

FACT

Alerts from TeV Gamma-Rays in the Context of
Multi-Messenger Astrophysics



Daniela Dorner for the FACT Collaboration



First G-APD Cherenkov Telescope

Major Goals

Proof of principle:

Silicon based photo sensors (G-APDs*) in Cherenkov Telescopes



Successful operation since October 2011

Observing Strategy:
Longterm monitoring

Sample:
Few bright TeV Blazars

- Variability studies
- Multi-wavelength correlations
- Flare alerts to other instruments



* Geiger-mode Avalanche Photodiodes

Facts about FACT

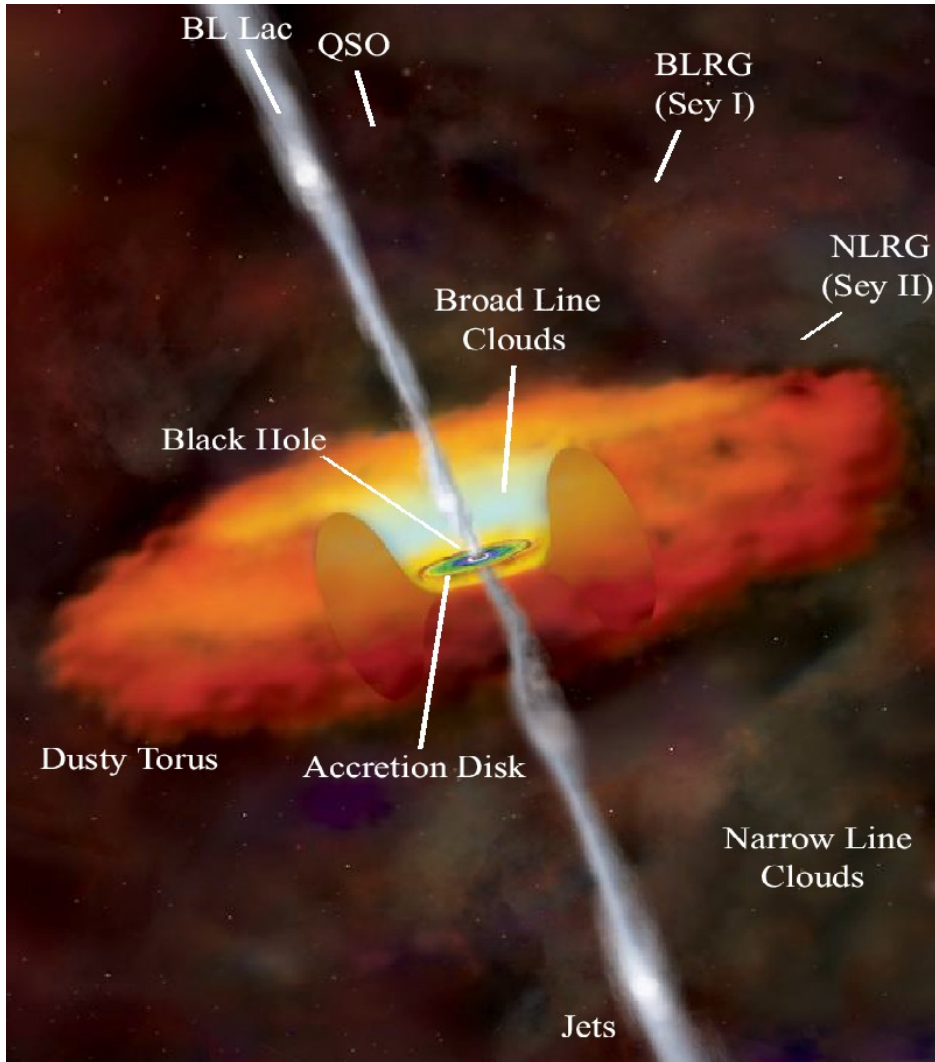


- 2200 m a.s.l.
Observatorio del Roque de los Muchachos, La Palma
- HEGRA CT3 mount
- 9.5 m² mirror area
- G-APD camera
 - Silicon based photosensors
 - 4.5° FoV, 1440 pixels à 0.11°
 - *H Anderhub et al 2013 JINST 8 P06008*
- Operational since Oct 2011
- Excellent performance
 - *A Biland et al 2014 JINST 9 P10012*

Photo: Thomas Krähenbühl



Targets for FACT: Blazars



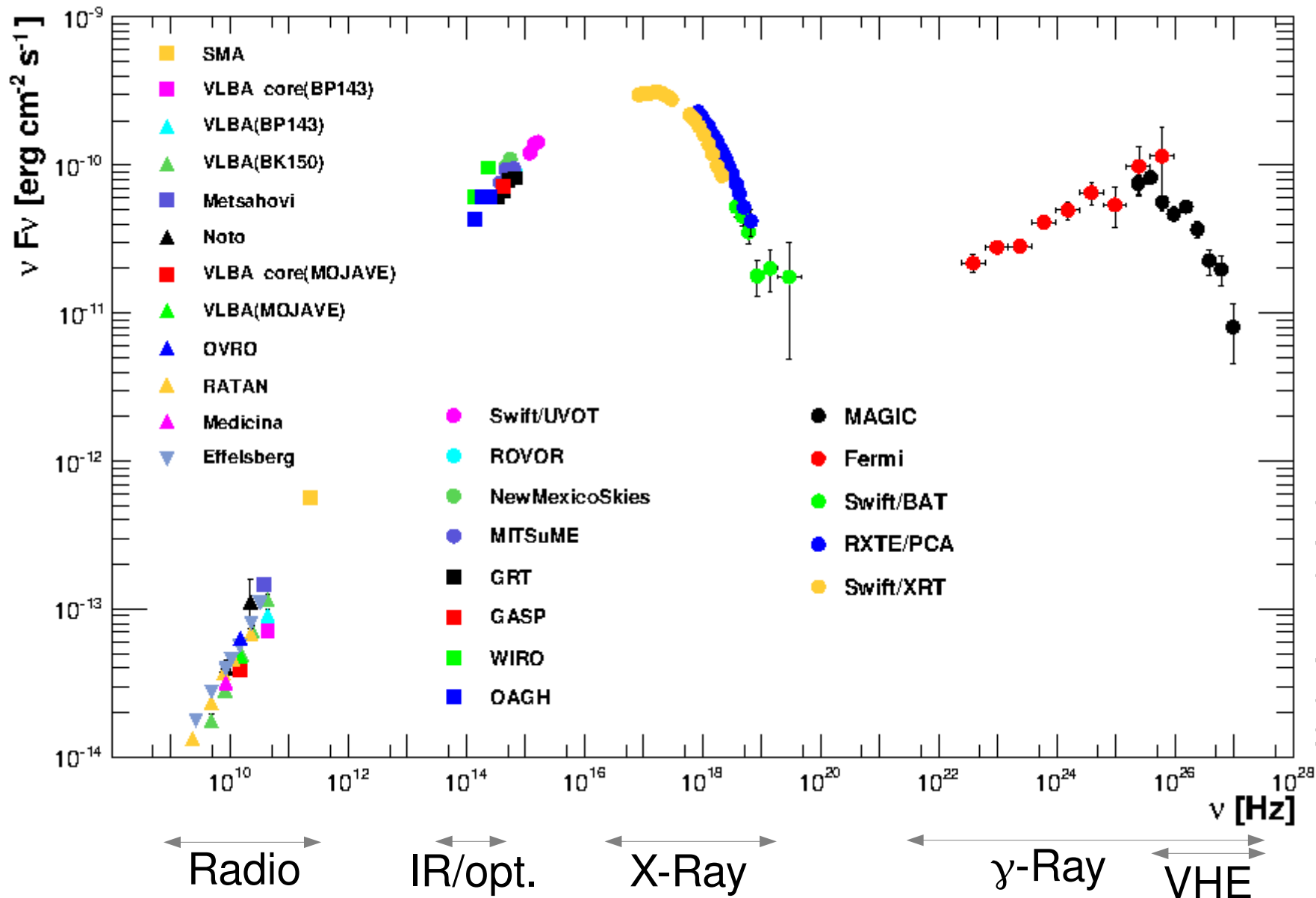
- Spectral energy distribution: Two-peak structure
→ Multi-wavelength observations crucial

<http://chandra.harvard.edu/resources/illustrations/quasar.html>



Spectral Energy Distribution

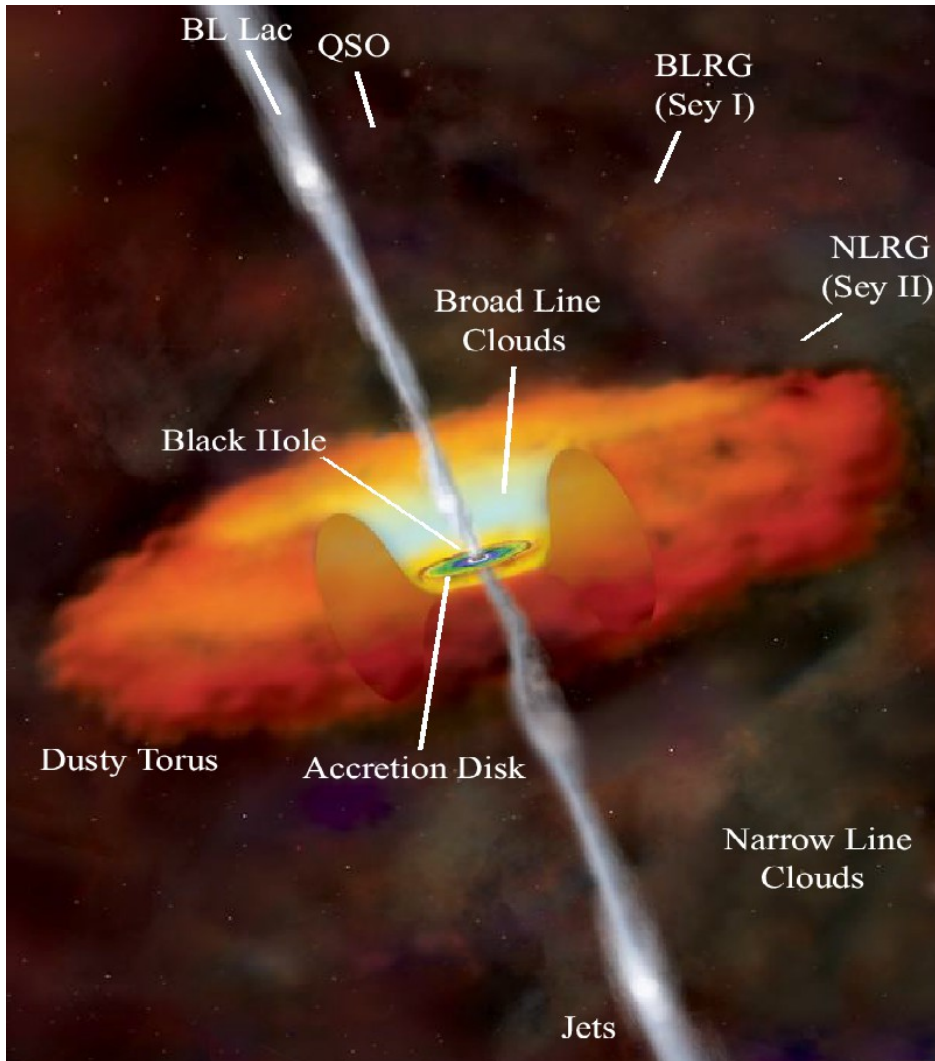
Mrk 421



Abdo A. et al. ApJ 736, 131, 2011



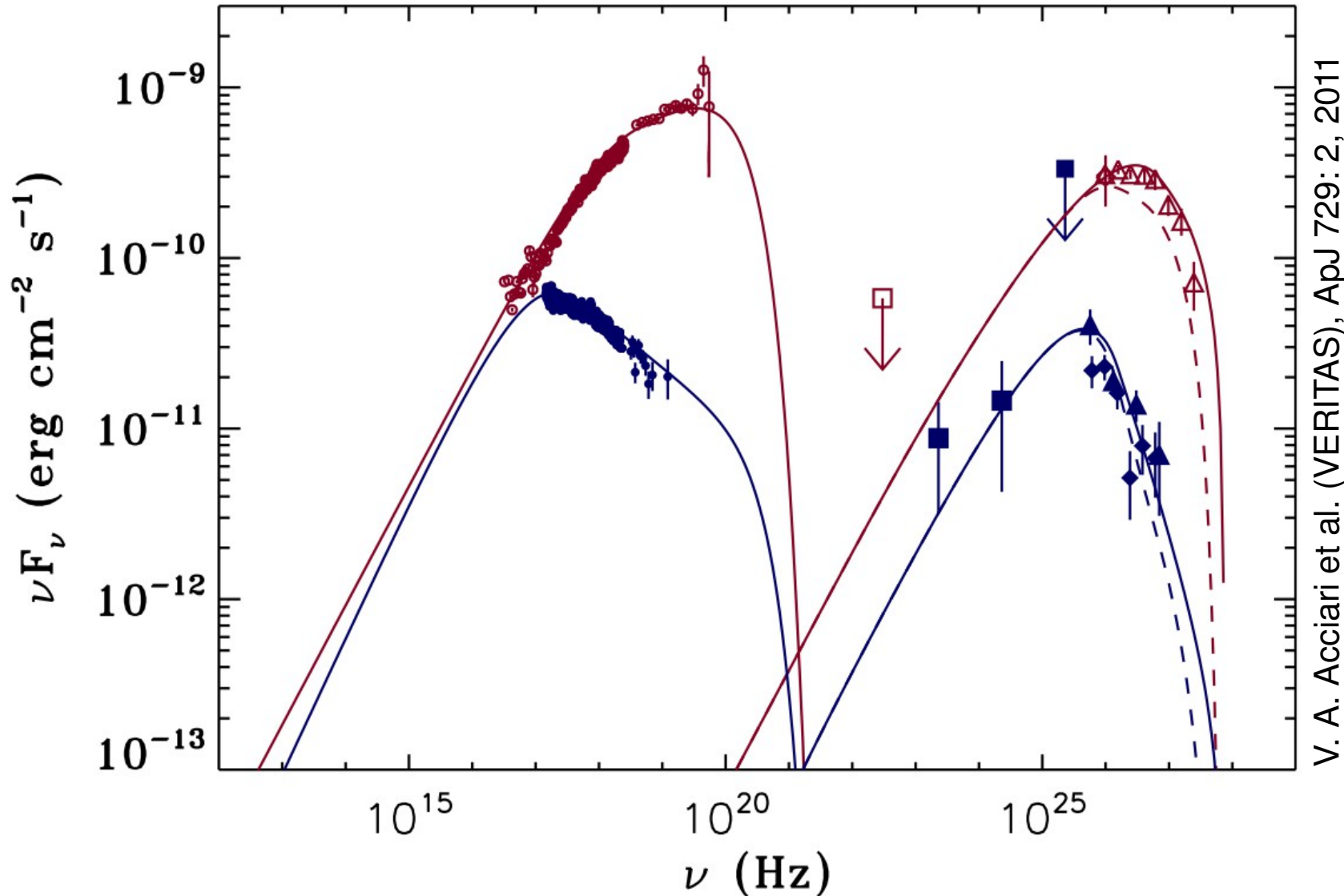
Targets for FACT: Blazars



<http://chandra.harvard.edu/resources/illustrations/quasar.html>

- Spectral energy distribution: Two-peak structure
→ Multi-wavelength observations crucial
- Extreme variability on different time scales
→ unbiased data sampling vital

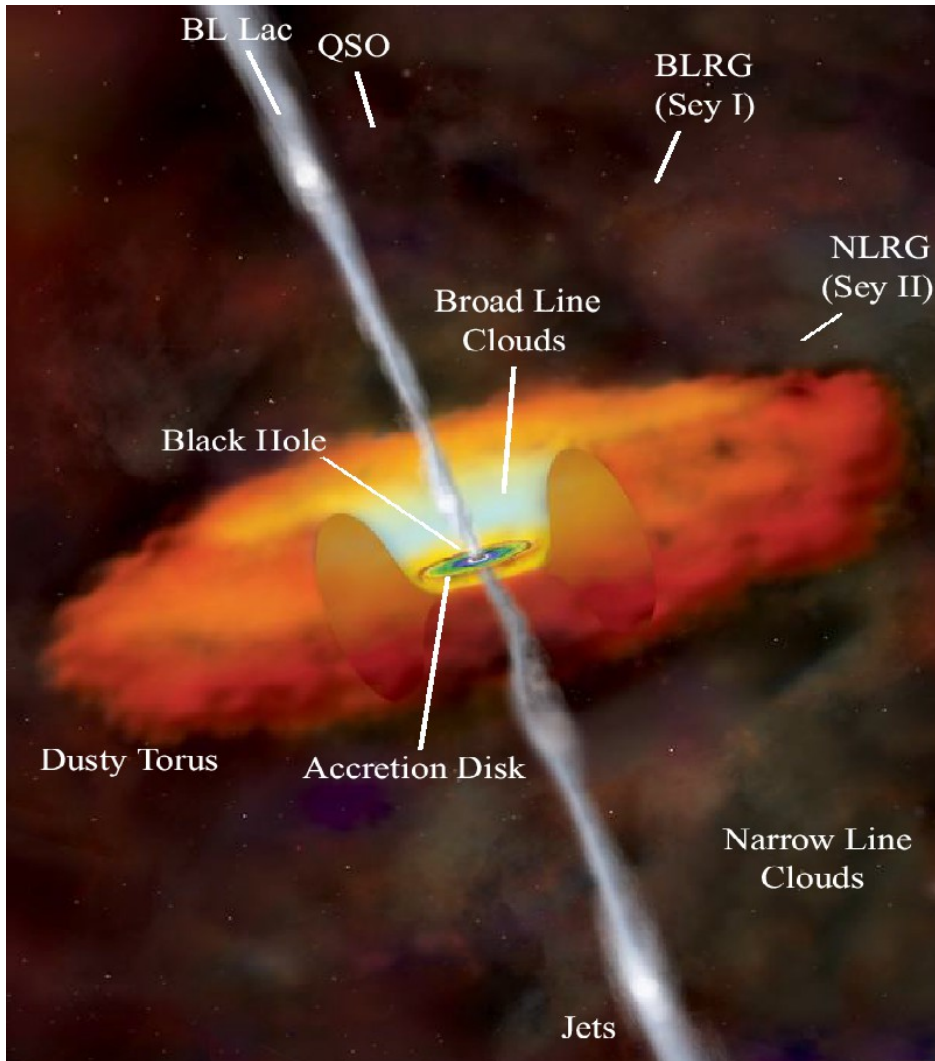
Extreme Variability



Mrk 501

V. A. Acciari et al. (VERITAS), ApJ 729: 2, 2011

Targets for FACT: Blazars

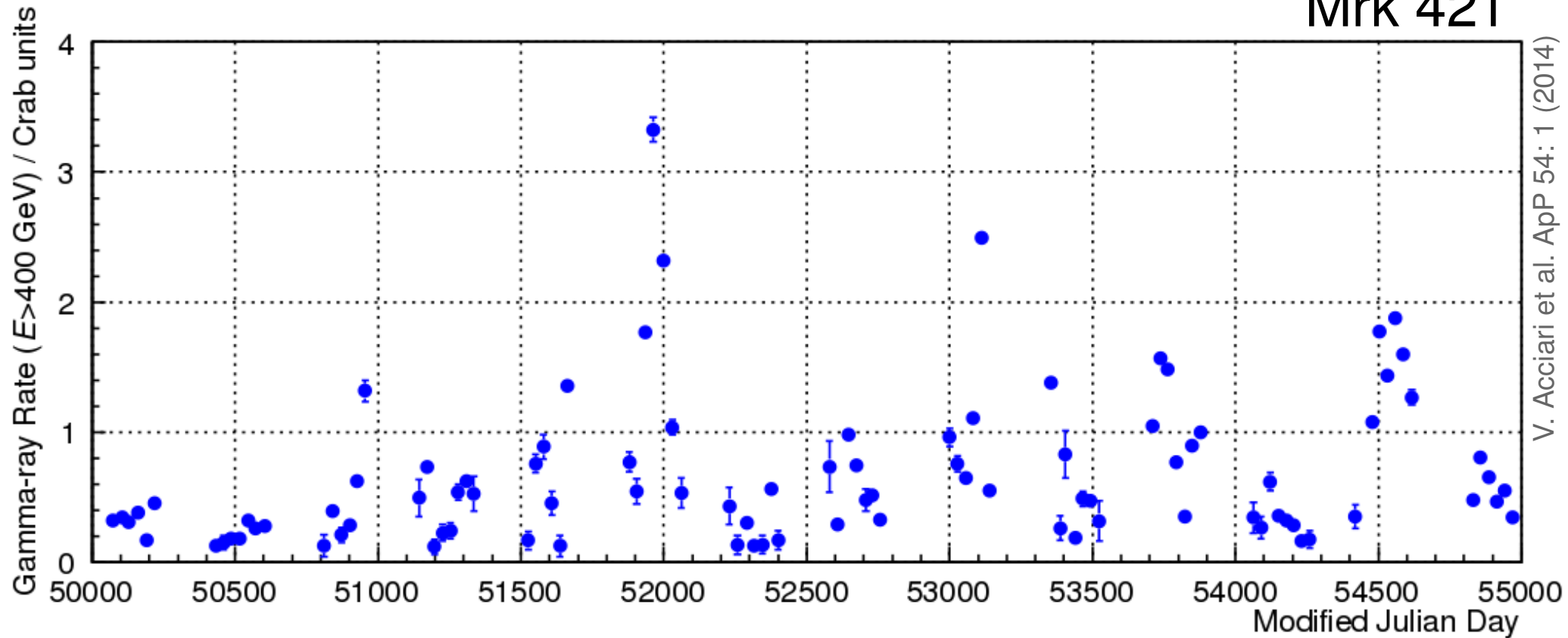


<http://chandra.harvard.edu/resources/illustrations/quasar.html>

- Spectral energy distribution: Two-peak structure
→ Multi-wavelength observations crucial
- Extreme variability on different time scales
→ unbiased data sampling vital

Long-term Variability

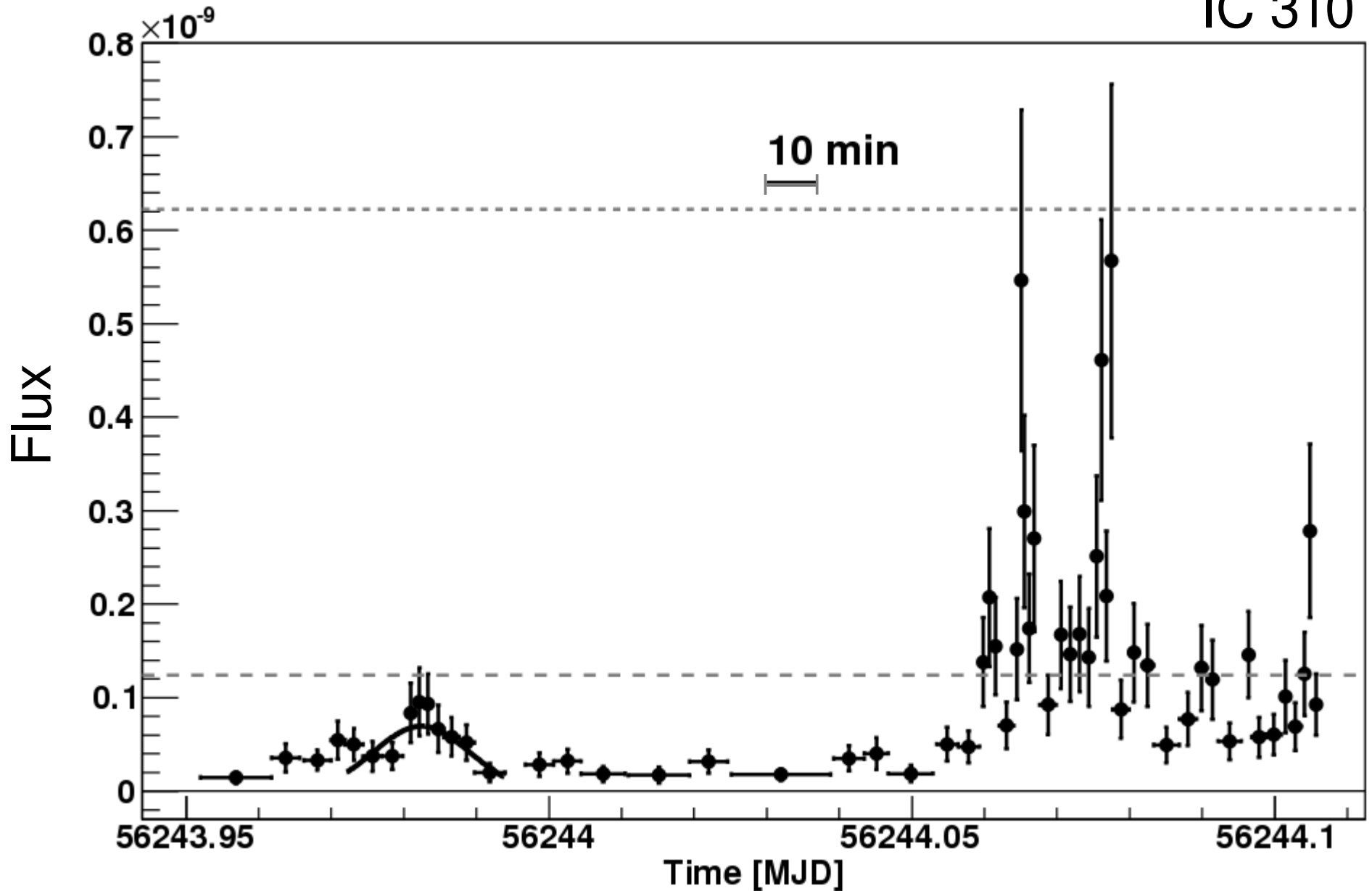
Mrk 421



Whipple Monitoring 14 years, monthly binning

Short-term Variability

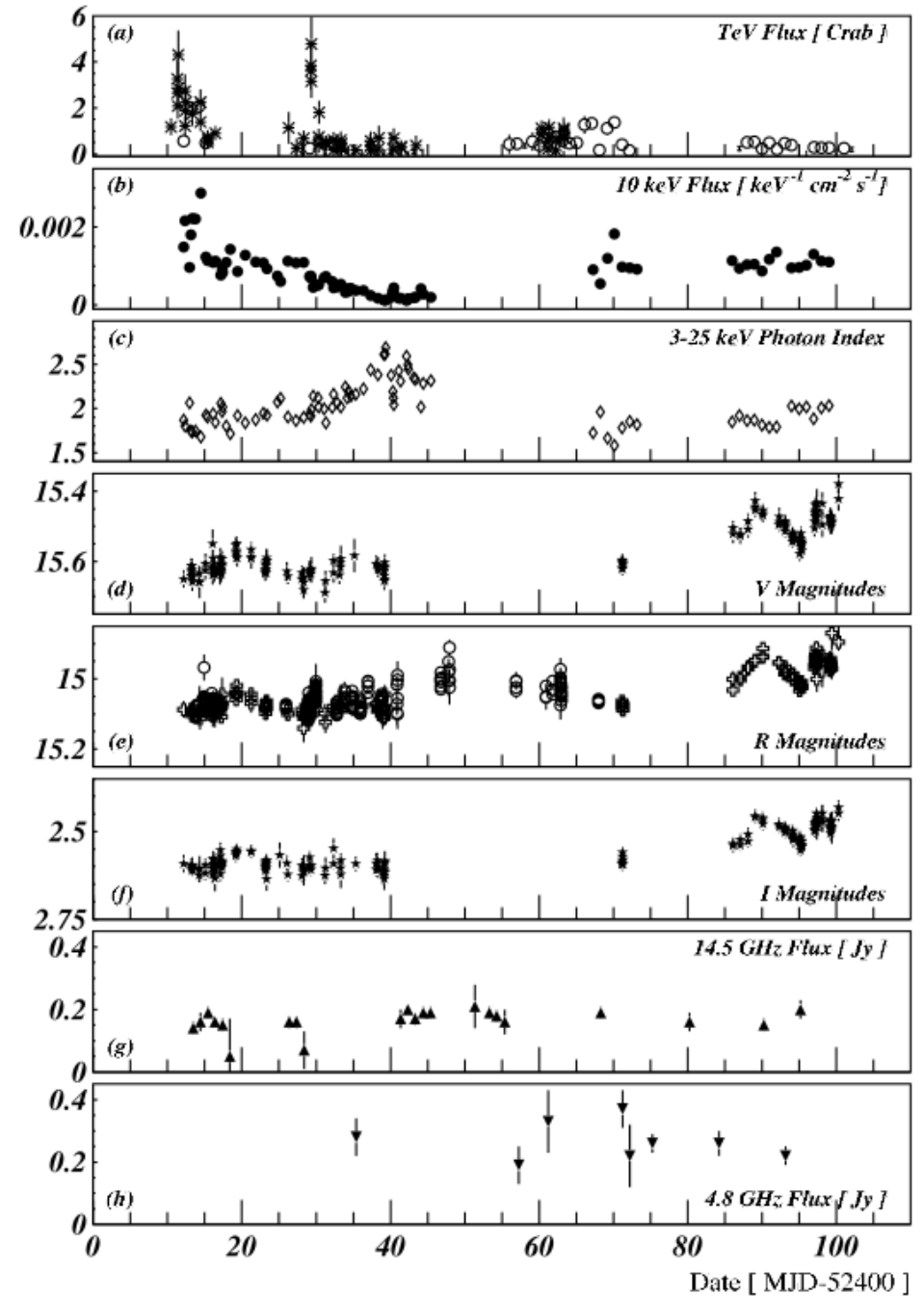
IC 310



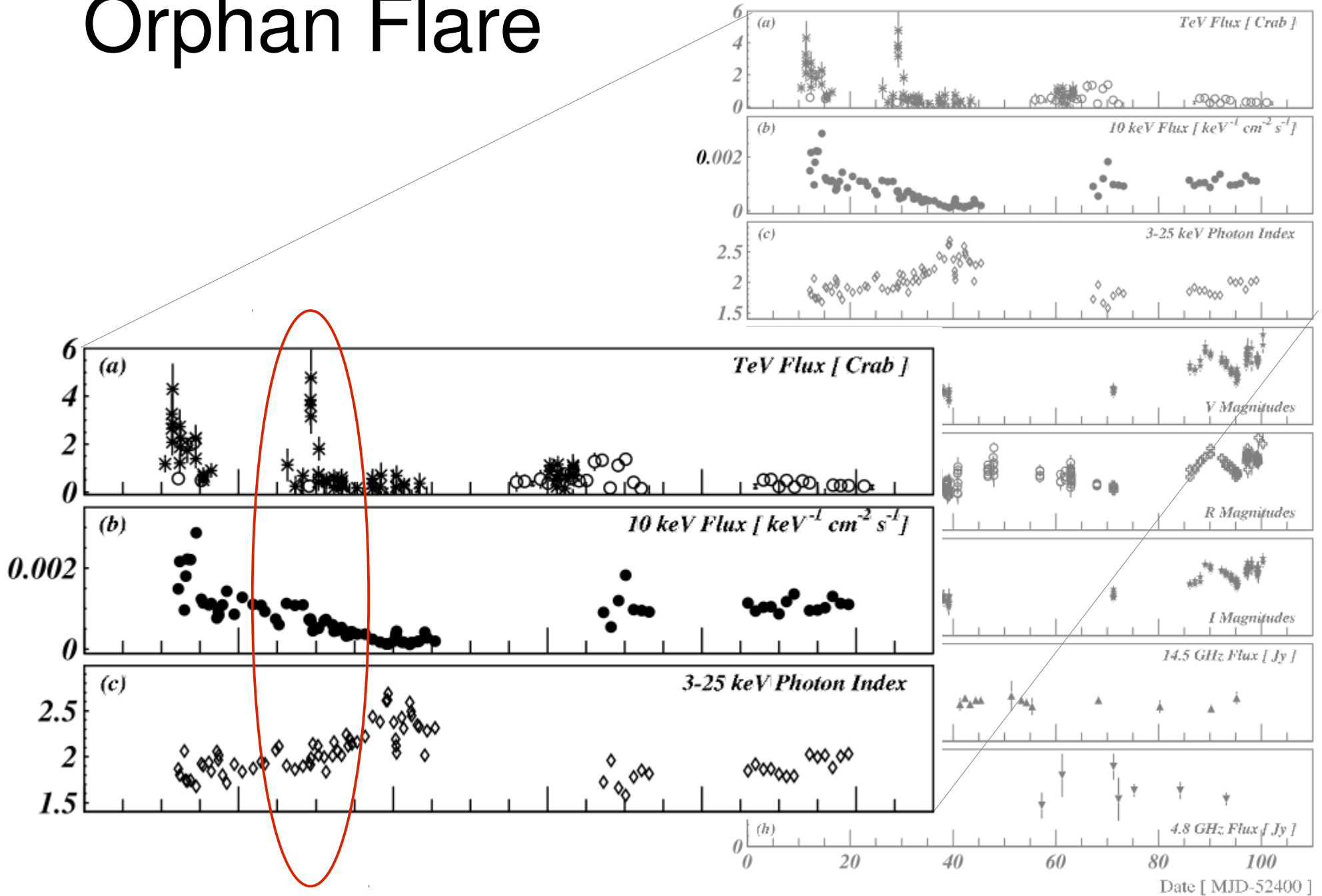
Aleksic, J. et al., Science p (2014)



Orphan Flare



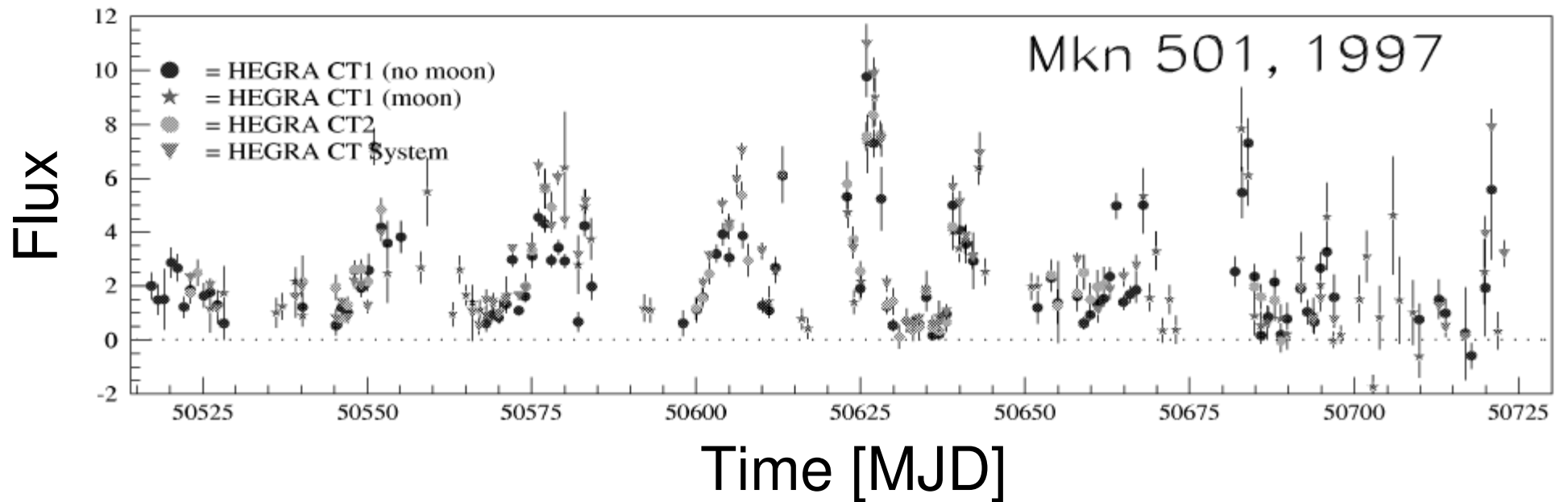
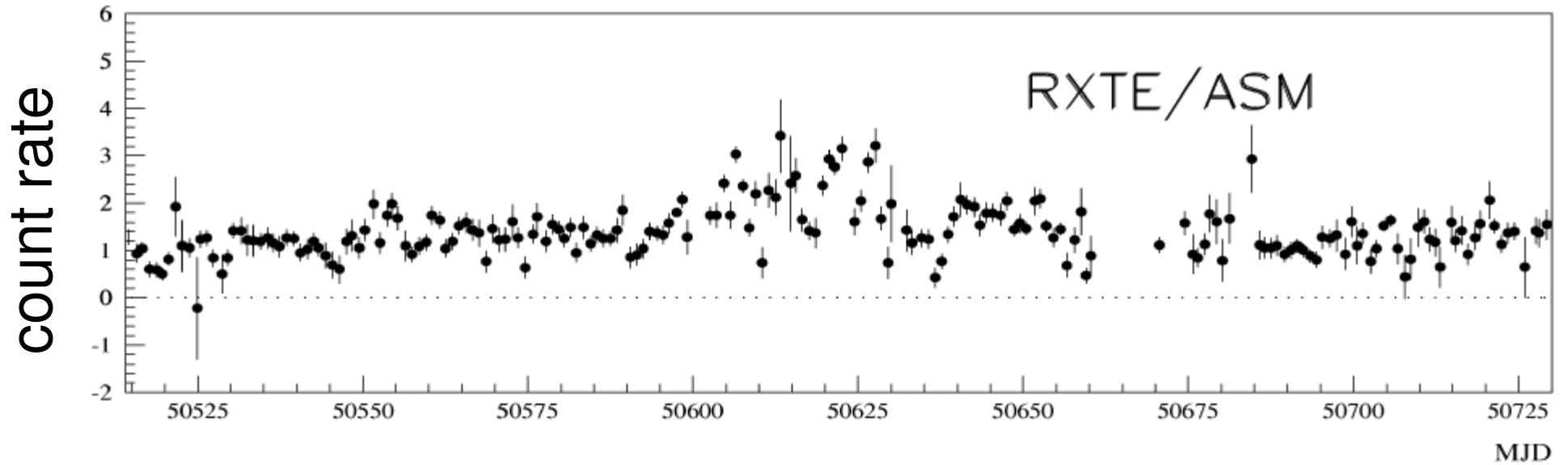
Orphan Flare



Krawczynski, H. et al., ApJ 601 p151-164 (2004)



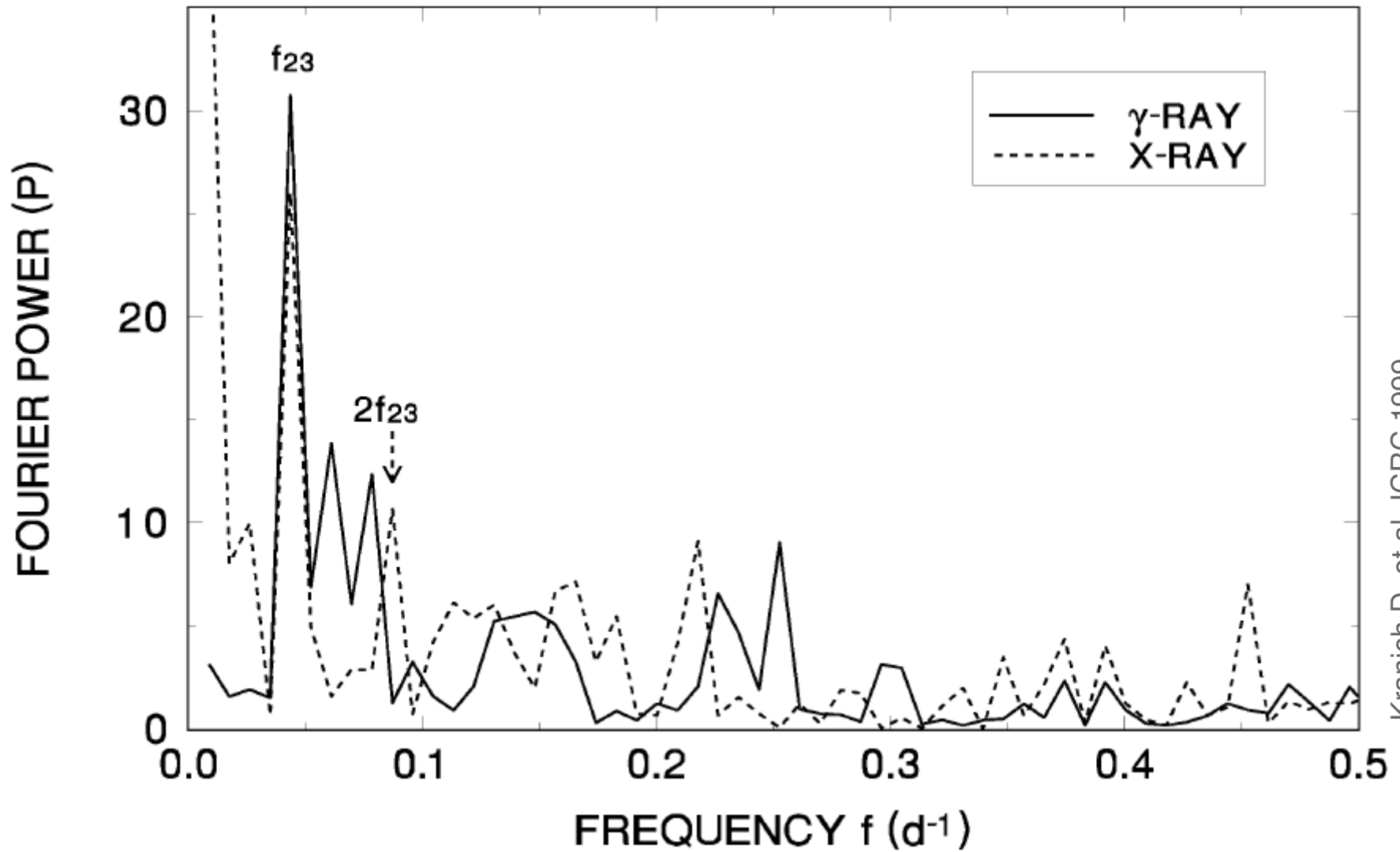
Periodicity



Kranich D. et al, ICRC 1999



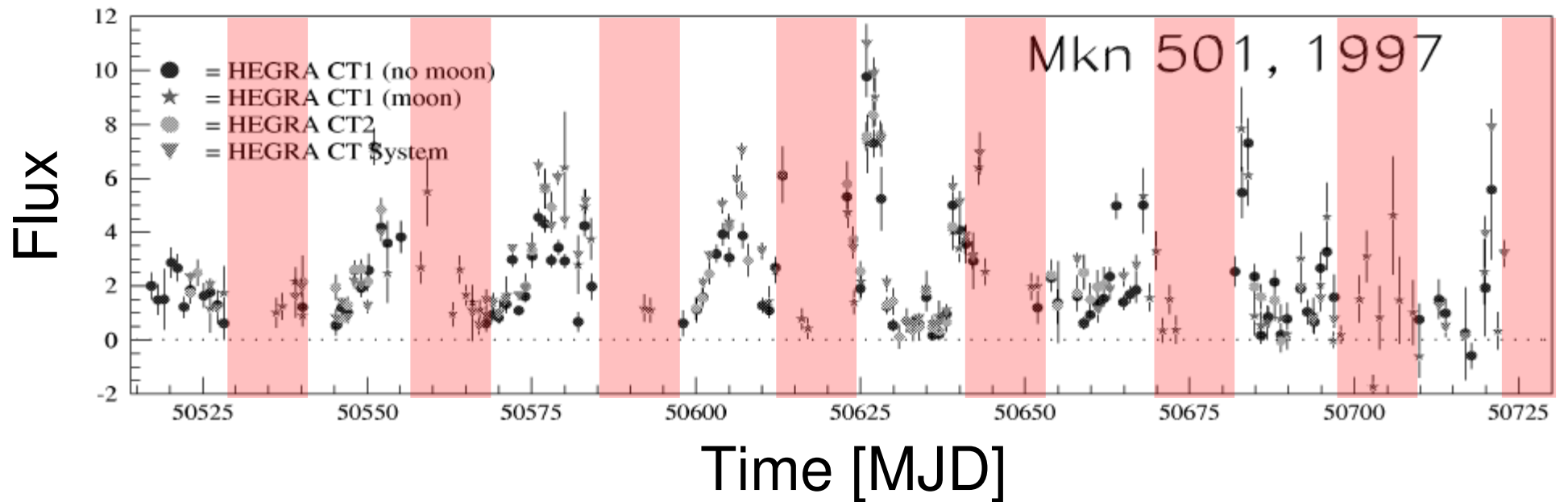
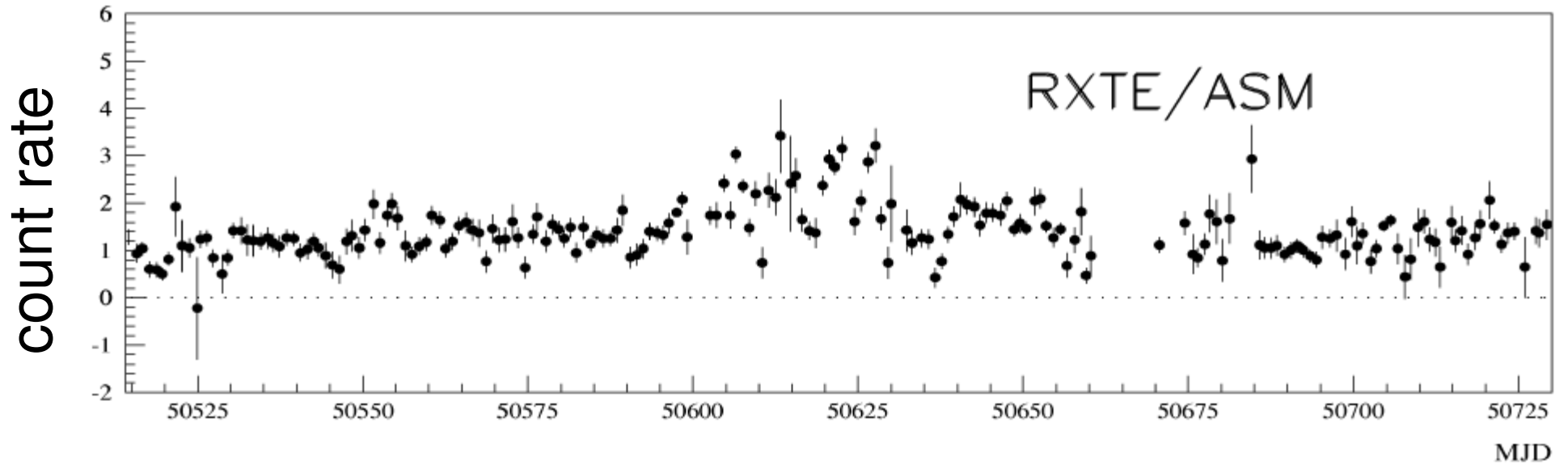
Periodicity



Kranich D. et al, ICRC 1999



Periodicity



Kranich D. et al, ICRC 1999



Requirements Monitoring

- No gaps
- Unbiased sample
- Stable monitoring



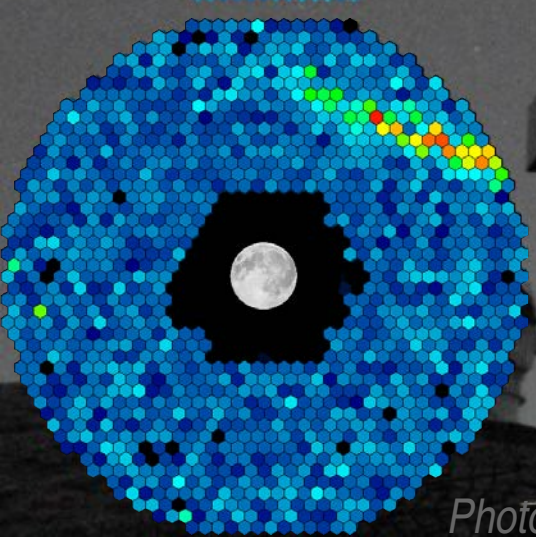
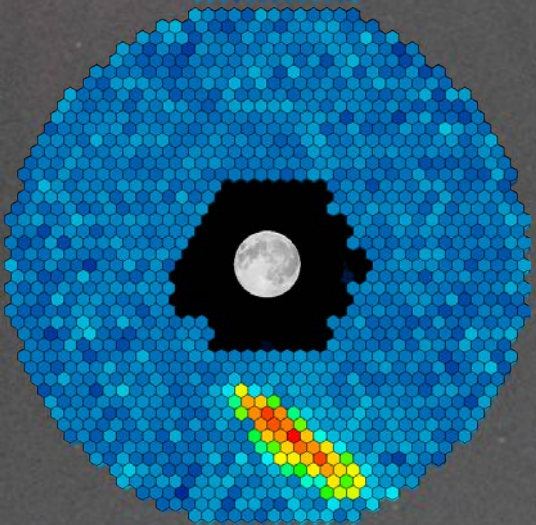
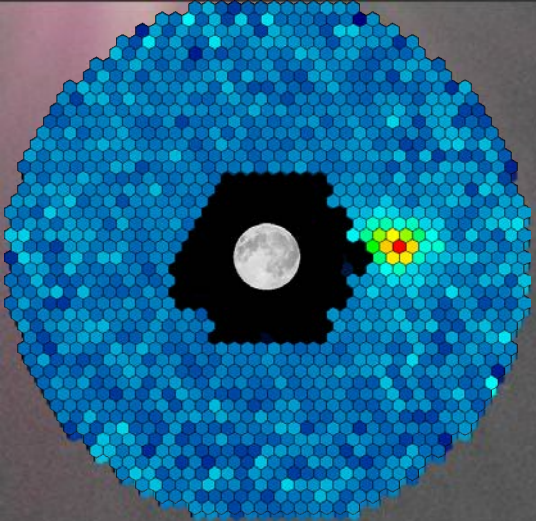
FACT – Ideal Monitoring Telescope



- Observations during strong moon light

Photo: Daniela Dorner





Showers images recorded pointing to brightest full moon in 2013.

Photos: D. Dörner, T. Krähenbühl

FACT – Ideal Monitoring Telescope



- Observations during strong moon light
 - Larger duty cycle
 - More complete data sample

Photo: Daniela Dorner

FACT – Ideal Monitoring Telescope

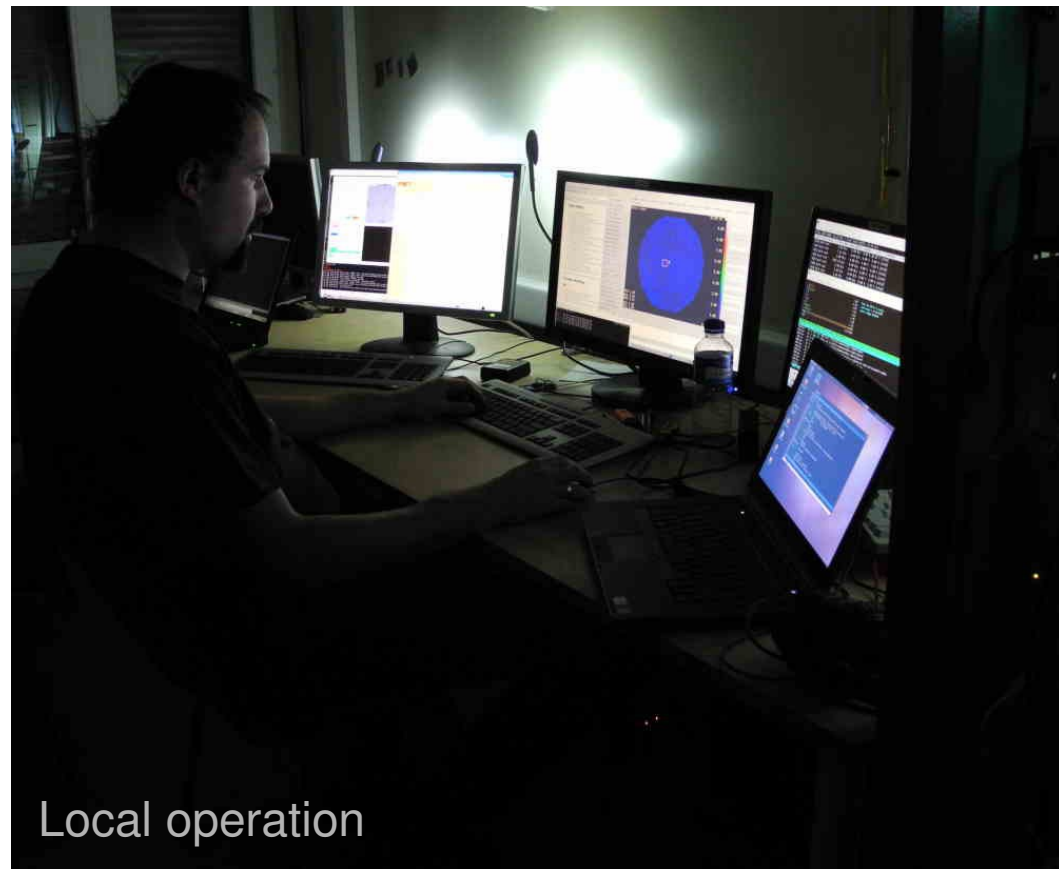


- Observations during strong moon light
 - Larger duty cycle
 - More complete data sample
- G-APDs robust and stable
 - Stable telescope

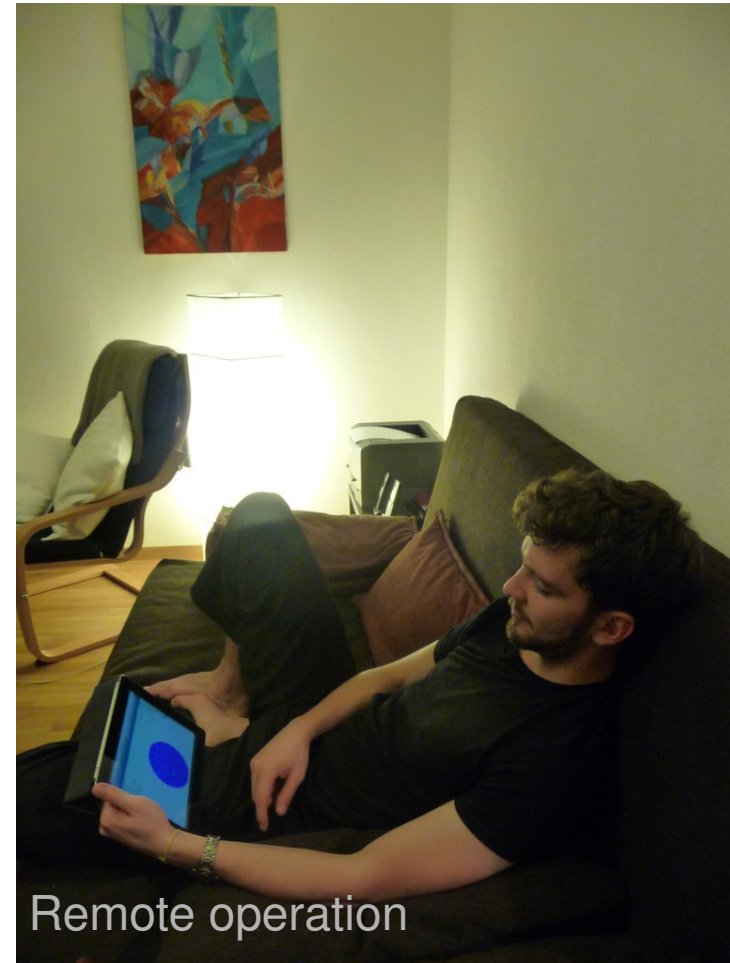
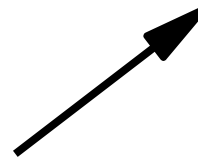
Photo: Daniela Dorner



Towards Robotic Operation



Local operation



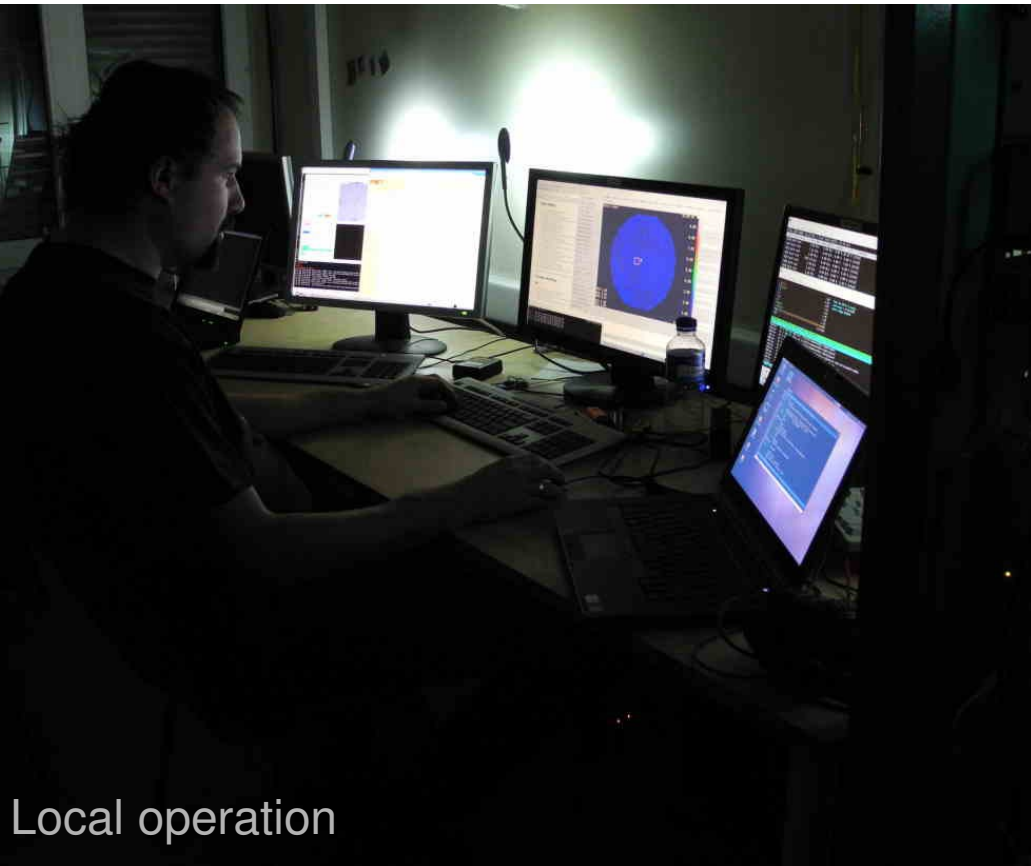
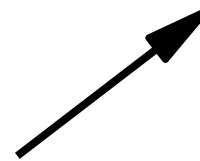
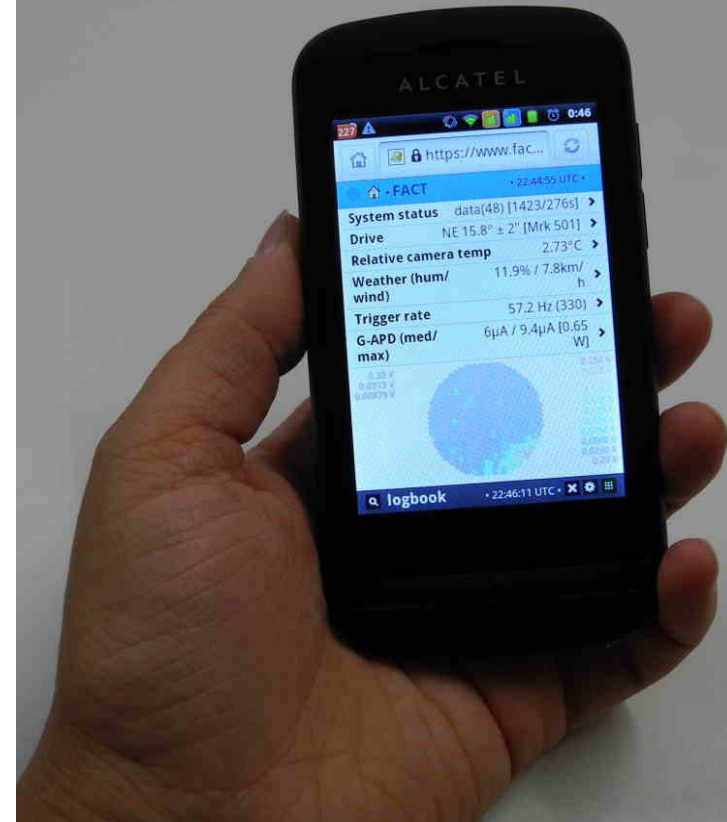
Remote operation

<http://www.fact-project.org/smartfact>



Towards Robotic Operation

Operation via
smartphone



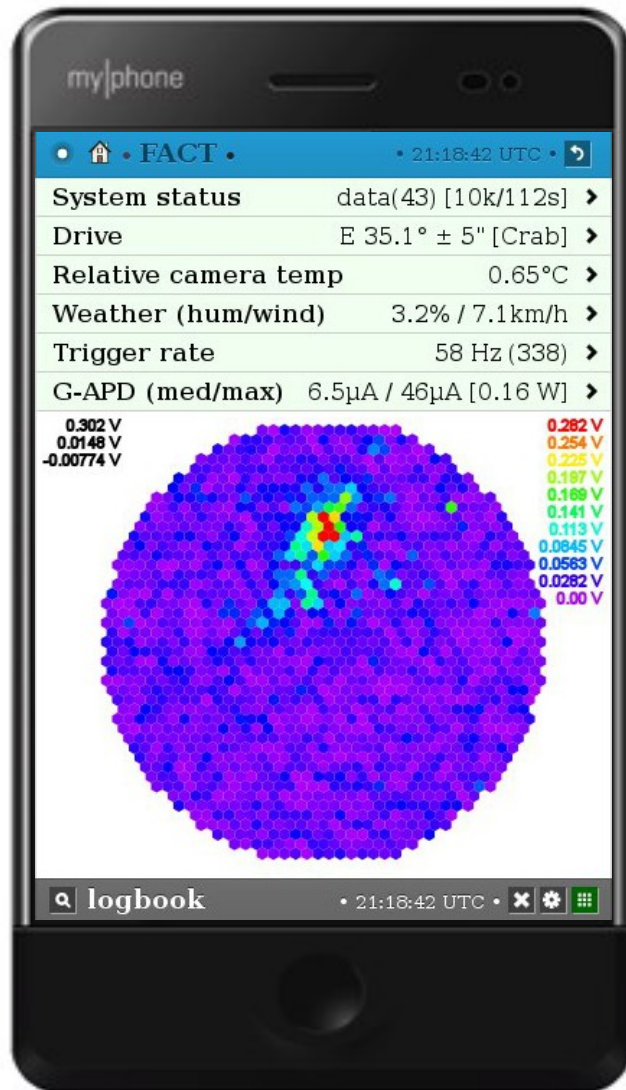
Local operation

Automatic Operation

<http://www.fact-project.org/smartfact>

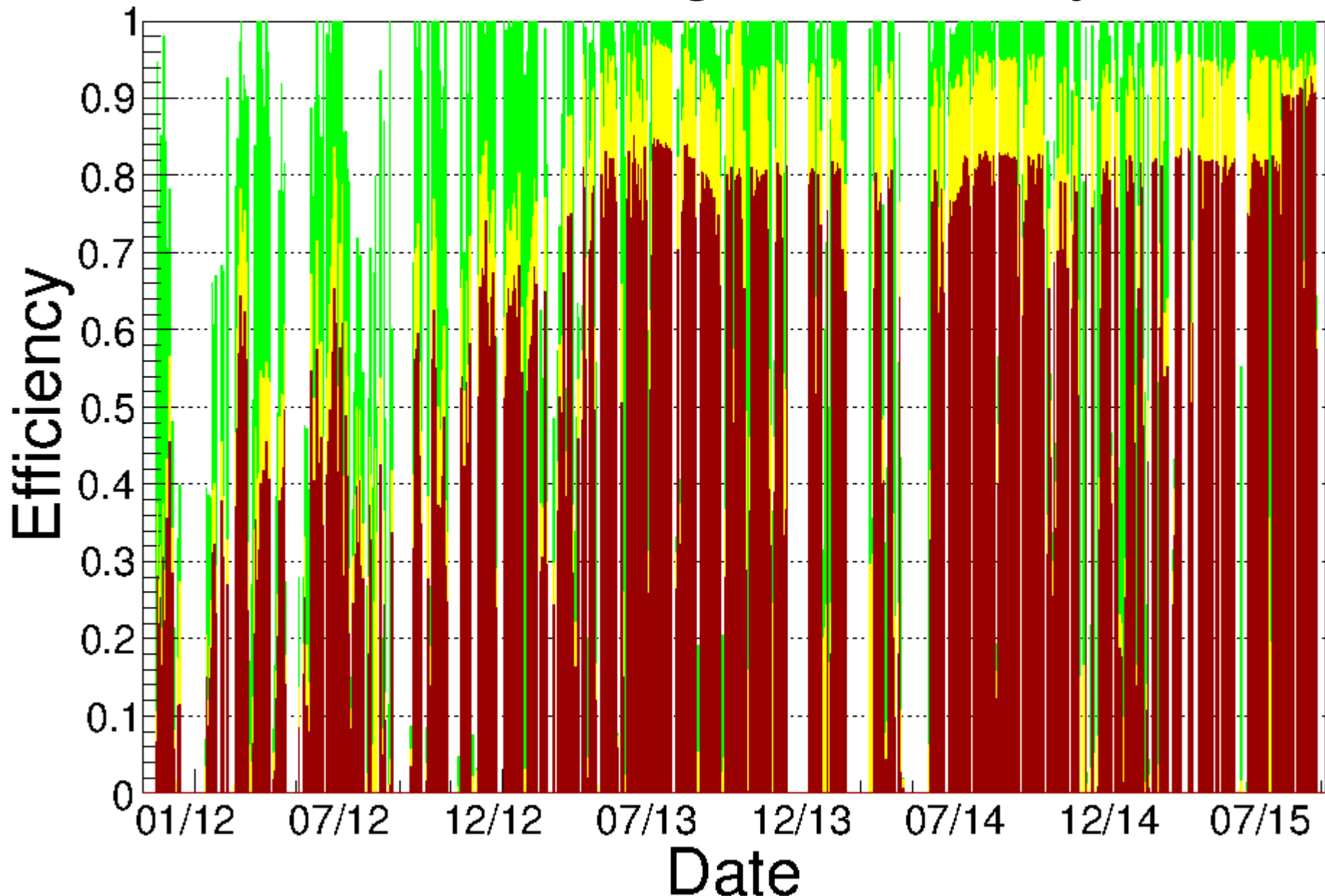


FACT – Ideal Monitoring Telescope

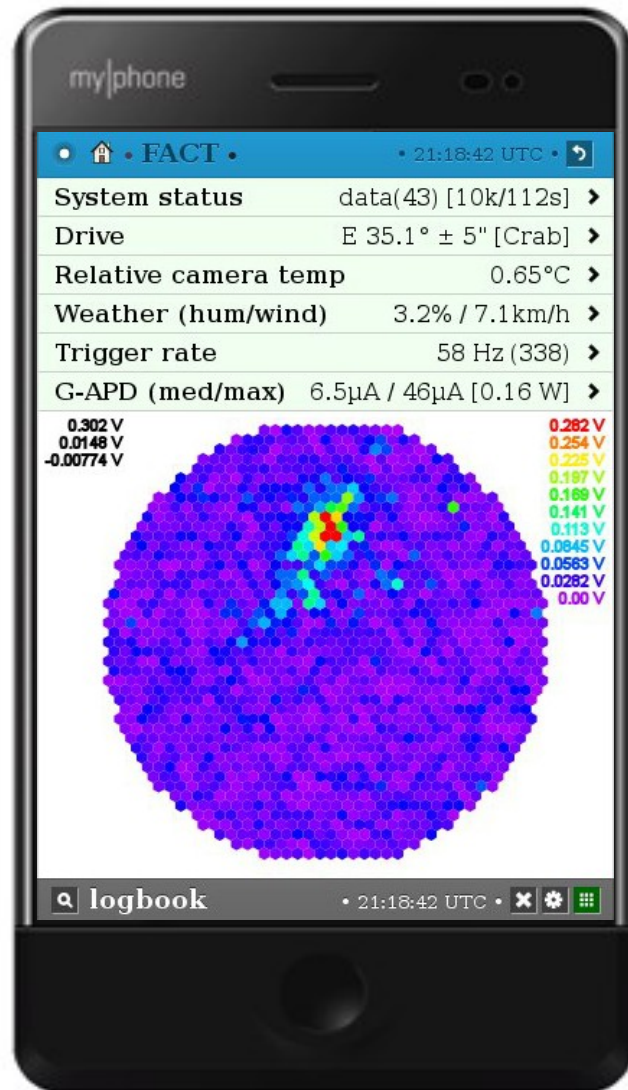


- Observations during strong moon light
 - Larger duty cycle
 - More complete data sample
- G-APDs robust and stable
 - Stable telescope performance
 - Remote and automatic operation

Data Taking Efficiency



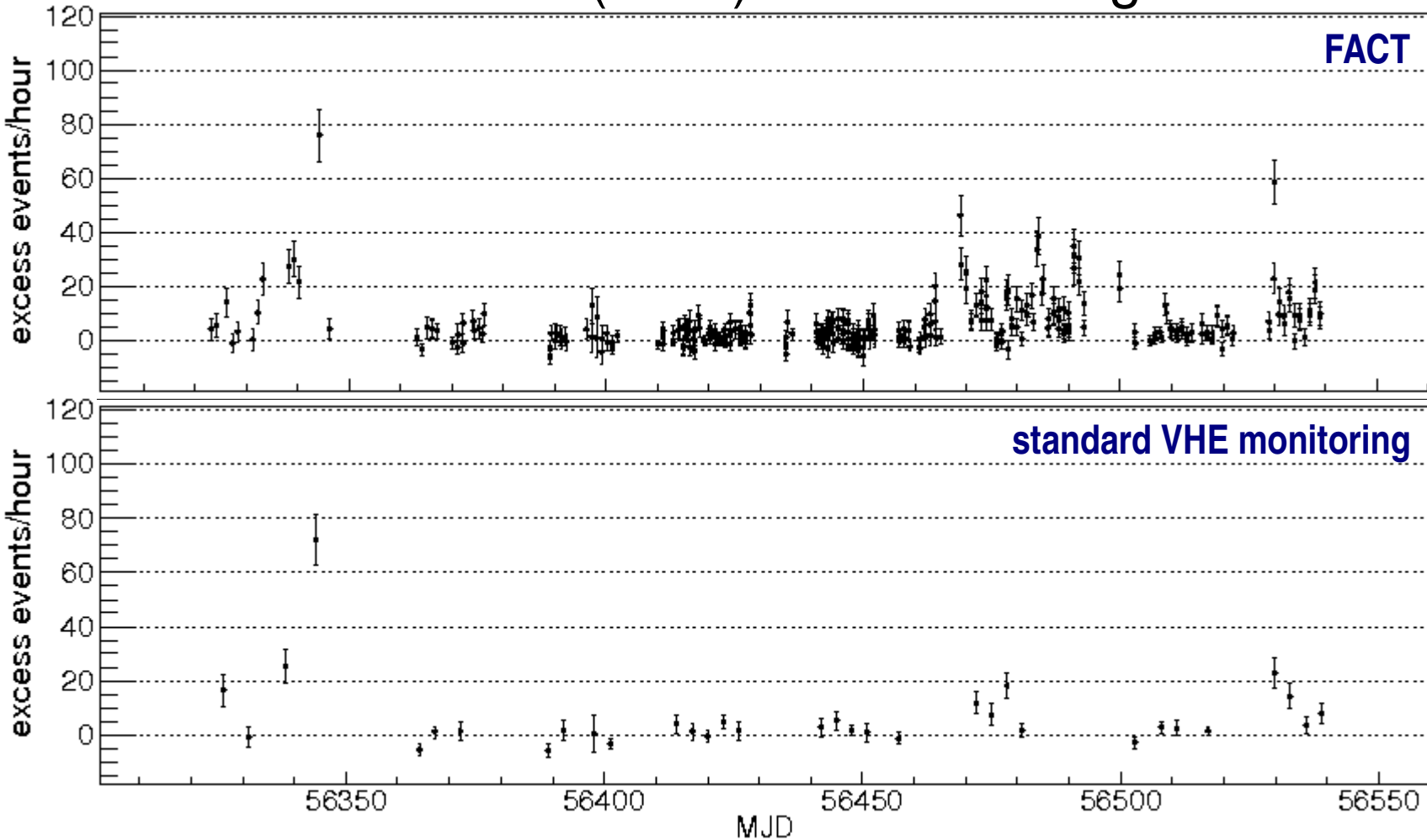
FACT – Ideal Monitoring Telescope



- Observations during strong moon light
 - Larger duty cycle
 - More complete data sample
- G-APDs robust and stable
 - Stable telescope performance
 - Remote and automatic operation
 - High data taking efficiency

Monitoring at TeV Energies

Mrk 501 (2013) 1-hour-binning

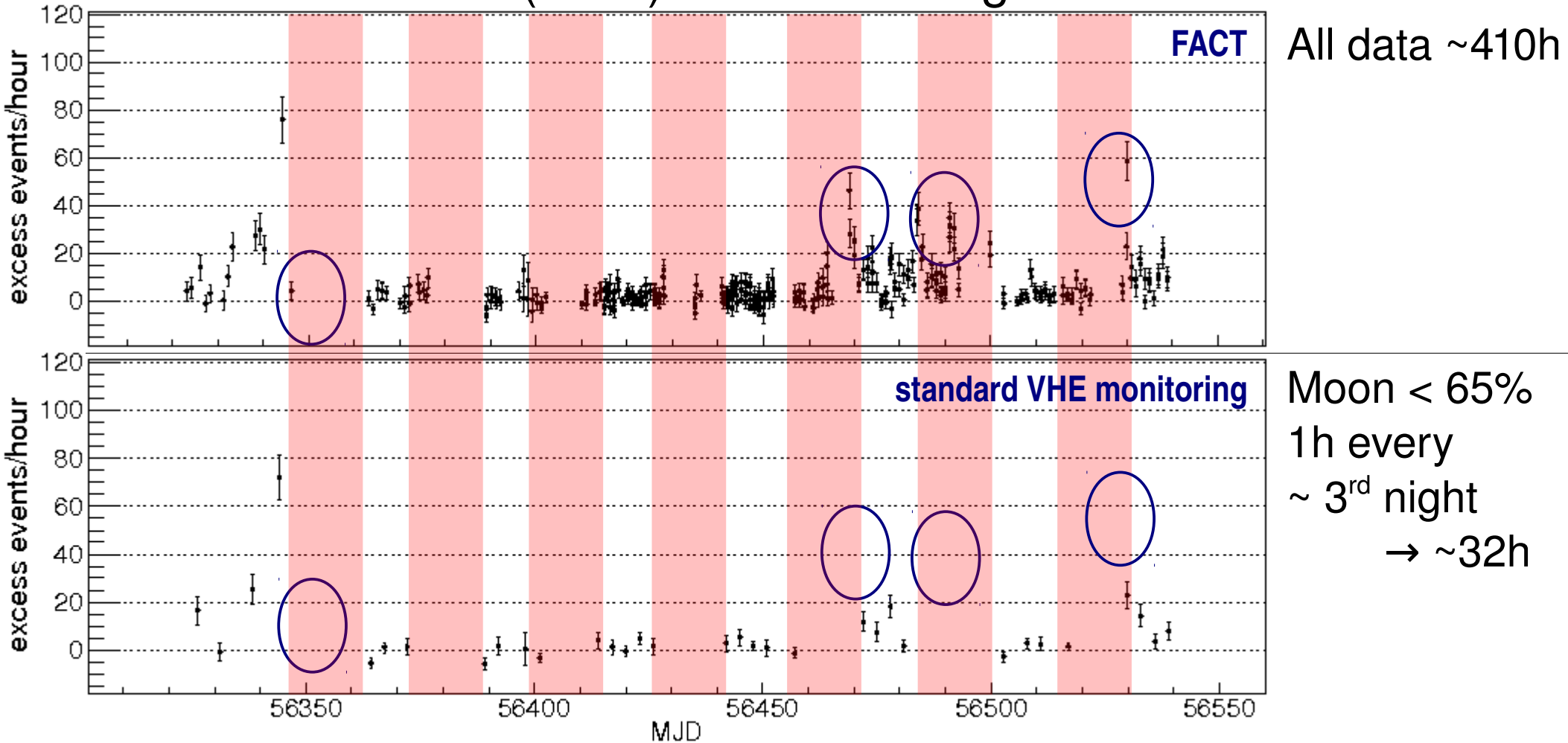


All data ~410h

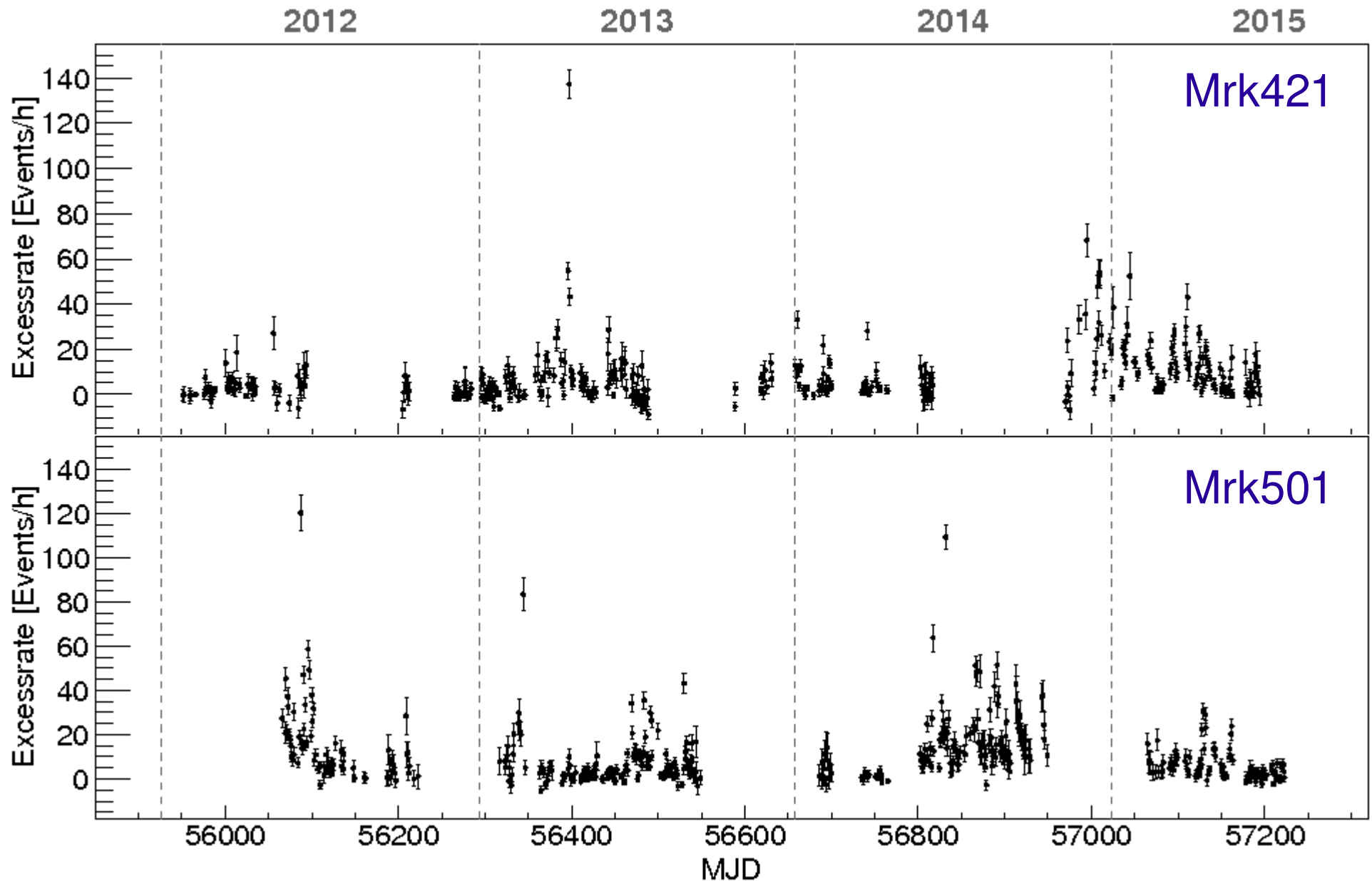
Moon < 65%
1h every
~ 3rd night
→ ~32h

Monitoring at TeV Energies

Mrk 501 (2013) 1-hour-binning



3.5 Years of Monitoring



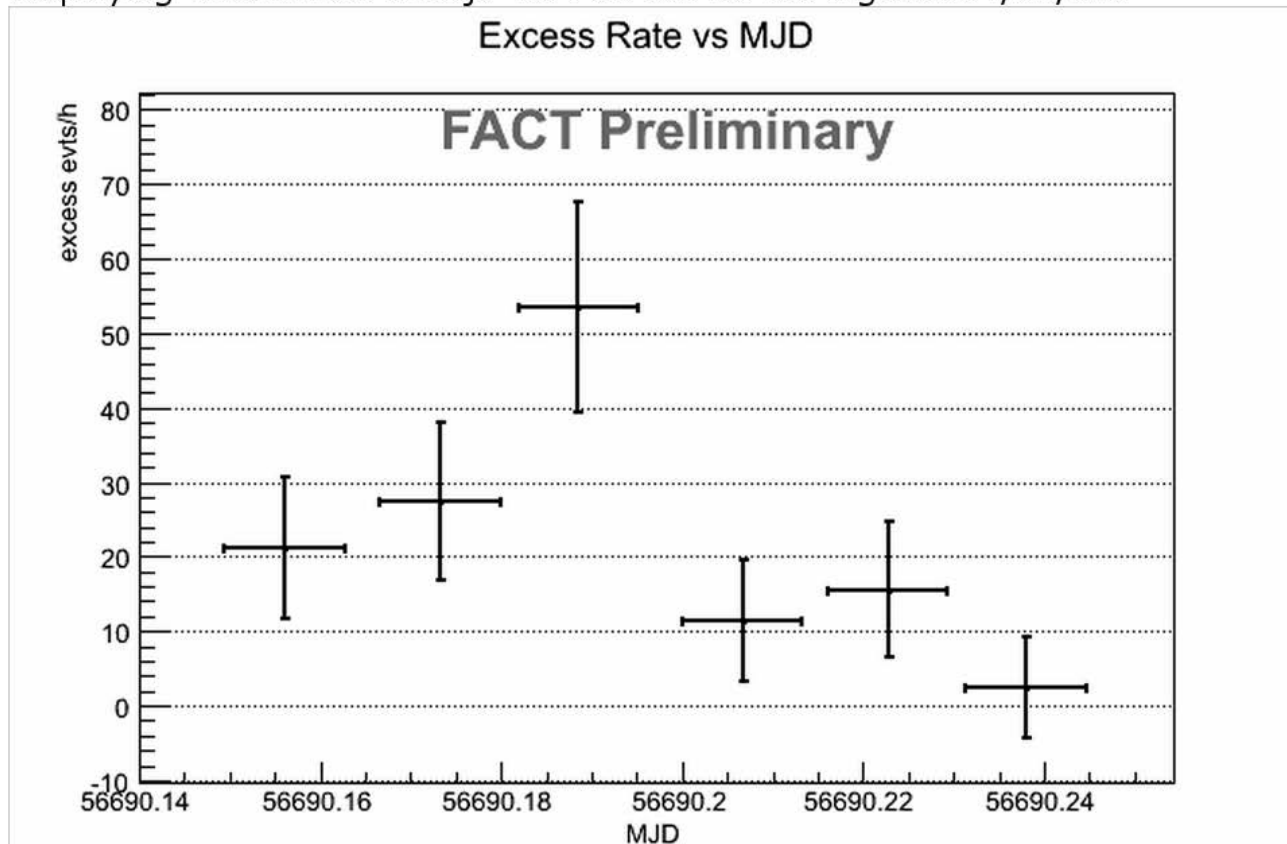
Quick Look Analysis

<http://www.fact-project.org/monitoring>

FACT Quick Look Analysis

Select date source
Select time binning and range

Displaying 'excess rate vs mjd' for Mrk 421 for the night 2014/02/01.



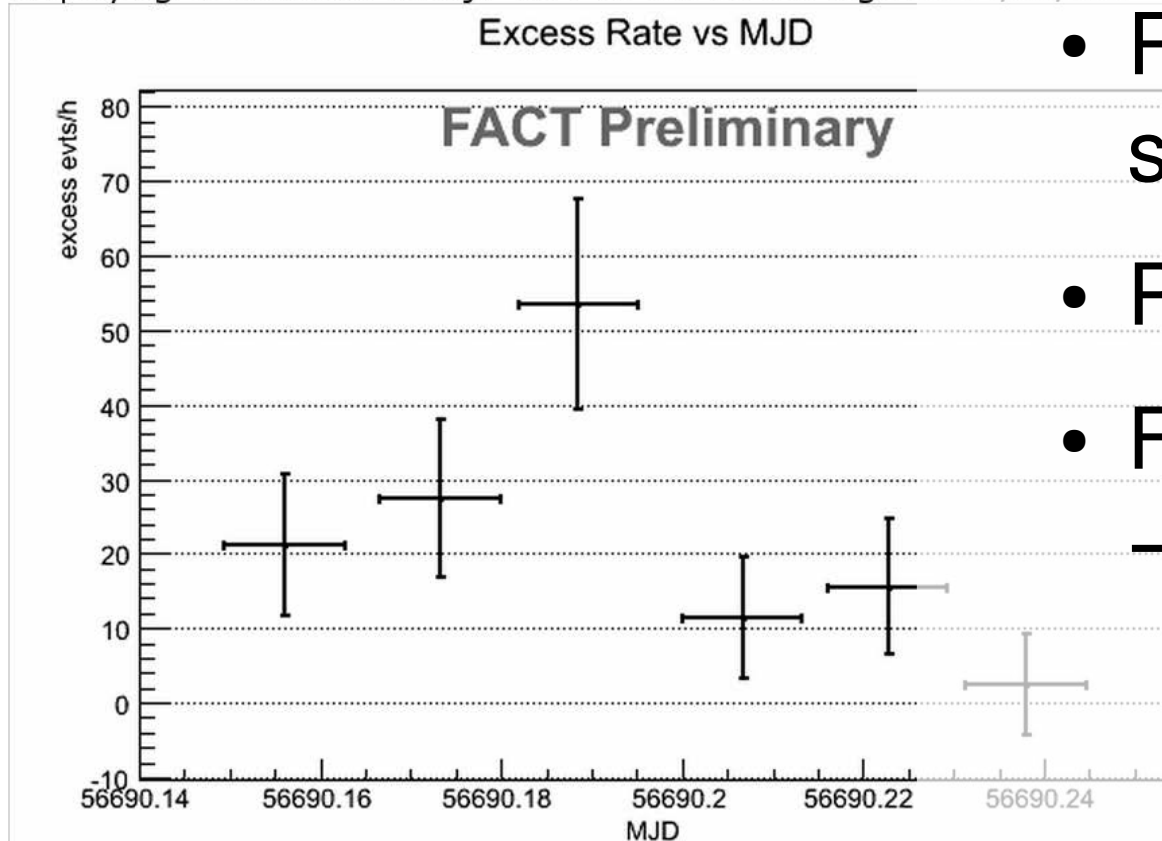
Quick Look Analysis

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FACT Quick Look Analysis

Select date source
Select time binning and range

Displaying 'excess rate vs mjd' for Mrk 421 for the night 2014/02/01.



- Since 2012/12/12
- Results within the same night
- Publicly available
- Flare alerts
→ MWL and ToO observations



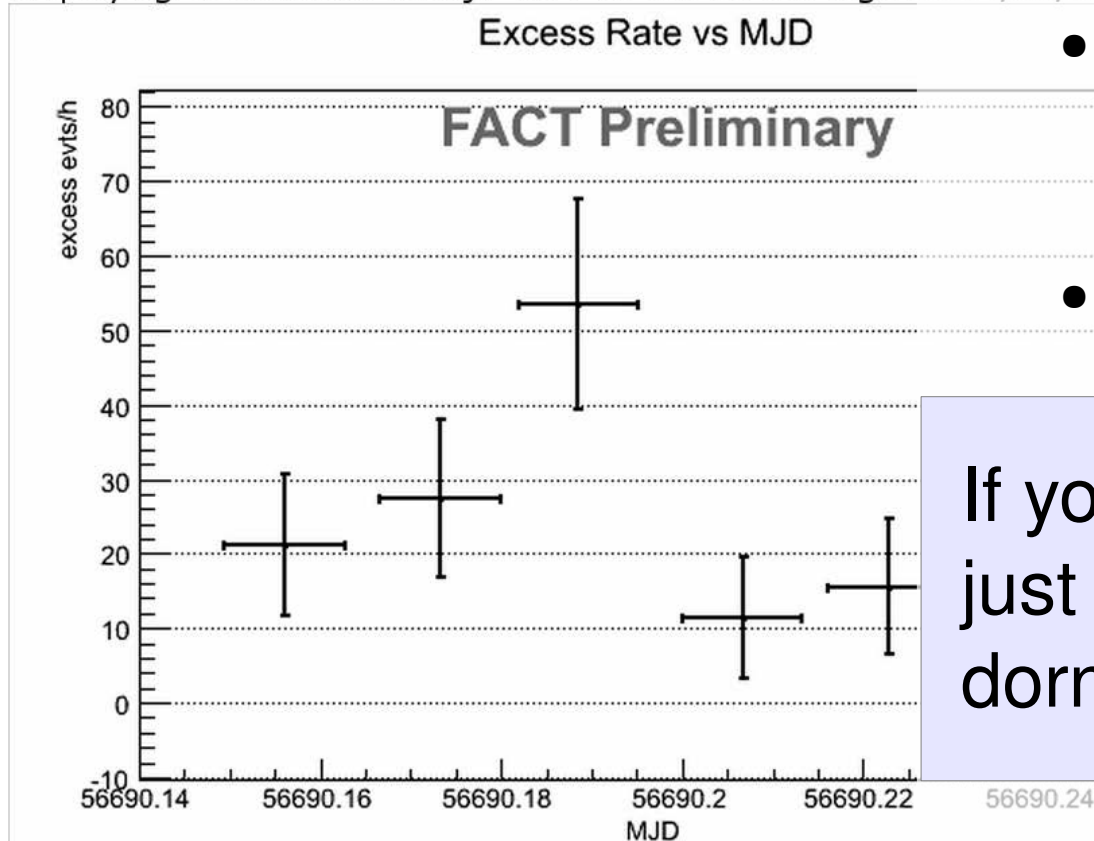
Quick Look Analysis

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FACT Quick Look Analysis

Select date source
Select time binning and range

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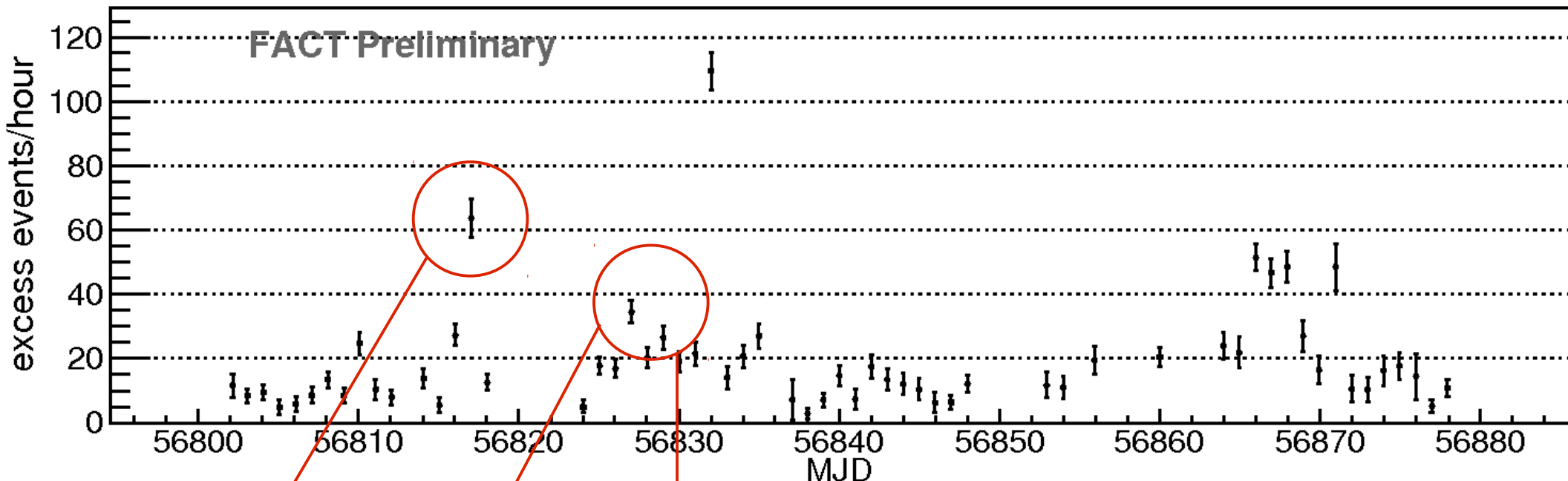
- Since 2012/12/12
- Results within the same night
- Publicly available

If you want to have the data,
just contact me:
dorner@astro.uni-wuerzburg.de



Mrk 501 – Flare Alerts

Excess rate curve from QLA: 1.6.-10.8.2014



Flare Alert
8./9. June 2014

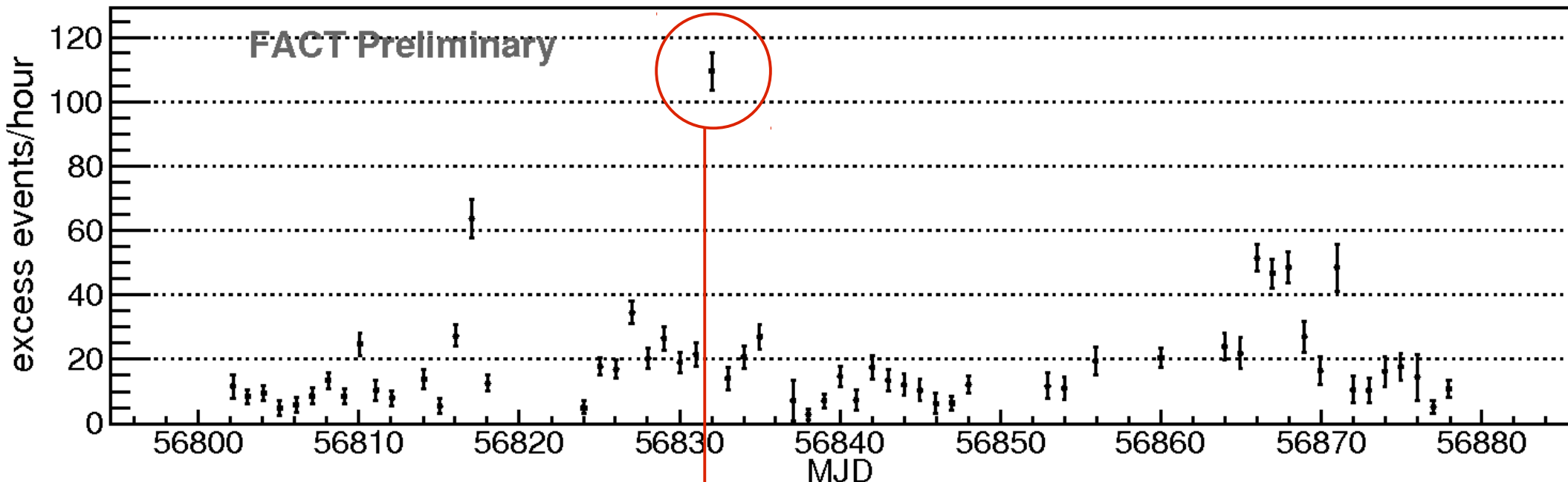
Flare Alert
20./21. June 2014

Flare Alert
18./19. June 2014

→ Observations with other instruments

Mrk 501 – Flare Alerts

Excess rate curve from QLA: 1.6.-10.8.2014

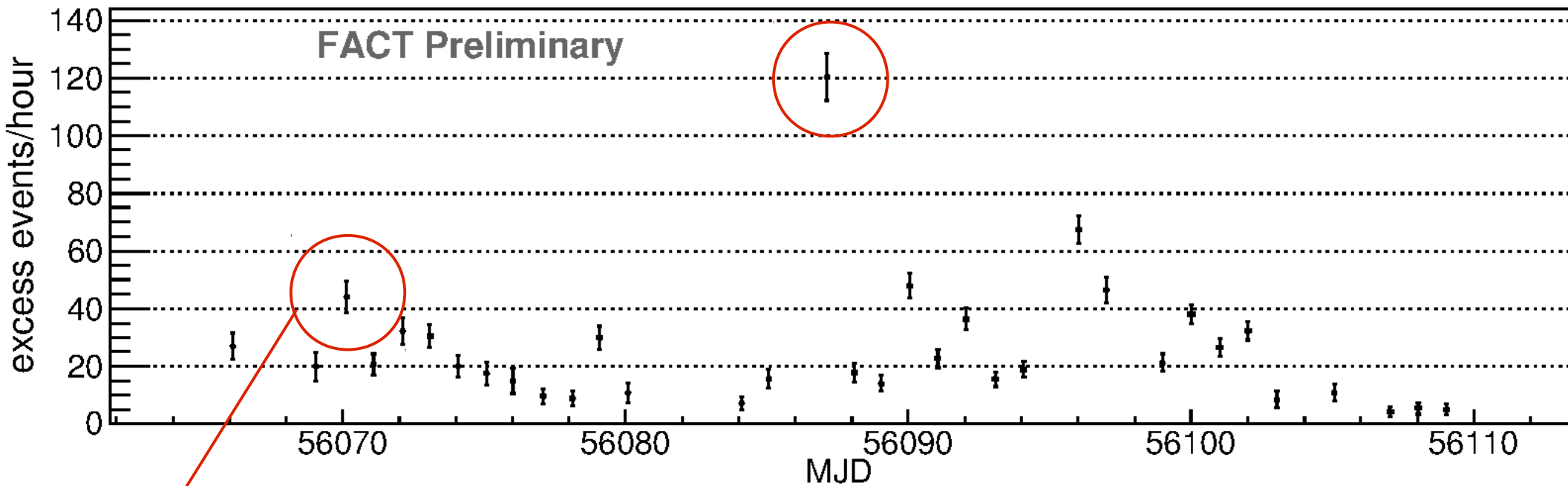


Flare Alert
23./24. June 2014

→ Astronomer's
Telegram #6268

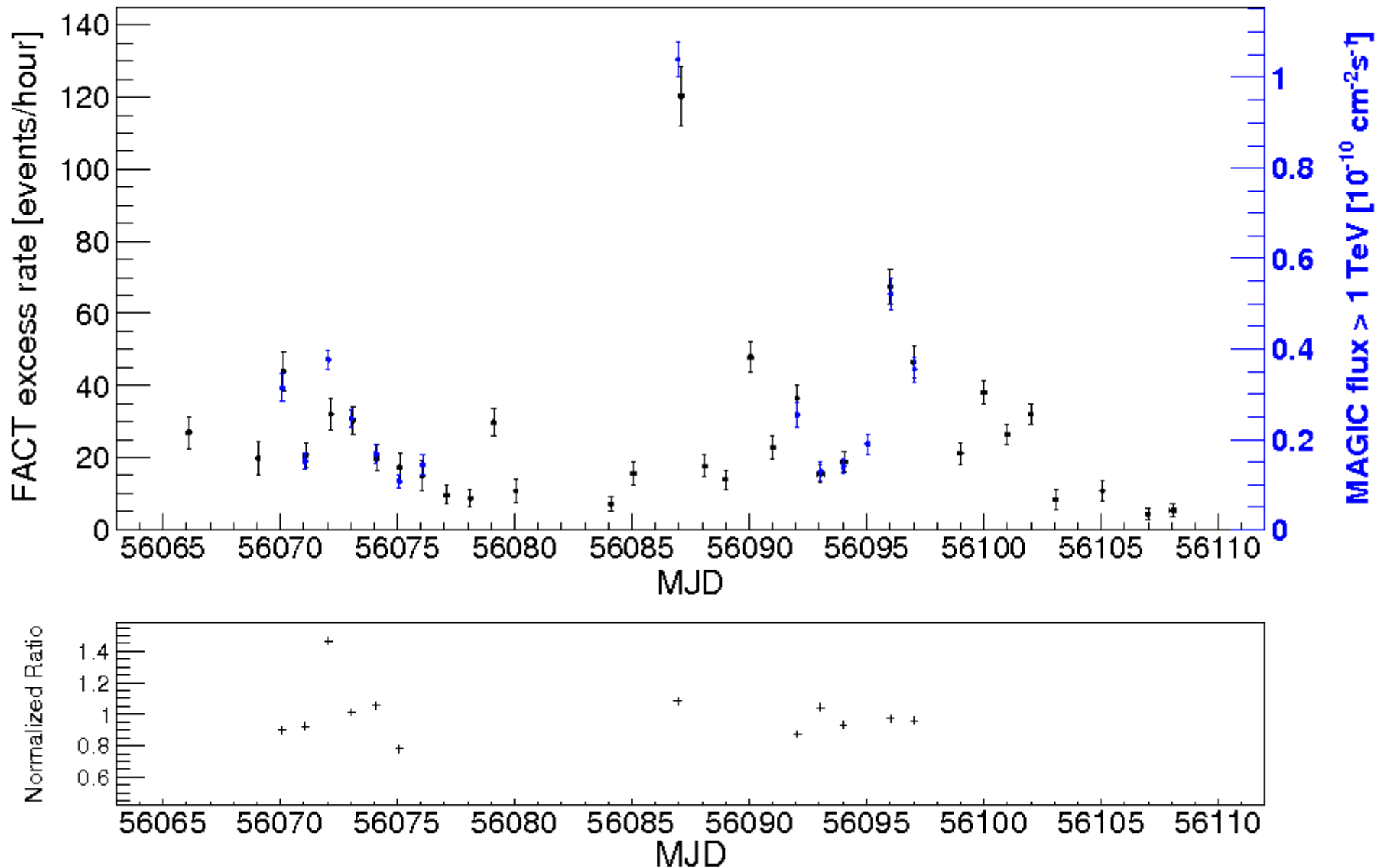
Mrk 501 – Flares May/June 2012

18.5.-30.6.2012

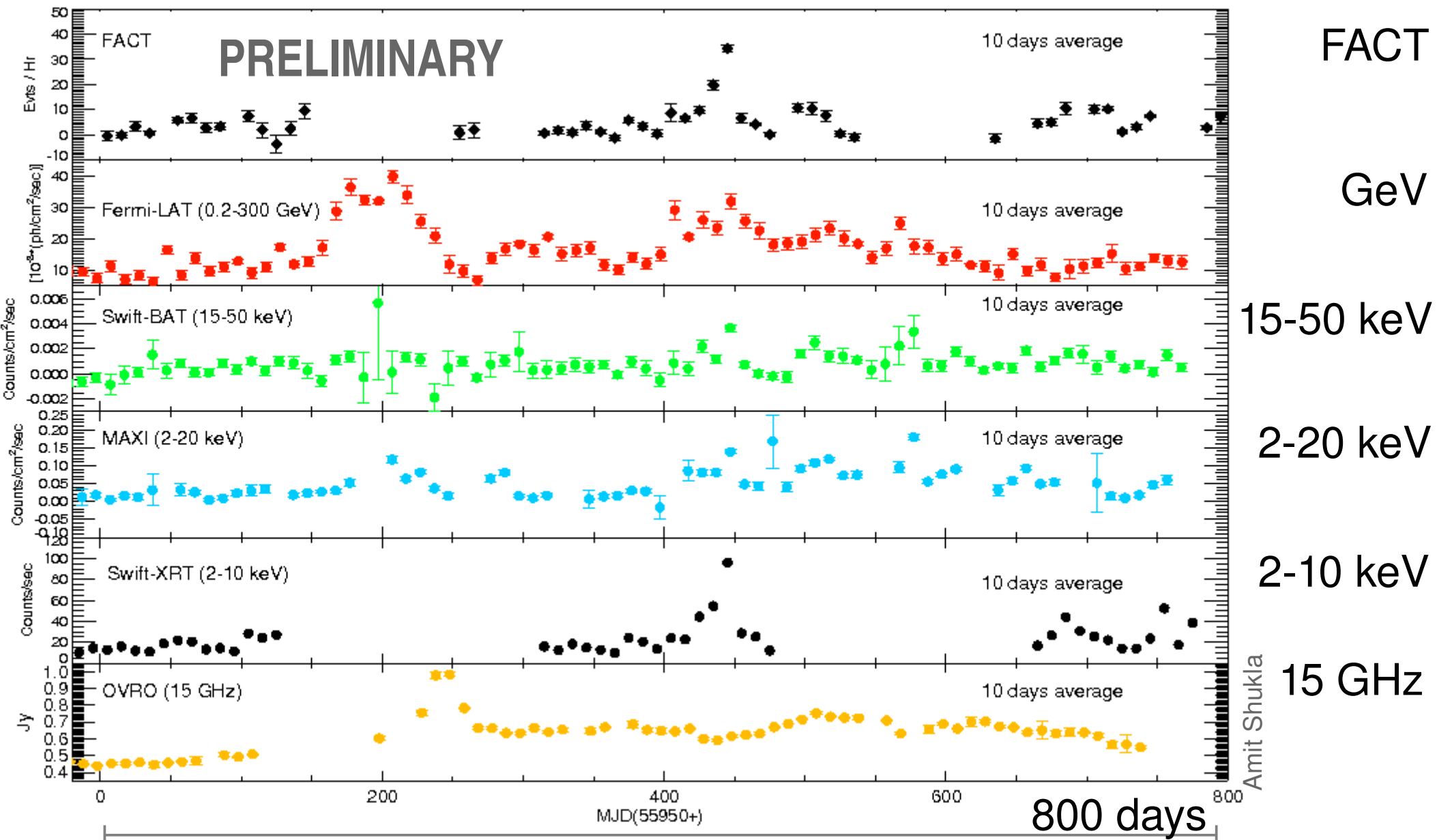


Triggered MAGIC observations

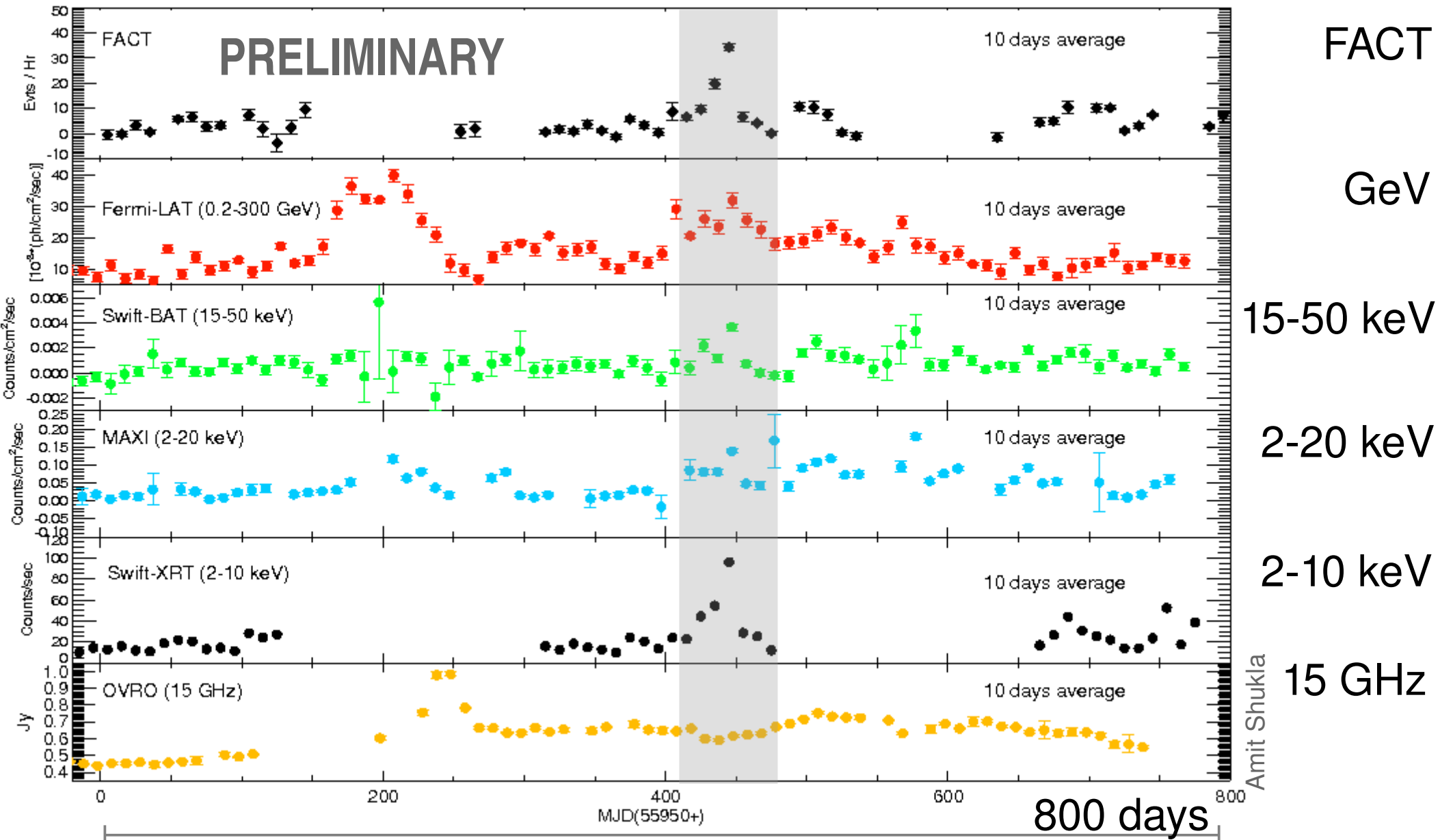
Mrk 501 – Flares May/June 2012



Multi-Wavelength Campaign Mrk 421

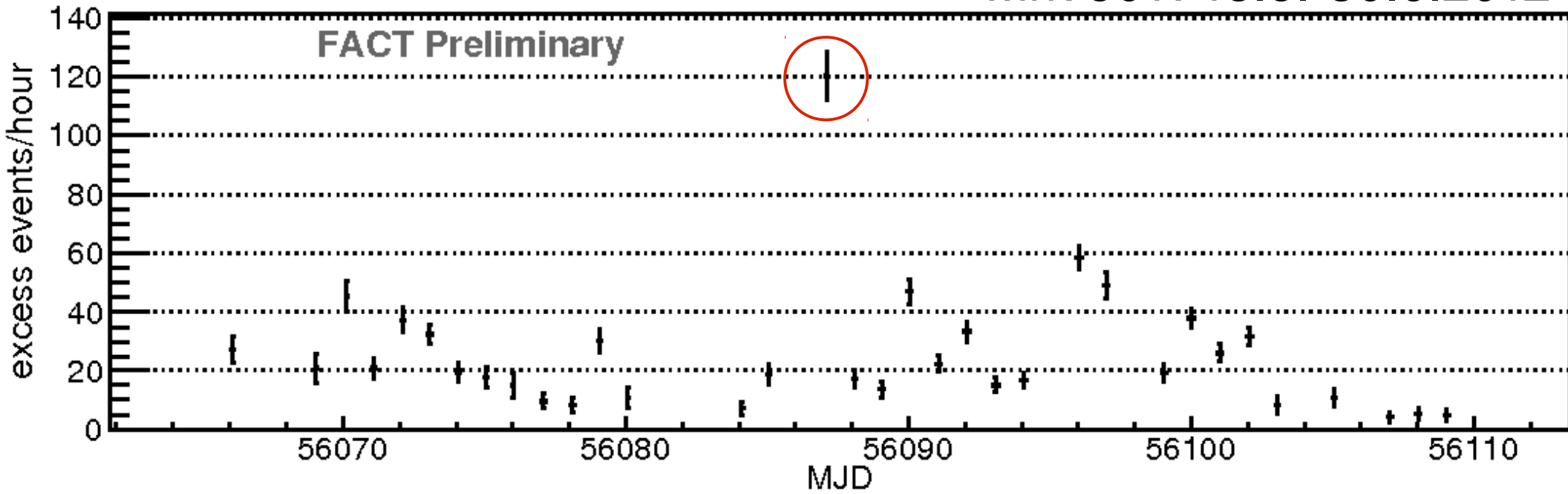


Multi-Wavelength Campaign Mrk 421



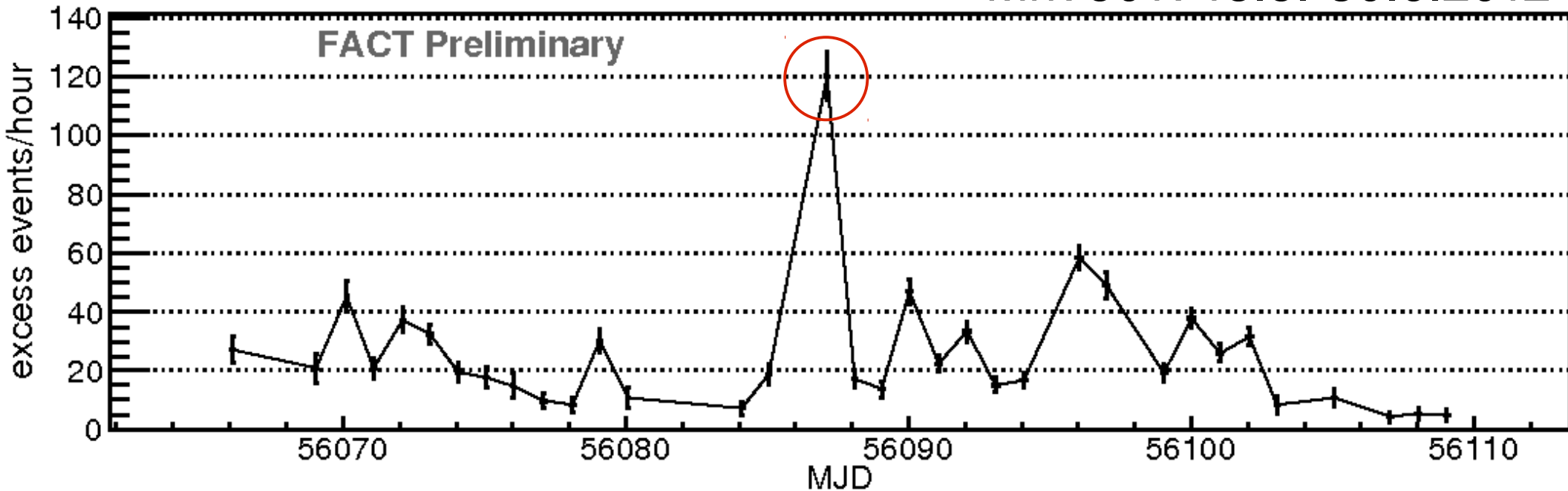
Outlook

Mrk 501: 18.5.-30.6.2012



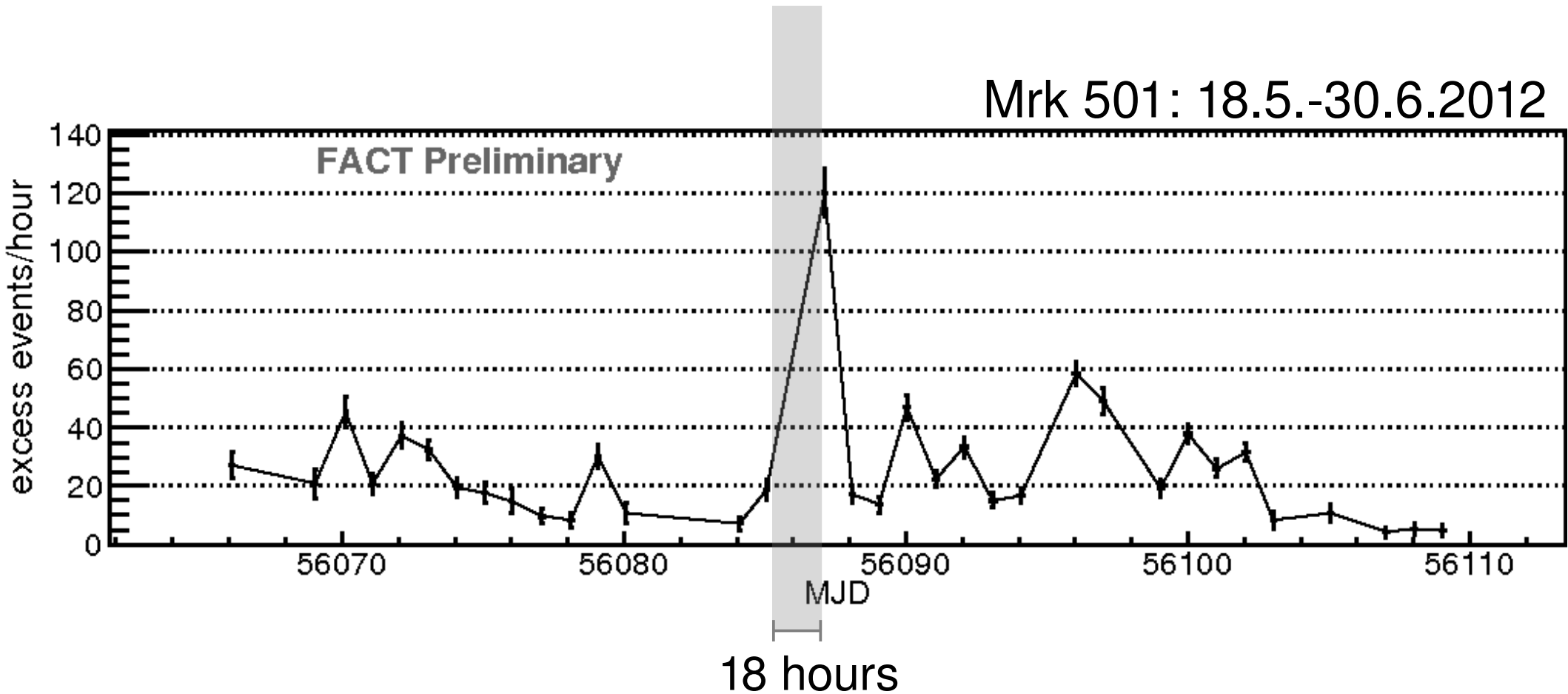
Outlook

Mrk 501: 18.5.-30.6.2012



Outlook

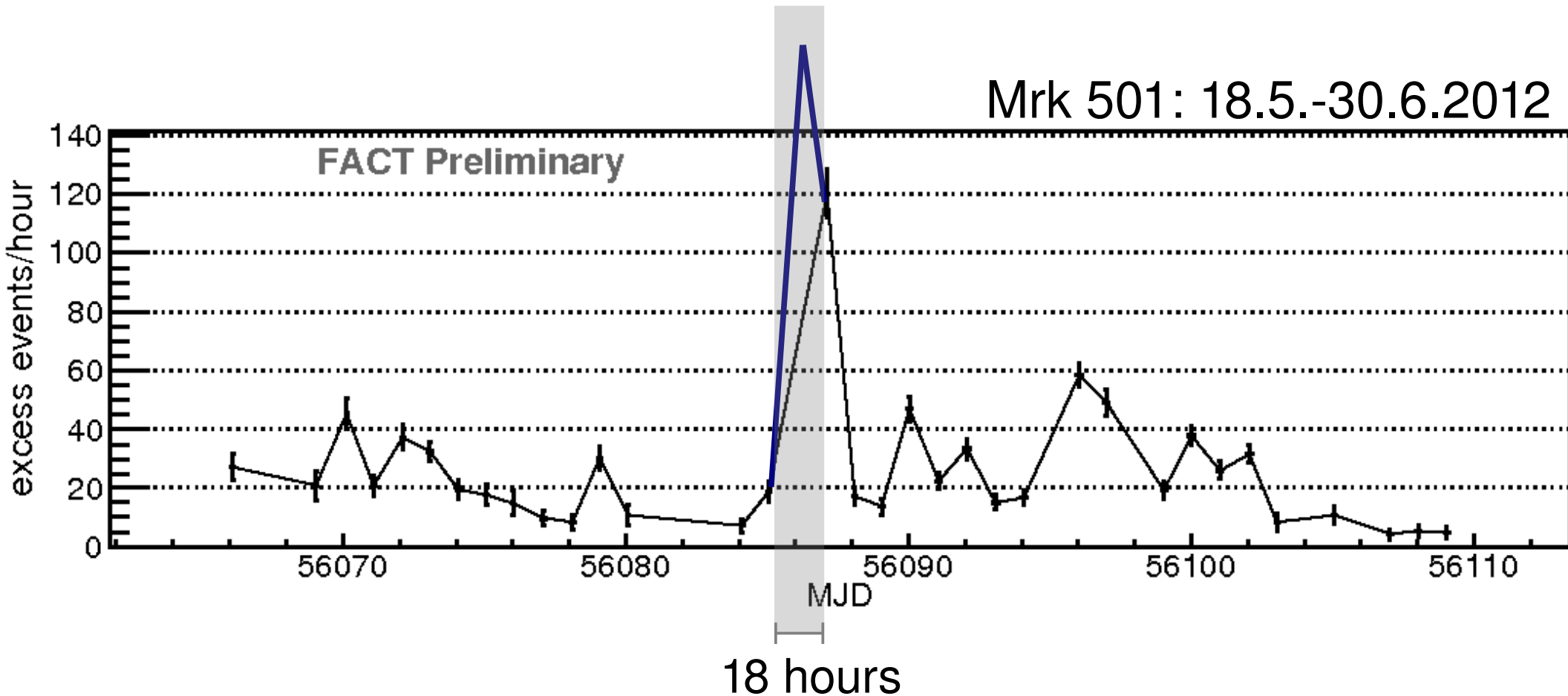
Mrk 501: 18.5.-30.6.2012



Gaps due to daytime

Outlook

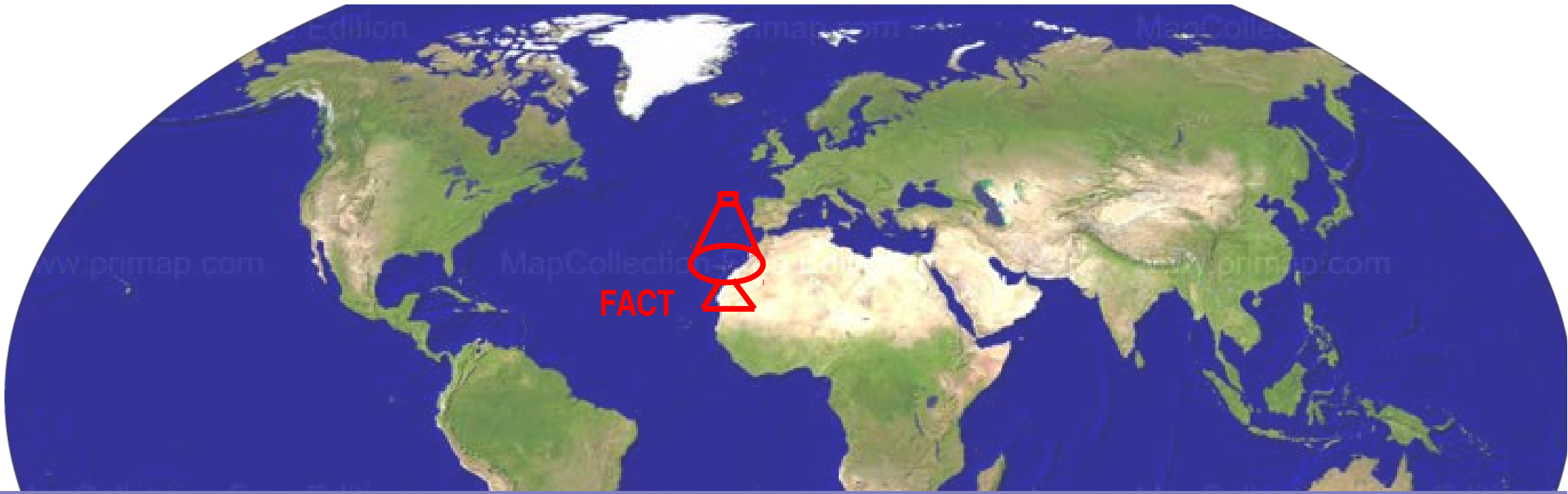
Mrk 501: 18.5.-30.6.2012



Gaps due to daytime

→ continuous monitoring around the globe needed

Continuous Monitoring

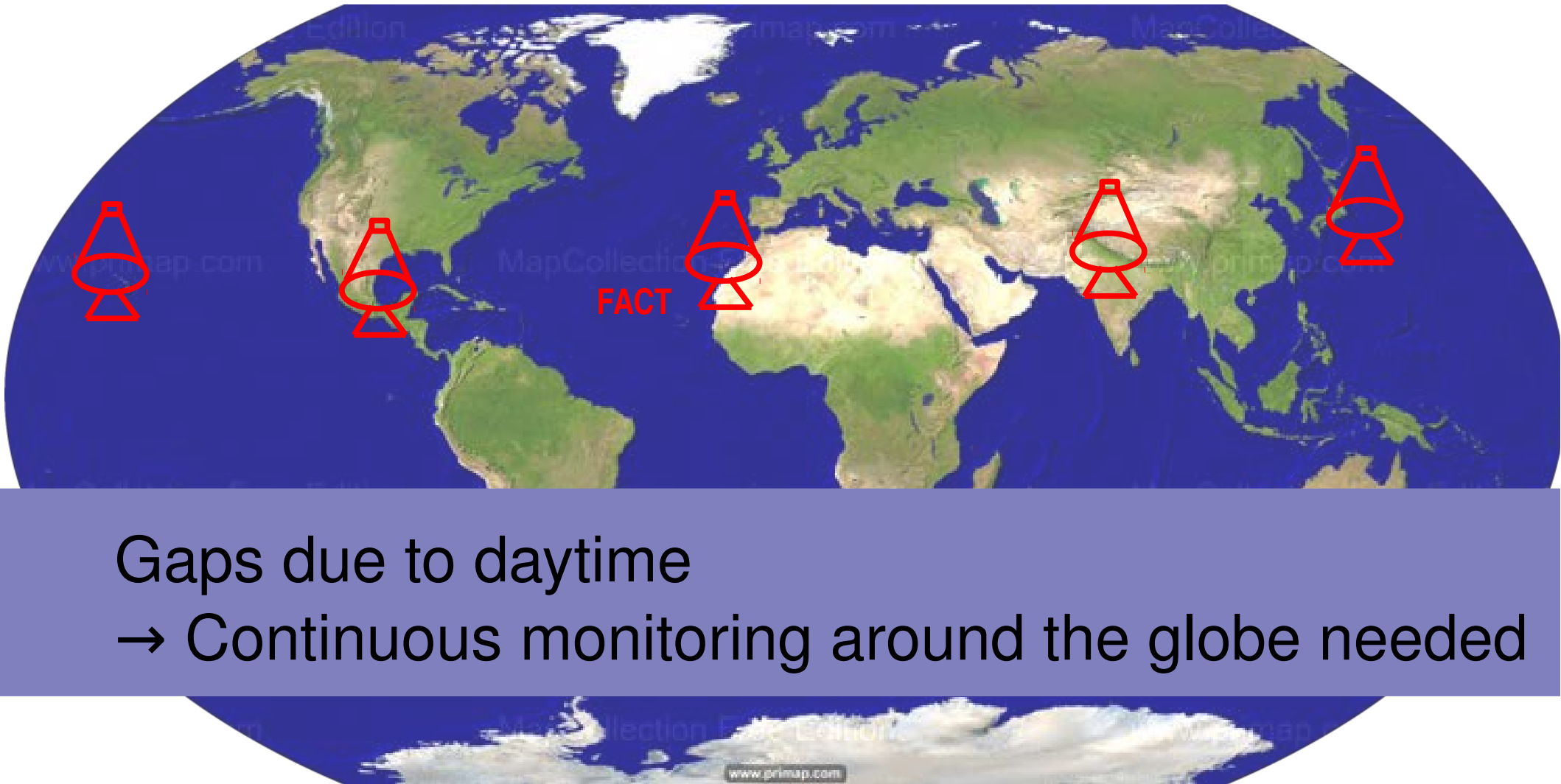


Gaps due to daytime

DWARF Network (M. Backes et. al ICRC 2009)



Continuous Monitoring



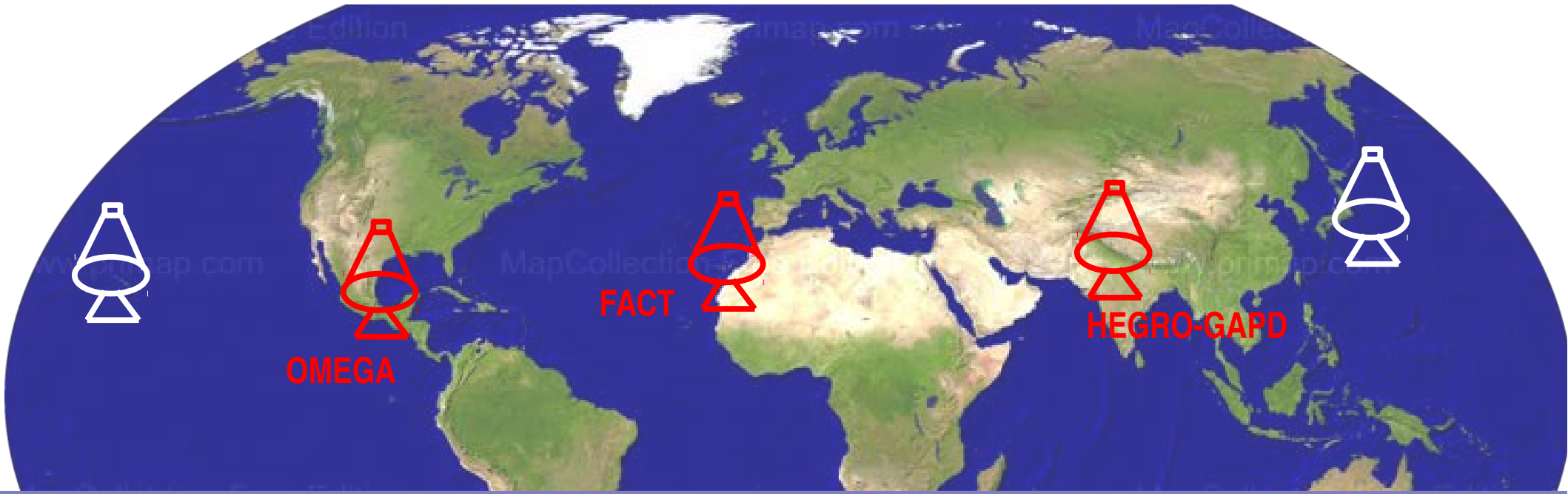
Gaps due to daytime

→ Continuous monitoring around the globe needed

DWARF Network (M. Backes et. al ICRC 2009)



Continuous Monitoring



Gaps due to daytime

→ Continuous monitoring around the globe needed



Continuous Monitoring



OMEGA



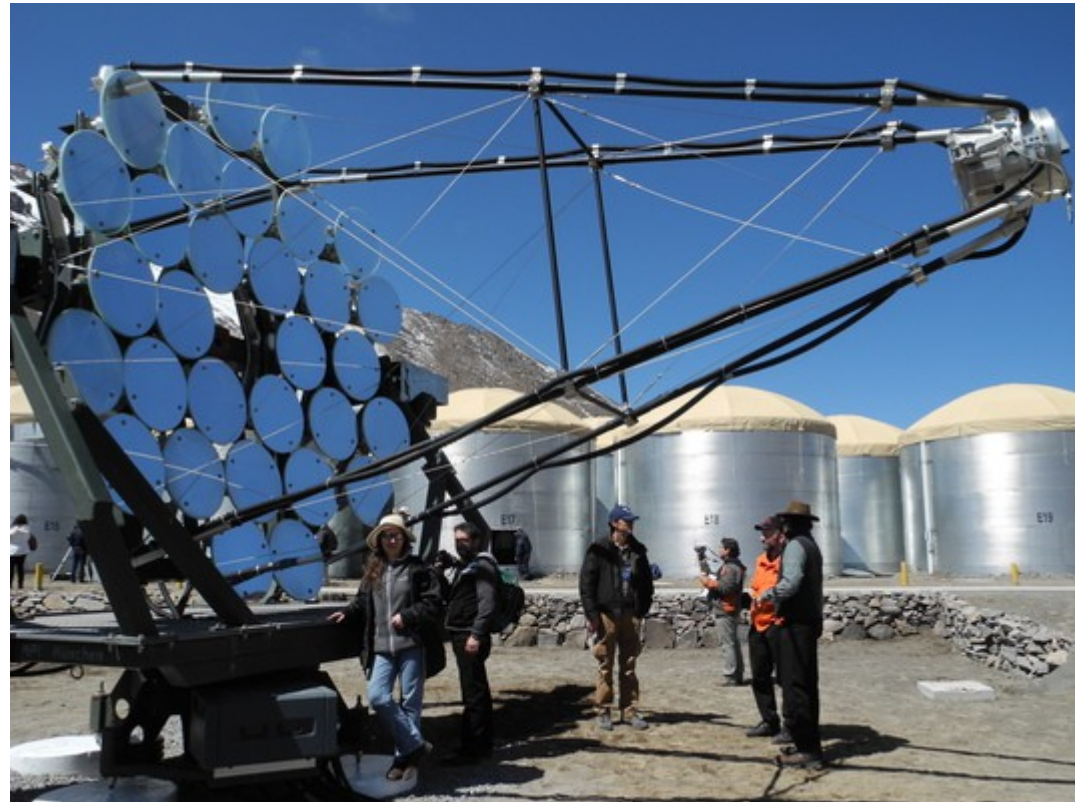
Gaps due to data
→ Continuous

ded

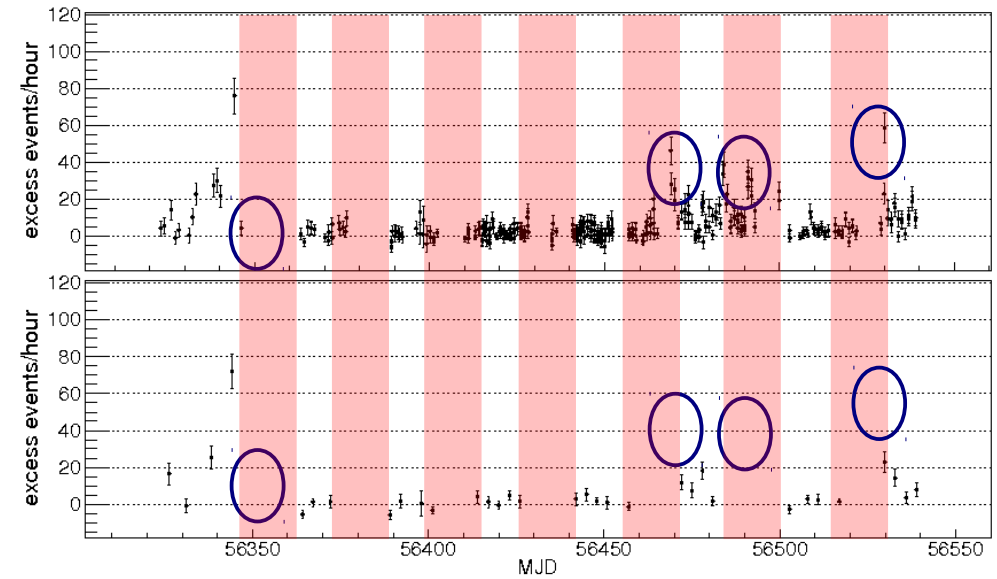
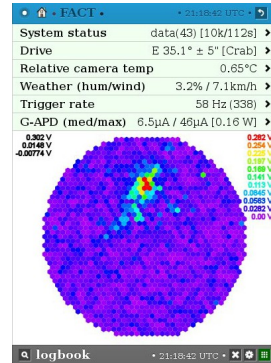


Second Telescope: M@TE

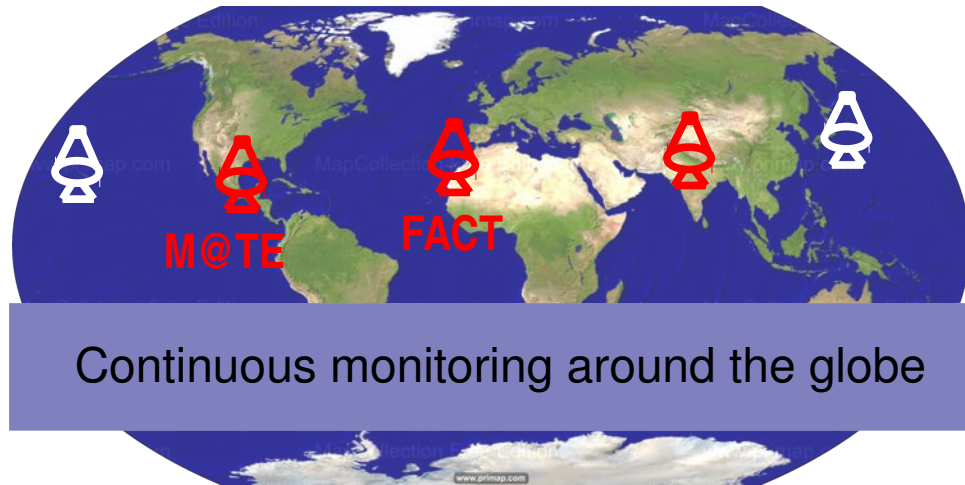
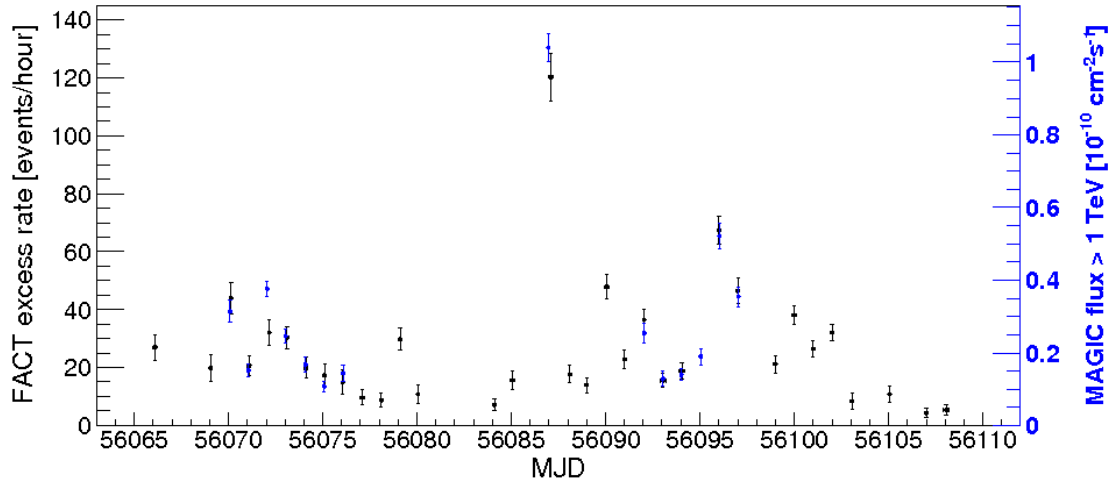
- M@TE:
Monitoring at TeV Energies
- 2 mounts from OMEGA project available
- Mexico:
two possible sites:
5 or 7 hours from La Palma
- Goal: Equip mount with improved SiPM camera
→ close more gaps in
TeV monitoring



Summary

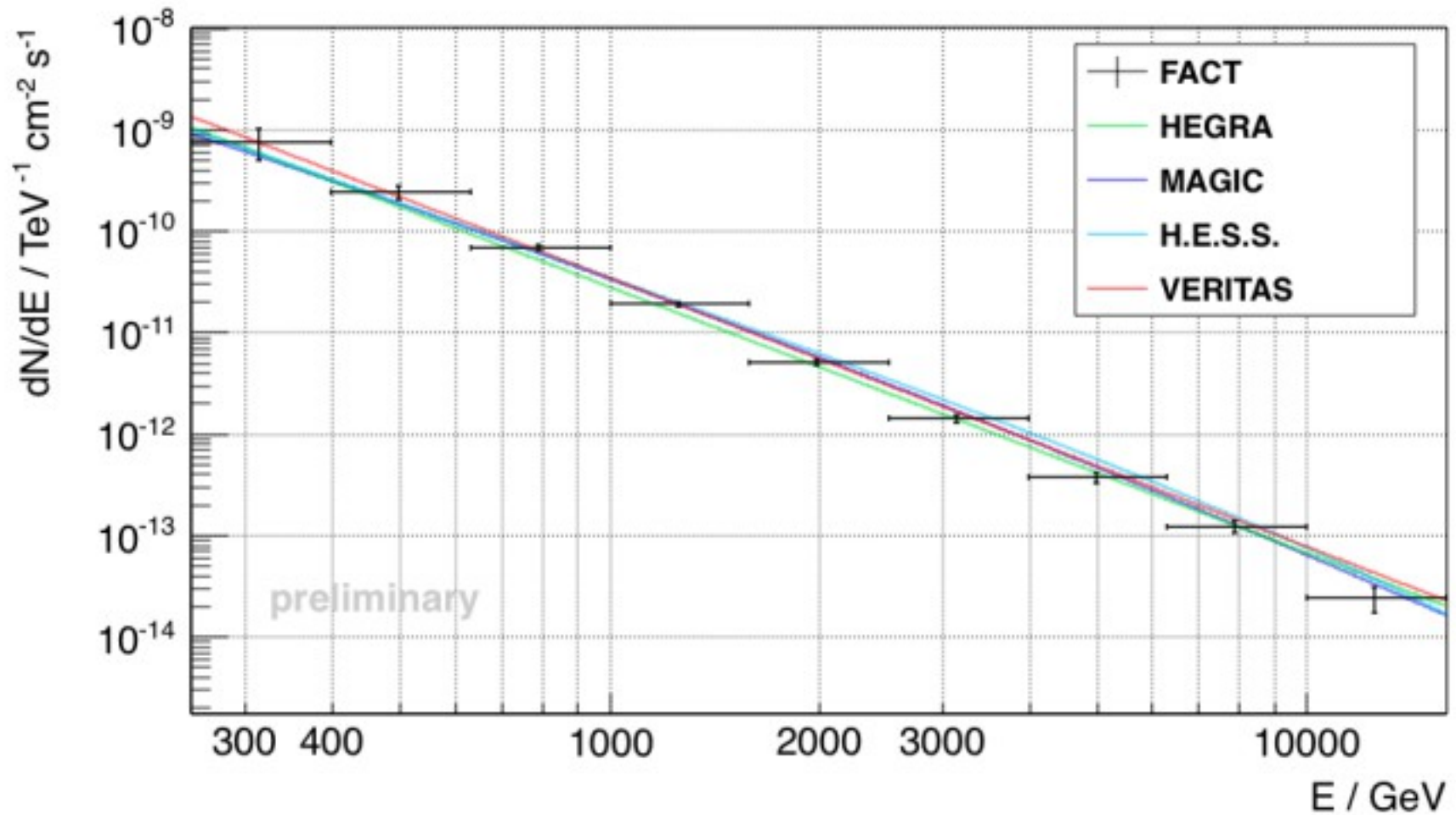


<http://www.fact-project.org/monitoring>



Backup

Crab Spectrum



Quick Look Analysis

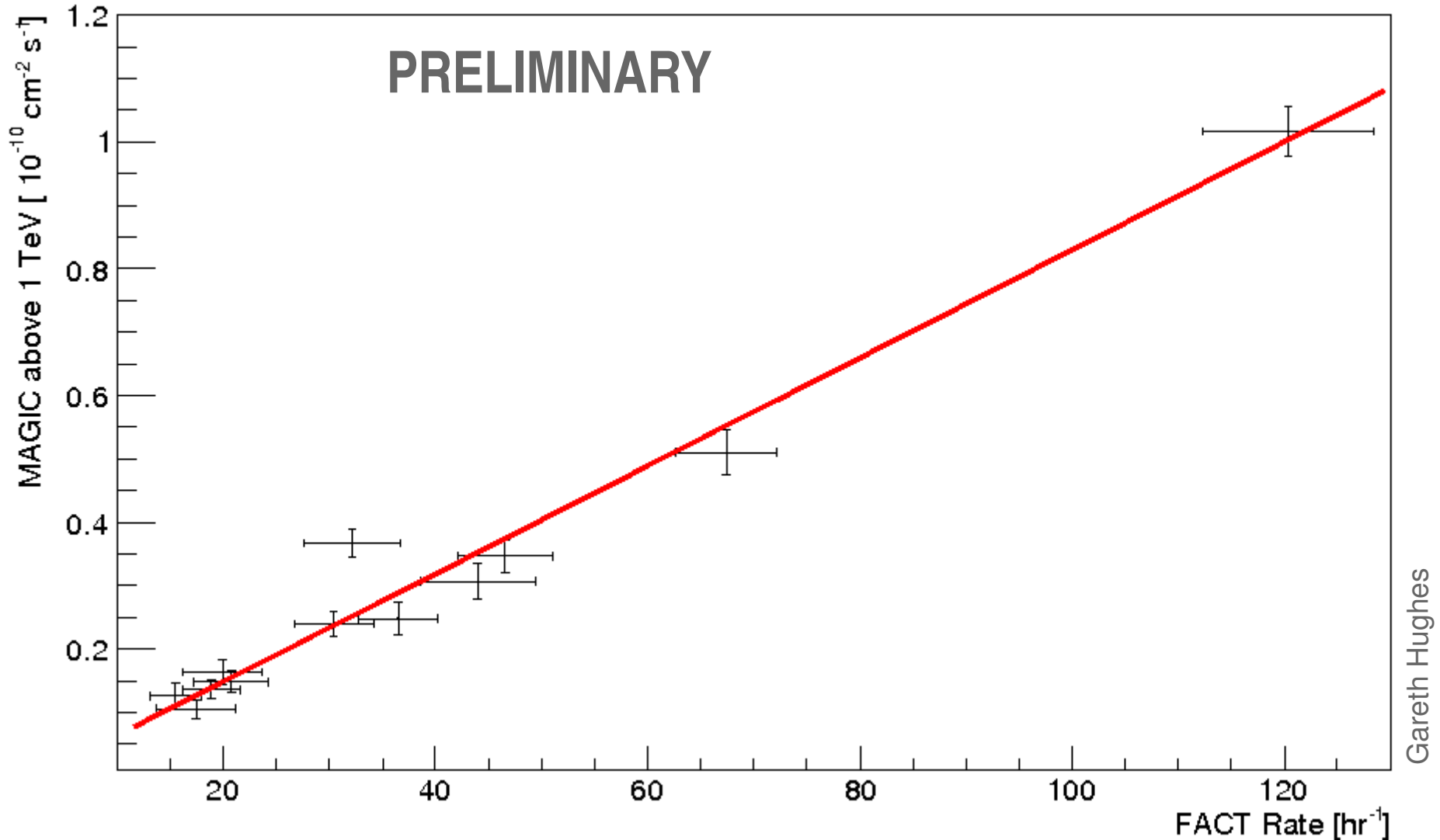
- Quick look analysis since 2012/12/12
- Results publicly available since 2013/09/01
- Official flare alerts since 2014/03/18

<http://www.fact-project.org/monitoring>

- Results available within the same night
- Flare alerts to other instruments
- Target-of-Opportunity programs



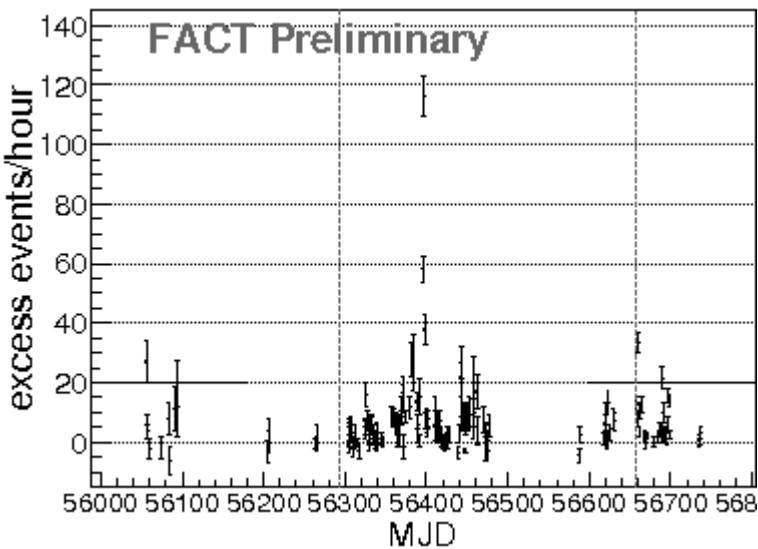
Mrk 501 – Flares May/June 2012



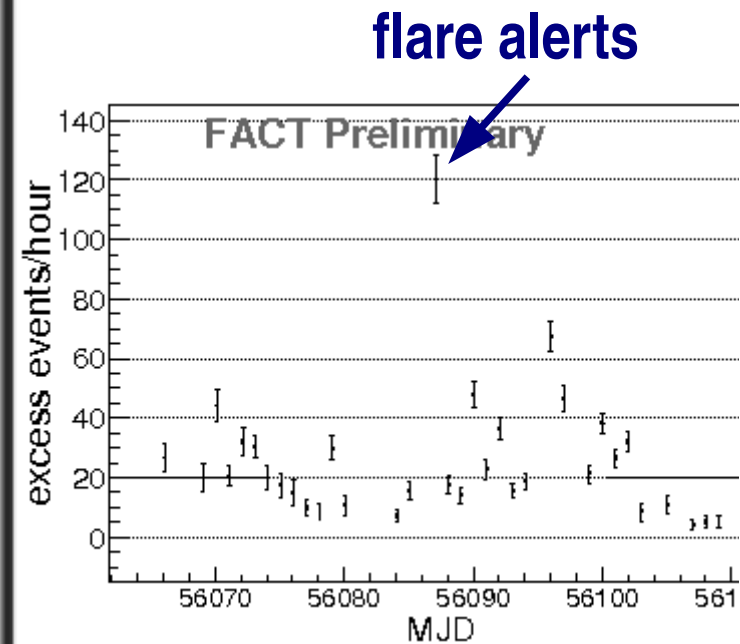
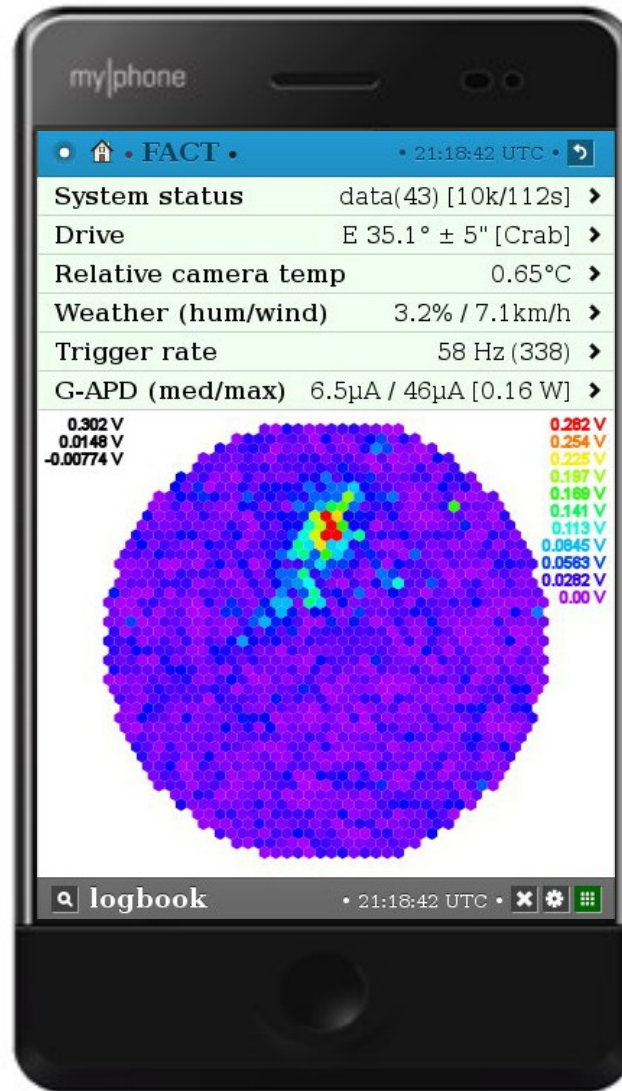
MAGIC: <http://143.107.180.38/indico/contributionDisplay.py?contribId=934&sessionId=3&confId=0>

Check out our monitoring results!

<http://www.fact-project.org/monitoring>



3.5 years monitoring



[JINST 9(2014) P10012]
arXiv:1502.02582