

Gauge-gravity duality and high-energy collisions

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New Trends in HERA Physics 2008
Ringberg Castle, Tegernsee

- Gauge/Gravity correspondence and AdS/CFT
An Introduction
- Gravity dual of an expanding medium
A medium: Hydrodynamics and the sQGP
- Gravity dual of a high-energy collision
Scattering at strong gauge coupling
- Prospects on QCD & AdS/CFT (tentative)
A growing subject

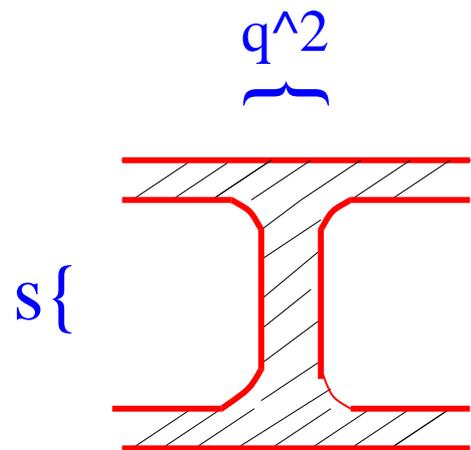
Strong Interactions and Strings

Historical Remark

Strings \nleftrightarrow QCD \leftrightarrow Strings

1968 \Rightarrow 1974 \Rightarrow 1998 \Rightarrow 2008 ...

Open String

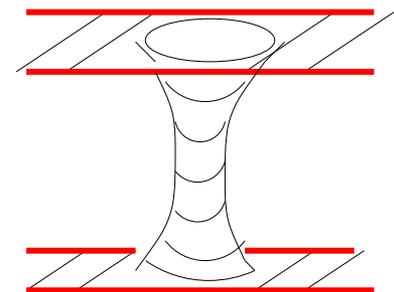


Veneziano Amplitude

$$A_R(s, q^2)$$

Gauge

Closed String



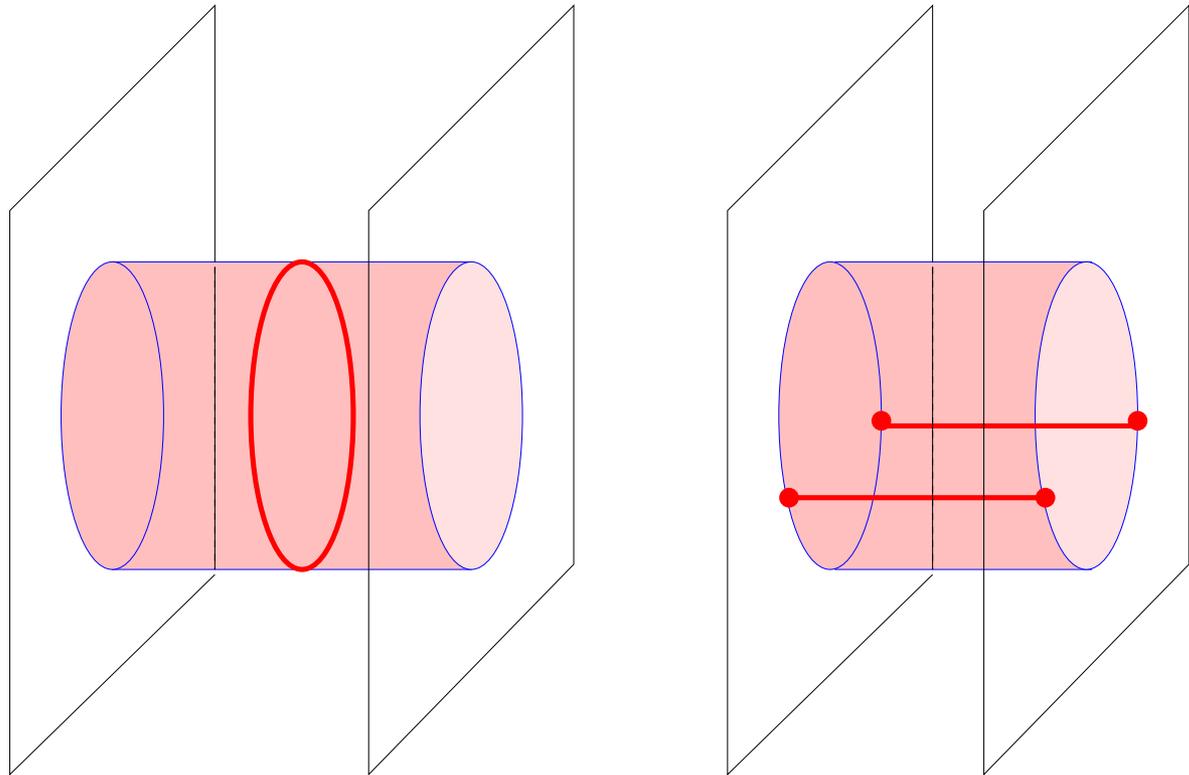
Shapiro-Virasoro Amplitude

$$A_P(s, q^2)$$

Gravity

The Gauge-Gravity Correspondence

“Duality”: Open String \Leftrightarrow Closed String



Schomerus, 2006

Closed String \Leftrightarrow *1 – loop Open String*

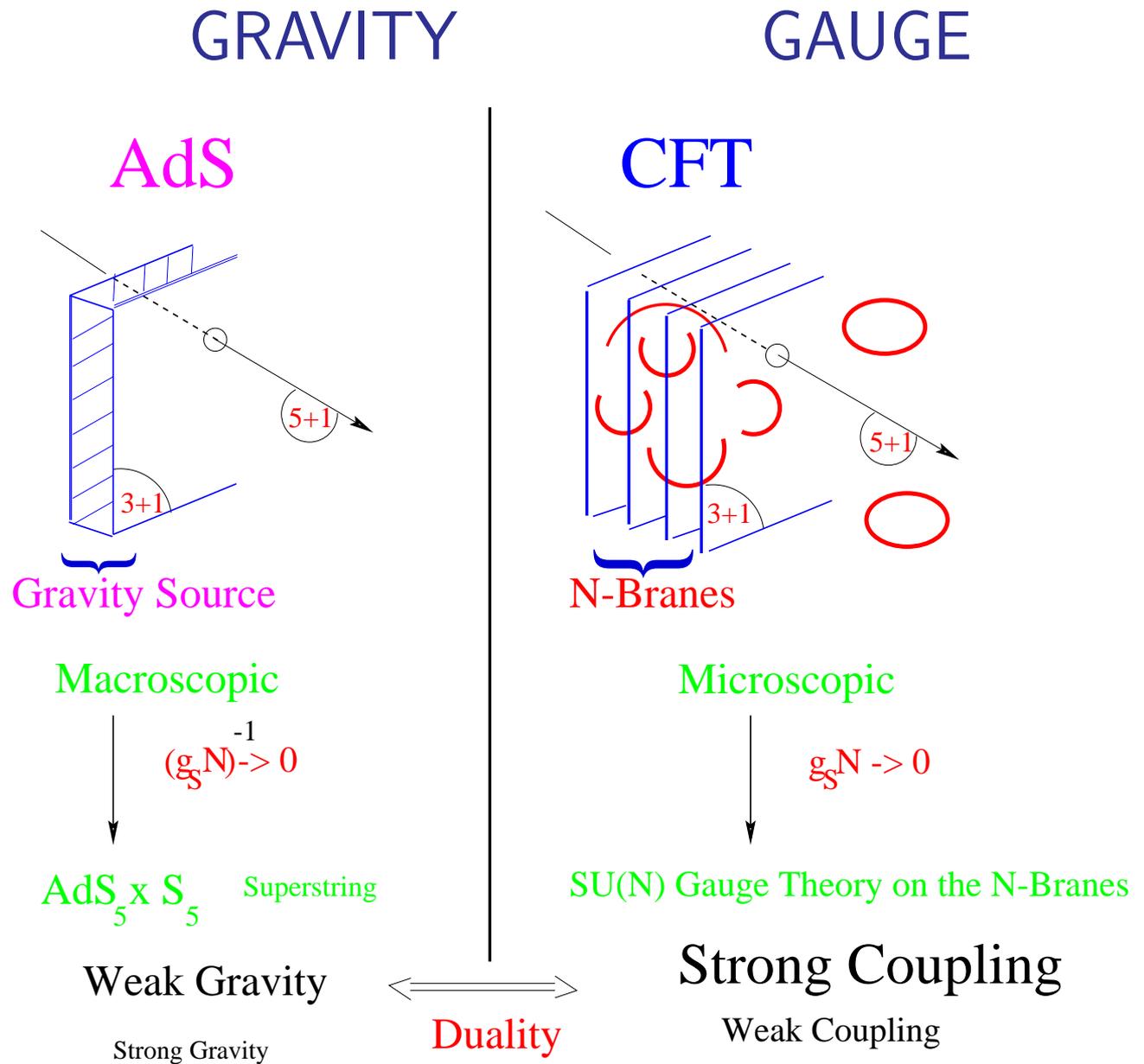
D – Brane “Universe” \Rightarrow *Open String Ending*

Gravity \Leftrightarrow *Gauge*

Large/Small Distance \Rightarrow *Gravity/Gauge Correspondence*

AdS/CFT Correspondence

J. Maldacena (1998)



WHY $AdS_5 \otimes S_5$?

- Solution of Gravity for D_3 Branes: Horowitz, Strominger, 1991

$$ds^2 = f^{-1/2}(-dt^2 + \sum_1^3 dx_i^2) + f^{1/2}(dr^2 + r^2 d\Omega_5)$$

“Physical” Branes (d=1+3) + Extra-Dimensions (d=6)

$$f = 1 + \frac{R^4}{r^4} ; R^4 = 4\pi\alpha'^2 g_{YM}^2 N_c$$

- “Maldacena breakthrough” : Maldacena, 1998

$$\frac{\alpha'(\rightarrow 0)}{r(\rightarrow 0)} \rightarrow z , R \text{ fixed} \Rightarrow g_{YM}^2 N_c \rightarrow \infty$$

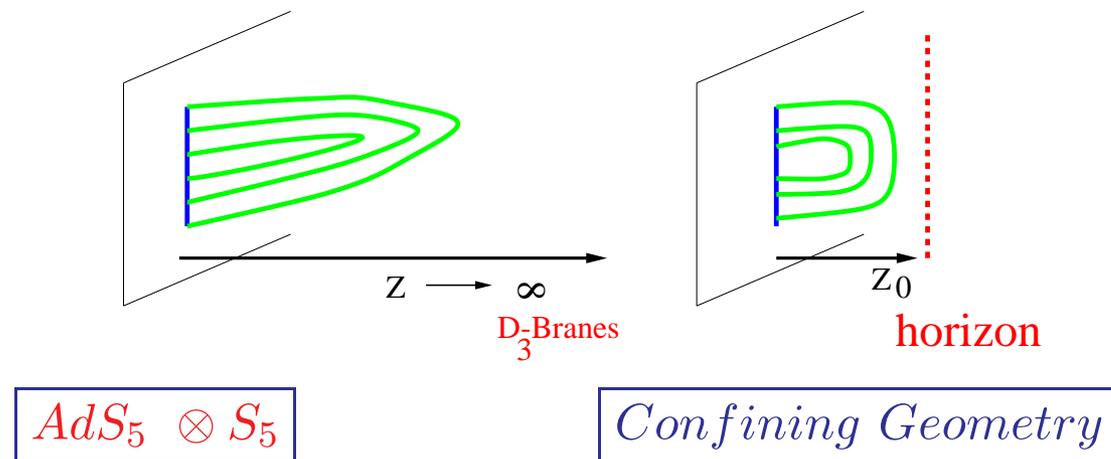
Strong coupling limit

$$ds^2 = \frac{1}{R^2 z^2}(-dt^2 + \sum_{1-3} dx_i^2 + dz^2) + R^2 d\Omega_5$$

Background Structure: $AdS_5 \otimes S_5$ (same R^2)

HOLOGRAPHY

- Holographic Principle: Brane/Bulk correspondence



- Brane \rightarrow Bulk: Holographic Renormalization

K. Skenderis (2002)

$$ds^2 = \frac{g_{\mu\nu}(z) dx^\mu dx^\nu + dz^2}{z^2}$$

$$g_{\mu\nu} = g_{\mu\nu}^{(0)} (= \eta_{\mu\nu}) + z^2 g_{\mu\nu}^{(2)} (= 0) + z^4 \langle T_{\mu\nu} \rangle + z^6 \dots +$$

$+ z^6 \dots +$: from Einstein Eqs.

EMERGENCE of the 5d BLACK HOLE

Balasubramanian, de Boer, Minic (2002)

- 4d Perfect Fluid “on the brane”

$$\langle T_{\mu\nu} \rangle \propto g_{\mu\nu}^{(4)} = \begin{pmatrix} 3/z_0^4 = \epsilon & 0 & 0 & 0 \\ 0 & 1/z_0^4 = p_1 & 0 & 0 \\ 0 & 0 & 1/z_0^4 = p_2 & 0 \\ 0 & 0 & 0 & 1/z_0^4 = p_3 \end{pmatrix}$$

- Holographic Renormalisation (Resummed) Janik, R.P. (2005)

$$ds^2 = -\frac{(1 - z^4/z_0^4)^2}{(1 + z^4/z_0^4)z^2} dt^2 + (1 + z^4/z_0^4) \frac{dx^2}{z^2} + \frac{dz^2}{z^2}$$

- \Rightarrow 5d Black Brane with horizon at $z_0 \sim T_0^{-3}$

$$ds^2 = -\frac{1 - \tilde{z}^4/\tilde{z}_0^4}{\tilde{z}^2} dt^2 + \frac{dx^2}{\tilde{z}^2} + \frac{1}{1 - \tilde{z}^4/\tilde{z}_0^4} \frac{d\tilde{z}^2}{\tilde{z}^2}$$

$$z \rightarrow \tilde{z} = z / \sqrt{1 + \frac{z^4}{z_0^4}}$$

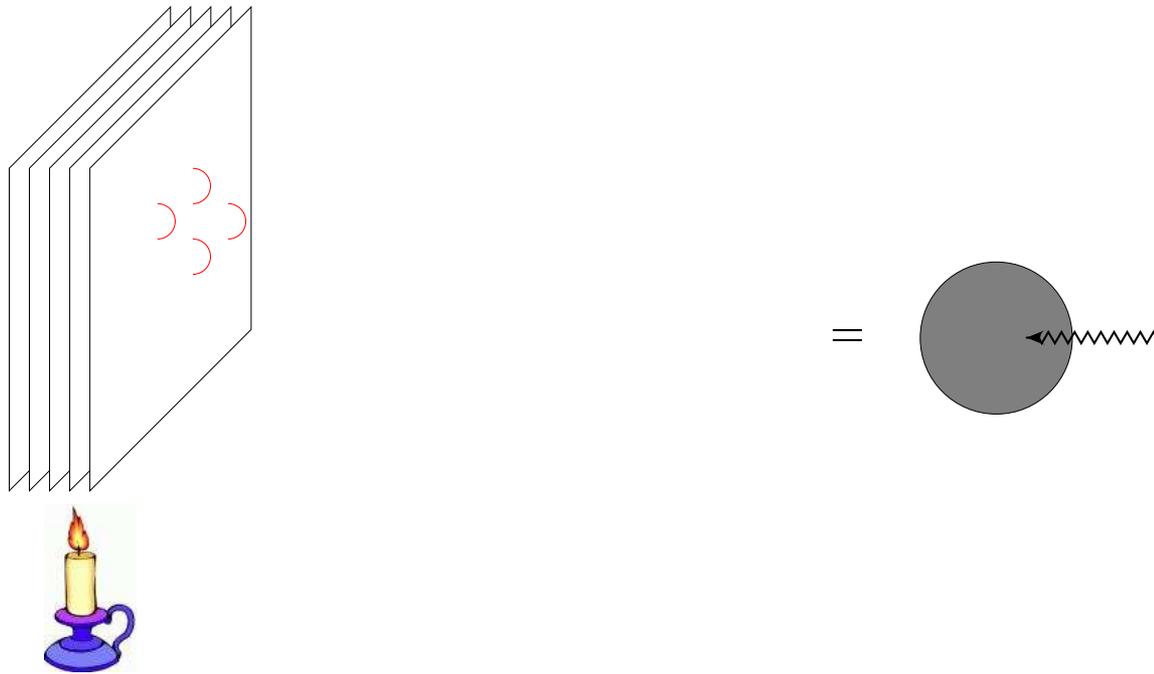
HYDRODYNAMICS *vs.* GRAVITY

Viscosity on the light of duality

Consider a graviton that falls on this stack of N D3-branes

Will be absorbed by the D3 branes.

The process of absorption can be looked at from two different perspectives:



Absorption by D3 branes (\sim viscosity) = absorption by black hole

$$\sigma_{abs}(\omega) \propto \int d^4x \frac{e^{i\omega t}}{\omega} \langle [T_{x_2x_3}(x), T_{x_2x_3}(0)] \rangle \Rightarrow \frac{\eta}{s} \equiv \frac{\sigma_{abs}(0)/(16\pi G)}{A/(4G)} = \frac{1}{4\pi}$$

AdS/CFT correspondence and the Quark-Gluon Plasma

Policastro, Son, Starinets (2001)

Gauge/Gravity: From QGP Statics to Dynamics

Janik, R.P.

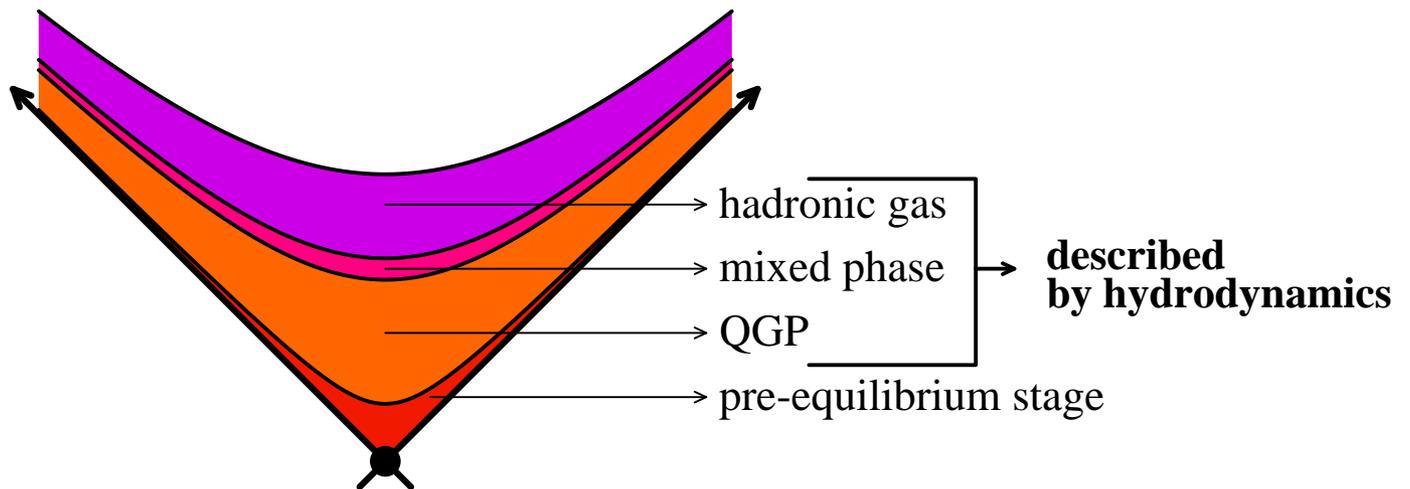
Janik, Heller, Benincasa, Buchel...

Kovchegov, Taliotis, Albacete,...

Nakamura, Sin, Kinoshita, Mukoyama, Nakamura, Oda, Natsuume, Okamura,...

Bhattacharyya, Hubeny, Minwalla, Ranganami, Loganayagam,...

.....



$$\tau = \sqrt{x_0^2 - x_1^2} ; y = \frac{1}{2} \log \frac{x_0 + x_1}{x_0 - x_1} ; x_T = x_2, x_3$$

Questions

- What is the Gravity Dual of a Flow?
- QGP: (almost) Perfect fluid behaviour, why?
- Universal $\frac{\eta}{S}$, Transport coefficients, Navier-Stokes,...
- Fast Pre-equilibrium stage, why?

AdS/CFT \Rightarrow Perfect Fluid at large τ

R.Janik, RP (2005)

- Boost-invariant T_{ν}^{μ} (Bjorken, 1983)

$$T_{\mu\nu} = \begin{pmatrix} f(\tau) & 0 & 0 & 0 \\ 0 & -\tau^3 \frac{d}{d\tau} f(\tau) - \tau^2 f(\tau) & 0 & 0 \\ 0 & 0 & f(\tau) + \frac{1}{2} \tau \frac{d}{d\tau} f(\tau) & 0 \\ 0 & 0 & 0 & \dots \end{pmatrix}$$

- Proper-time evolution

$$f(\tau) \propto \tau^{-s} : T_{\mu\nu} t^{\mu} t^{\nu} \geq 0 \Rightarrow 0 < s < 4$$

$$f(\tau) \propto \tau^{-\frac{4}{3}} : \text{Perfect Fluid}$$

$$f(\tau) \propto \tau^{-1} : \text{Free streaming}$$

$$f(\tau) \propto \tau^{-0} : \text{Full Anisotropy } \epsilon = p_{\perp} = -p_L$$

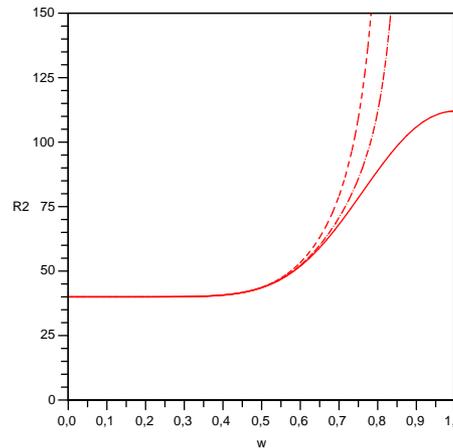
- Holographic renormalization:
 \Rightarrow Existence of a Dynamical Scaling

$$v = \frac{z}{\tau^{1/3}}$$

The Moving Black Hole in 5-d

- Holographic Renormalization:
 \Rightarrow Regularity Criterium

$$\mathfrak{R}^2 = R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$$



$$s = \frac{4}{3} \pm .1$$

Nonsingular Dual Geometry \Leftrightarrow Perfect Fluid

- Asymptotic metric
 \Rightarrow Black Hole (Brane) Moving off in the 5th dimension

$$\text{Horizon : } z_h(\tau) = \left(\frac{3}{e_0} \right)^{\frac{1}{4}} \cdot \tau^{\frac{1}{3}} .$$

$$\text{Temperature : } T(\tau) \sim \frac{1}{z_h} \sim \tau^{-\frac{1}{3}}$$

$$\text{Entropy : } S(\tau) \sim \text{Area} \sim \tau \cdot \frac{1}{z_h^3} \sim \text{const}$$

Some recent Results

- Going beyond perfect fluid

In-flow Viscosity, Relaxation time, Transport Coeff., etc...

Janik, Heller, Bak, Benincasa, Buchel, Nakamura, Sin,.....
Kinoshita, Mukoyama, Nakamura, Oda, Natsuume, Okamura,...

$$\partial_\tau \epsilon = -\frac{4}{3} \frac{\epsilon}{\tau} + \frac{\eta}{\tau^2} + \dots \Rightarrow \frac{\eta}{s} = \frac{1}{4\pi}$$

- Going beyond boost-invariance

General hydrodynamic equations from AdS/CFT

Bhattacharyya, Hubeny, Minwalla, Ranganami, Loganayagam,...

$$T_{rescaled}^{\mu\nu} = \underbrace{(\pi T)^4 (\eta^{\mu\nu} + 4u^\mu u^\nu)}_{\text{perfect fluid}} - \underbrace{2(\pi T)^3}_{\text{viscosity}} + \underbrace{(\pi T^2) \left(\log 2 T_{2a}^{\mu\nu} + 2T_{2b}^{\mu\nu} + (2 - \log 2) \left(\frac{1}{3} T_{2c}^{\mu\nu} + T_{2d}^{\mu\nu} + T_{2e}^{\mu\nu} \right) \right)}_{\text{second order hydrodynamics}}$$

- Going beyond hydrodynamics?

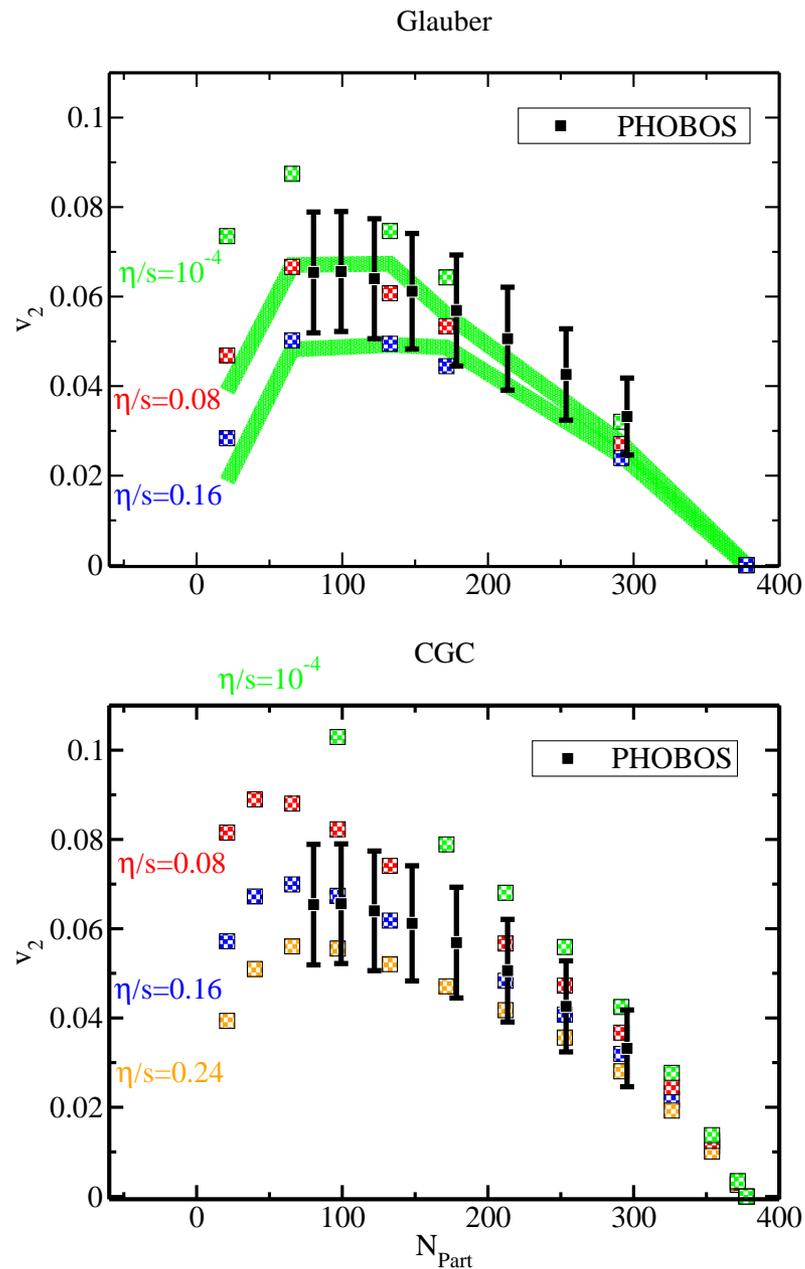
Physics in the expanding plasma: Fundamental flavours, mesons,...
Janik, Heller, Große, Surowka ...

Modeling heavy-ion collisions: see next...

Strings *vs.* Reality: Elliptic Flow

Ollitrault (1992)

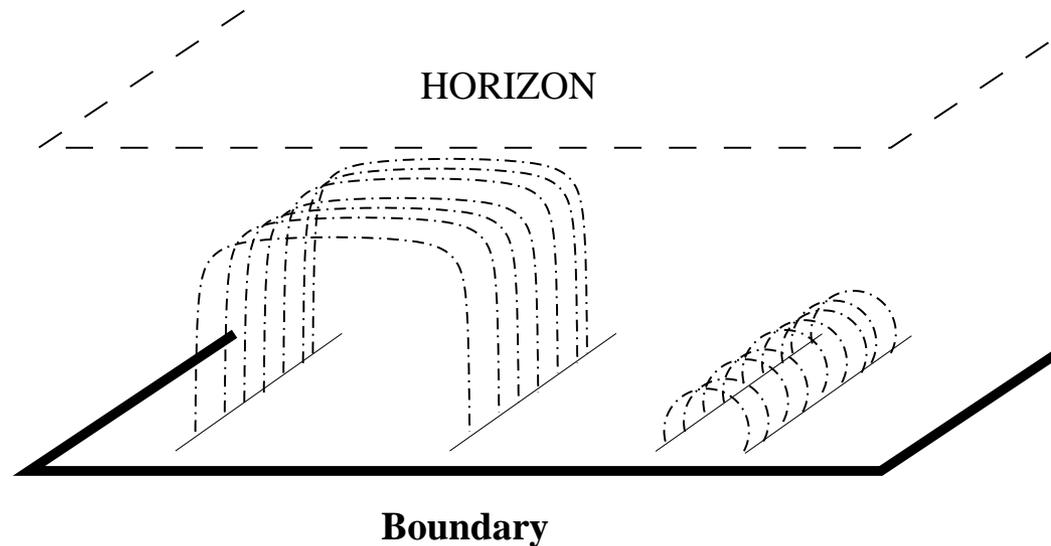
$$\frac{\partial N}{\partial \Phi} \propto 1 + 2 v_2 \cos 2\Phi$$



Luzum, Romatschke (2008)

Gravity dual of high-energy collisions

- **Conditions:**
Defining Probes and confining *vs.* conformal geometry
- **One Holographic method**
Wilson Lines \Rightarrow Minimal Surfaces



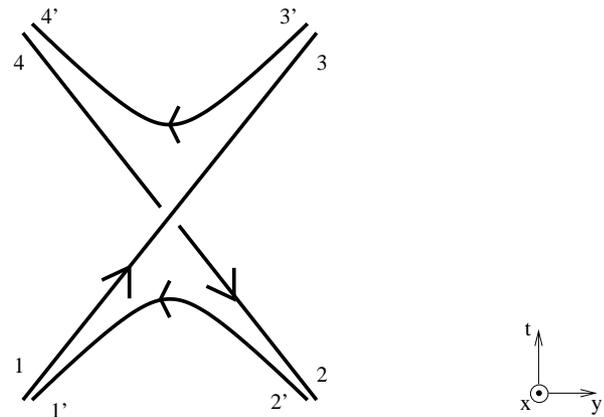
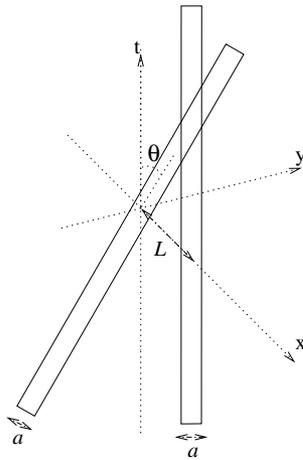
$$\langle e^{iP \int_C \vec{A} \cdot d\vec{l}} \rangle = \int_{\Sigma} e^{-\frac{\text{Area}(\Sigma)}{\alpha'}} \approx e^{-\frac{\text{Min. Area}}{\alpha'}} \times \text{Fluctuations}$$

- **Exemple: The Perimeter *vs.* Area law**
Conformal case : $\langle \text{Wilson Lines} \rangle = e^{T^*V(L)} \sim e^{T \times 1/L}$
Confining case : $\langle \text{Wilson Lines} \rangle = e^{T^*V(L)} \sim e^{T \times L}$

AMPLITUDES at STRONG COUPLING

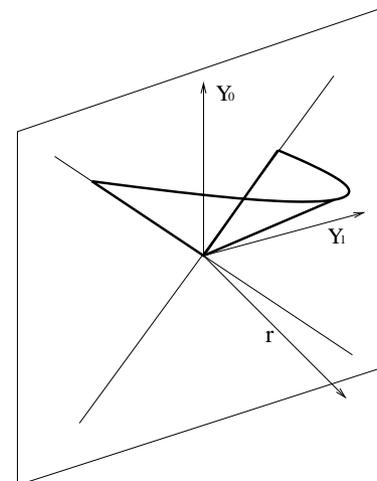
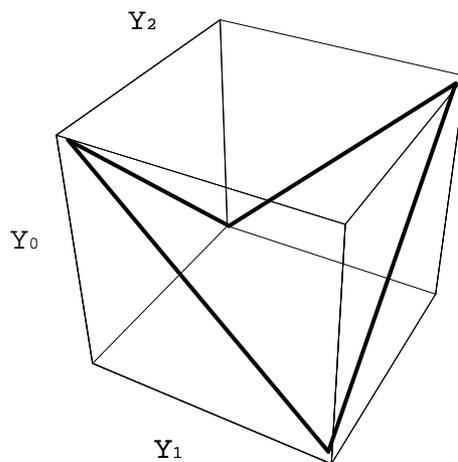
- SOFT AMPLITUDES

Regge trajectories: R.Janik, RP (2000)



- HARD AMPLITUDES

S^4 YM amplitudes: Alday, Maldacena (2007)



Prospects

In progress:

- **Gauge-Gravity Correspondence**
A promising way towards QCD at strong coupling
- **Results on AdS/CFT \rightarrow S^4 QCD Hydrodynamics**
Einstein *vs.* Navier-Stokes, Thermalization,...
- **Other studies**
Jet Quenching, Quark Dragging, and many others...

What in front of us?

- **More Theoretical Work**
Gauge/Gravity beyond AdS/CFT
- **More “Translation” Work**
Formulating Observables: V_2 , Diffraction, Soft/Hard Interplay
- **From S^4 QCD to S^0 QCD ?**
Construct the “Gravity Dual” of QCD

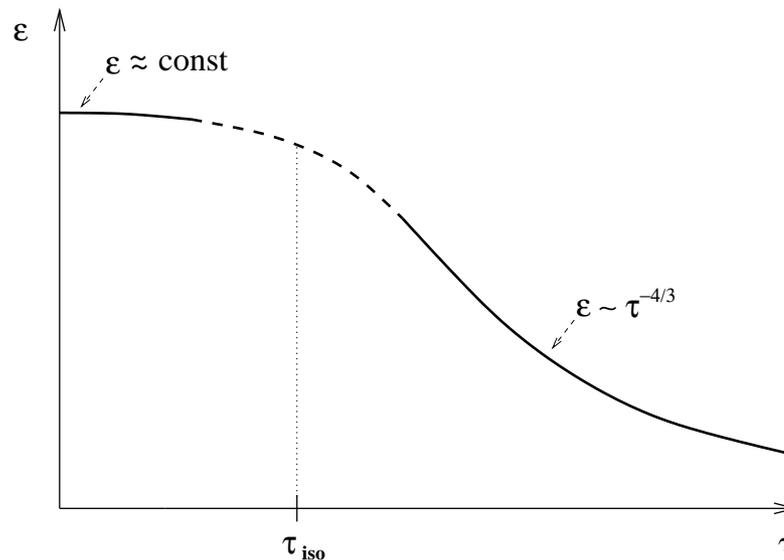
Why all that seems to work?

■

EXTRA SLIDES

AdS/CFT: Anisotropy at small τ

Kovchegov, Taliotis arXiv:0705.1234



Evaluation of The Isotropization/Thermalization time

$$\text{Matching : } z_h^{\text{late}}(\tau) = \left(\frac{3}{e_0}\right)^{\frac{1}{4}} \equiv z_h^{\text{early}}(\tau) = \tau$$

$$\text{Isotropization : } \tau_{\text{iso}} = \left(\frac{3N_c^2}{2\pi^2 e_0}\right)^{3/8}$$

$$\text{Evaluation : } \epsilon(\tau) = e_0 \tau^{4/3} \Big|_{\tau=.6} \sim 15 \text{ GeV fermi}^{-3}$$

$$\Rightarrow \boxed{\tau_{\text{iso}} \sim .3 \text{ fermi}}$$