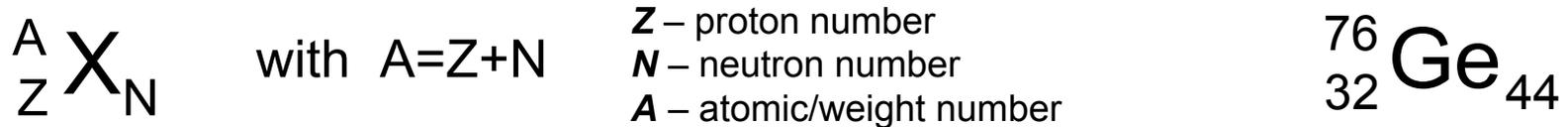


INTRODUCTION TO NUCLEAR MODELS

Daniel Kollar – Friday Physics Session

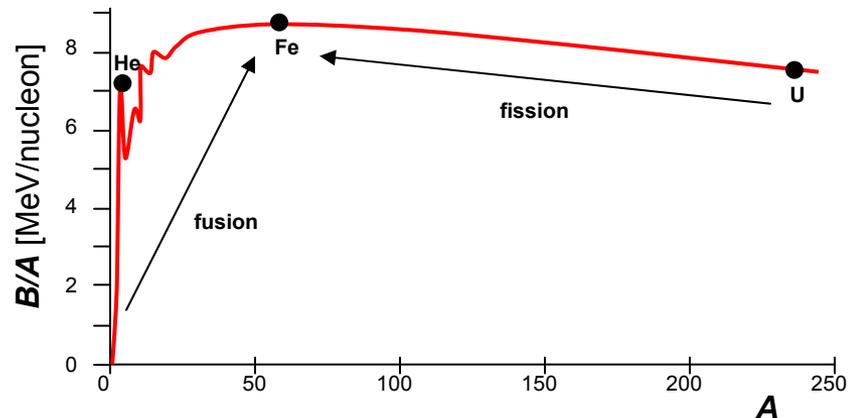
NUCLEUS BASICS



size: $R \cong R_0 A^{\frac{1}{3}}$ $R_0 \cong 1.2 \text{ fm}$ $R \cong 5.1 \text{ fm}$

mass: $m(X) < Zm_p + Nm_n$ (that's why it holds together)

binding energy: $B(Z, N) = [Zm_p + Nm_n - m(Z, N)] \cdot c^2$
 $B({}^{76}\text{Ge}) = [36m_p + 44m_n - m({}^{76}\text{Ge})] \cdot c^2 = 661.6 \text{ MeV}$



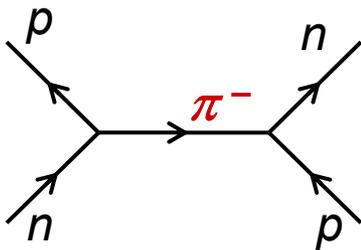
NUCLEAR FORCES

short range, spin-orbital character

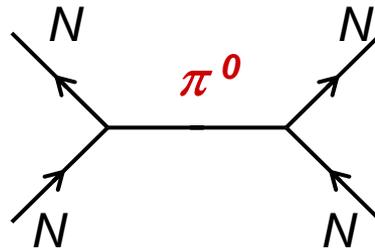
YUKAWA THEORY OF NUCLEAR FORCES

based on exchange of π^0 , π^+ , and π^-

$$p + n \rightarrow p + n$$



$$N + N \rightarrow N + N$$

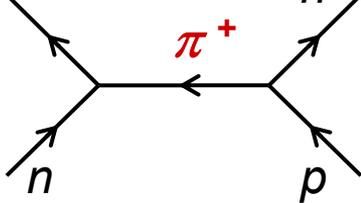


consider range of 1.4 fm
 \Rightarrow from uncertainty principle

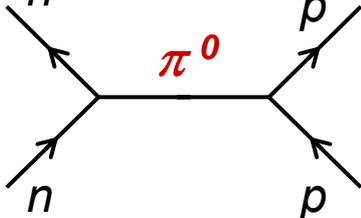
$$m c r \approx \hbar$$

$$\rightarrow m_{\pi} \approx 140 \text{ MeV}$$

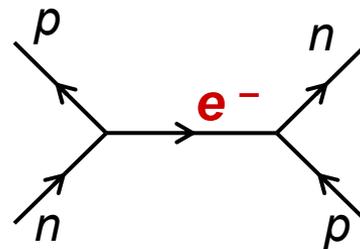
$$p + n \rightarrow n + p$$



$$n + n \rightarrow n + n$$

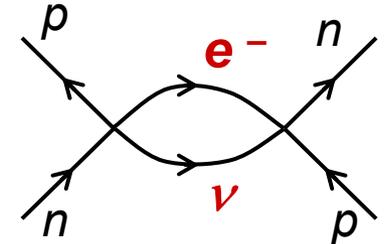


Heisenberg



violates angular
momentum conservation

Fermi



too weak

NUCLEAR MODELS

- total *wavefunction* of the nucleus is far too complicated to be useful even if it was possible to calculate it (only possible for the lightest nuclei)

⇒ we make use of models and use simple analogies

Types of nuclear models

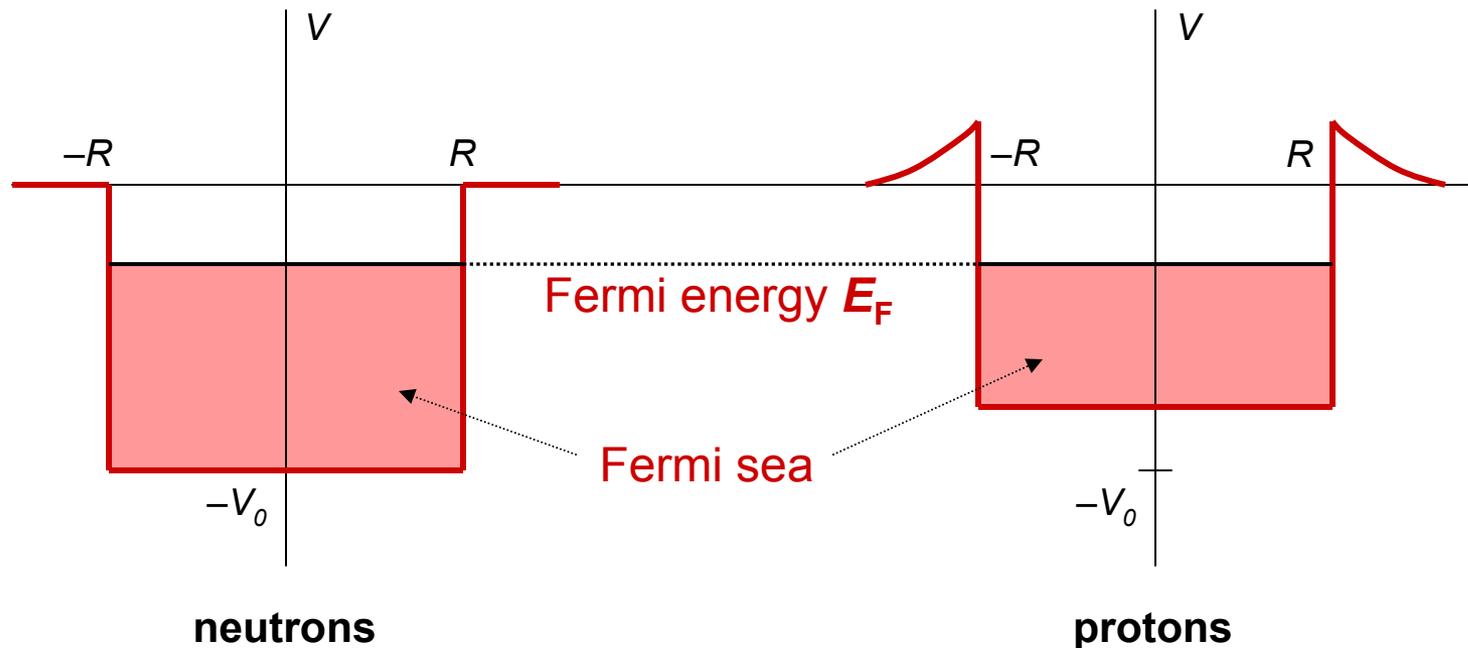
	Semiclassical	Quantum mechanical
Independent particle	Fermi gas	Shell
Collective	Liquid drop	Rotational Vibrational

FERMI GAS MODEL

Built on analogy between nucleus and ideal gas

- particles don't interact
- particles move independently in the mean field of the nucleus

Ground state → particles occupy lowest energy states allowed by the Pauli principle



FERMI MODEL

Distribution of nucleon momentum states: $dn = \frac{2Vd^3p}{(2\pi\hbar)^3}$

- total number of states up to E_F : $n = \int_0^{E_F} dn$
 - momentum \rightarrow energy
 - $n_p = Z$; $n_n = A - Z$
 - volume $\Rightarrow V = 4/3 \pi r_0^3 A$

- Fermi energy:

$$E_F^p = \frac{\hbar^2}{2mr_0^2} \left(\frac{9\pi}{4} \cdot \frac{Z}{A} \right)^{\frac{2}{3}}$$

$$E_F^n = \frac{\hbar^2}{2mr_0^2} \left(\frac{9\pi}{4} \cdot \frac{A-Z}{A} \right)^{\frac{2}{3}}$$

\rightarrow for $Z = A - Z = A/2$

$$E_F = \text{const} \cdot \frac{1}{m}$$



equal for all nuclei

$$E_F \approx 30 \text{ MeV} \quad (V_0 \approx 40 \text{ MeV})$$

NUCLEAR LIQUID DROP MODEL

Weizsäcker formula for the binding energy ($A \geq 30$)

$$B(A, Z) = a_V \cdot A$$

condensation energy $\propto V$
holding nucleus together

$$- a_S \cdot A^{\frac{2}{3}}$$

surface tension $\propto S$
near-surface nucleons are bound less

$$- a_C \cdot Z(Z - 1) \cdot A^{-\frac{1}{3}}$$

Coulomb potential

$$- a_A \cdot (A - 2Z)^2 \cdot A^{-1}$$

asymmetry

$$- \Delta \quad \Delta = \begin{cases} -\delta & \text{even-even} \\ 0 & \text{even-odd} \\ +\delta & \text{odd-odd} \end{cases}$$

pairing energy

$$\delta \propto A^{-\frac{1}{2}}$$

$$R \propto A^{\frac{1}{3}}$$

NUCLEAR LIQUID DROP MODEL

Weizsäcker formula for the mass of the nucleus

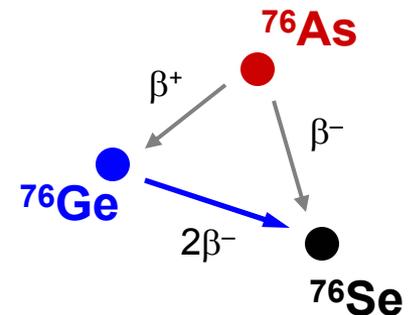
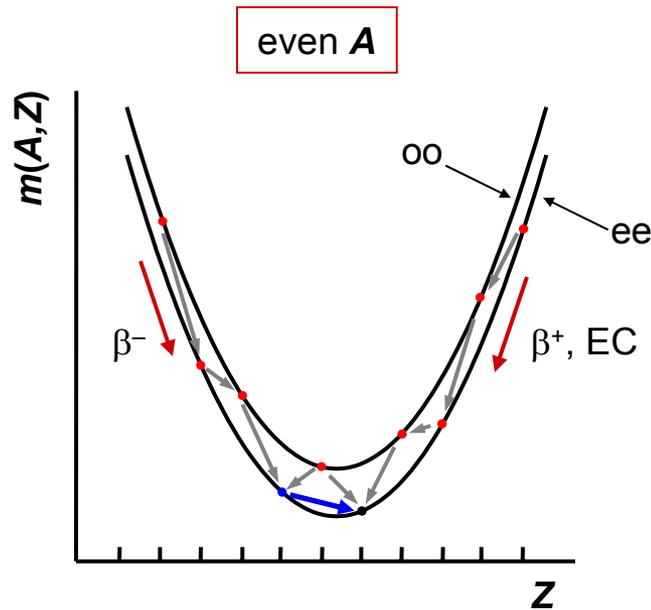
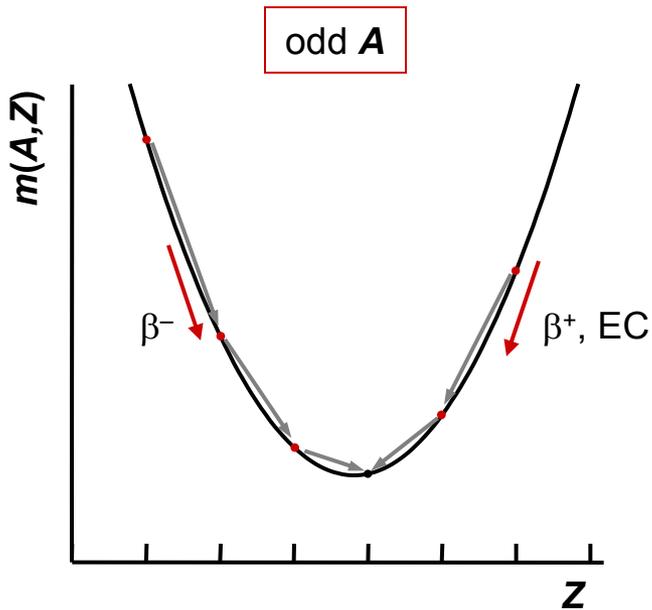
$$m(A, Z) = Zm_p + (A - Z)m_n - B(A, Z)$$

$$m(A, Z) = Zm_p + (A - Z)m_n - a_V A + a_S A^{\frac{2}{3}} + a_C Z(Z - 1)A^{-\frac{1}{3}} + a_A (A - 2Z)^2 A^{-1} + \Delta$$

for constant $A \Rightarrow m(A, Z)$ is quadratic in Z

$$\Delta = \begin{cases} -\delta & ee \\ 0 & oe \\ +\delta & oo \end{cases}$$

$\delta \propto A^{-\frac{1}{2}}$

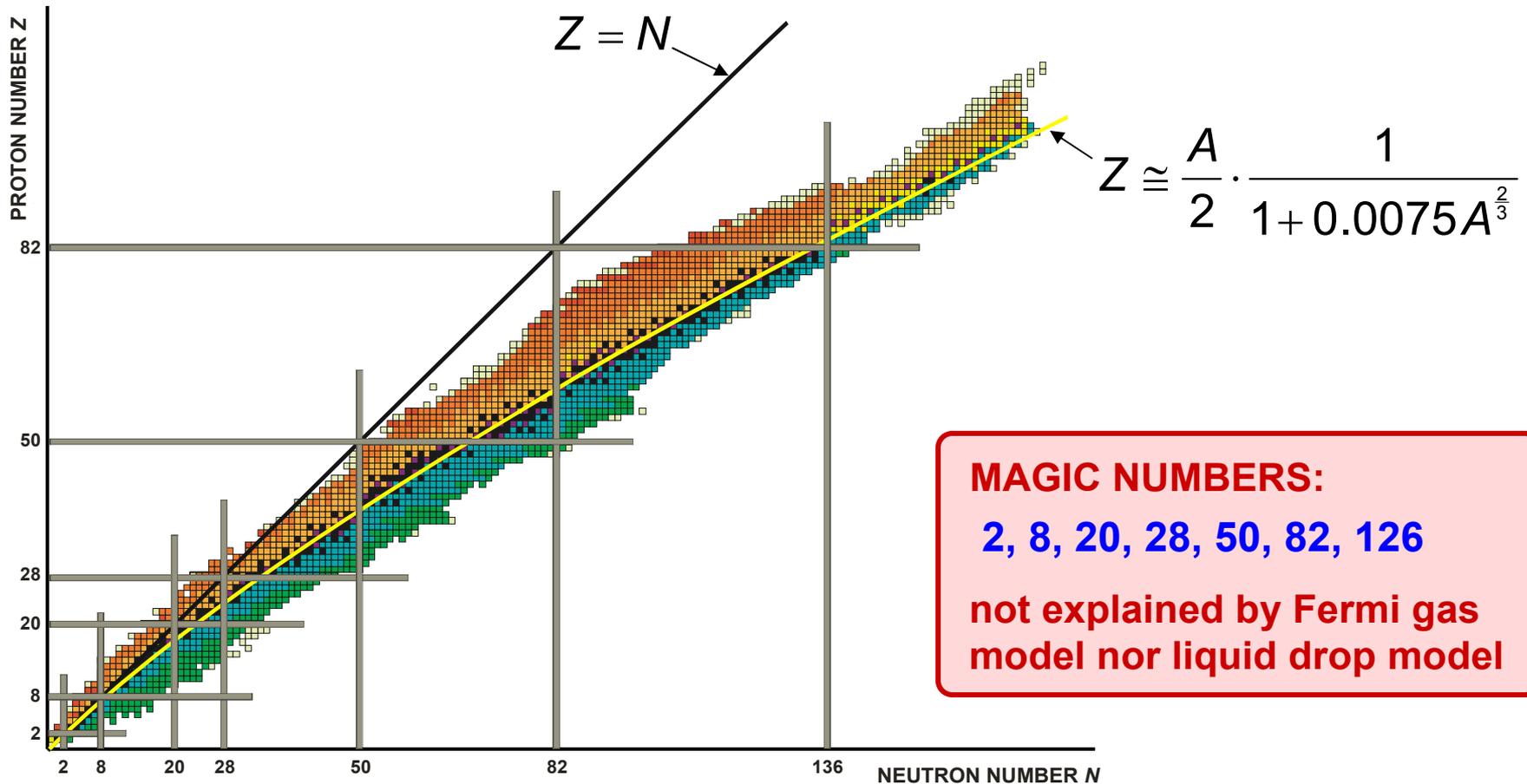


NUCLEAR LIQUID DROP MODEL

Valley of stability

$$m(A, Z) = Zm_p + (A - Z)m_n - a_V A + a_S A^{\frac{2}{3}} + a_C Z(Z - 1)A^{-\frac{1}{3}} + a_A (A - 2Z)^2 A^{-1} + \Delta$$

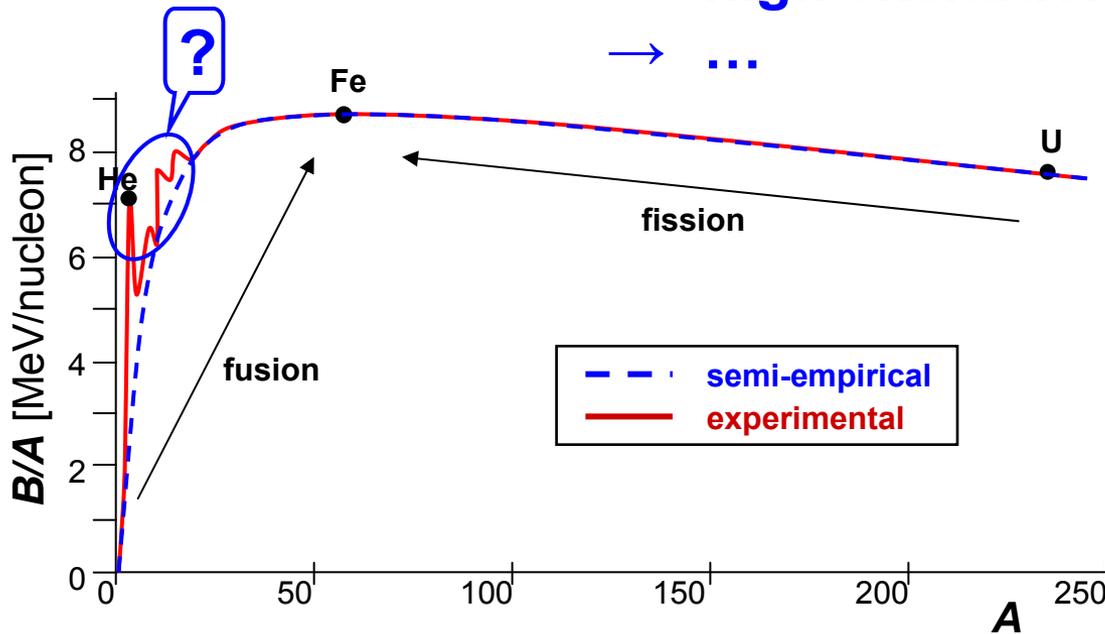
For fixed A the most stable Z is obtained by differentiating $m(A, Z)$



NUCLEAR SHELL MODEL – WHY?

New model needed to explain discontinuities of several nuclear properties

- binding energy
- high relative abundances
- low n-capture cross section
- high excitation energies
- ...



MAGIC NUMBERS

Magic number
=
closed shell

Magic numbers indicate similarity of nucleus to electron shells of atom, **BUT** still different from “Atomic magic numbers” (2, 10, 18, 36, 54, 86)

NUCLEAR SHELL MODEL

AIM

→ Explain the magic numbers

ASSUMPTION

→ Interactions between nucleons are neglected

→ Each nucleon can move independently in the nuclear potential

STEPS

→ Find the potential well that resembles the nuclear density

→ Consider the spin-orbit coupling

NUCLEAR SHELL MODEL

Potential well

Hamiltonian of a nucleus:

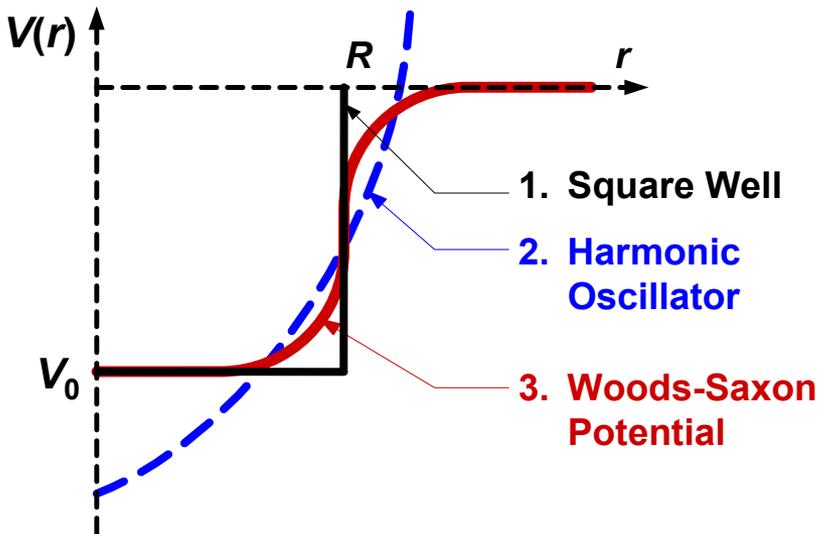
$$H = \sum_i (T_i + V(r_i)) + \lambda \left[\sum_{i,j,i \neq j} v(r_{ij}) - \sum_i V(r_i) \right]$$

central potential

residual potential

Central potential \gg Residual potential $\Rightarrow \lambda \rightarrow 0$

Potential well candidates



Solve Schrödinger equation

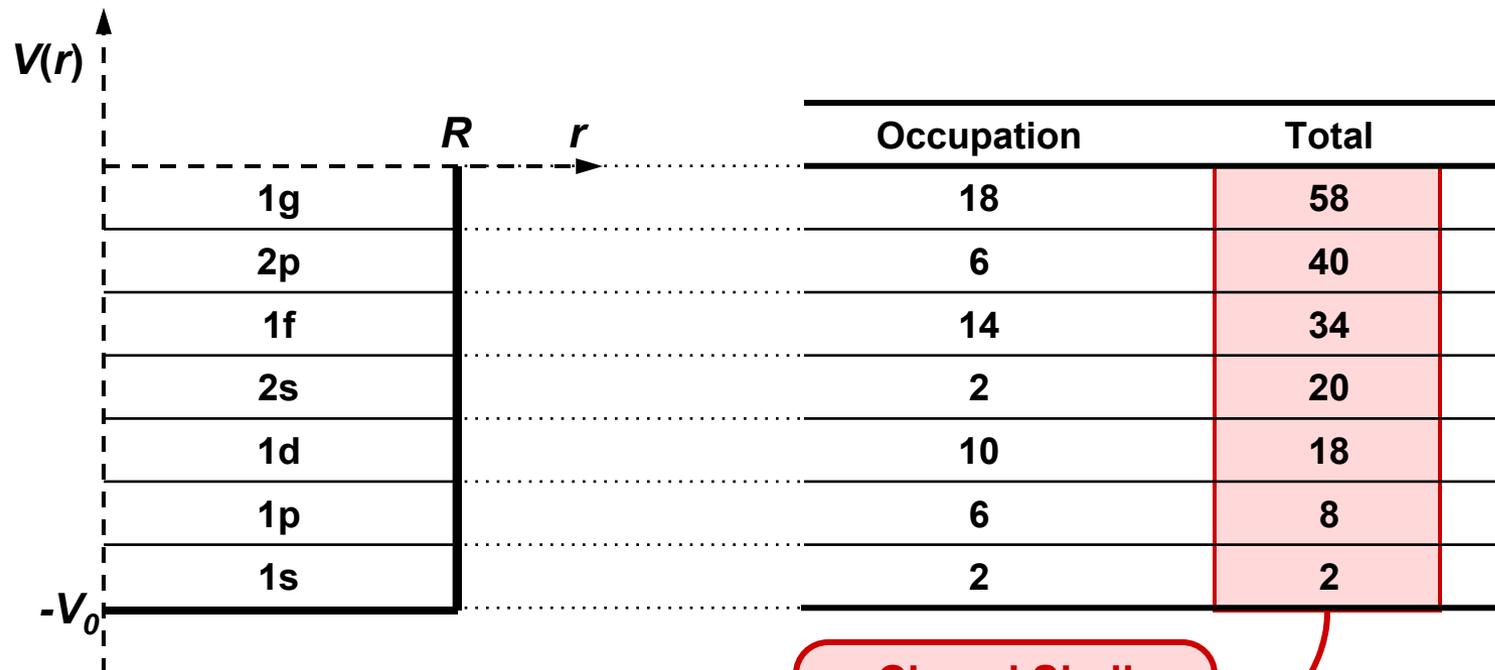
$$\frac{d^2}{dr^2} R_{nl} + \frac{2M}{r\hbar^2} \left(E_{nl} - V(r) - \frac{l(l-1)\hbar^2}{2Mr^2} \right) R_{nl} = 0$$

NUCLEAR SHELL MODEL

Square well potential

$$V_0(r) = \begin{cases} -V_0 & r \leq R \\ 0 & r > R \end{cases}$$

⇒ no analytical solution



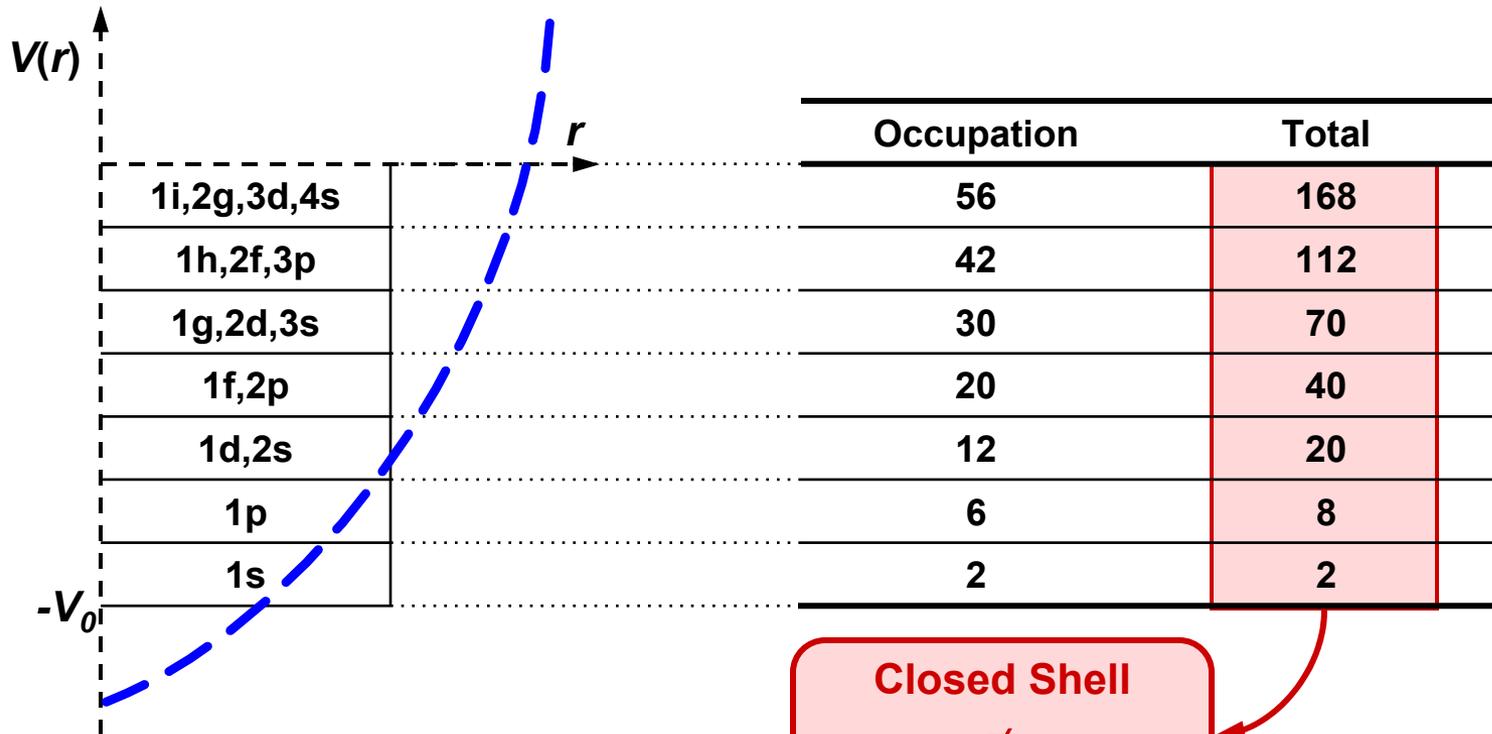
Closed Shell
≠
Magic Number

NUCLEAR SHELL MODEL

Harmonic potential

$$V_0(r) = -V_0 + \frac{1}{2}M\omega^2 r^2$$

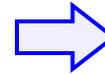
⇒ analytical solution possible



Closed Shell
≠
Magic Number

NUCLEAR SHELL MODEL

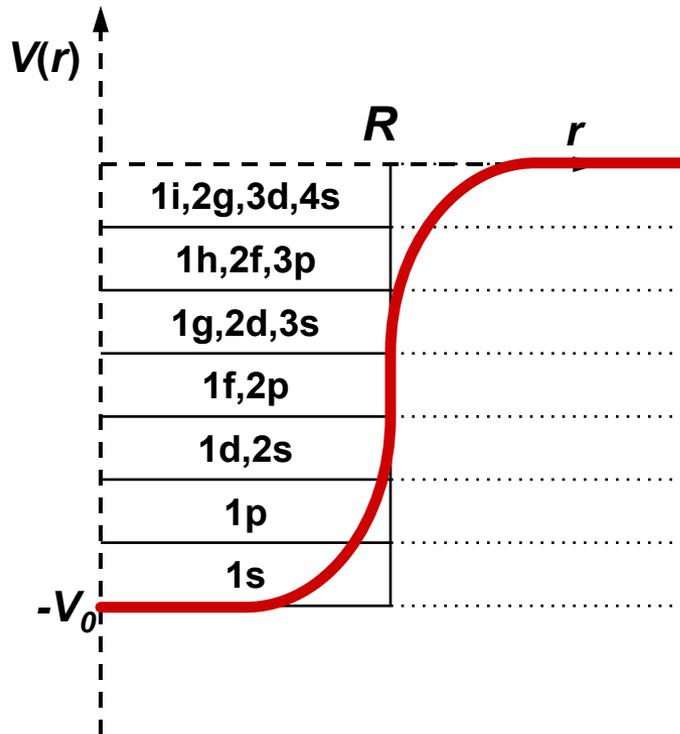
Woods-Saxon potential



resembles the nuclear density from scattering measurements

$$V_0(r) = -\frac{V_0}{1 + \exp\left(\frac{r-R}{a}\right)}$$

⇒ no analytical solution



Occupation	Total
56	168
42	112
30	70
20	40
12	20
6	8
2	2

Closed Shell
≠
Magic Number

NUCLEAR SHELL MODEL

Spin-orbit coupling contribution

Maria Mayer (*Physical Review* 78 (1950), 16) suggested:

1. There should be a non-central component
2. It should have a magnitude which depends on **S** & **L**

$$V(r) = V_0(r) - V_s(r) \mathbf{L} \cdot \mathbf{s}$$

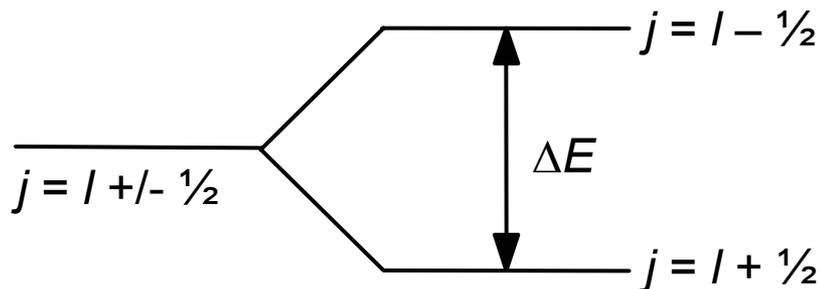
non-central potential

with $V_s(r) = V_{0s} \frac{1}{r} \frac{d}{dr} f(r)$

Woods-Saxon shape

Results in **energy splitting** of individual levels for given **J** (angular momentum)

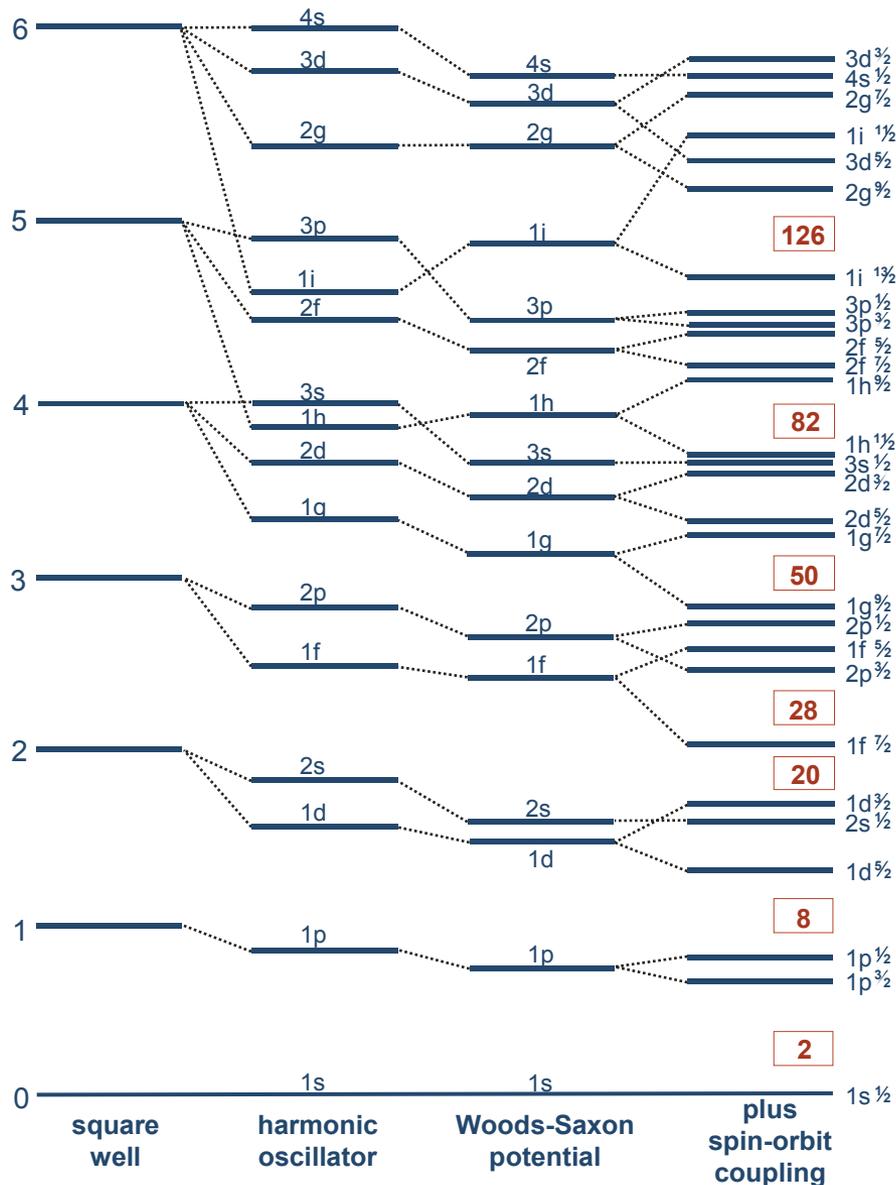
for $l > 0$



e.g. \Rightarrow $1d \begin{cases} 1d \ 3/2 \\ 1d \ 5/2 \end{cases}$

NUCLEAR SHELL MODEL

Level splitting



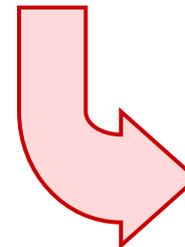
REMARKS

The essential features are given by any potential of the form

$$V(r) = V_0(r) + V_s(r)\mathbf{L} \cdot \boldsymbol{\sigma}$$

Energies of levels are parameter dependent

Shell model fails when dealing with deformed nuclei, i.e., nuclei far from magic numbers



**Collective models:
rotational, vibrational**

Other models

Close to **CLOSED-SHELL** nuclei well described by shell model

However, most of the nuclear properties are indeed determined by nucleons outside the closed shells

Collective models → **treating the closed shells as inert and only dealing with the rest**

Models not mentioned (but used):

- 1. rotational model** → rotations of permanently deformed nuclei
- 2. vibrational model** → excitations within shell – multipole account
- 3. Nilsson model** → shell model with deformed potential
- 4. α -particle model** → α -particle clusters inside the nucleus
- 5. interacting boson model** → considering pairs of nucleons as bosons