

# DarkSide: recent results with underground argon and outlook

**Davide Franco**

**APC**

on behalf of the  
DarkSide Collaboration

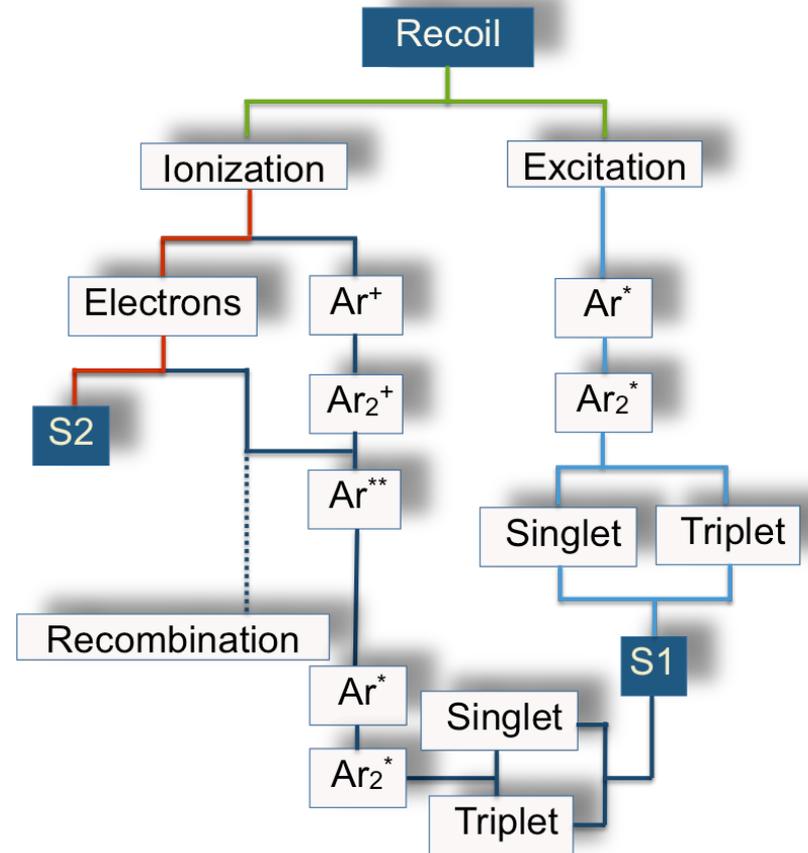
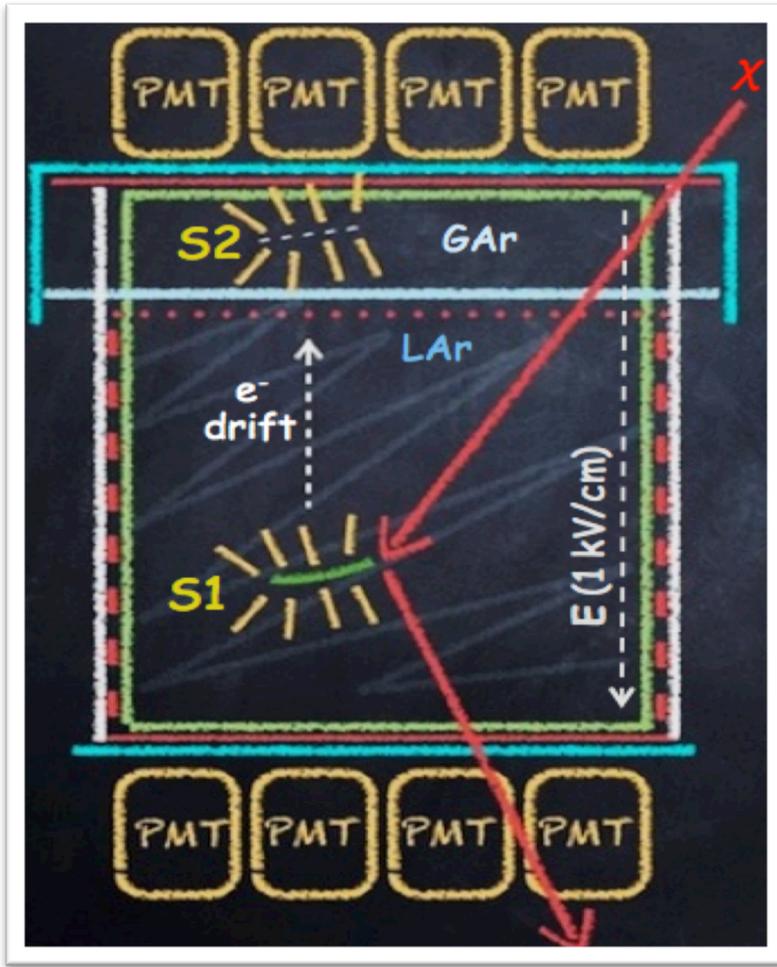
**Prospects in Low Mass Dark Matter**

30<sup>th</sup> November and 1<sup>st</sup> December 2015

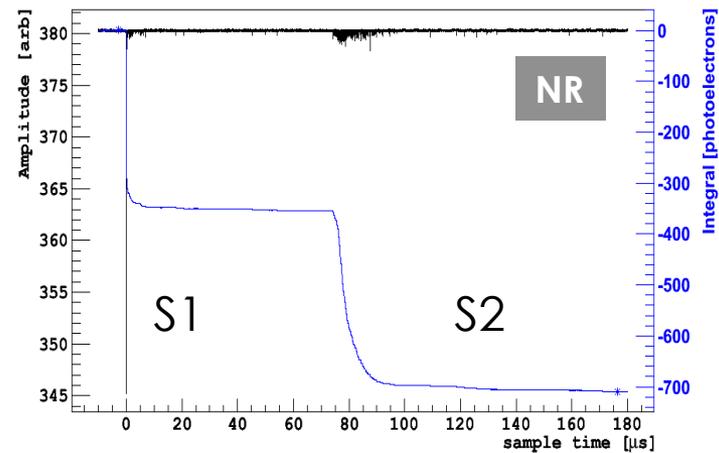
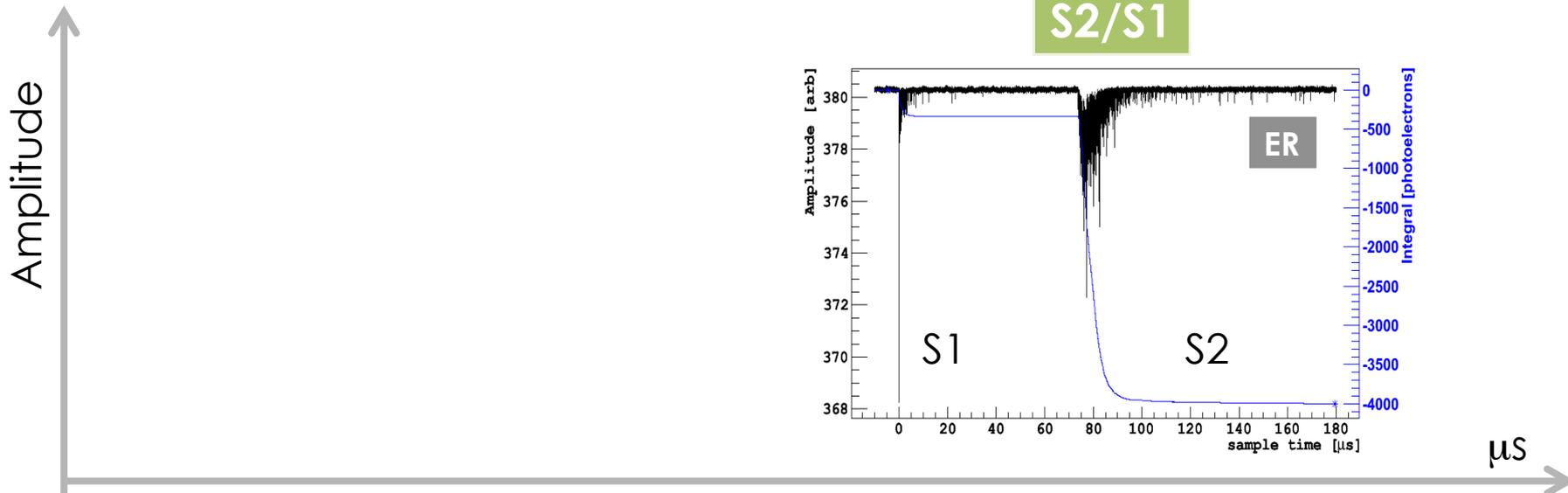
MPI-P Munich

- WIMP dark matter search using **direct detection**
- **Dual-phase Liquid Argon** Time Projection Chamber (LAr TPC)
- Ultra low **background**
  - Deep underground (LNGS)
  - Low-background materials, including **Ar target**
- **Active neutron** and muon **veto**s
  - in situ background measurement
- Powerful **background rejection**
  - **Pulse shape discrimination** (PSD)
  - Ionization to scintillation ratio ( $S2/S1$ )
  - Surface rejection using 3D position reconstruction
  - Multiple scatter event rejection

# Dual-Phase Liquid Argon TPC



# Discrimination Power



$10^2 - 10^3$

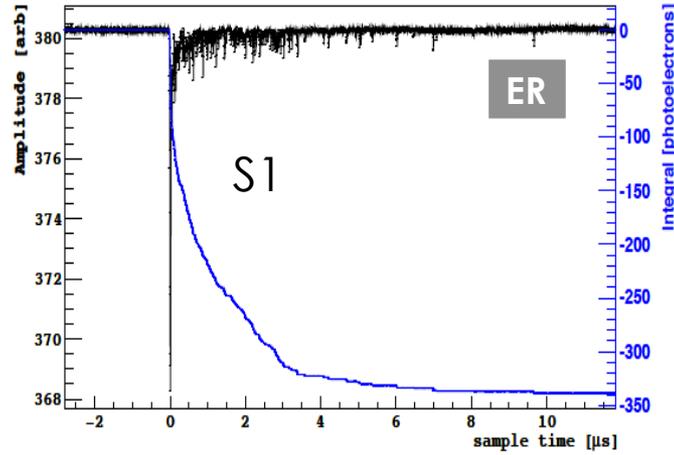
Rejection Factors

Benetti et al. (WARP) 2006

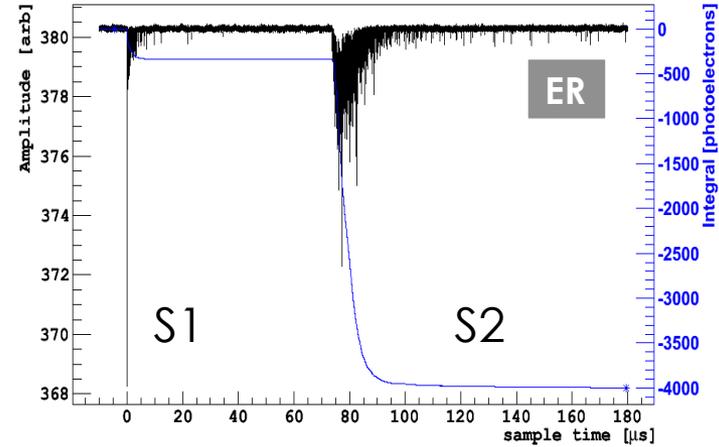
# Discrimination Power

## Pulse Shape Discrimination

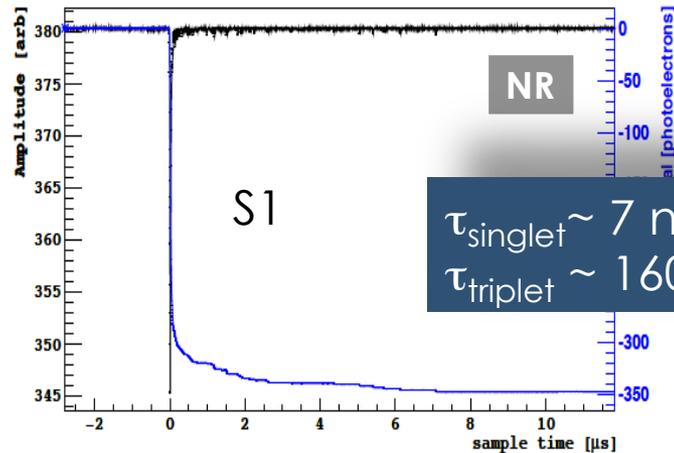
Amplitude ↑



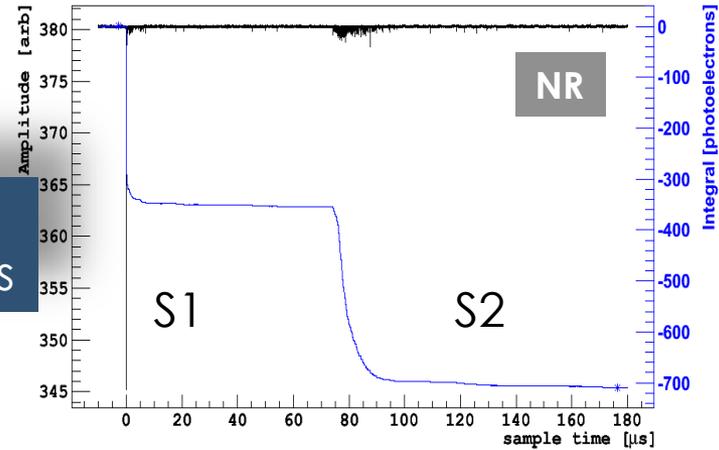
## S2/S1



$\mu$ s →



$\tau_{\text{singlet}} \sim 7 \text{ ns}$   
 $\tau_{\text{triplet}} \sim 1600 \text{ ns}$



Rejection Factors

$10^7 - 10^8$

$10^2 - 10^3$

DarkSide Phys. Lett. B 743 (2015)

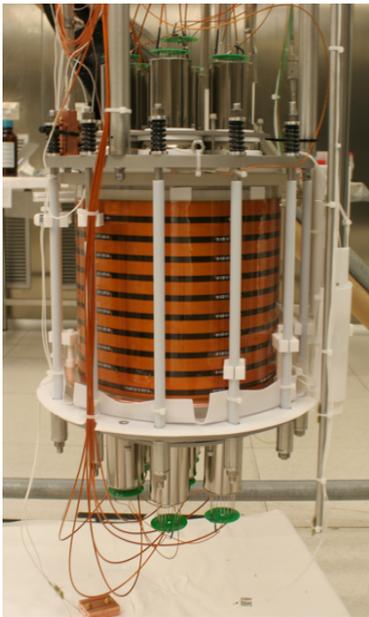
Benetti et al. (WARP) 2006

# The DarkSide Program

## DarkSide-10

2011-2013

10 kg



## DarkSide-50

2013-2016

150 kg  
50 kg FM

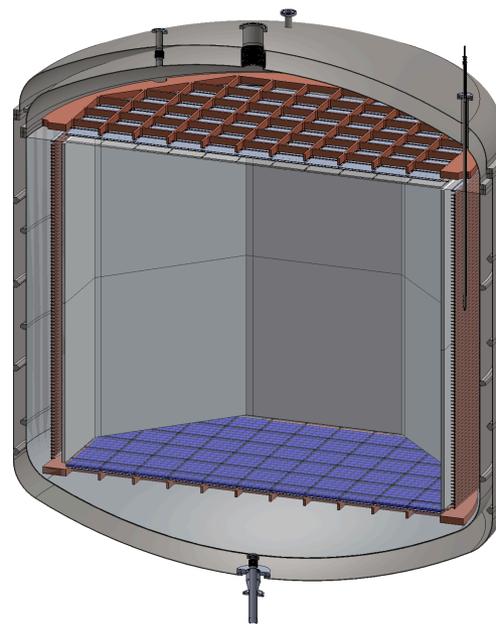


$\sim 10^{-45} \text{ cm}^2$

## DarkSide-20k

2020-2025

30 ton  
20 ton FM



$\sim 10^{-47} \text{ cm}^2$

## ARGO

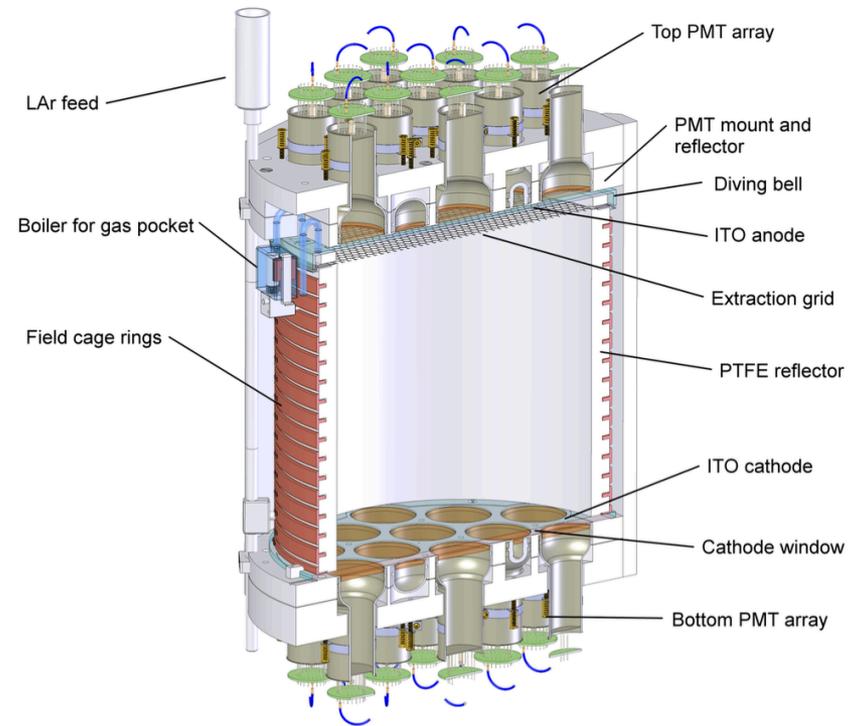
2025-2030

150 ton  
100 ton FM

“neutrino floor”  
 $\sim 10^{-48} \text{ cm}^2$

## The TPC

- 50 kg active mass of **UAr**
- 19 top + 19 bottom R11065 HQE 3'' PMTs
- 36 cm height, 36 cm diameter
- Low field of **0.2 kV/cm drift**
- Lateral walls covered by high reflectivity PTFE
- Cold pre-amps
- 2.8 kV extraction field

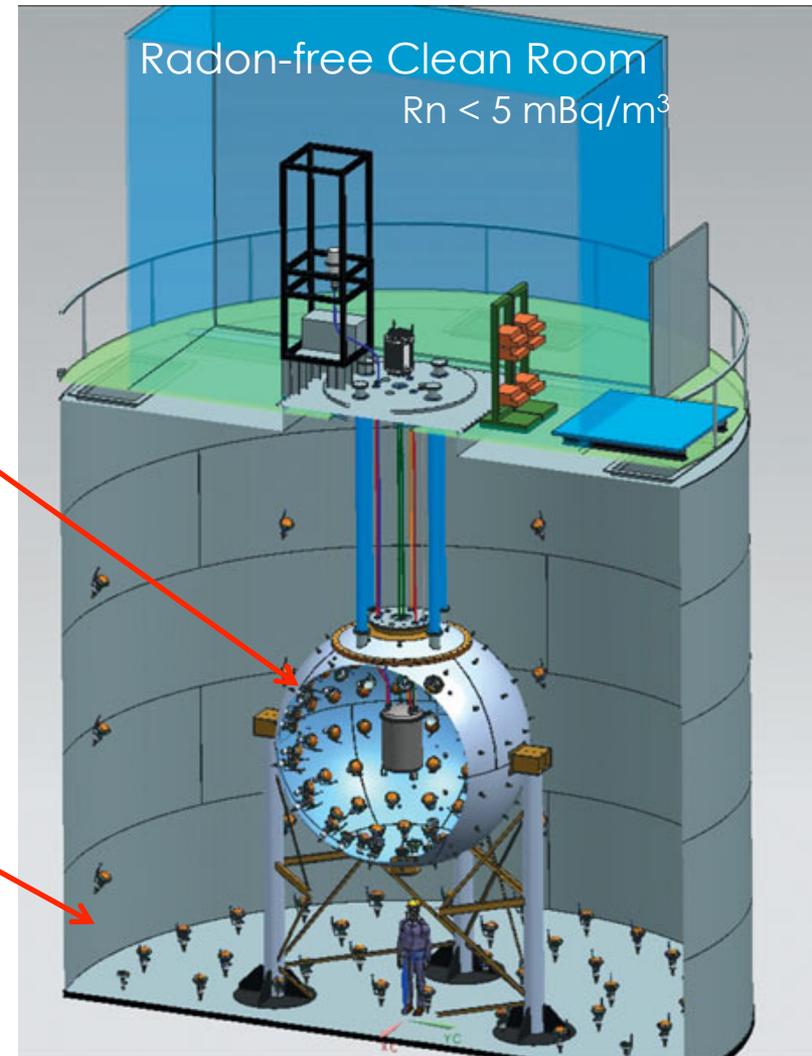


## Liquid Scintillator Veto (against neutrons)

- 4 m diameter sphere
- **30 tonne** of boron-loaded scintillator (PC+PPO+TMB)
- 110 8" PMTs
- LY  $\sim$  500 pe/MeV

## Cherenkov Water Detector

- 11 m diam. x 10 m
- **1000 tonne** water Cherenkov detector
- 80 PMTs



# The Detectors



# The $^{39}\text{Ar}$ Issue Before DS50

## Atmospheric Ar:

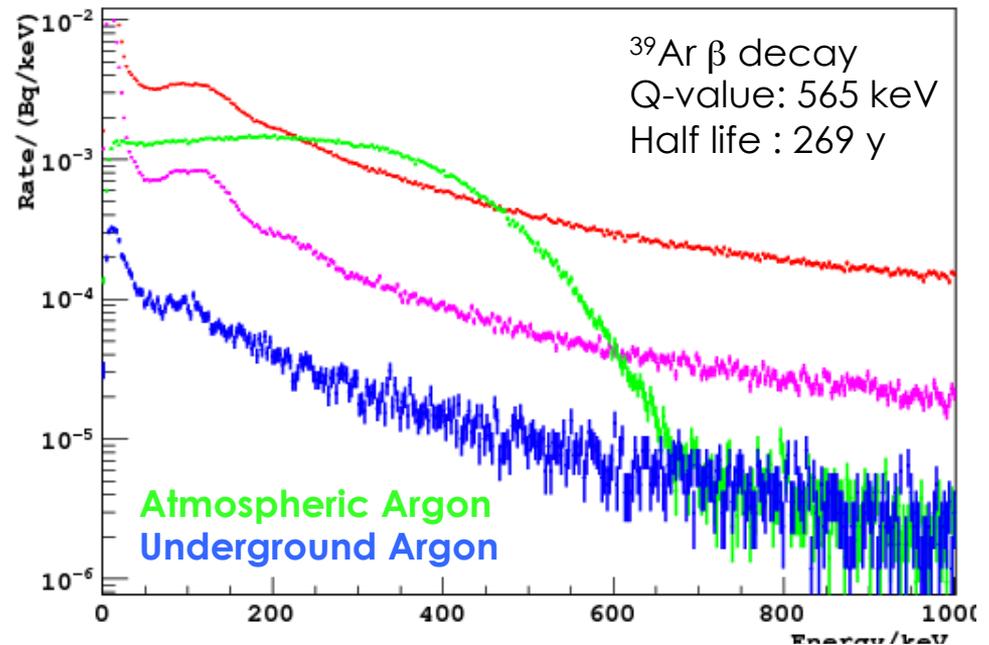
$$^{39}\text{Ar}/^{40}\text{Ar} = 8 \times 10^{-16}$$

Rate ~ **1 Bq/kg**

## Underground Ar:

$$^{39}\text{Ar} < \mathbf{6.5 \text{ mBq/kg}}$$

(arXiv:1204.6011)



	Total Rate [mBq /100 keV]	Estimated BG Rate [mBq/100keV]	BG Subtracted Rate [mBq/100keV]
Atmospheric Argon	$108.8 \pm 0.4$	$1.5 \pm 0.2$	$107.2 \pm 1.9$
Underground Argon	$1.87 \pm 0.06$	$1.5 \pm 0.2$	$0.32 \pm 0.23$
$^{39}\text{Ar}$ Suppression Factor	$58.2 \pm 1.9$		<b>&gt; 153 (95%)</b>

arXiv:1204.60111

**Depletion Factor > 150**

# Internal Calibrations

Atmospheric argon:

$\sim 1$  Bq / kg of  $^{39}\text{Ar}$

$1.5 \times 10^7$  events in 47.1 days

Q-value = 565 keV

Injected gaseous  $^{83\text{m}}\text{Kr}$

$E = 41.5$  keV

Average **Light Yield**:

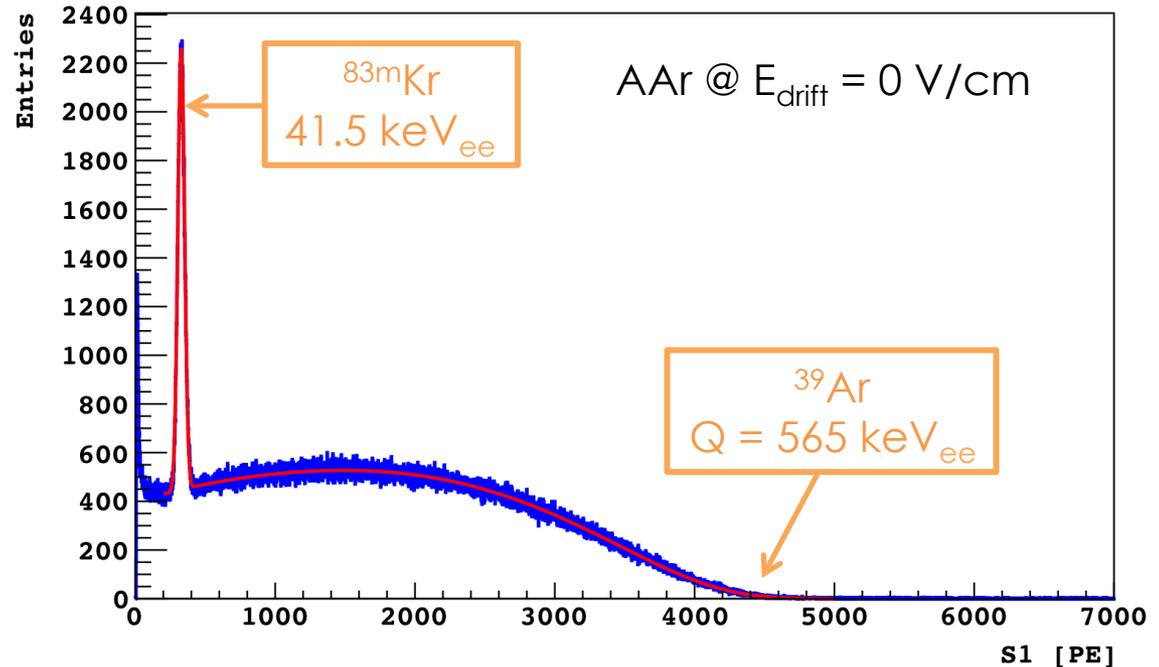
$\sim 7.9$  pe/keV at null field

$\sim 7.0$  pe/keV at 200 V/cm

**Electron lifetime**:  $\sim 10$  ms

Maximum drift time in the TPC:  $375$   $\mu\text{s}$  at 200 V/cm

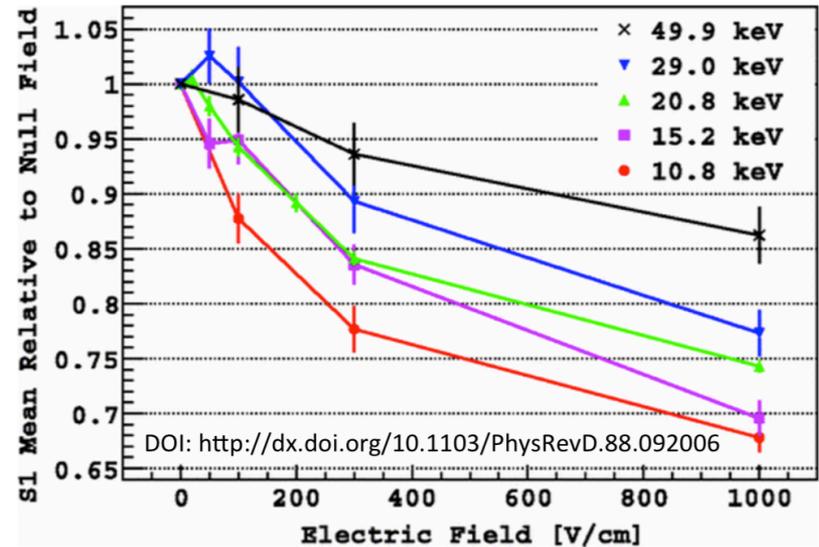
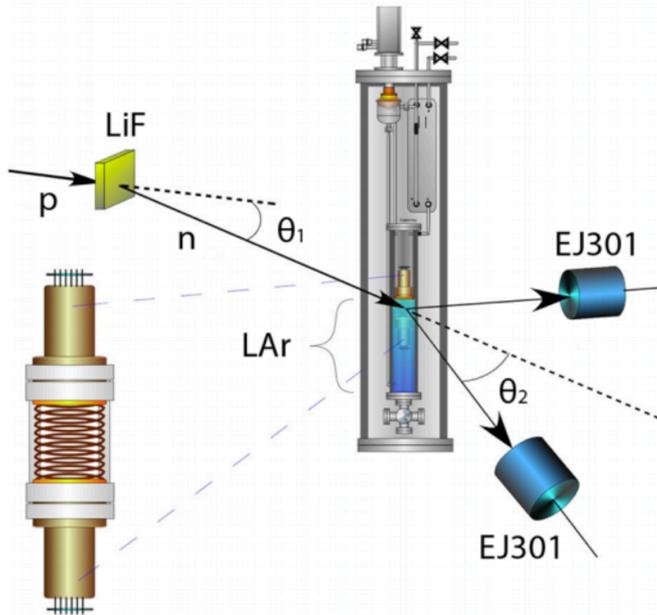
Drift velocity  **$0.93$  mm/ $\mu\text{s}$**



# Nuclear Recoil Energy Scale

## SCENE

### Scintillation Efficiency of Nuclear Recoils in Noble Elements

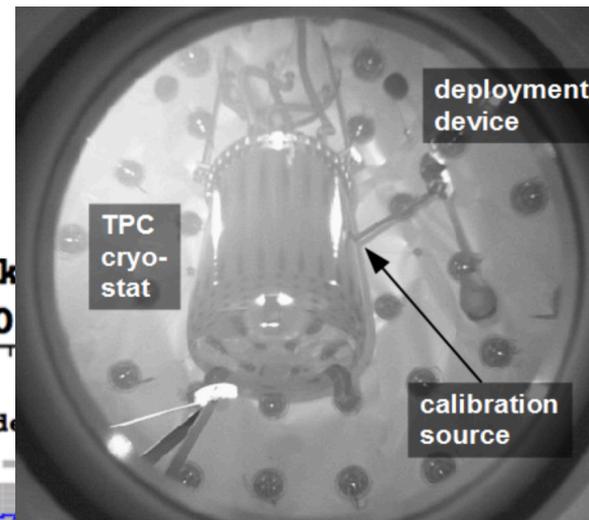
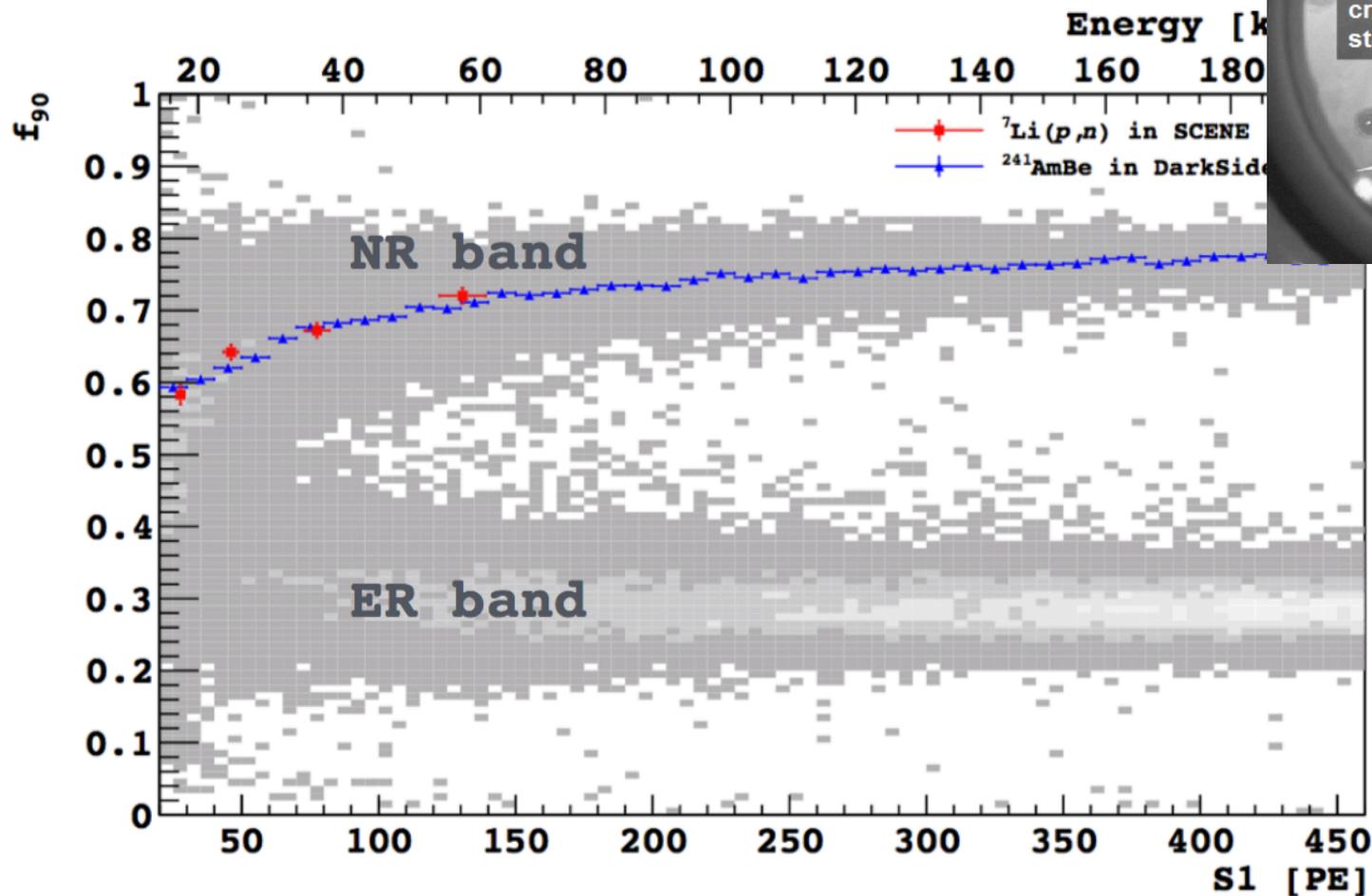


**Neutron** calibration in large detectors affected by neutron **multiple scatterings**

SCENE has collected extremely **pure samples of single nuclear recoils** in a small scale TPC

# External Calibrations

Neutron source: **AmBe** w/ and w/o collimator

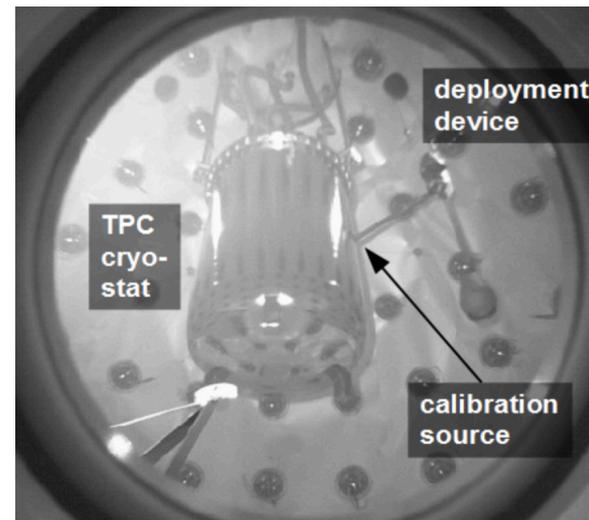


# External Calibrations

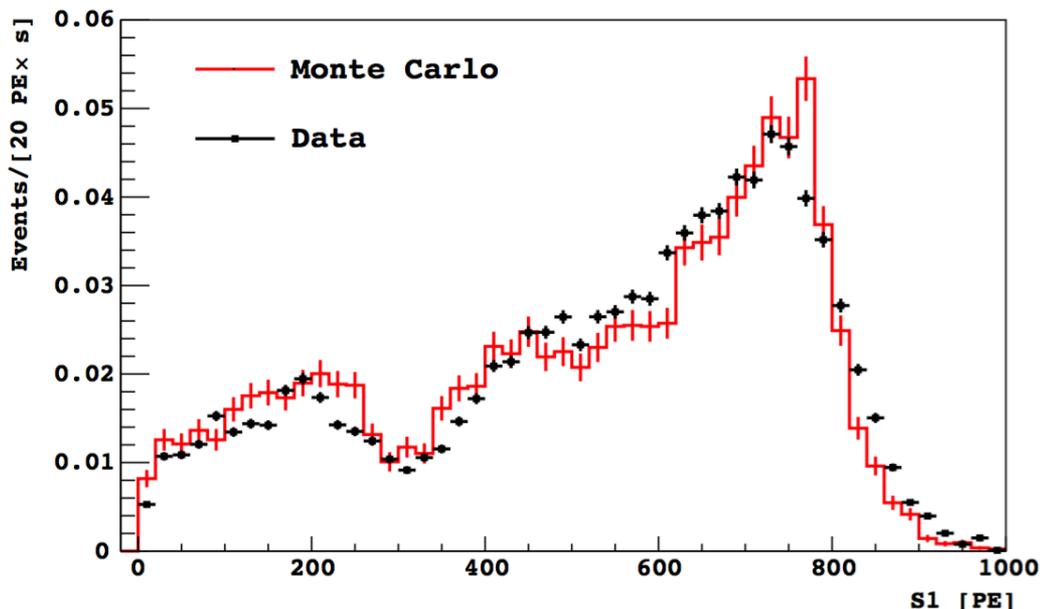
**Gamma sources:**  $^{57}\text{Co}$  (122 keV),  $^{133}\text{Ba}$  (356 keV),  $^{137}\text{Cs}$  (663 keV)

Full **MC** description of all the DS detectors, including:

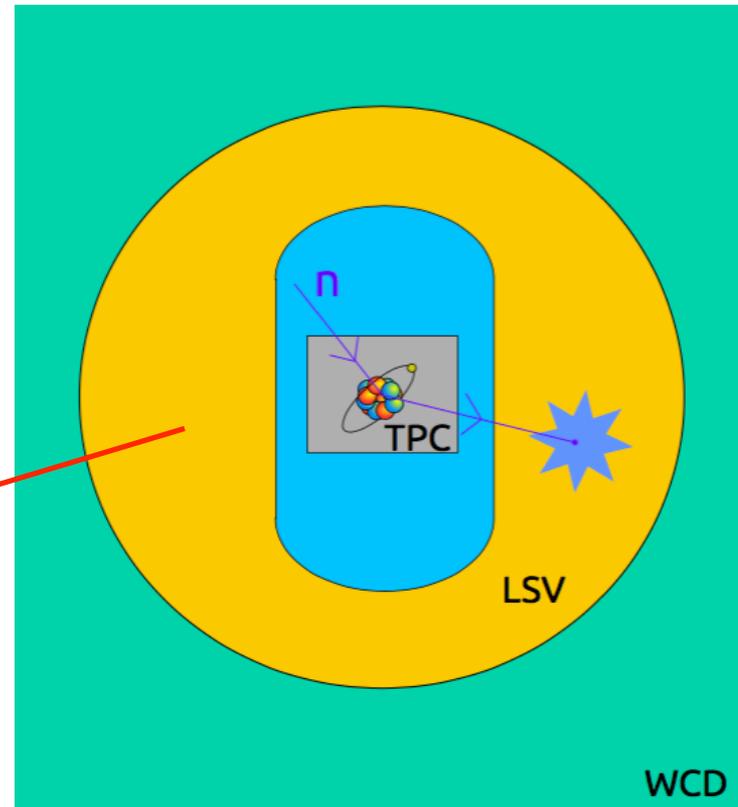
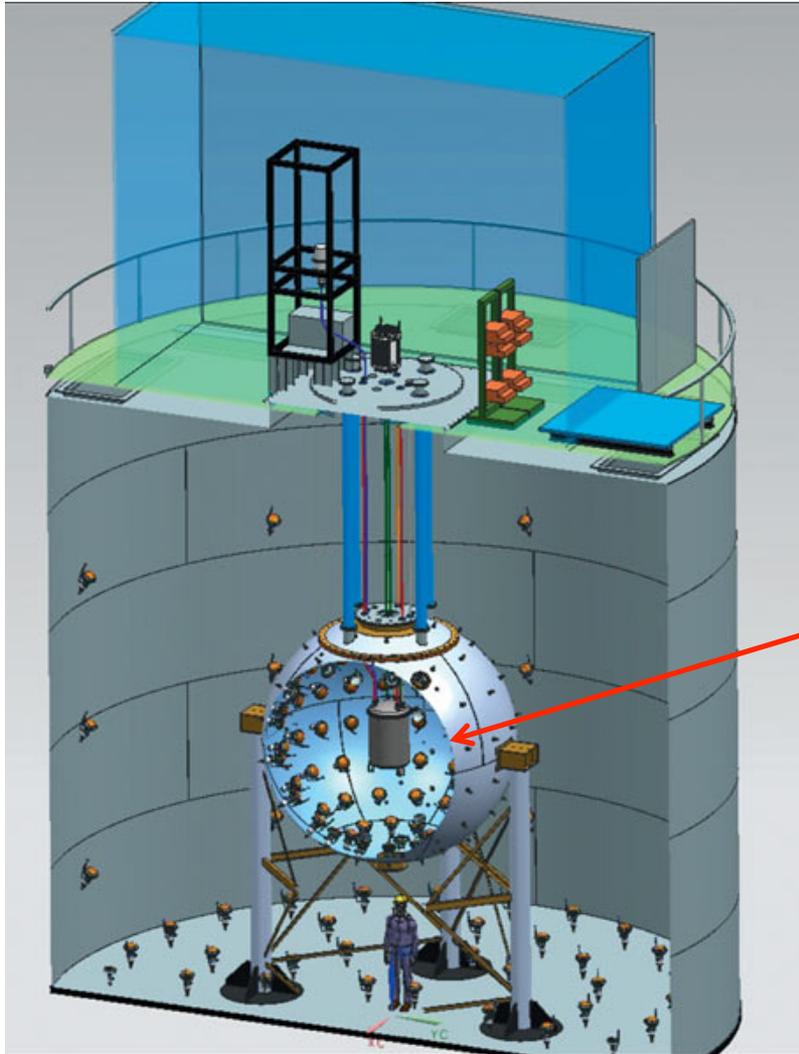
- Custom-made ionization, excitation, and recombination model
- Full photon generation and tracking
- Optical tuning of the TPC: accuracy at a few percent level
- Nuclear quenching model



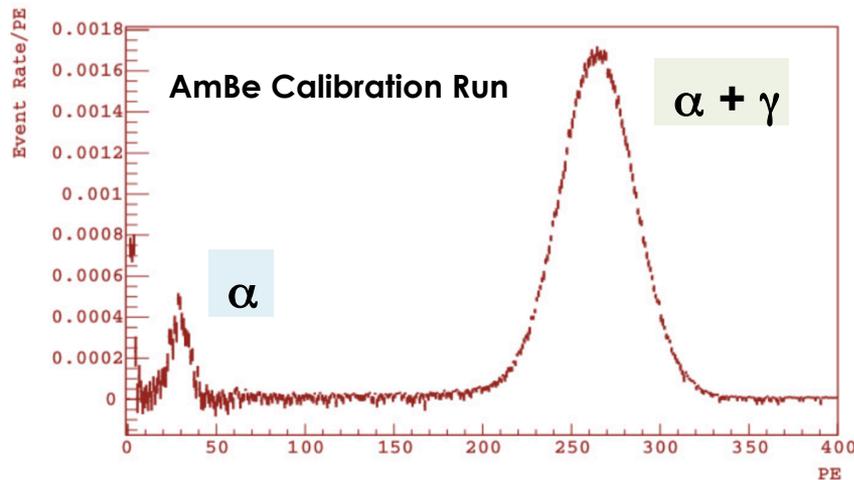
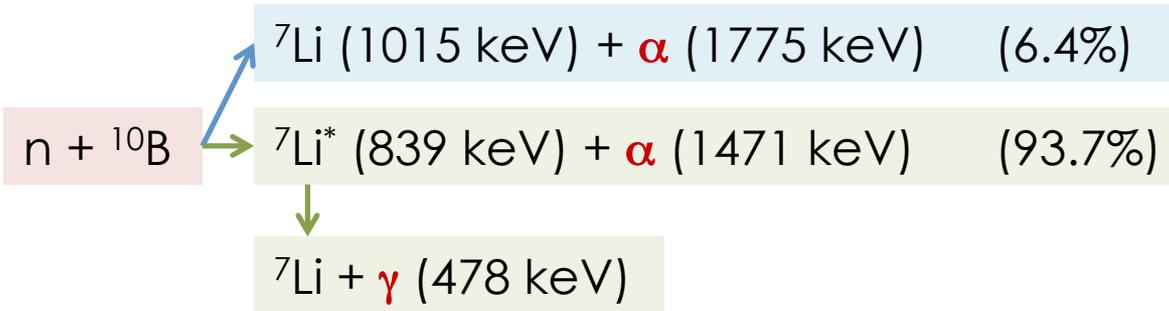
**DATA-MC comparison:  $^{57}\text{Co}$  source next to the cryostat**



# The Neutron Veto



# The Neutron Veto



## Neutron Veto Efficiency

**Efficiency from capture signal alone at > 99%**  
(from calibrations and simulations)

- ~0.6% of lost neutrons because of escaping proton capture gamma
- ~0.05% of neutrons leave no signal in LSV at all

**Larger total efficiency due to thermalization signal**

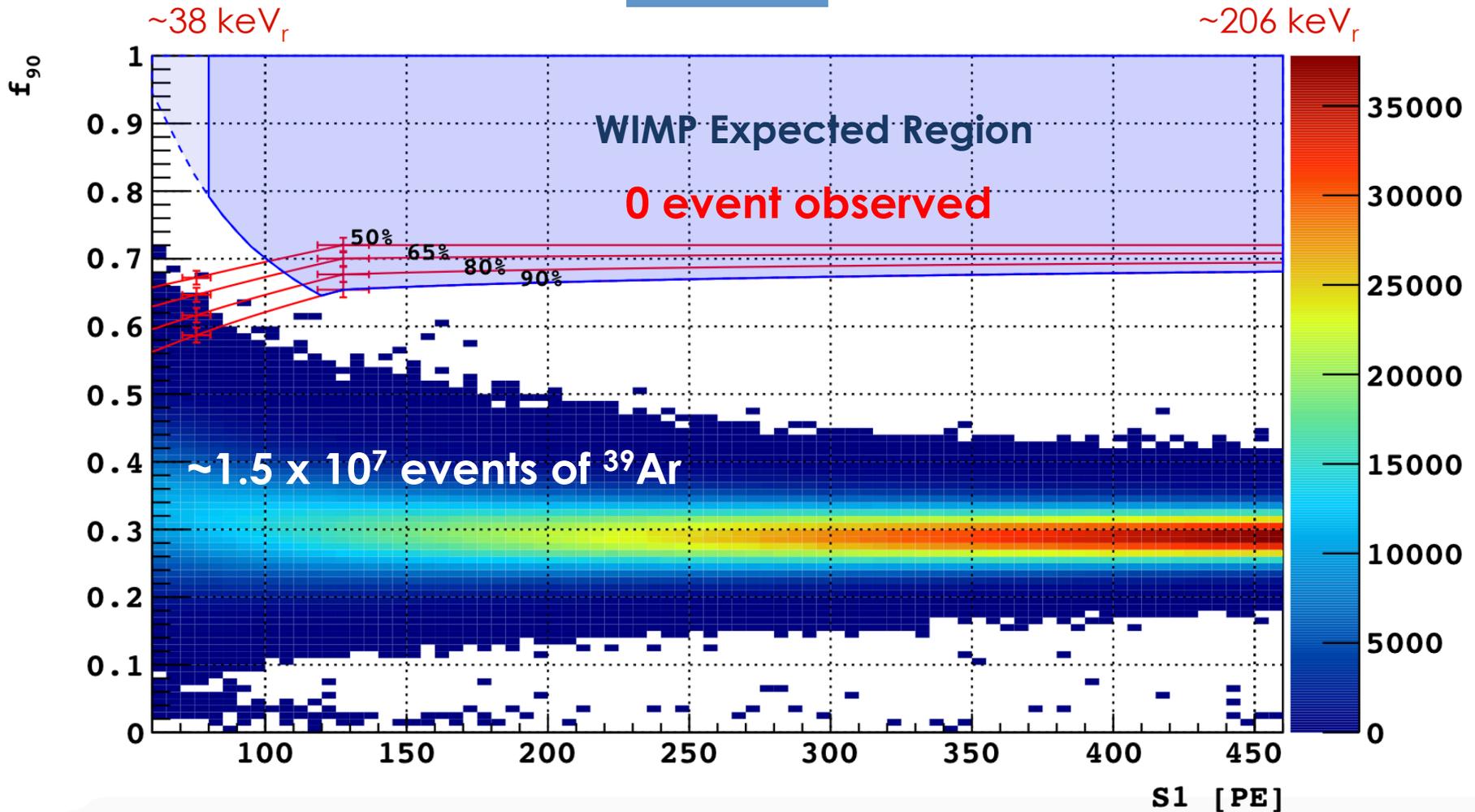
Cut at 1 PE threshold: ~0.9% acceptance loss

- **Oct 2013:** LArTPC, Neutron Veto and Muon Veto commissioned
  - TPC filled with **atmospheric argon** (AAr)
- **Up to June 2014:** data taken with high  $^{14}\text{C}$  content in LSV
  - **47.1 live days** (1422 kg day fiducial) for the first physics result
  - TMB ( $^{14}\text{C}$ ) was removed to reduce the  $^{14}\text{C}$  rate
- **Oct to Dec 2014:** **Calibration** of TPC with radioactive sources
- **Jan 2015:** Add **radiopure TMB** at 5% concentration
- **Mar to Apr 2015:** filling with **underground argon** (UAr) and re-commissioning the detector
- **Apr to Aug 2015:** Accumulate data with UAr for **dark matter search**

# Atmospheric Argon Run: the PSD

No S2/S1

Phys. Lett. B 743 (2015) 456.



47.1 days livetime with AAr  
 $1422 \pm 67 \text{ kg day}$

# Underground Argon



## 1. Extraction at Colorado (CO<sub>2</sub> Well)

Extract a crude argon gas mixture (Ar, N<sub>2</sub>, and He)

## 2. Purification at Fermilab

Separate Ar from He and N<sub>2</sub>

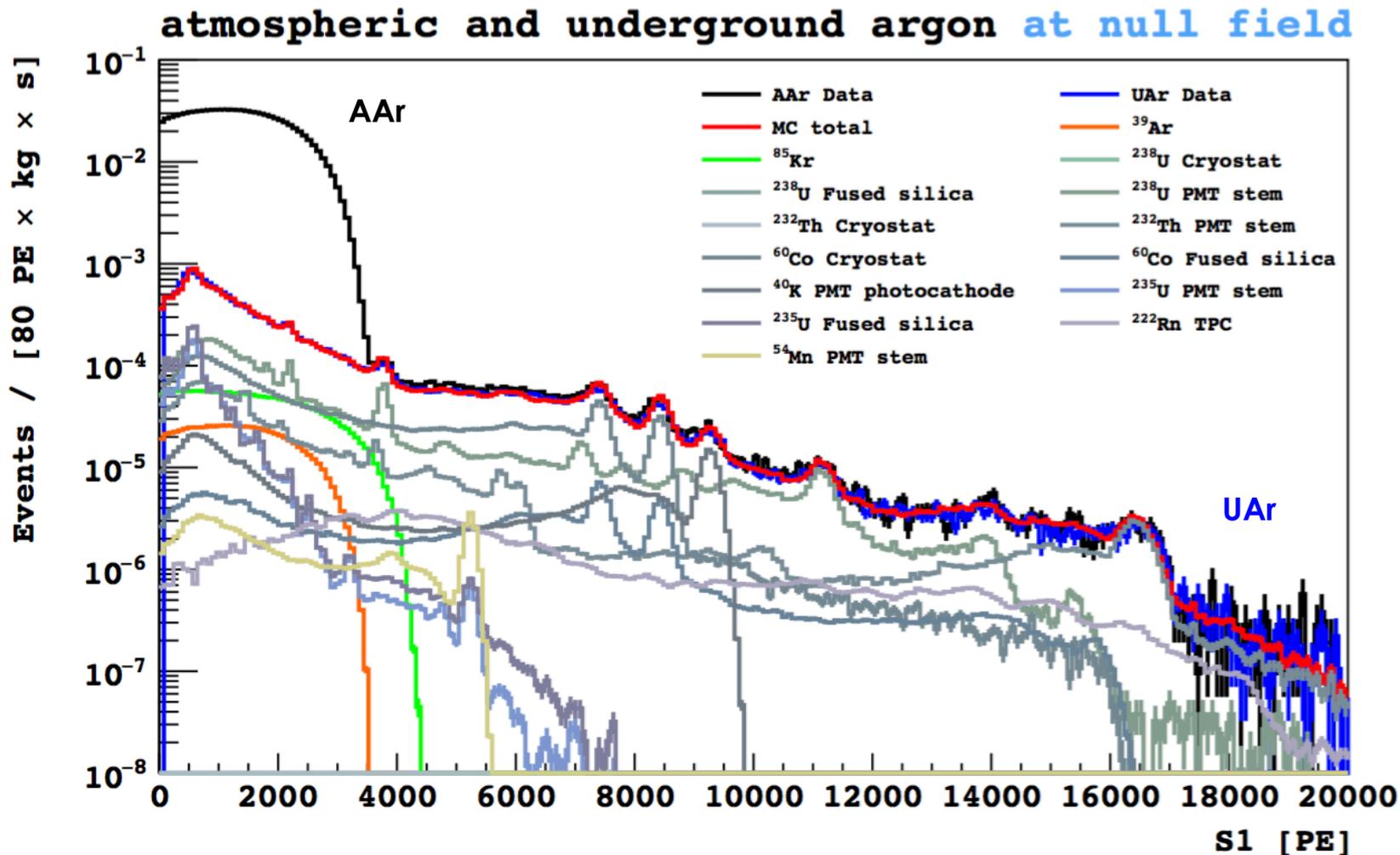


## 3. Ship to LNGS

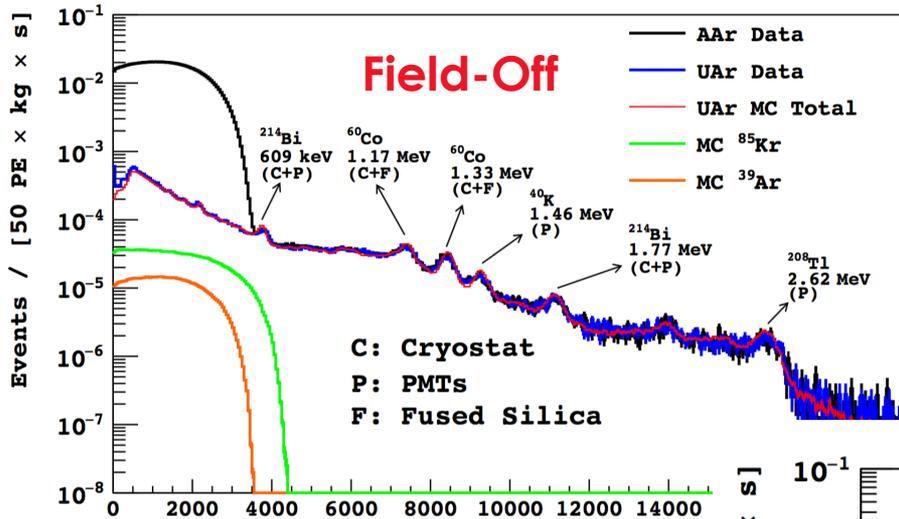
Ready to fill into DS-50



# The Underground Argon Run



# The Underground Argon Run



Three-dimensional likelihood fit based on MC pdf's using:

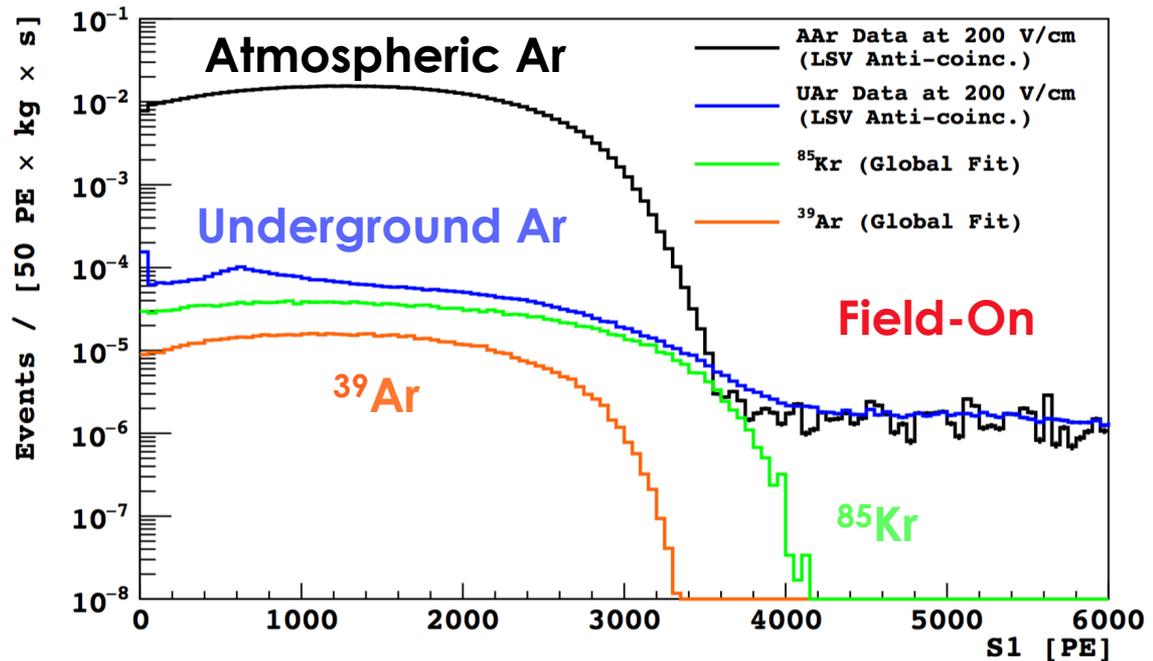
- **S1 field-off**
- **S1 field-on**
- **Drift time**

## Underground Ar:

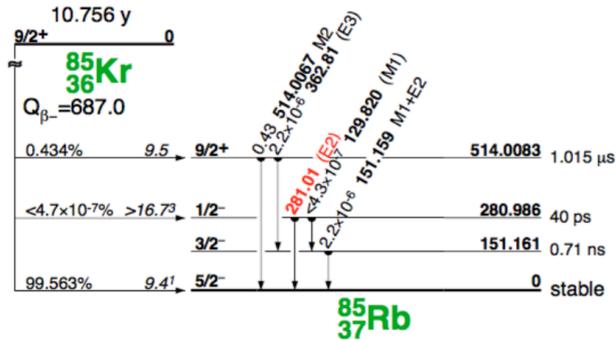
Naturally shielded against cosmic rays

Rate ~ **0.7 mBq/kg**

**Depletion factor ~ 1400**



# The $^{85}\text{Kr}$ Contamination



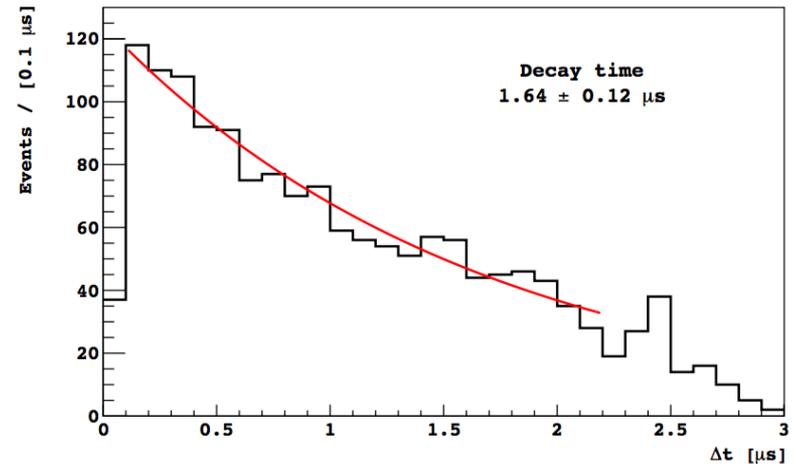
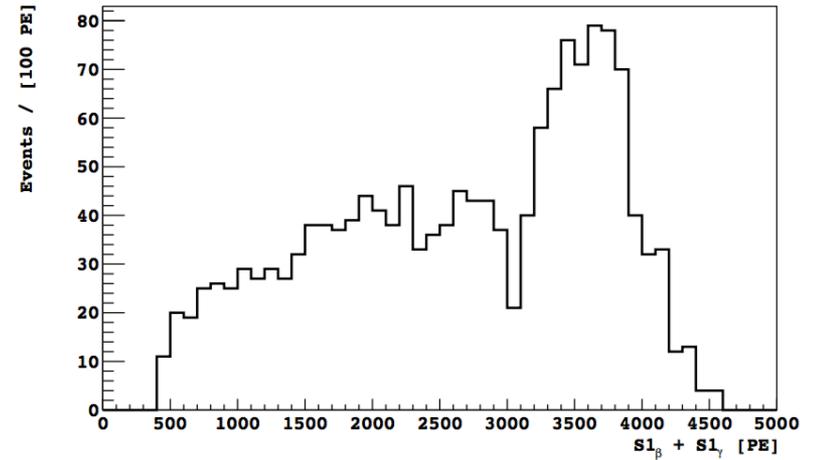
**99.57%** BR via  $\beta$  (687 keV)

**0.43%** BR via  $\beta$  (173 keV) + delayed  $\gamma$  (514 keV)  
Coincidence mean time:  $\sim 1.46 \mu\text{s}$

Deriving the small BR rate from the main branch:

**$35.3 \pm 2.2$  cpd**

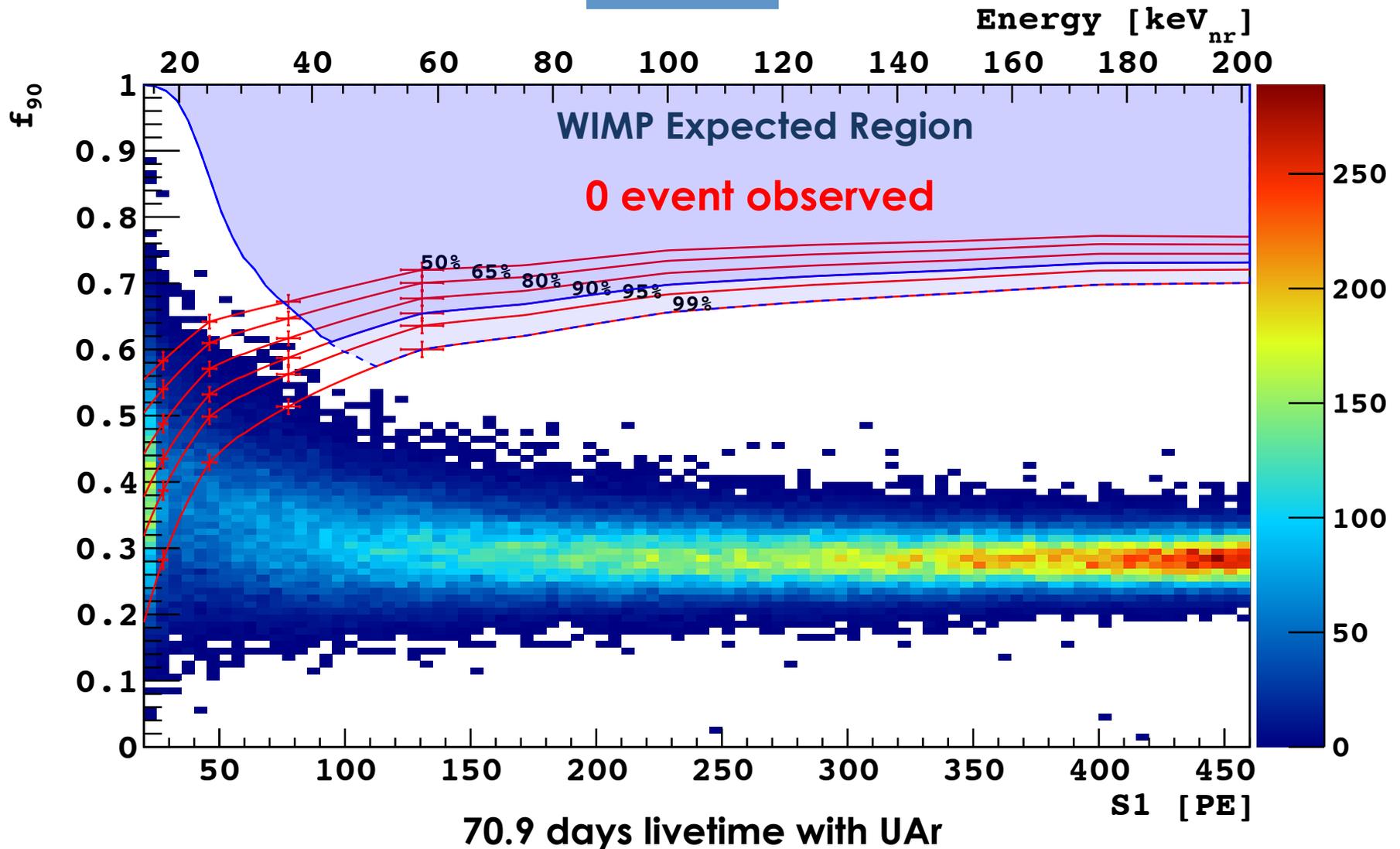
**$33.1 \pm 0.9$  cpd**



# Underground Argon Run: the PSD

No S2/S1

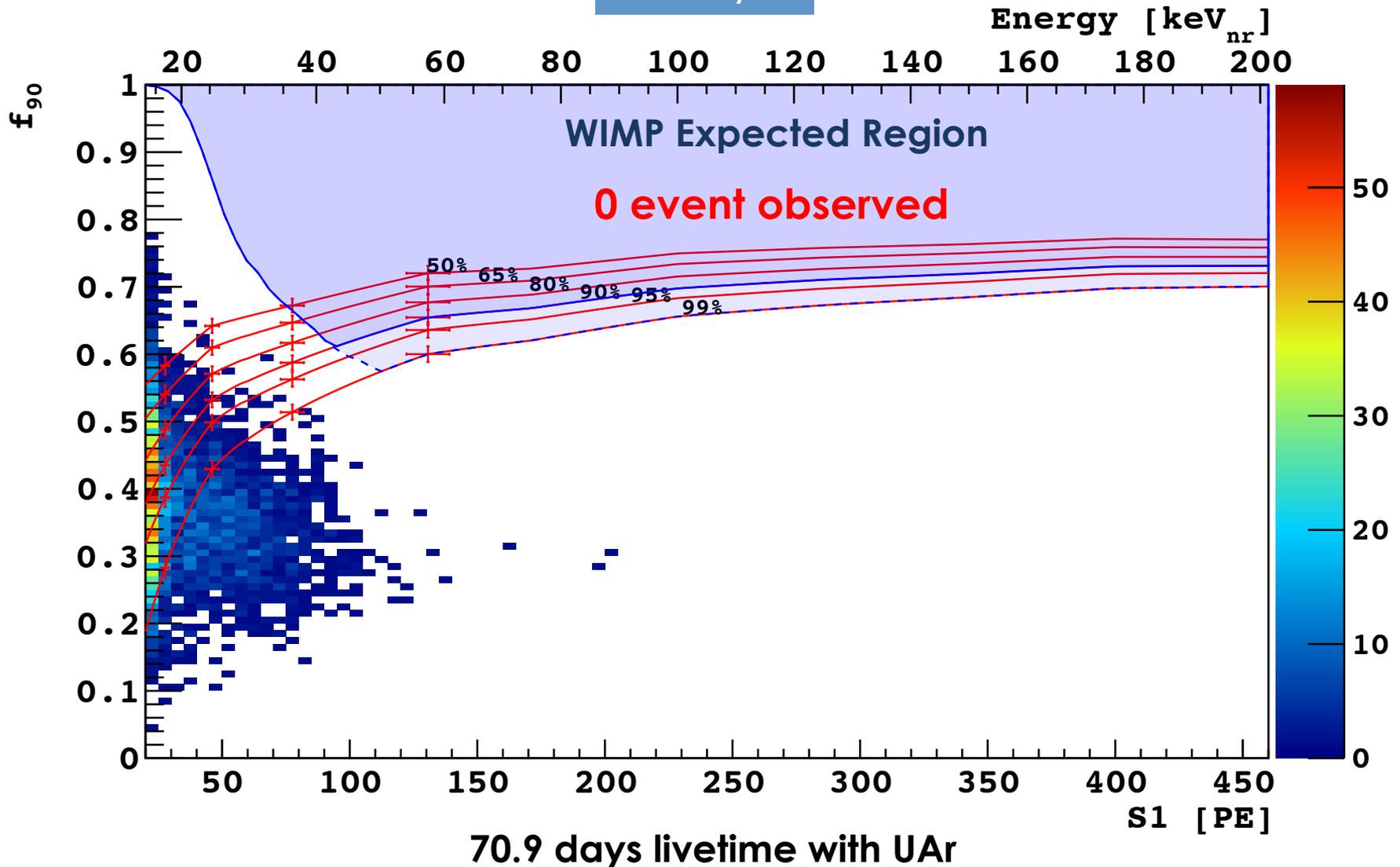
arXiv:1510.00702



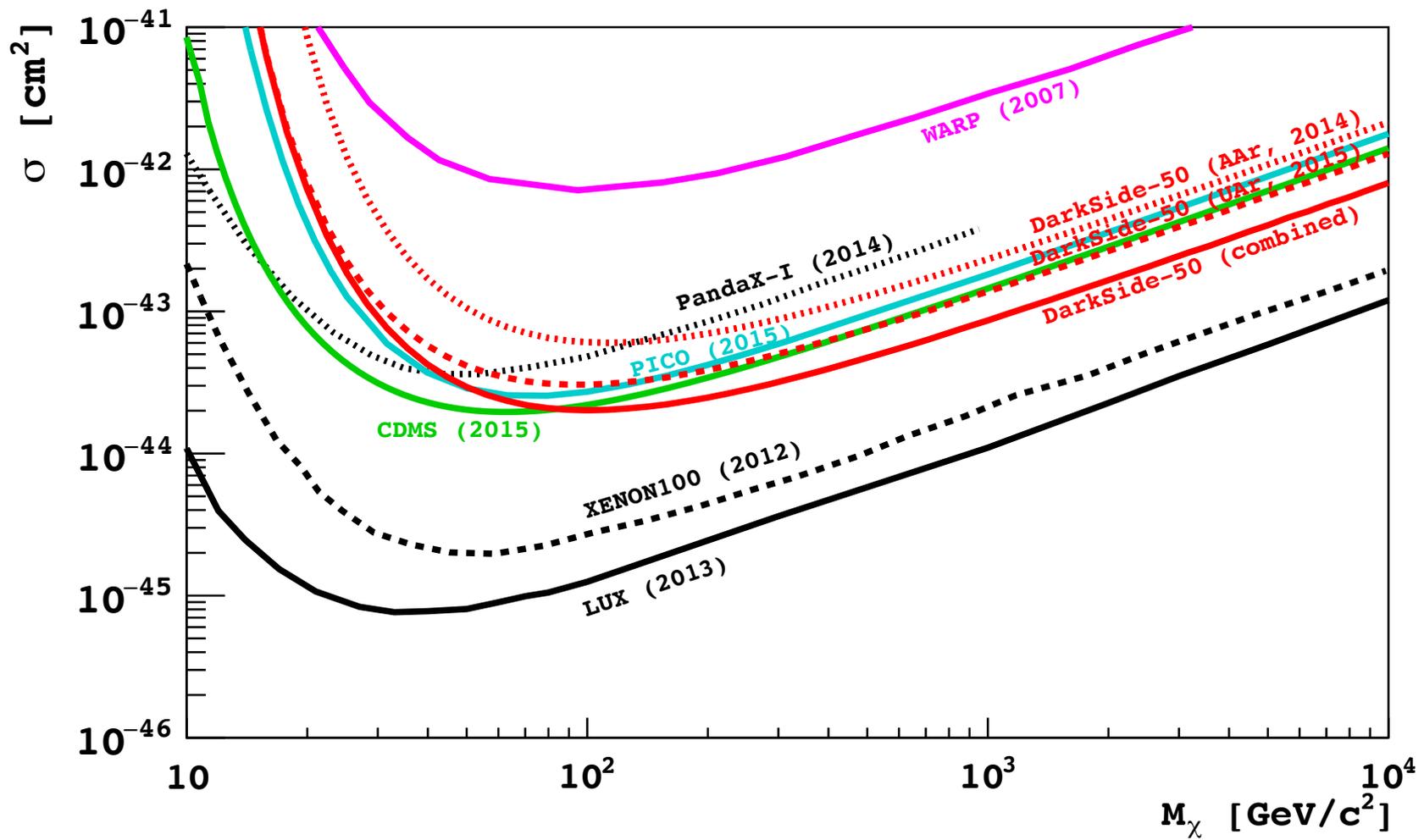
# Underground Argon Run: the PSD + S2/S1

with S2/S1

arXiv:1510.00702



# WIMP Limits



DarkSide arXiv:1510.00702

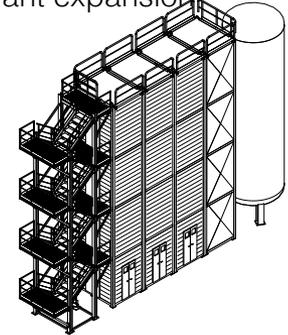
## Depleted Ar: the URANIA and ARIA projects

### URANIA

Replacement of the Ar extraction plant in Colorado to reach capacity of **100 kg/day** of UAr



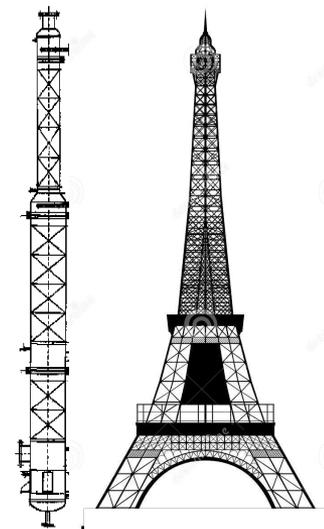
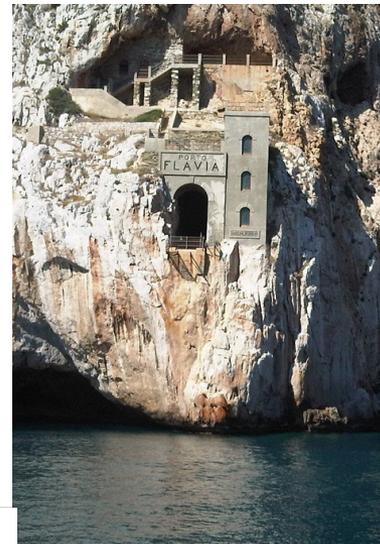
Plant expansion



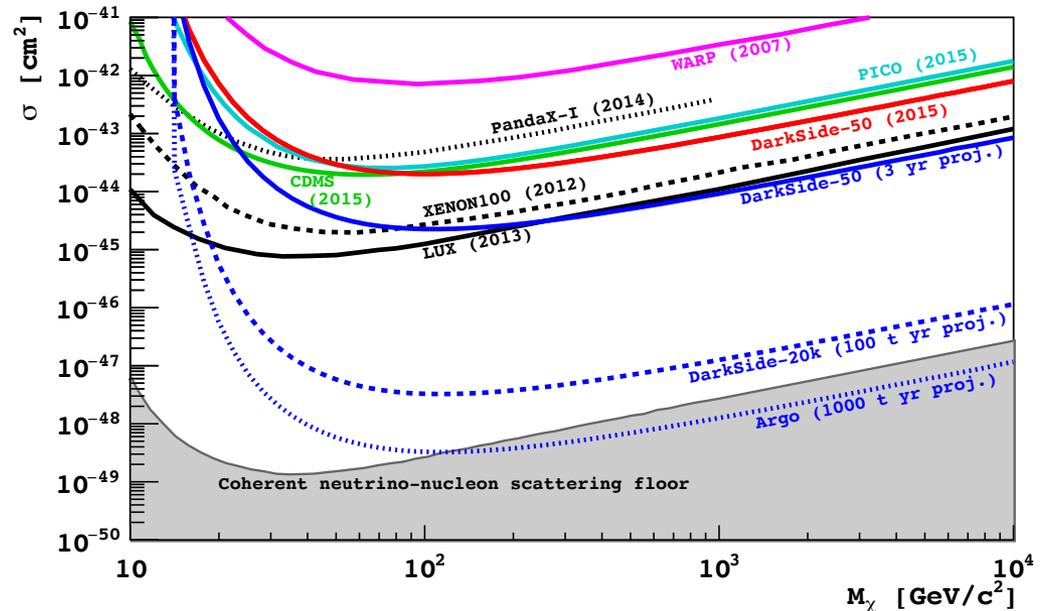
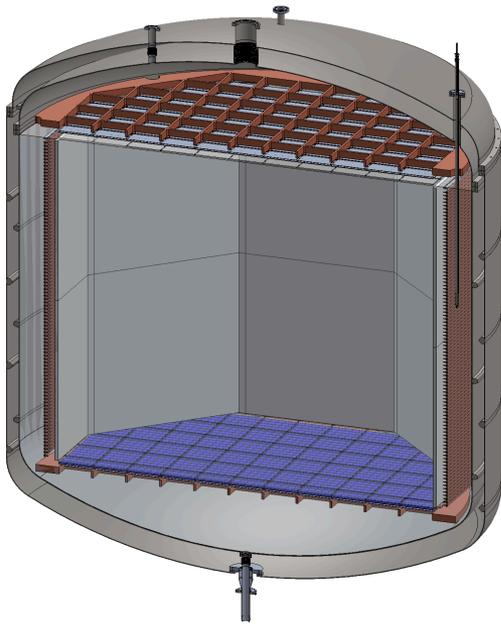
### ARIA

**Very tall distillation column** in Seruci mine (Sardinia) for chemical and isotopic purification of UAr

Exploits finite vapor pressure difference between  $^{39}\text{Ar}/^{40}\text{Ar}$ :  $^{39}\text{Ar}$  reduction factor of 10 per pass at the rate of **100 kg/day**



- Octagonal TPC (~3 m diameter – ~3 m height)
- Target mass: **23 ton UAr**
- Equipped with 15 m<sup>2</sup> of **SiPM**: more radio-pure, smaller total mass, higher efficiency
- **Veto** similar to the DS50 one
- Lol submitted in April
- Technical proposal under submission

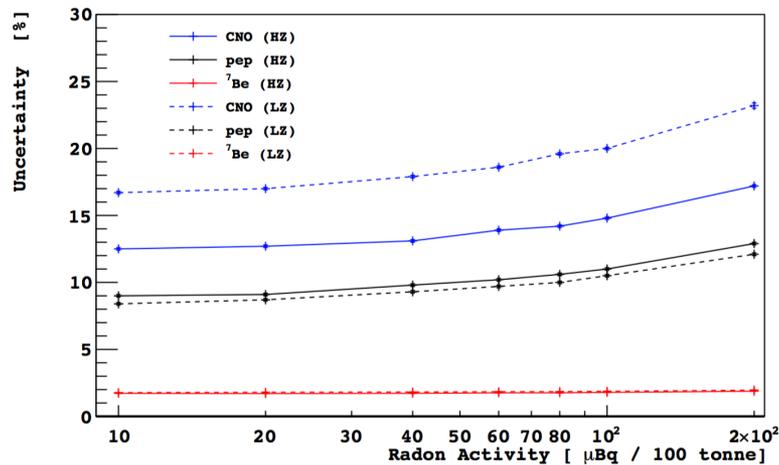


all projected sensitivities are evaluated by assuming 0 bg

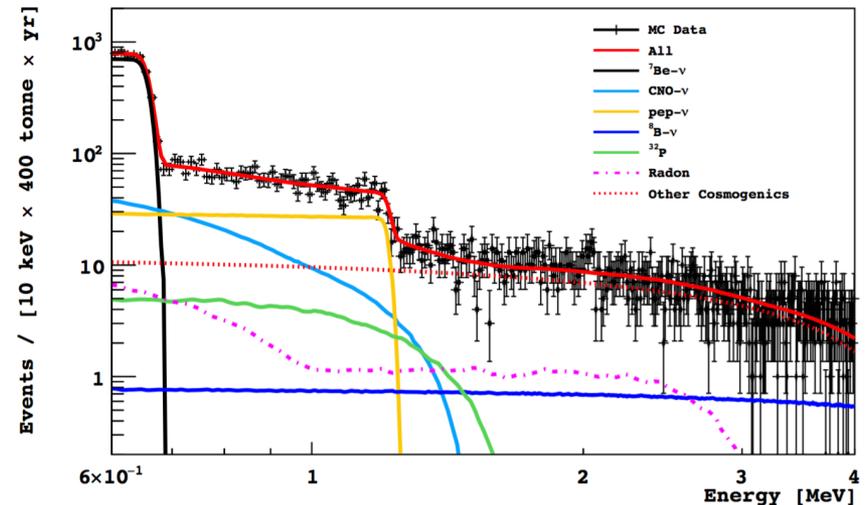
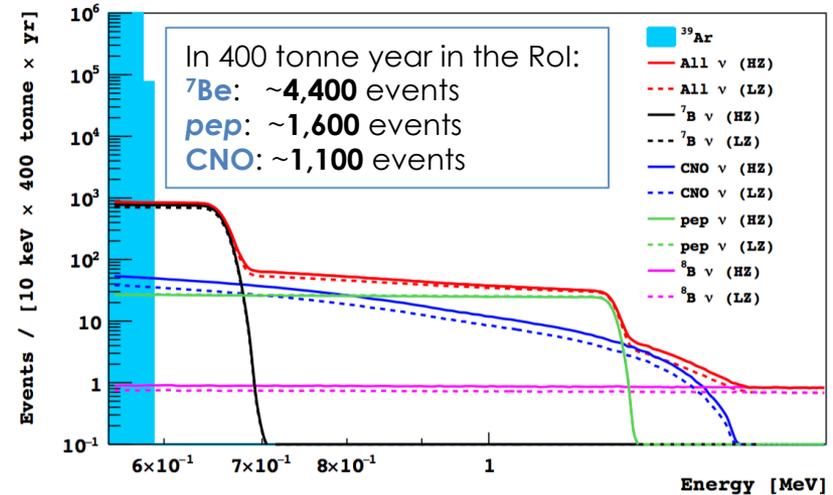
# ARGO: the Solar Neutrino Channel

## Goals:

- **Observation of CNO** neutrinos in 400 tonne year exposure (accuracy at ~15%)
- **Metallicity** solar models discrimination
- **$^7\text{Be}$**  accuracy at 2%
- **Pep** accuracy at ~10%



## Elastic Scattering



arXiv:1510.04196

$^{39}\text{Ar}$  BG from **47.1 live days** (1422 kg day fiducial) of AAr corresponds to that expected in **38.7 years of UAr** DS-50 run

Concentration of  $^{39}\text{Ar}$  in UAr is **~1400 times lower than in AAr**

Pulse shape discrimination **rejects ER with a factor larger than  $1.5 \times 10^7$**

## **Future detectors are planned:**

Letter of Intent submitted to LNGS April 27 2015

Technical Proposal under submission to INFN and NSF

DarkSide is demonstrating that  
**a background free large volume LAr TPC is possible**