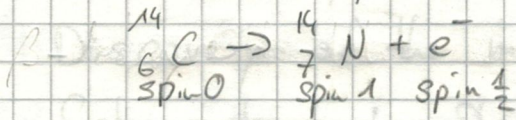


History

1896 Discovery of Radio activity

1911 Continuous spectrum of e^- in β -decay measured



\Rightarrow conservation laws of energy & momentum violated?

1920 \Rightarrow Big confusion among physicists at that time

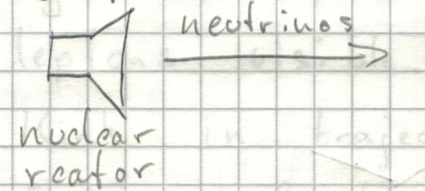
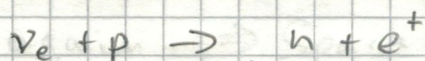
- even Bohr believed that conservation laws could be violated.

1930 New particle postulated by Pauli to solve this problem

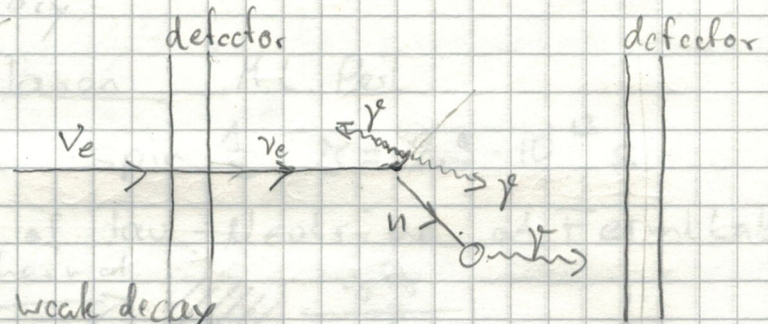
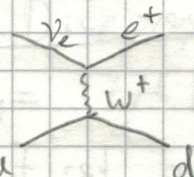
1933 Fermi names the new particle Neutrino

\rightarrow Fermi's theory of β -decay: neutrino assumed to be massless and chargeless

Cowan and Reines neutrino experiment 1956



tank with water containing cadmium chloride placed between large liquid scintillation detectors



- $\bar{\nu}_e$ decays into e^+ via weak decay

- e^+ slowed down in water and annihilated producing two gammas

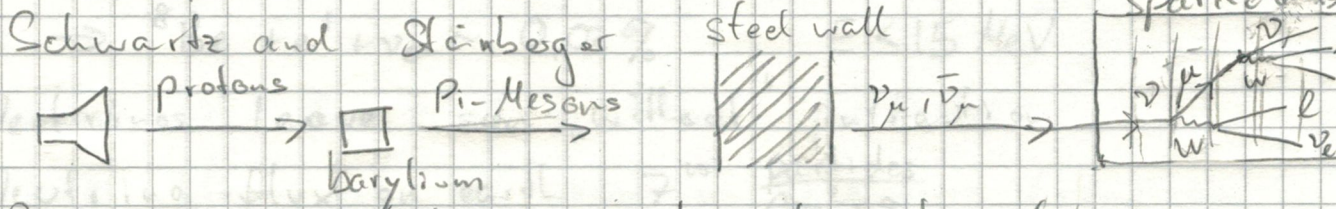
- the neutron is also slowed down and eventually captured by cadmium nucleus \rightarrow one gamma (reaches detector about a microsecond after the two gammas from positron annihilation)

Moon neutrino

1936 Moon discovered in Cosmic Radiation

$$m_\mu = 105 \text{ GeV} \quad q = -e \quad \text{spin } \frac{1}{2} \quad \tau = 2.2 \mu\text{s}$$

1962 Discovery of the moon neutrino by Lederman

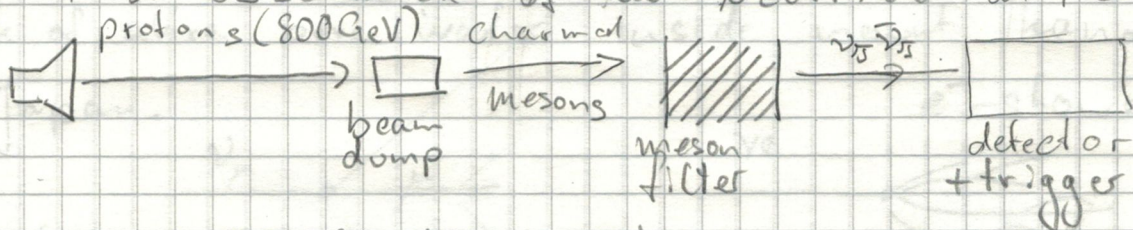


- Protons are shot onto a barium target.
- Pi Mesons are created which decay mainly into $\mu^- + \bar{\nu}_\mu$ and $\mu^+ + \nu_\mu$. Decay into electrons is suppressed by helicity.
- Muons are filtered out by steel wall.
- Neutrinos eventually decay into muons in spark chambers and then muons can decay into neutrinos again. \Rightarrow Spark chamber makes trajectories of leptons visible. Decay of muon can be seen as kink in trajectory.

1975 Discovery of the Tauon M.L. Perl

$$m_\tau = 1776 \text{ MeV} \quad q = -e \quad \text{spin } \frac{1}{2} \quad \tau = 2.9 \cdot 10^{-13} \text{ s}$$

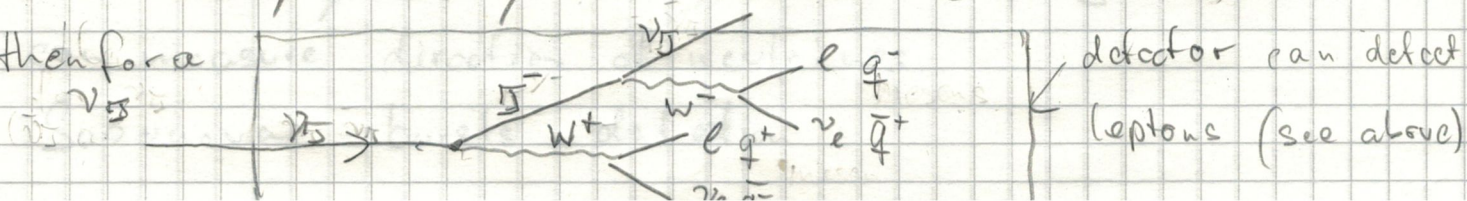
2000 First observation of Tau-Neutrino at Fermi Lab



\rightarrow protons get shot on beam dump (cooled block of metal because protons are very energetic)

\rightarrow Charmed mesons eg. D^+ are created which decay mainly into Tauons (see above)

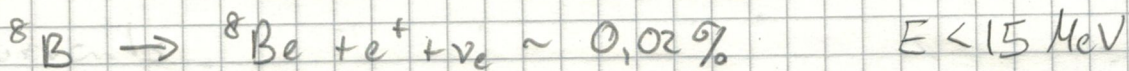
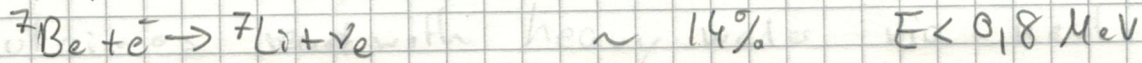
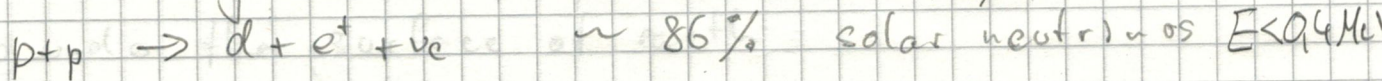
therefore



detector can detect leptons (see above)

Solar Neutrinos

Neutrinos get produced in the sun by nuclear fusion



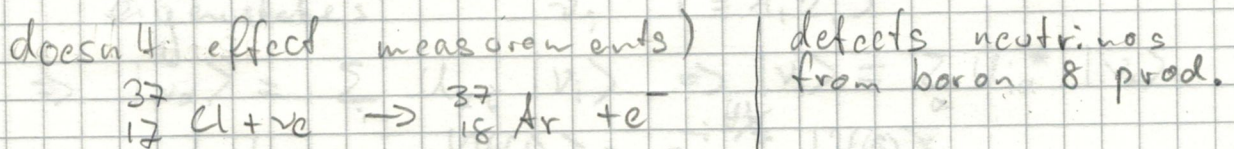
→ Neutrinos leave sun "without" interaction

→ Neutrino flux on earth 7^{10} particles $\text{cm}^{-2} \cdot \text{s}$

1968 R Davis found that there are less

neutrinos coming from the sun than there should be

⇒ Homestake experiment: tank filled with C_2Cl_4 placed about 1.5 km under the surface (so cosmic radiation doesn't effect measurements)



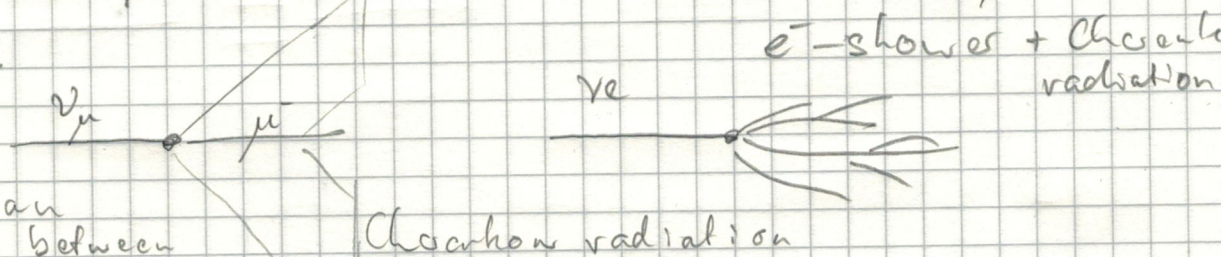
→ Radioactive Argon summons in tank

→ Measurement of radioactivity gives number of neutrino reactions

→ Davis found that only $1/3$ of the predicted neutrino flux from sun is measured.

Super Kamiokande

Tank of ultra pure water inside mount Kanoyama in Japan.



→ Detector can distinguish between ν_μ and ν_e

Cherenkov radiation

On the walls of the tank photon multipliers are placed

⇒ Measures ν_μ ν_e flux from sun → $\frac{2}{3}$ flux expected

⇒ can measure direction of neutrinos

Supernova bursts detection

SNO

Tank filled with heavy water also placed deep under the surface of earth.

Neutrinos react with heavy water via neutral currents.

⇒ Cherenkov radiation of scattered particles

⇒ Measures flux of all neutrino flavours from sun

In agreement with prediction!

Neutrino oscillations

flavour eigenstates ν_α $\alpha = e, \mu, \tau$

mass eigenstates ν_i $i = 1, 2, 3$

$$\Rightarrow |\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle \Rightarrow |\nu_i\rangle = \sum_\alpha U_{i\alpha}^* |\nu_\alpha\rangle$$

⇒ time evolution $|\nu(t)\rangle = e^{-i\frac{Ht}{\hbar}} |\nu(0)\rangle$

⇒ probability of finding a neutrino in some other flavour eigenstate after time t :

$$|\langle \nu_\beta(t) | \nu_\alpha(0) \rangle|^2$$

$$\begin{aligned} \Rightarrow \langle \nu_\beta(t) | \nu_\alpha(0) \rangle &= \langle \nu_\beta(0) | e^{-i\frac{Ht}{\hbar}} | \nu_\alpha(0) \rangle \\ &= \sum_{i,j} U_{i\beta}^* U_{\alpha j} e^{-i(\frac{E_i t}{\hbar} - p_j x)} \langle \nu_i | \nu_j \rangle \\ &= \sum_i U_{i\beta}^* U_{\alpha i} e^{-i(\frac{E_i t}{\hbar} - p_i x)} \end{aligned}$$