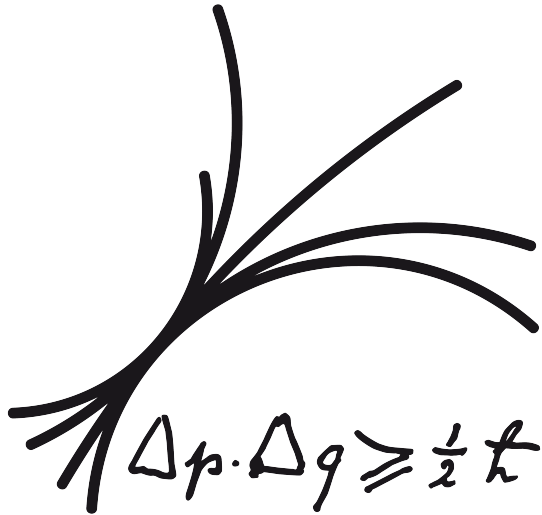


Project Review Talk: Neutrino Physics Leaving the Standard Path



Alexander Merle
MPP München



MAX-PLANCK-GESELLSCHAFT

Based on:

some “highlights” & work in progress

Project Review, 15-12-2014

1. Astroparticle Physics Group

- **Dr. habil. Georg Raffelt** (*group leader*)
- **Dr. Frank Daniel Steffen** (*permanent scientist*)
- **Dr. Sovan Chakraborty** (*postdoc*)
- **Dr. Alexander Kartavtsev** (*postdoc*)
- **Dr. Alexander Merle** (*postdoc*)
- **Dr. Antonio Palazzo** (*postdoc*)
- **Alexandra Dobrynina** (*Ph.D. student*)
- **Ignacio Izaguirre** (*Ph.D. student*)
- **Maximilian Totzauer** (*Ph.D. student*)
- **Hendrik Vogel** (*Ph.D. student*)
- **Jost Migenda** (*Master student*)
- **Moritz Platscher** (*Master student*)
- **Julia Stadler** (*Master student*)

1. Astroparticle Physics Group

- Dr. habil. Georg Raffelt (*group leader*)
- Dr. Frank Daniel Steffen (*permanent scientist*)
- Dr. Sovan Chakraborty (*postdoc*)
- Dr. Alexander Kartavtsev (*postdoc*)
- Dr. Alexander Merle (*postdoc*)
- Dr. Antonio Palazzo (*postdoc*)
- Alexandra Dobrynina (*Ph.D. student*)
- Ignacio Izaguirre (*Ph.D. student*)
- Maximilian Totzauer (*Ph.D. student*)
- Hendrik Vogel (*Ph.D. student*)
- Jost Migenda (*Master student*)
- Moritz Platscher (*Master student*)
- Julia Stadler (*Master student*)

In case you did not pay attention before:

This is the weird guy talking to you right now.

1. Astroparticle Physics Group

RESEARCH DONE IN THE ASTROPARTICLE GROUP:

- **supernova neutrinos**
- **(non-standard) Dark Matter candidates:
(s)axions, gravitinos, sterile neutrinos,...**
- **leptogenesis**
- **non-equilibrium QFT**
- **low energy neutrino phenomenology**
- **neutrino mass and flavour model building**
- **...**

1. Astroparticle Physics Group

RESEARCH DONE IN THE ASTROPARTICLE GROUP:

- supernova neutrinos
- (non-standard) Dark Matter candidates:
(s)axions, gravitinos, **sterile neutrinos**,...
- leptogenesis
- non-equilibrium QFT
- low energy neutrino phenomenology
- **neutrino mass and flavour model building**
- ...

I will concentrate on those two subjects.

At MPP, I work with 2 talented students:

↑
very



2. Sterile neutrino Dark Matter

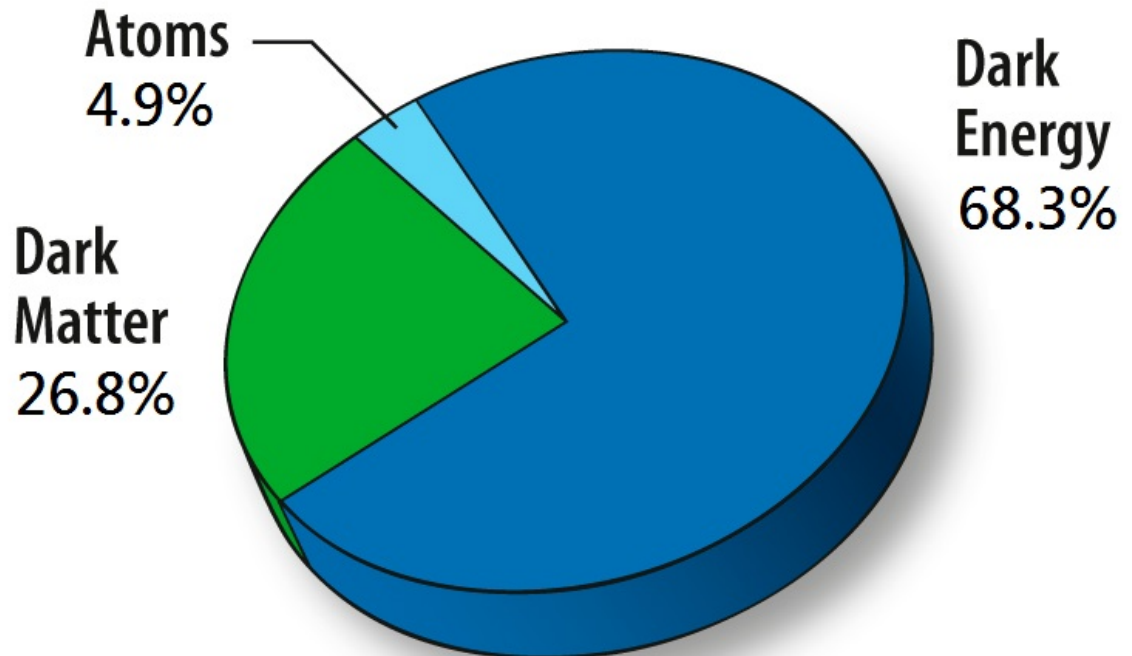
(work done in collaboration with Max Totzauer)

We know from various observations:

There is about 5 times more “Dark Matter” in the Universe than ordinary matter (the stuff we consist of).

BUT:

We don't know what it is...

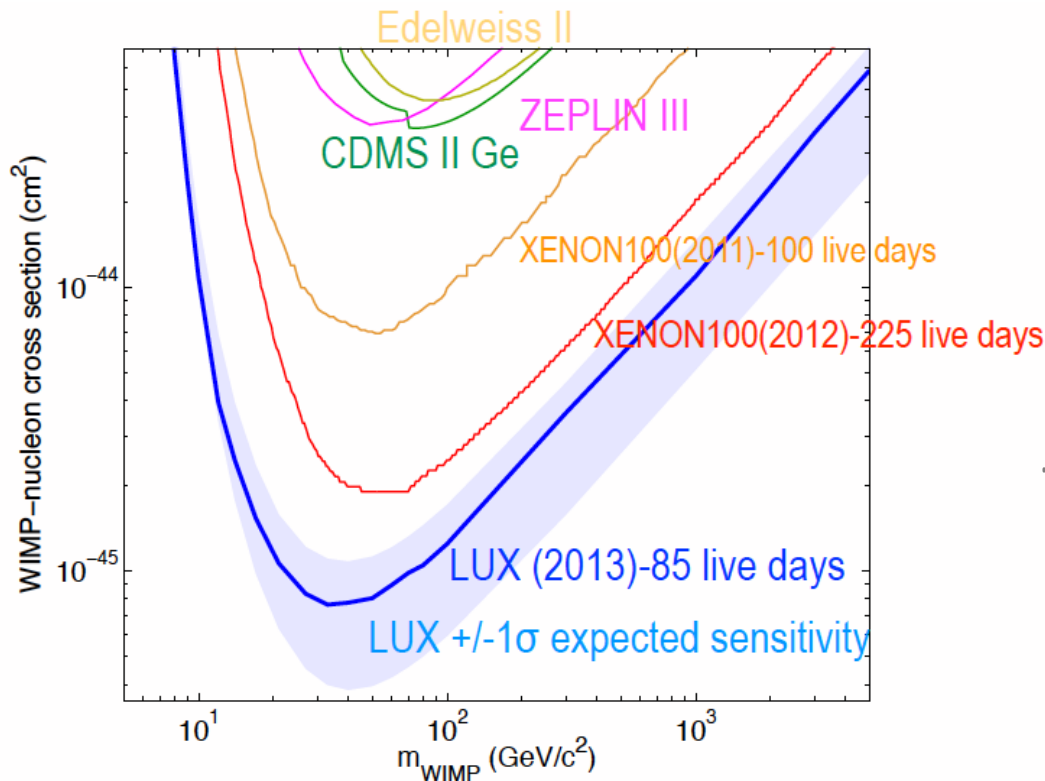


TODAY

2. Sterile neutrino Dark Matter

(work done in collaboration with Max Totzauer)

We had a very good guess for which particles could be behind the Dark Matter...



[McKinsey & Gaitskell:
Talk at Sanford
Research Facility, 30
October 2013]

BUT: They seem to WIMP out!!!

2. Sterile neutrino Dark Matter

(work done in collaboration with Max Totzauer)

A possible way out: **THINK OF ALTERNATIVES!**

- *axion, gravitino,...*
- *sterile neutrino with a mass of...*
 - *read here if you are an experimentalist: a few keV/c²*
 - *read here if you are a theorist: a few keV*
 - *read here if you are a string theorist: nearly zero*

in practice:

Like ordinary neutrino just a larger mass and suppressed interactions

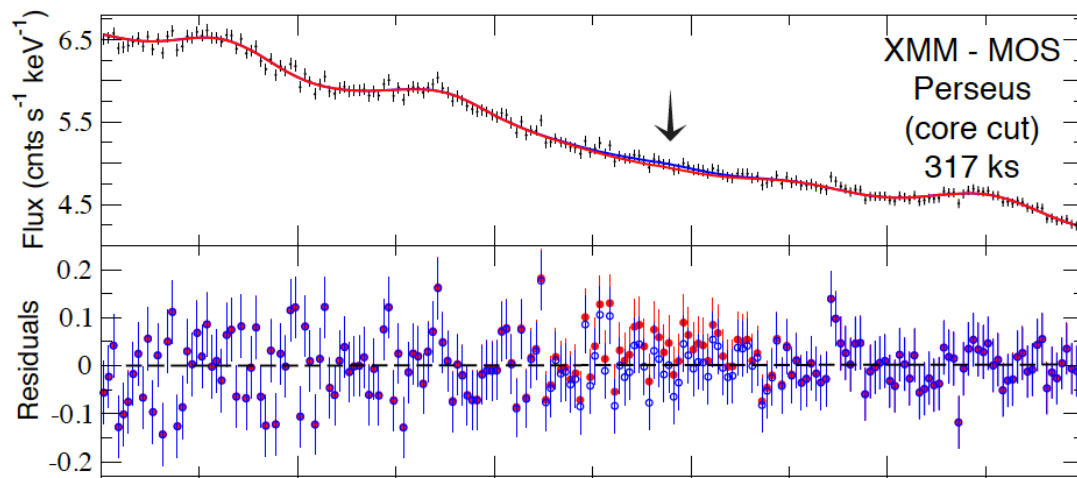
2. Sterile neutrino Dark Matter

(work done in collaboration with Max Totzauer)

Such a sterile neutrino could be the Dark Matter if:

- *it has the right mass and coupling*
- *we find a way to produce it in the early Universe*

This is a hot topic due to a (tentative, controversial, extensively discussed) signal seen earlier this year:



mass=7.1 keV?!?

(7.1 keV/c²?!?)

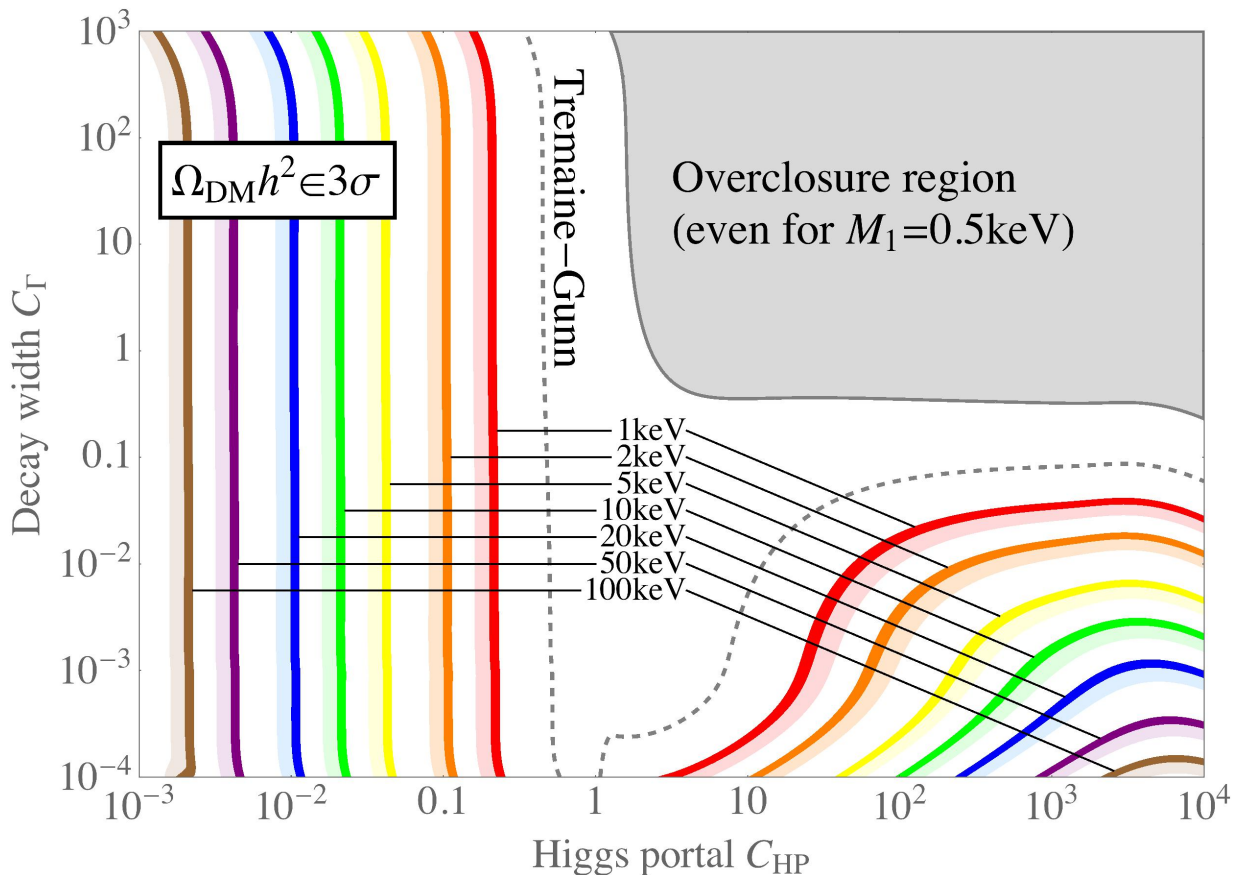
(nearly zero?!?)

Stay tuned...

2. Sterile neutrino Dark Matter

(work done in collaboration with Max Totzauer)

My work with Max: (non-thermal) production of keV sterile neutrinos by decays of e.g. scalars



We're not afraid to look into the details!

3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

We know that neutrinos have a tiny mass (<1 eV):

$$|m_{ee}| < 0.3-0.6 \text{ eV [KamLAN-Zen: Phys. Rev. C85 (2012) 045504]}$$

$$|m_{ee}| < 0.140-0.380 \text{ eV [EXO-200: Phys. Rev. Lett. 109 (2012) 032505]}$$

$$|m_{ee}| < 0.300-0.710 \text{ eV [CUORECINO: Astropart. Phys. 34 (2011) 822-831]}$$

$$m_{\beta} < 2.3 \text{ eV [MAINZ, Eur. Phys. J. C40 (2005) 447-468]}$$

$$\Sigma < 0.23 \text{ eV [Planck+WP+highL+BAO, Astron. Astrophys. 571 (2014) A16]}$$



BUT WHY IS IT SO SMALL?!?

(<http://imprinttrainingcenter.blogspot.co.uk/2010/12/understanding-and-controlling-anger.html>)

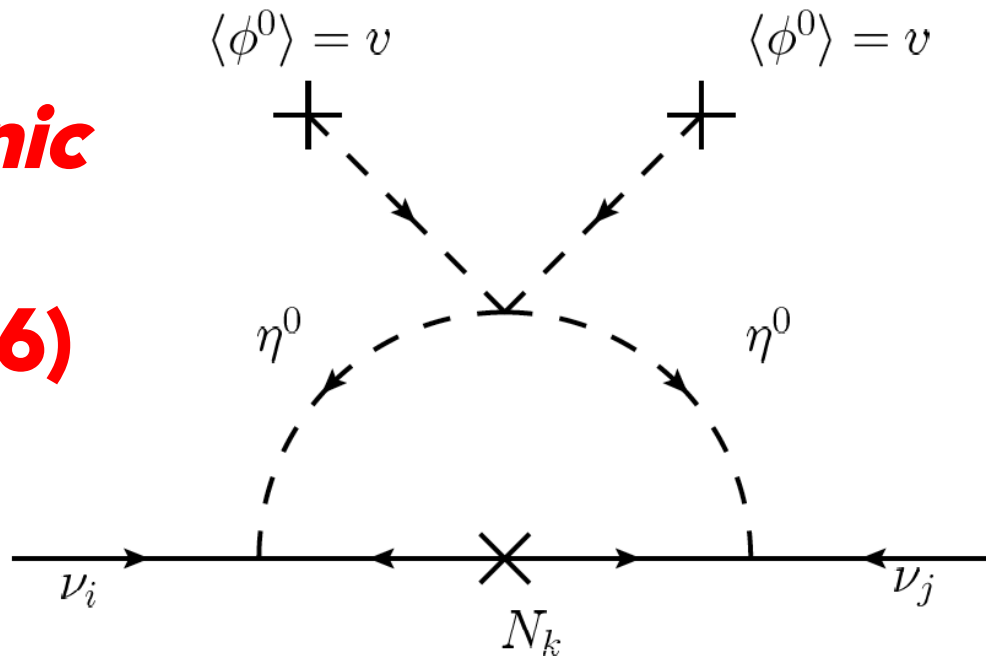
3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

A natural explanation:

Neutrinos have a **NATURAL MASS OF ZERO** but only **OBTAIN A NON-ZERO MASS** by a (small) **QUANTUM LOOP CORRECTION** induced by **NEW PHYSICS**

**Scotogenic
model
(Ma, 2006)**



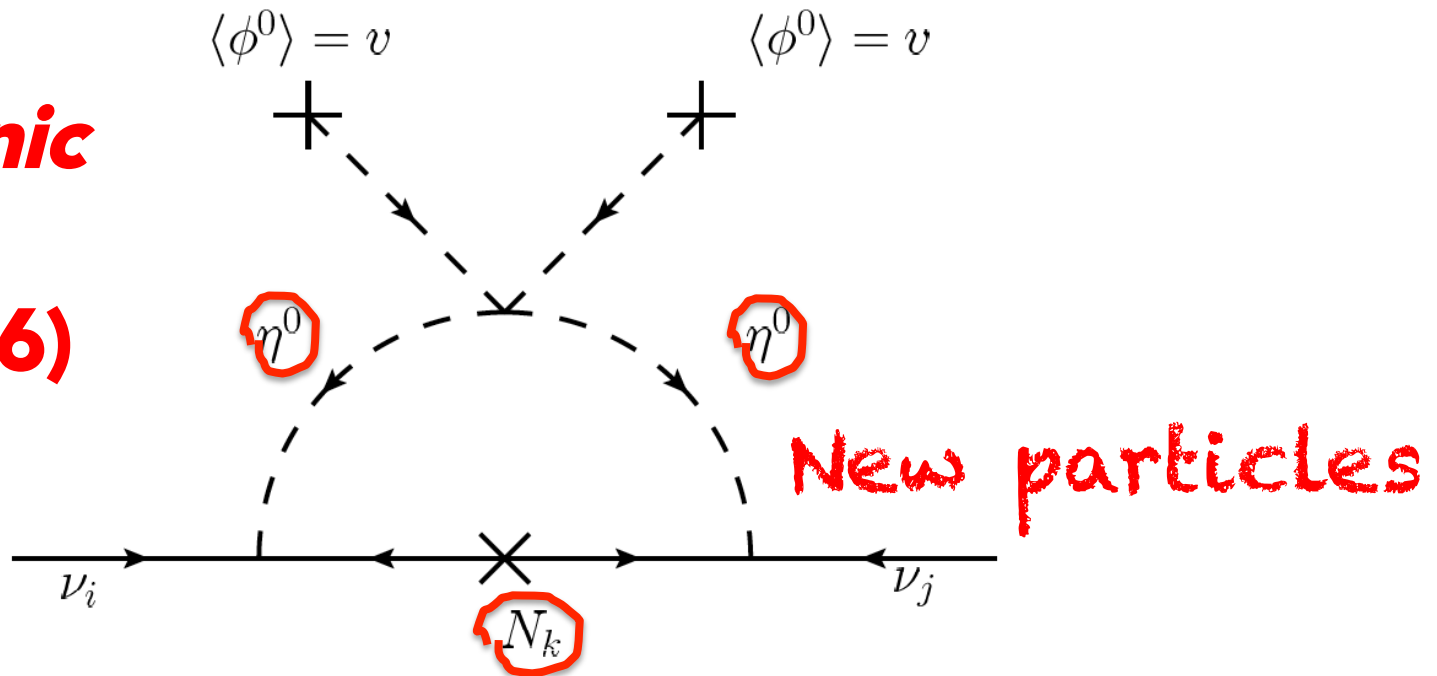
3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

A natural explanation:

Neutrinos have a **NATURAL MASS OF ZERO** but only **OBTAIN A NON-ZERO MASS** by a (small) **QUANTUM LOOP CORRECTION** induced by **NEW PHYSICS**

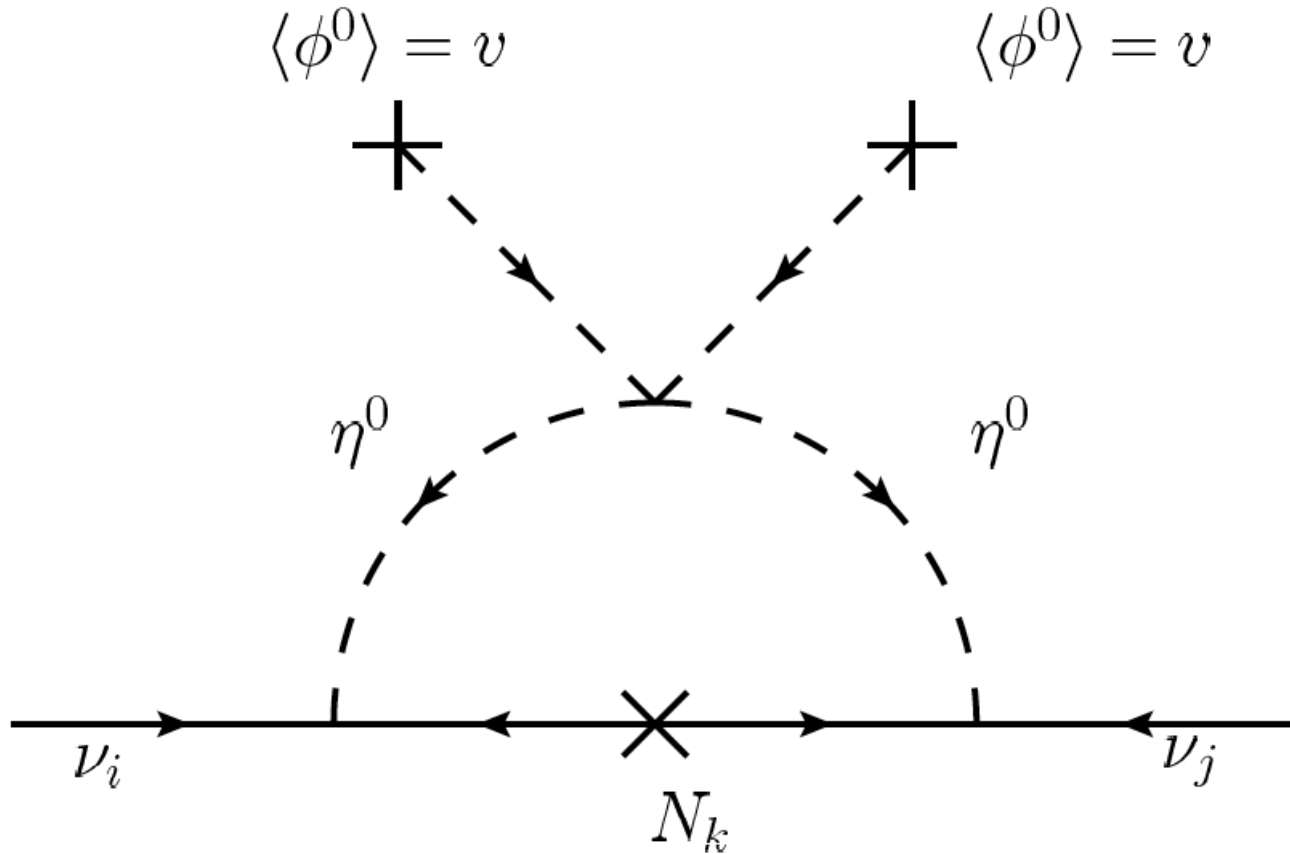
**Scotogenic
model
(Ma, 2006)**



3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

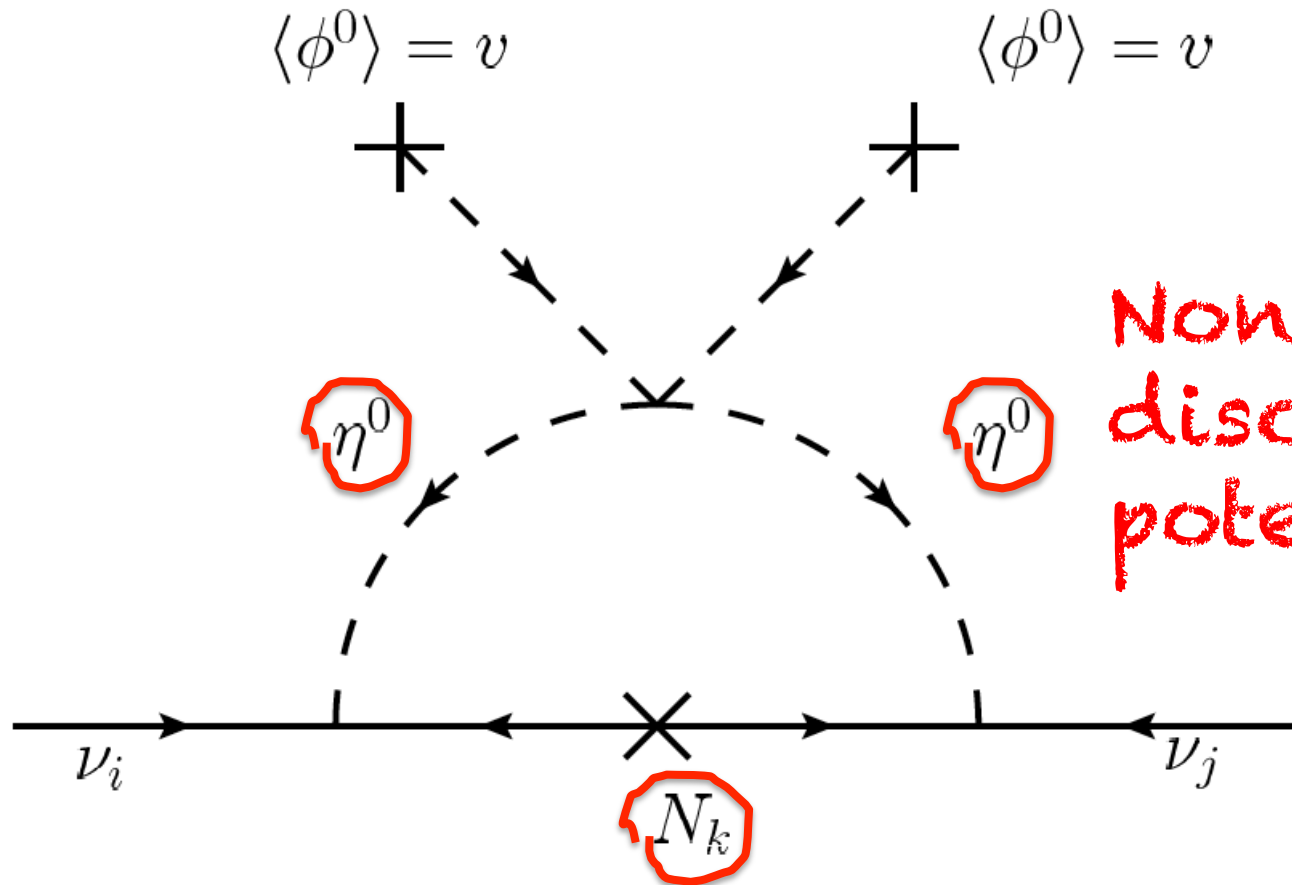
This stuff has many relations to other groups at MPP:



3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

This stuff has many relations to other groups at MPP:

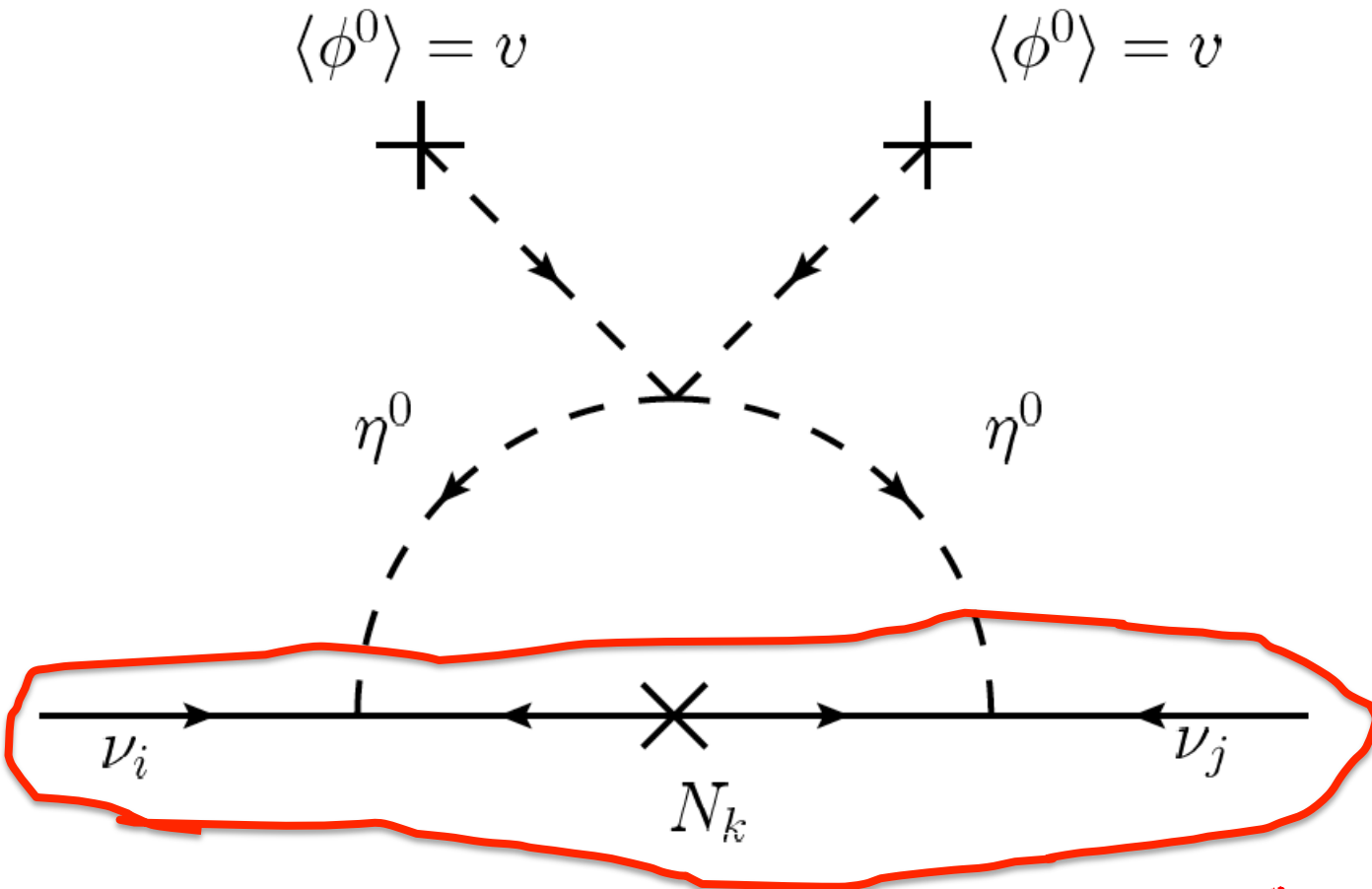


Non-negligible
discovery
potential at LHC

3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

This stuff has many relations to other groups at MPP:

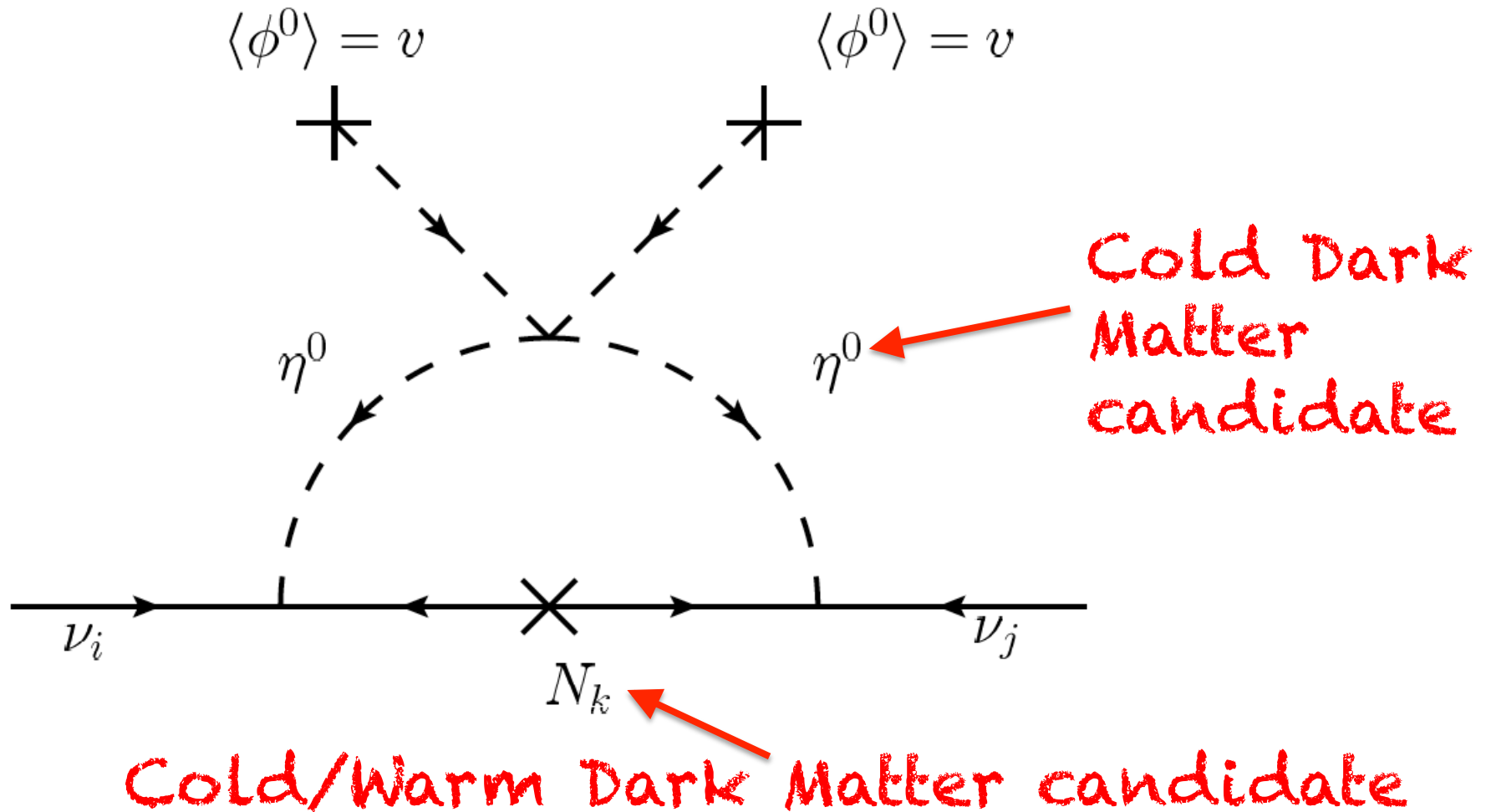


Majorana neutrinos hunted for in GERDA

3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

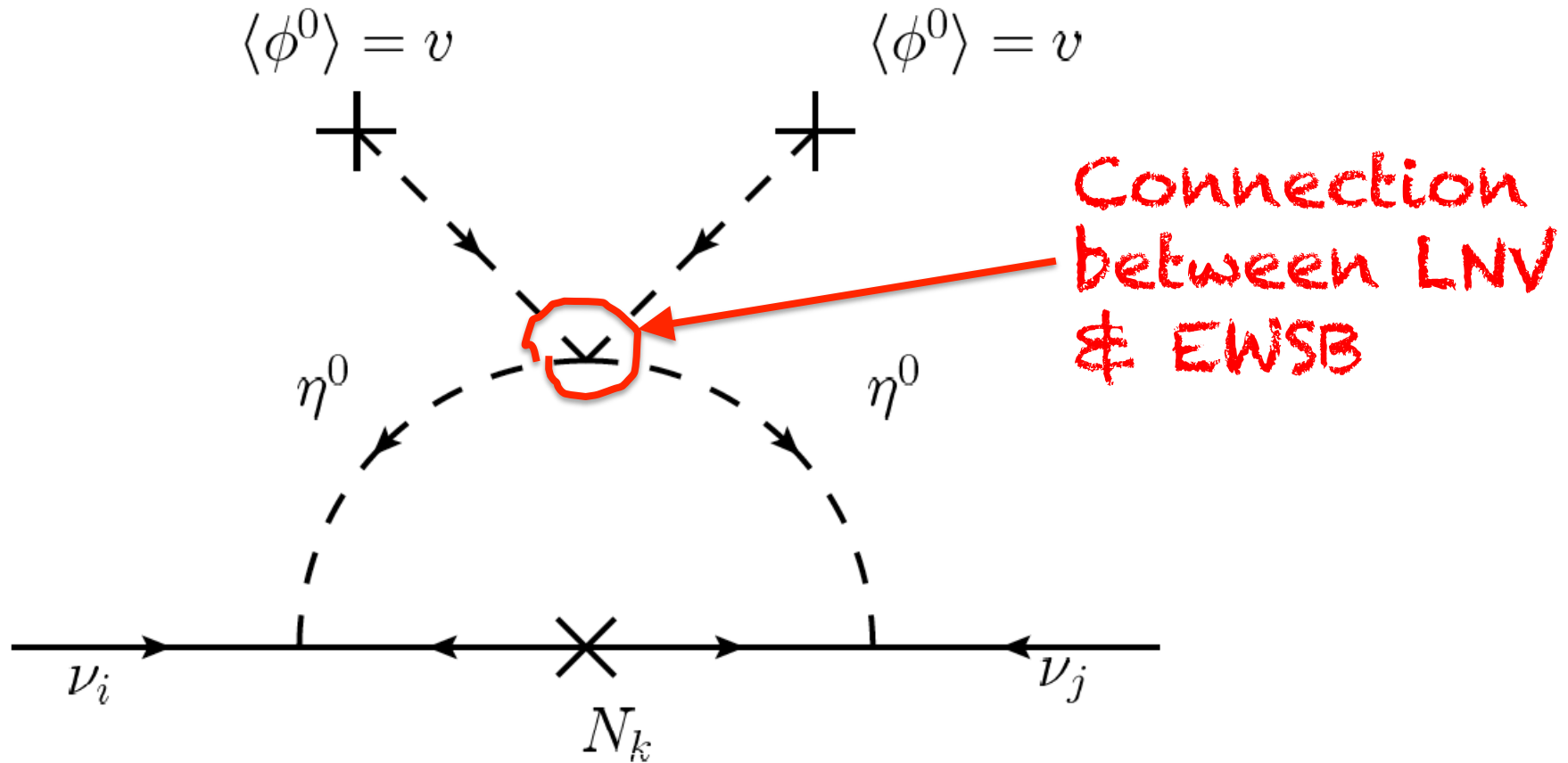
This stuff has many relations to other groups at MPP:



3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

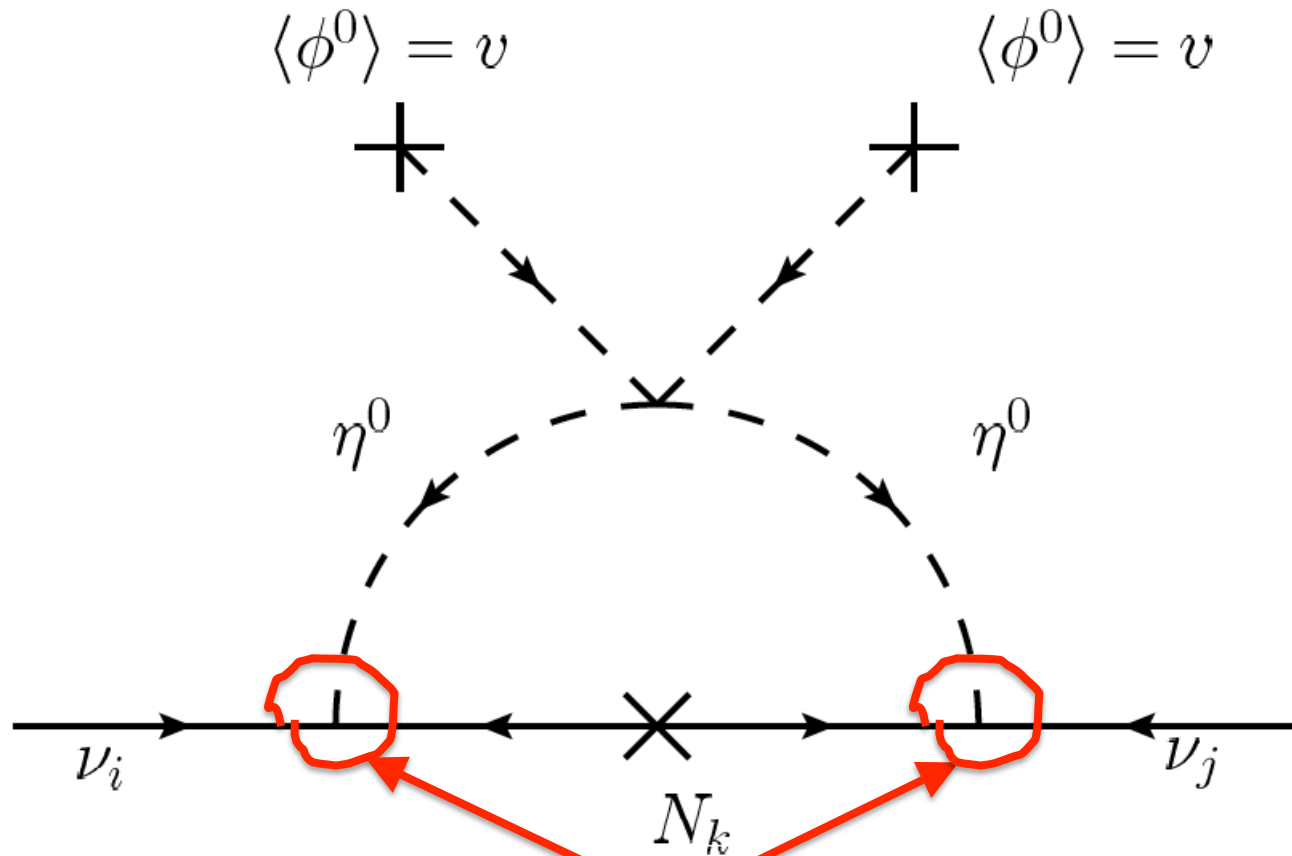
This stuff has many relations to other groups at MPP:



3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

This stuff has many relations to other groups at MPP:



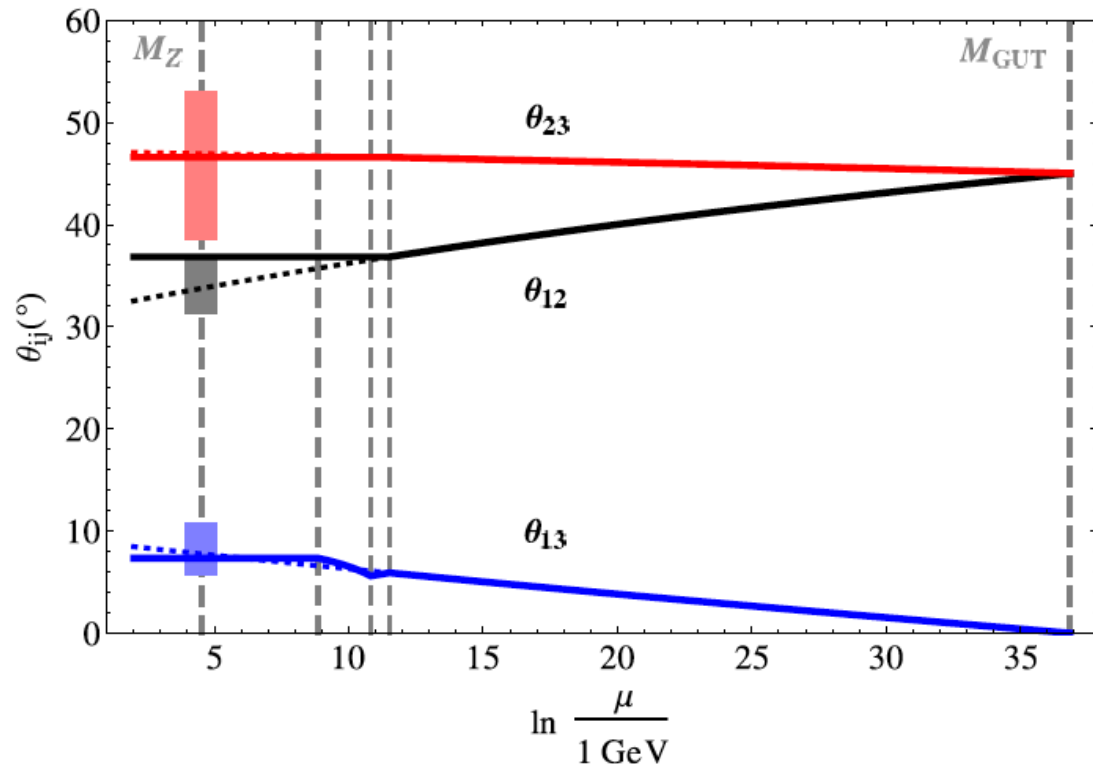
LFV at low and high energies

3. Radiative Neutrino Masses

(work done in collaboration with Moritz Platscher)

My work with Moritz: renormalisation of the scotogenic model & phenomenological investigation

Potentially strong running effects, not realised until 2012... → it's about time to put the model to the test!!!



[Bouchand, AM: JHEP 1207 (2012) 084]

4. Conclusions

- **there is a lot going on in the astroparticle physics group... sorry I couldn't cover everything!**
- **interesting times lie ahead of us... 2015 could bring a new physics discovery such as a clear Dark Matter signal or news from LHC**
- **but for now, the last thing I want to say is...**



**MERRY
X-MAS
& HAPPY
NEW
YEARS!!!**

