# Analytic de Sitter vacua in maximal Supergravity

Marco Gorghetto Master's thesis under the supervision of Prof. G. Dall'Agata

#### IMPRS EPP Workshop, Munich June 30, 2015

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• **Scope**: analysis of supersymmetry breaking patterns of a truncation of maximal Supergravity in D = 4

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- Motivations (1): understand the vacuum selection in String Theory and the general mechanisms of spontaneous supersymmetry breaking
  - focus on a relatively simple (although unrealistic) model
  - need to couple supersymmetric theories to gravity



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Motivations (2): reproduce (at least) a positive value of the cosmological constant Λ

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## Scope and motivations

- Motivations (2): reproduce (at least) a positive value of the cosmological constant  $\Lambda$ 
  - nowadays and during the inflation epoch  $\Lambda > 0$



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• In a supergravity theory with  $e^{-1}\mathcal{L}=\cdots-V(\phi)$ ,  $\Lambda=V(\phi_0)$ 

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• Local supersymmetric theory with N = 8 supercharges

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- Local supersymmetric theory with N = 8 supercharges
- Matter content completely fixed by supersymmetry
  - 70 scalars
  - 56 spinors
  - 28 vectors
  - ► 8 gravitinos
  - 1 graviton

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- Local supersymmetric theory with N = 8 supercharges
- Matter content completely fixed by supersymmetry
  - 70 scalars
  - ► 56 spinors
  - 28 vectors
  - 8 gravitinos
  - 1 graviton
- Sigma-model with target manifold the coset Lie group  $E_{7(7)}/SU(8)$
- Descends from M-theory
  - Iow energy limit
  - compactification of extra dimensions on  $\mathbb{T}^7$
- Does not include a scalar potential

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- A scalar potential can be introduced deforming pure supergravity by introducing a gauge group  $G_g$ 

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  - ▶ gauge covariant derivatives  $\partial_\mu o D_\mu = \partial_\mu + g A_\mu$
  - new terms in  $\delta_{SUSY}$  and in  $\mathcal L$
  - potential  $V(\phi)$  of order  $g^2$

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1.  $G_g = SO(8)$  gauged supergravity arising from compactification on  $S^7$  and truncation by the  $\mathbb{Z}_2 \times \mathbb{Z}_2 \times \mathbb{Z}_2$  orbifold

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- 2. Isotropic models, i.e. compactification on  $\mathbb{T}^7/(\mathbb{Z}_2)^3$

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- 2. Isotropic models, i.e. compactification on  $\mathbb{T}^7/(\mathbb{Z}_2)^3$ 
  - N = 1 supergravity model  $n_c = 7, \ \Phi = (S, T_1, T_2, T_3, U_1, U_2, U_3)$ and  $n_v = 0$

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  - N = 1 supergravity model  $n_c = 7$ ,  $\Phi = (S, T_1, T_2, T_3, U_1, U_2, U_3)$ and  $n_v = 0$

• 
$$K = -\log(S + \bar{S}) - \sum_{i=1}^{3} \log(T_i + \bar{T}_i) - \sum_{i=1}^{3} \log(U_i + \bar{U}_i)$$

 $\blacktriangleright W = W_0 + W_\alpha \Phi^\alpha + \frac{1}{2!} W_{\alpha\beta} \Phi^\alpha \Phi^\beta + \frac{1}{3!} W_{\alpha\beta\gamma} \Phi^\alpha \Phi^\beta \Phi^\gamma$ 

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## SO(8) gauged supergravity: results

	Residual sym	٨	Scalar square masses: $m_{(multiplicity)}^2$
(i)	SO(8)	-3	$-\frac{2}{3}(14)$
(ii)	SO(7)	-4	$2_{(1)}, -\frac{4}{5}_{(6)}, -\frac{2}{5}_{(7)}$
(iii)	SO(6)	$-\frac{25\sqrt{5}}{16}$	$2_{(2)}, -1_{(5)} - \frac{1}{4}_{(4)}, 0_{(3)}$
(iv)	SO(7)	$-5^{3/4}$	$2_{(1)}, -\frac{4}{5}_{(6)}, -\frac{2}{5}_{(7)}$
(v)	G <sub>2</sub>	$-\frac{108}{25}\sqrt{\frac{2}{5}}\sqrt[4]{3}$	$\frac{4+\sqrt{6}}{3}(1), -\frac{11+\sqrt{6}}{18}(6), \frac{4-\sqrt{6}}{3}(1), \frac{-11+\sqrt{6}}{18}(6)$

Table: Vacua of the truncated theory for  $G_g = SO(8)$ 

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1. Constraint:  $T_1 = T_2 = T_3 = T$ ,  $U_1 = U_2 = U_3 = U$ 

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  - $\blacktriangleright W = W_0 + W_\alpha \Phi^\alpha + \frac{1}{2!} W_{\alpha\beta} \Phi^\alpha \Phi^\beta + \frac{1}{3!} W_{\alpha\beta\gamma} \Phi^\alpha \Phi^\beta \Phi^\gamma$
  - conditions on the parameters for a vacuum located at the origin  $S = T = U = \overline{S} = \overline{T} = \overline{U} = 1$

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  - 1.1 General conditions for Minkowski supersymmetric stable vacua •  $W_0 = W_S = W_T = W_U = 0$
  - 1.2 General conditions for Minkowski no-scale vacua •  $W_S = \frac{1}{2}W_0, W_T = 0, W_U = \frac{3}{2}W_0, W_0 \neq 0$

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    - $W_S = \frac{1}{2}W_0, W_T = 0, W_U = \frac{3}{2}W_0, W_0 \neq 0$
  - 1.3 Perturb these conditions to search for de Sitter vacua:
    - $W_0 \to W_0(\epsilon, \lambda), W_S \to W_S(\epsilon, \lambda), W_T \to W_T(\epsilon, \lambda), W_U \to W_U(\epsilon, \lambda)$

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• Sample of perturbation parameters around a no-scale vacuum:

 $(W_0, W_S, W_T, W_U) = (1, \frac{1}{2} + \epsilon, (1+i)\epsilon, \frac{3}{2} + \lambda)$ 

- Blue: de Sitter vacua
- Purple: strictly stable vacua



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 Vacua in the unconstrained model (with 7 moduli) for the same value of the perturbation parameters



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 $(W_0, W_S, W_T, W_U) = (1, \frac{1}{2} + \epsilon, (1+i)\epsilon, \frac{3}{2} + \lambda)$ 

- Blue: de Sitter vacua
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- Vacua in the unconstrained model (with 7 moduli) for the same value of the perturbation parameters
  - Just some anti de Sitter vacua become unstable
  - All vacua become marginally stable (with 4 flat directions)



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Conclusions and outlook

## Conclusions and outlook

- We determined analytically a set of de Sitter vacua in a truncation of maximal supergravity.
- How can the vacua be reproduced in the N = 8 landscape?
- Are all of them upliftable to D = 11 supergravity?

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