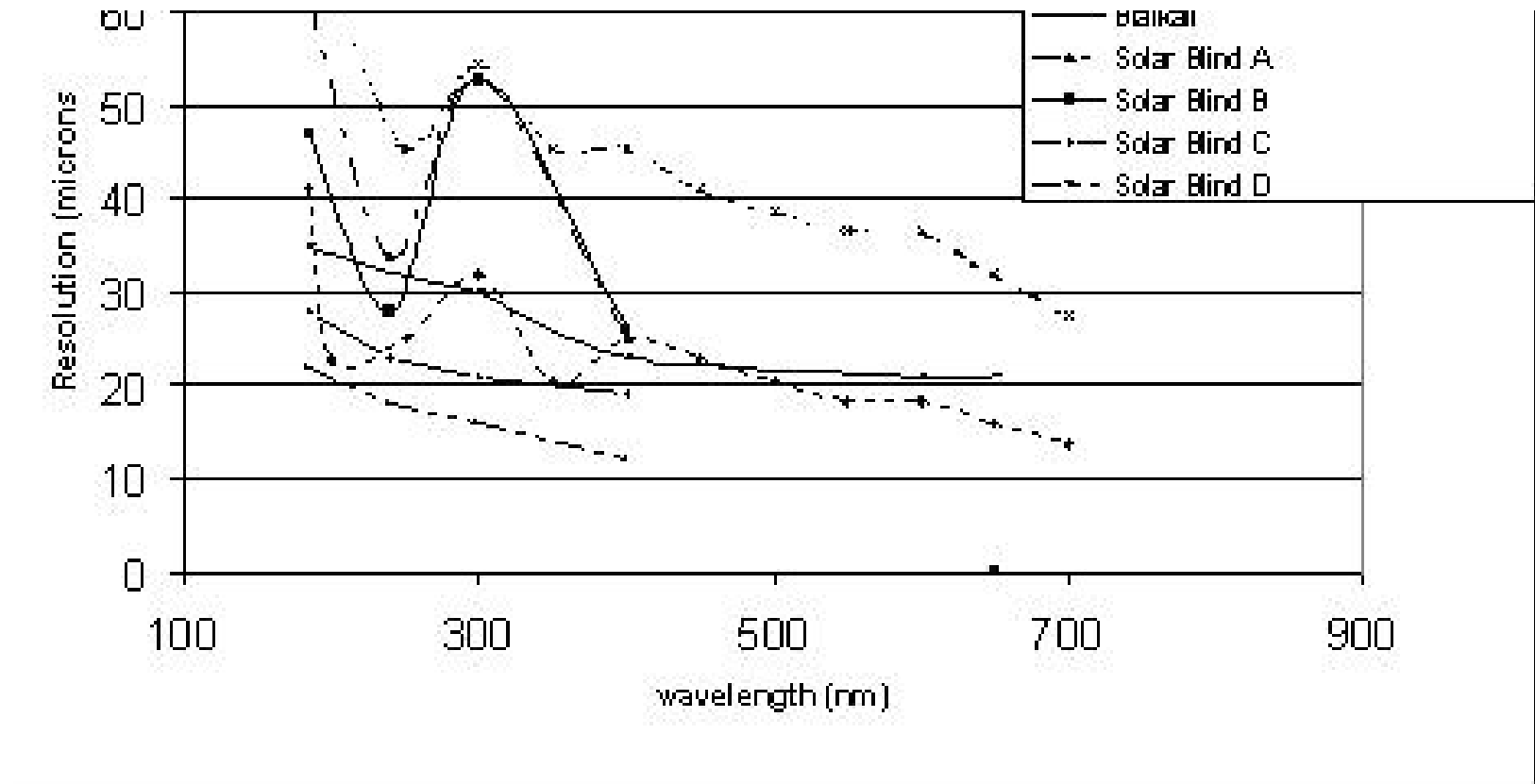
CdZnTe Imager (CZTI)



	UVIT / OPT	SXT	LAXPC	CZTI	SSM
Detector	UV: photon counting CCD Opt: CCD photometer	X-ray CCD (at the focal plane)	Proportional counter	CdZnTe detector array	Position-sensitive proportional counter
Imaging / non-imaging	imaging	imaging	non-imaging	imaging	imaging
Optics	Twin Ritchey Chretien 2 mirror system	Conical foil (~Wolter-I) mirrors	Collimator	2-D coded mask	1-D coded mask
Bandwidth	1300-3200 Ang	0.3 - 8 keV	3 - 100 keV	10 - 100 keV	2 - 10 keV
Geometric Area (cm²)	1250	250	10800	1000	180
Effective Area (cm²)	60 (depends on filter)	125 @ 0.5 keV 200 @ 1-2 keV 25 @ 6 keV	6000 @ 5-30 keV (E > 10 keV)	1000 90 @ 5 keV (Xe gas)	~40 @ 2 keV 90 @ 5 keV (Xe gas)
Field of View	0.50° dia	0.35° (FWHM)	19° x 19°	17° x 17°	6° x 90°
Energy Resolution	<1000 Å (depends on choice of filters)	2% @ 6 keV	9% @ 22 keV	3% at 10 keV	19% @ 6 keV
Angular Resolution	1.8 arcsec	3 - 4 arcmin (HPD)	~(1-5) arcmin (in scan mode only)	8 arcmin	~10 arcmin
Time resolution	10 ms	2.6μ, 0.3μ, 1ms	10 microsec	1 ms	1 ms
Typical observation time per target	30 min	0.5 - 1 day	1 - 2 days	2 days	5 min
Sensitivity (Obs. Time)	21 st magnitude (5σ) (1.800x)	10 microCrab (5σ) (10000x)	0.1 milliCrab (3σ) (1000x)	0.5 milliCrab (3σ) (1000x)	~30 milliCrab (3σ) (~300x)

Multiwavelength images courtesy HEASARC : IR image of galaxy (credit : Univ. Of Massachusetts and IPAC/CALTECH)

Resolution/Wavelength



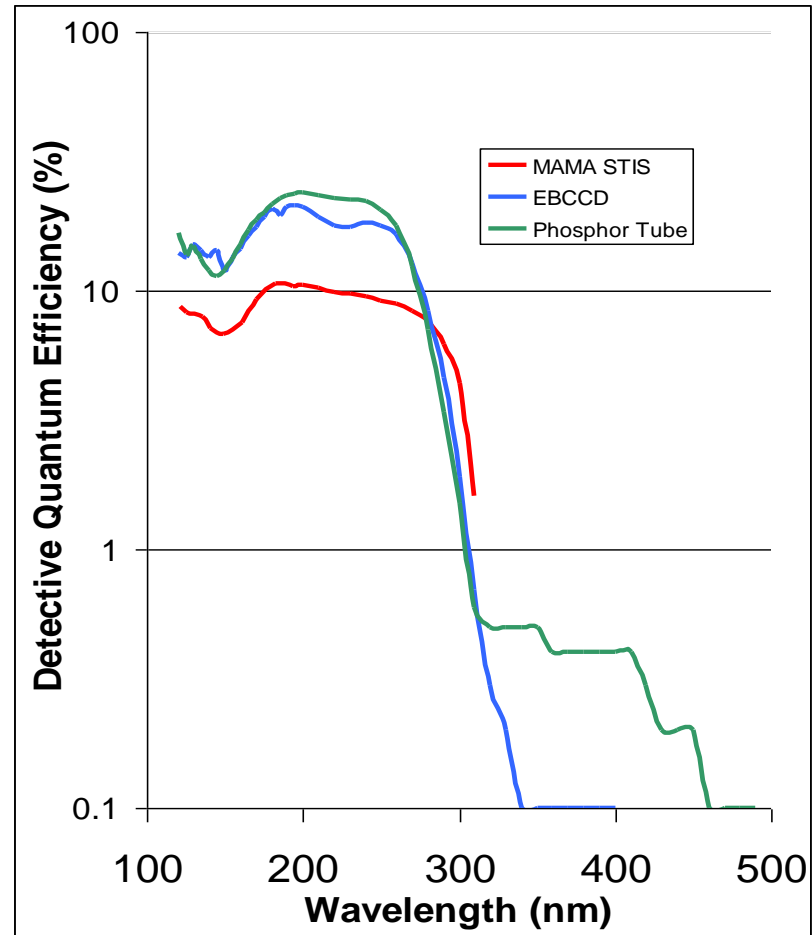
Summary for S-20

- Electron Energy Distribution-marked change about 270nm
- Optical Absorption peaks at 370nm and a secondary peak at 270nm
- Not an Interference Effect!
- QE peaks at 370nm (45%) and 260nm
- Spatial Resolution improves at 260nm

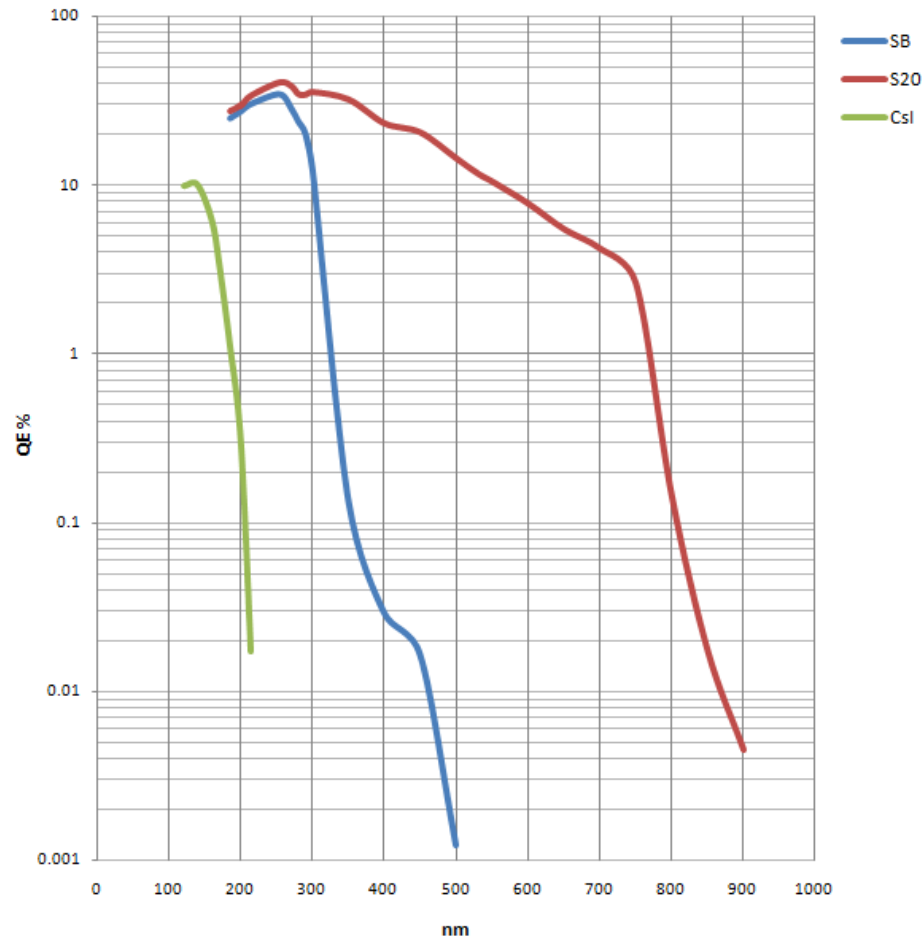
Do Other Cathodes show a similar Effect

- SB cathodes show a dip in QE
- Our Electron energy distribution and Spatial Resolution measurements were inconclusive

Photek 2007



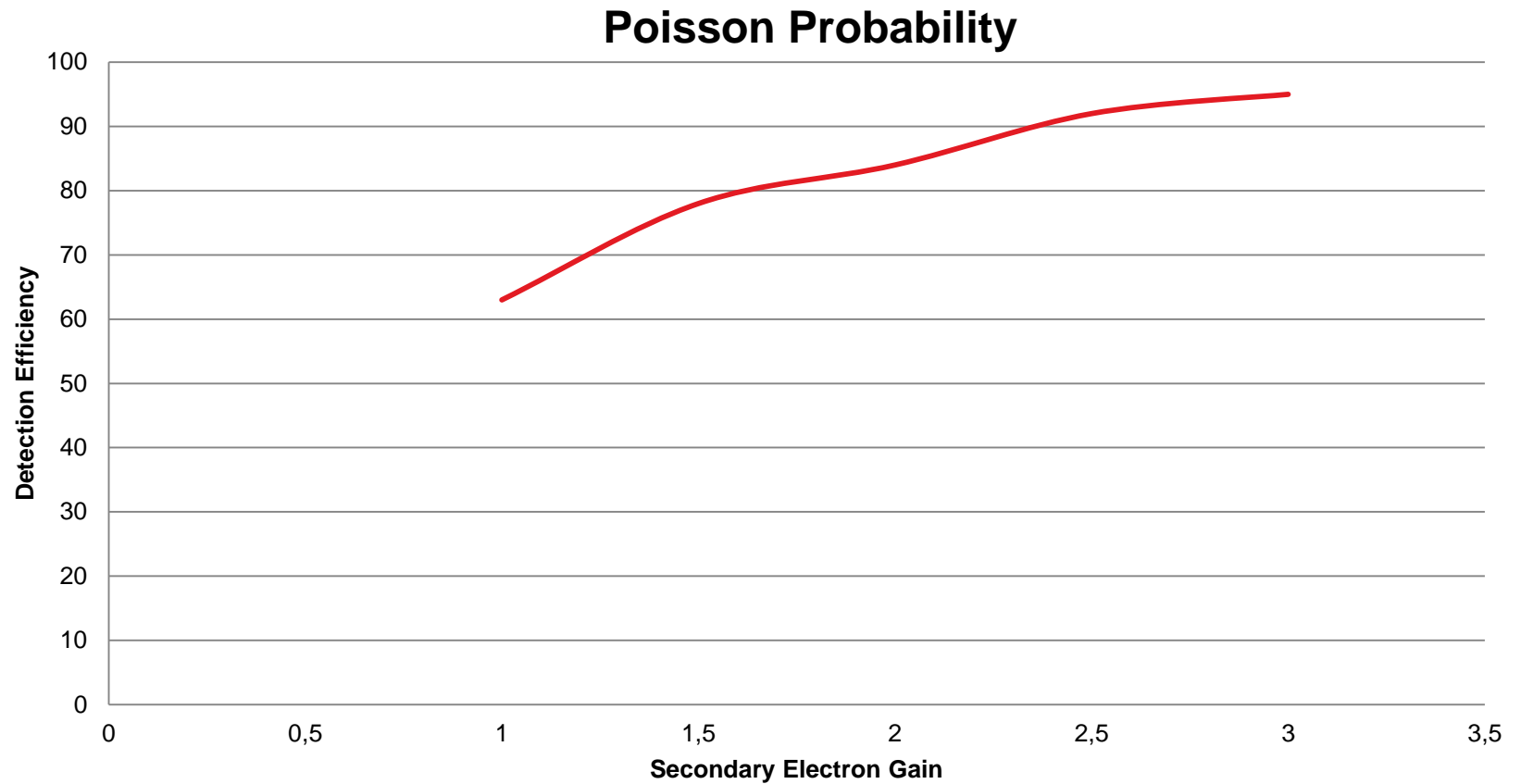
Standard Spectral Responses



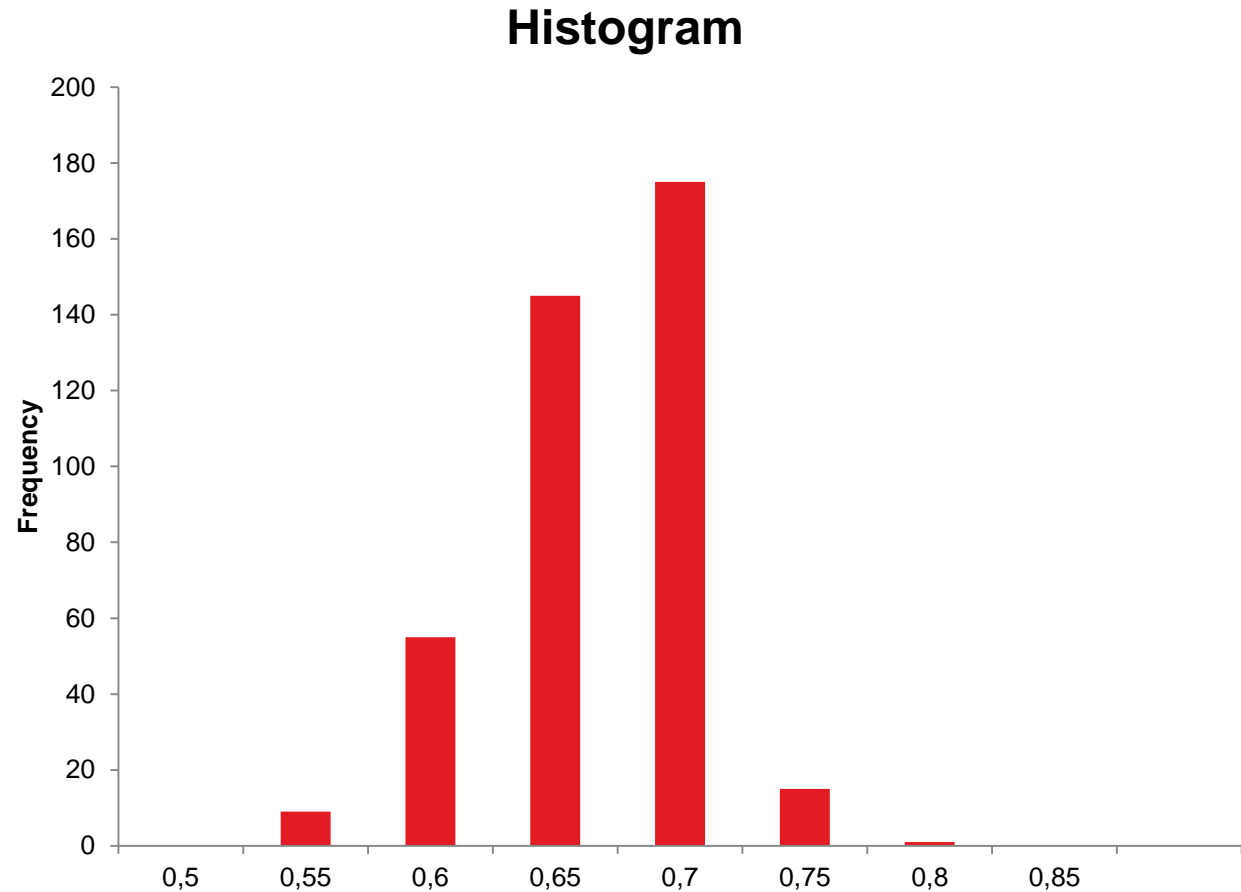
Counting Efficiency

1. Some Electrons are reflected (depends on Z)
2. Poisson Distribution
3. Geometrical Losses (e.g. OAR)

Counting Efficiency



Counting Efficiency of MCPs



Does It Matter?

- $DQE = QE \times CE \times \text{number of electrons/photon}$
- If CE is 50%, and we have 2 electrons/photon, the net result is similar to a photocathode with a cathode that only gives one electron/photon
- If CE is high, the same cathode that reads high QE because it is 2 electrons/photon has not the expected superior SNR

Thank you for listening

Photek Limited

26 Castleham Road, St Leonards on Sea, East Sussex, TN38 9NS, UK

T +44 (0)1424 850555

F +44 (0)1424 850051

E sales@photek.co.uk

W www.photek.co.uk