

Projects of the Electronics Division

Project Review 2008



Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)

- Projects in 2008
- Requests for 2009
- Status of Selected Projects
 - HEC-I
 - HEC-II
 - MAGIC-I Sum-Trigger
 - MAGIC-II Camera

- Main projects
 - HEC Hadronic Endcap Calorimeter (EA, EE)
 - HEC-II HEC Electronics Upgrade for the SLHC (EA, EE)
 - MAGIC-II Air Cherenkov Telescope Camera (EA, EE, EP)
 - MDT Monitored Drift Tube Chambers (EA)
 - MDT-II MDT Electronics Upgrade for the SLHC (EA, EE)
- Additional projects
 - Cresst (EP)
 - Gerda (EA, EE, EP)
 - ILC / SiPM (EE)
 - Muon Cooling (EP)
 - SCT (EA)
 - Support for the Semiconductor Laboratory (EP)

Group naming

EA: Elektroanlagen

EE: Elektronik Entwicklung

EP: Elektronikproduktion

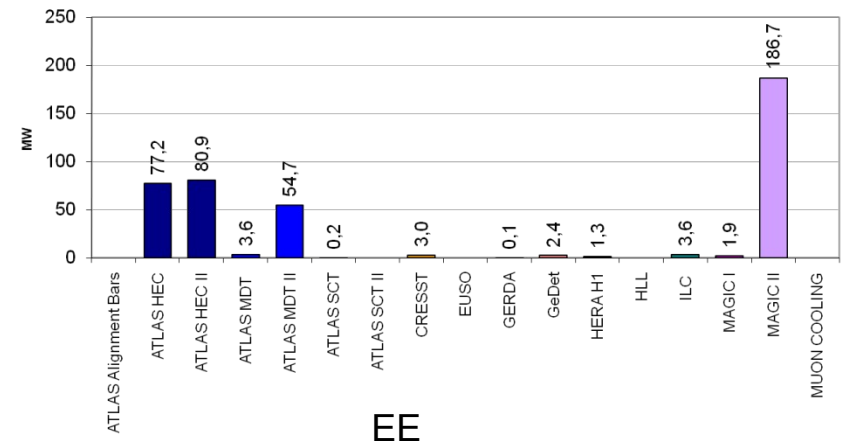
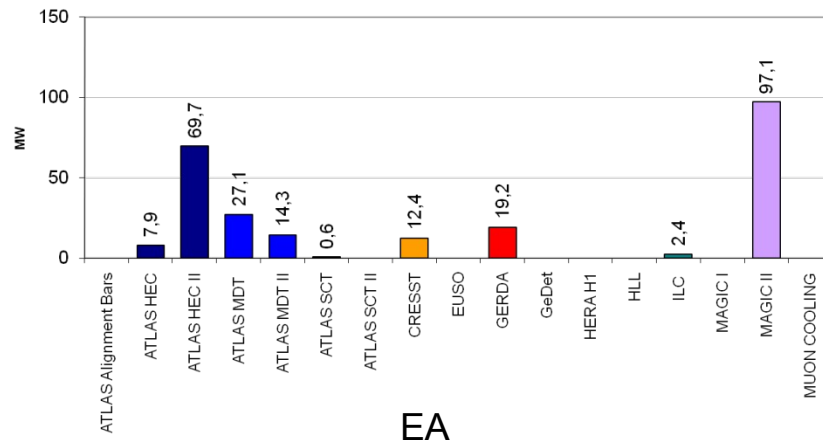
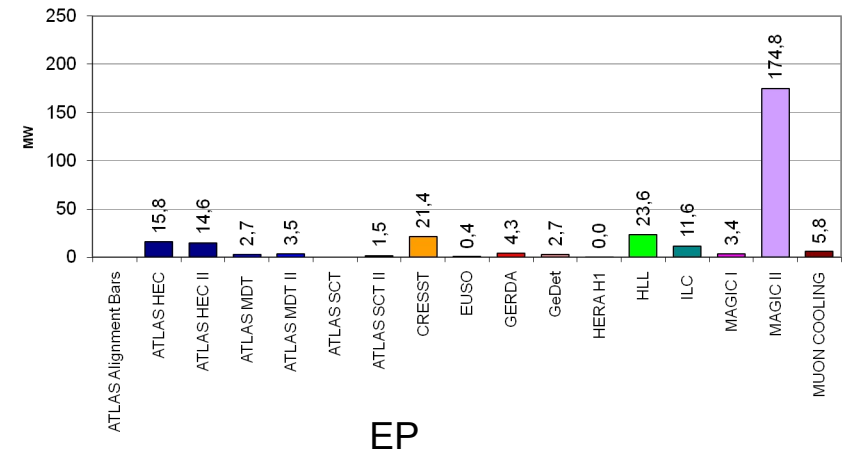
Projects in 2008

Average manpower/quarter (Nov. 2008):

Elektronikproduktion (EP):
78MW/Q (Nominal 65MW/Q)

Elektroanlagen (EA):
68 MW/Q (Nominal 65MW/Q)

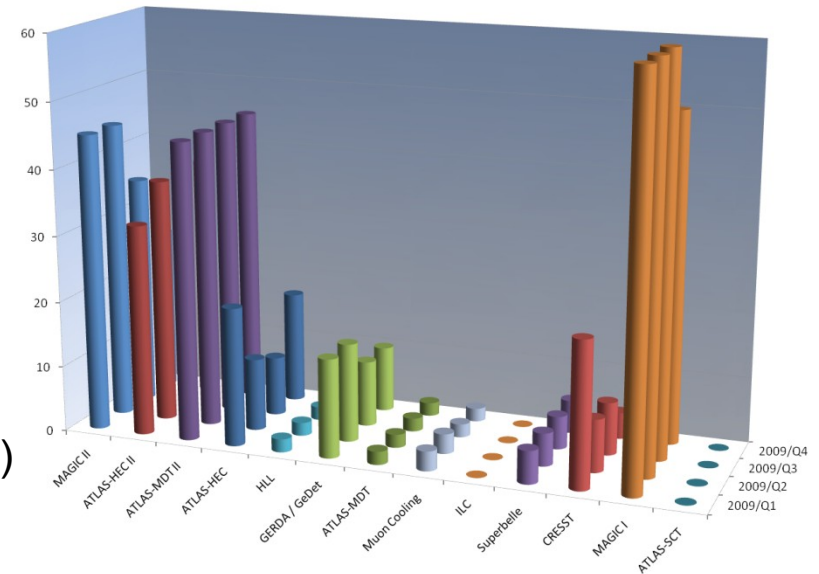
Elektronik Entwicklung (EE):
110MW/Q (Nominal 90MW/Q)



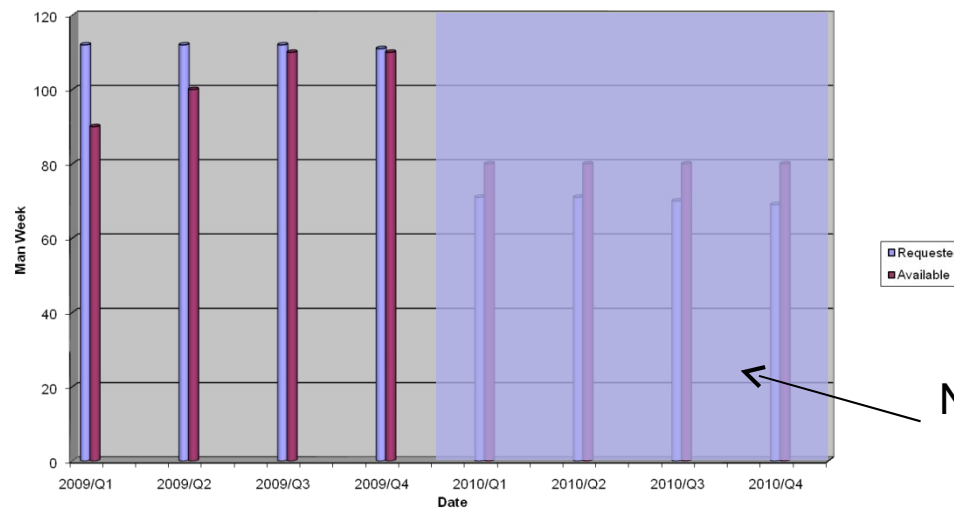
Requests for 2009

Main tasks in 2009 (requested):

- New MAGIC-I camera (MAGIC-II clone)
- Improved sum-trigger
- Chip development for SLHC-HEC (HEC-II)
- Low-voltage development for SLHC-HEC
- Upgrade for MDT-I (CSM-Chip)
- Chip development for SLHC-MDT (MDT-II)
- CSM module development for SLHC-MDT (MDT-II)



Demands on EE
(Without new Projects in 2009)



Some requests are underestimated!!!

-> We will do our best!!!

Not complete
for 2010

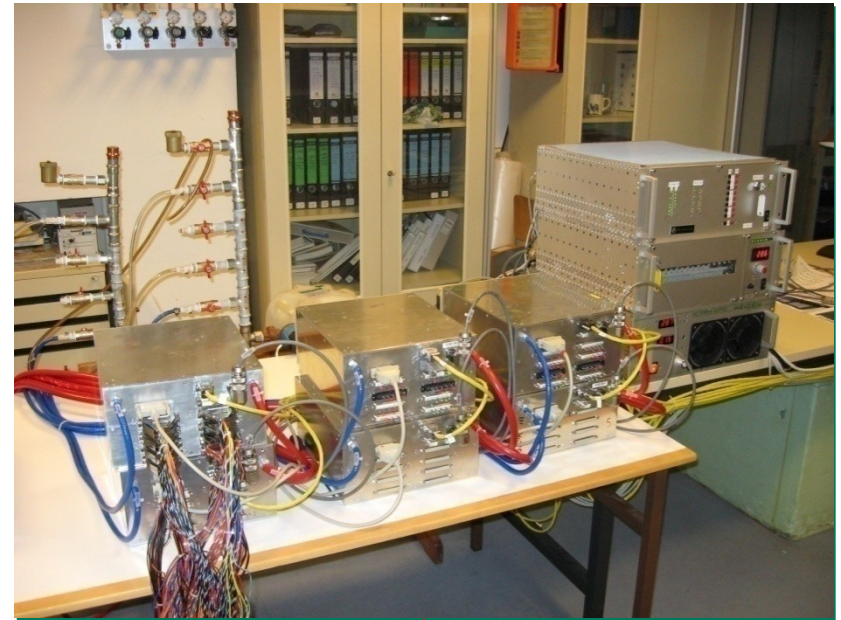
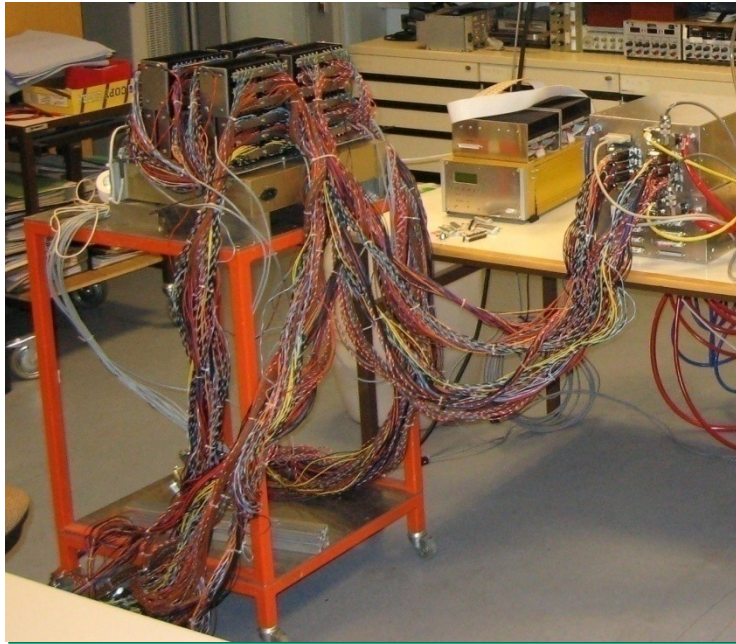


Status of Selected Projects

HEC-I / Low-Voltage System

The HEC low-voltage system is installed in the ATLAS-detector and works well!

(One low-voltage box was changed
-> Problem with DC-DC converter)

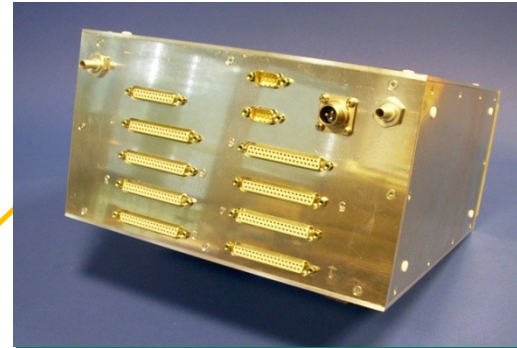


Test system in our lab

HEC-I / Some Data about Low-Voltage

Low-voltage system supplies the power for the HEC-amplifiers

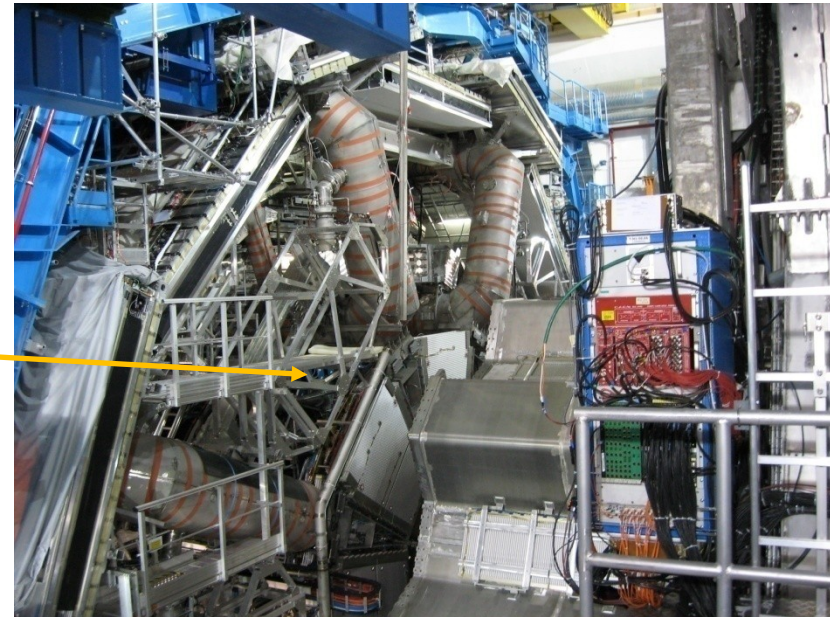
- One 280V supply
- One control system
- 8 low-voltage boxes
(Mounted between the tile fingers)
- One box per quadrant
- Each box is for 40 preamplifier-boards
- Full control and monitoring



LV-box



Assembled into the ATLAS-detector



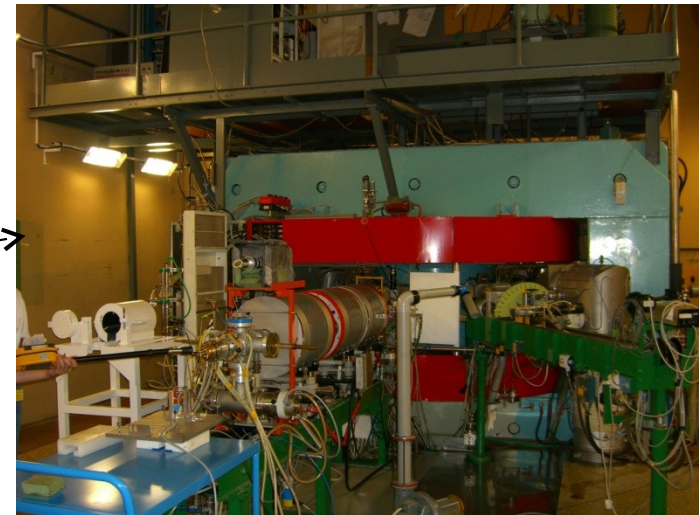
Upgrade of the Hadronic Endcap Calorimeter (HEC-II)

SLHC luminosity upgrade leads to increased particle rates

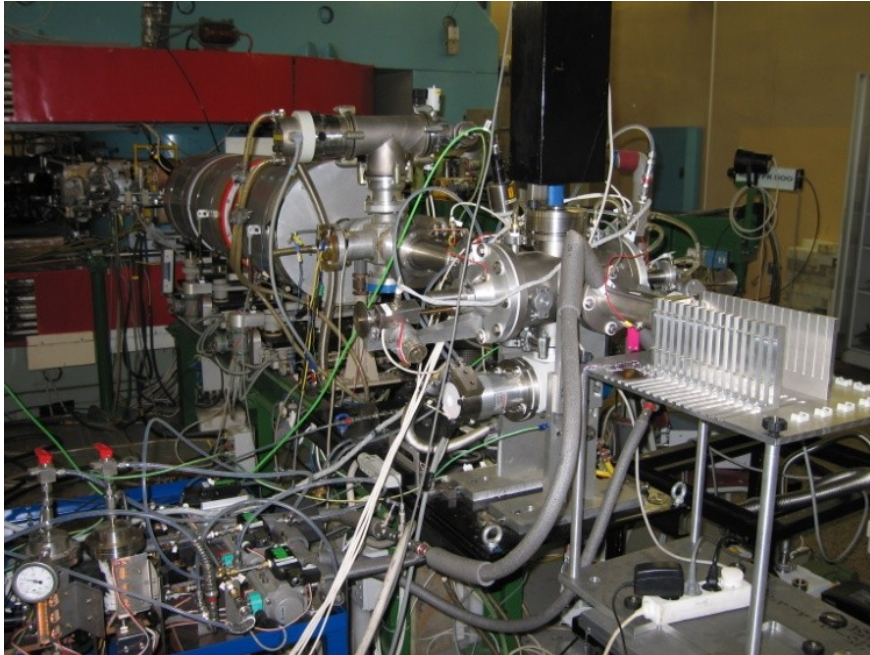
- > Improved amplifiers for the ATLAS-HEC (Factor 10 higher radiation hardness)
- > Reduced structure size in amplifier chips (e.g. 250nm or less)
- > Possible use of a different technology (SiGe instead of GaAs)

Investigation of technologies from different partners:

- Institute for Semiconductor Physics (Frankfurt/Oder) (SiGe, CMOS)
 - Triquint (GaAs)
 - IBM (SiGe)
 - AMS (SiGe)
-
- Radiation test (neutrons) at cyclotron in Rez (near Prague)
 - Selecting the technology from results (CMOS, SiGe or GaAs)

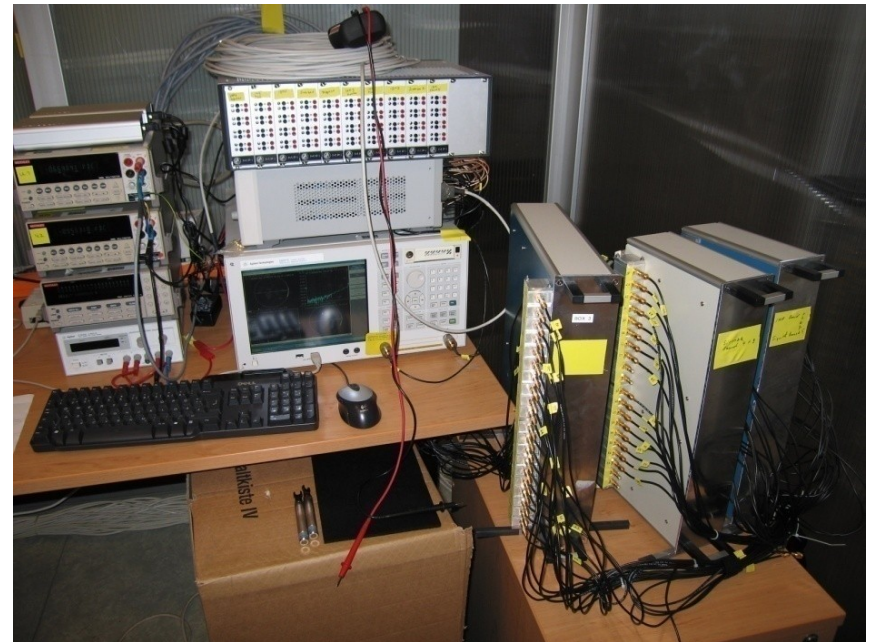


Some Words about the Radiation Measurement

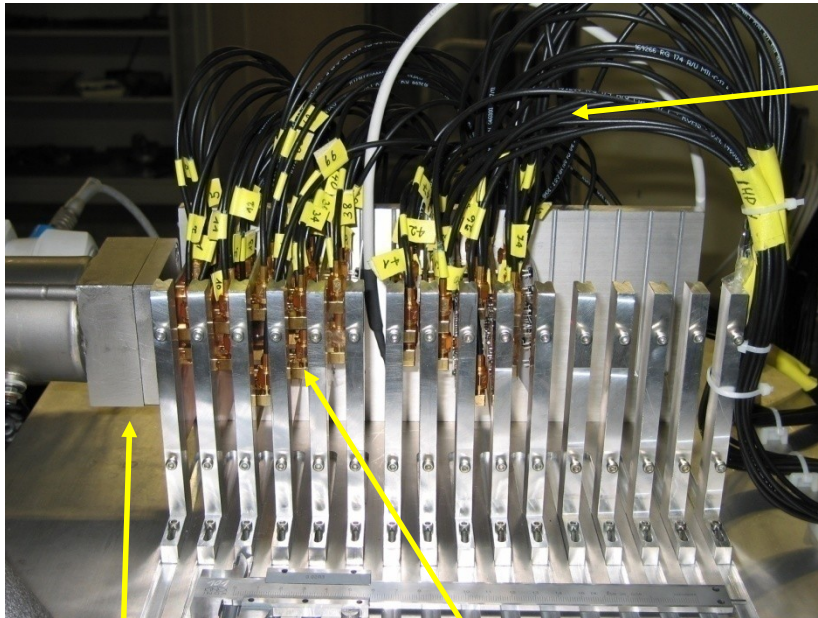


About 40m cable between the transistor and the measurement system.

- Testing up to 37 devices (transistors)
- Measuring DC-values and S-parameters
- Measurement during neutron irradiation



Some Words about the Radiation Measurement



Cabling for RF/DC-signals

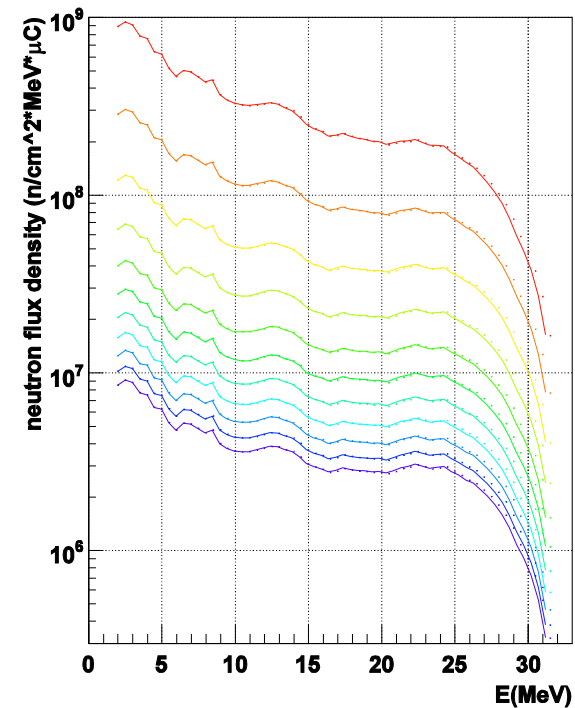
Deuterium target
(neutron source)

Up to 11 boards
with devices under test

Up to four devices per board located in the beam

Neutron flux density spectrum
at different positions:

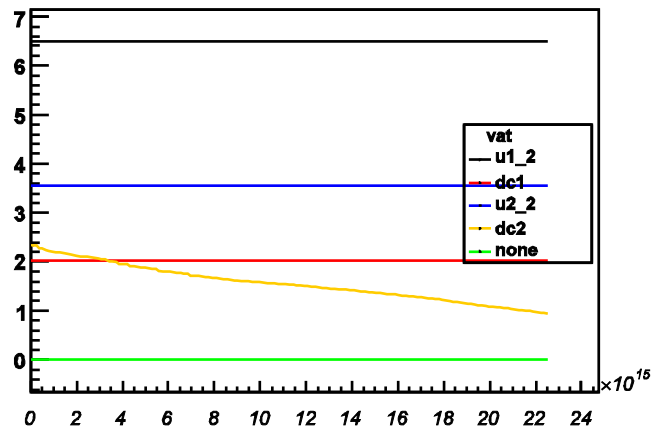
slot NFD .vs. energy



Example: SiGe-Bipolartransistor

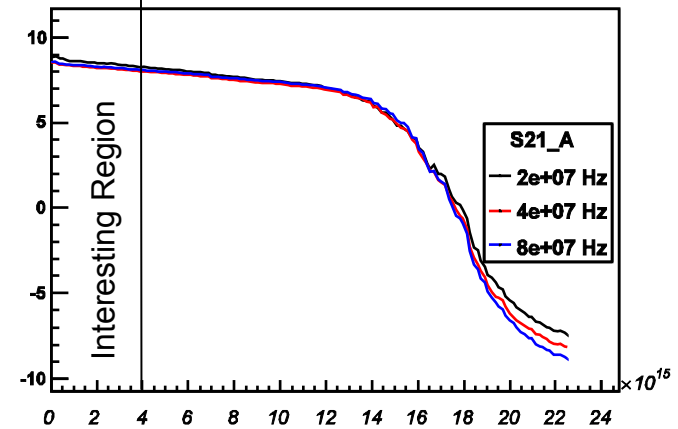
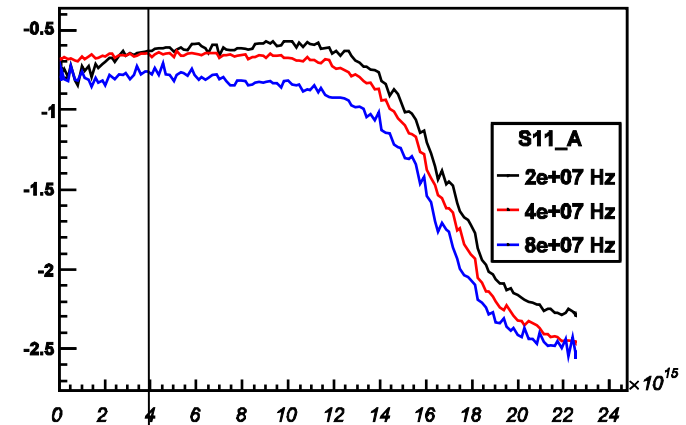
- npn-bipolartransistor from IHP
- SiGe-technology
- $0.42 \times 0.84 \mu\text{m}^2$ structure size
- 2 elementary cells in parallel
- Included ESD-protection

- Positioned in slot 1



Currents and voltages

Parameter S11 ("Impedance")



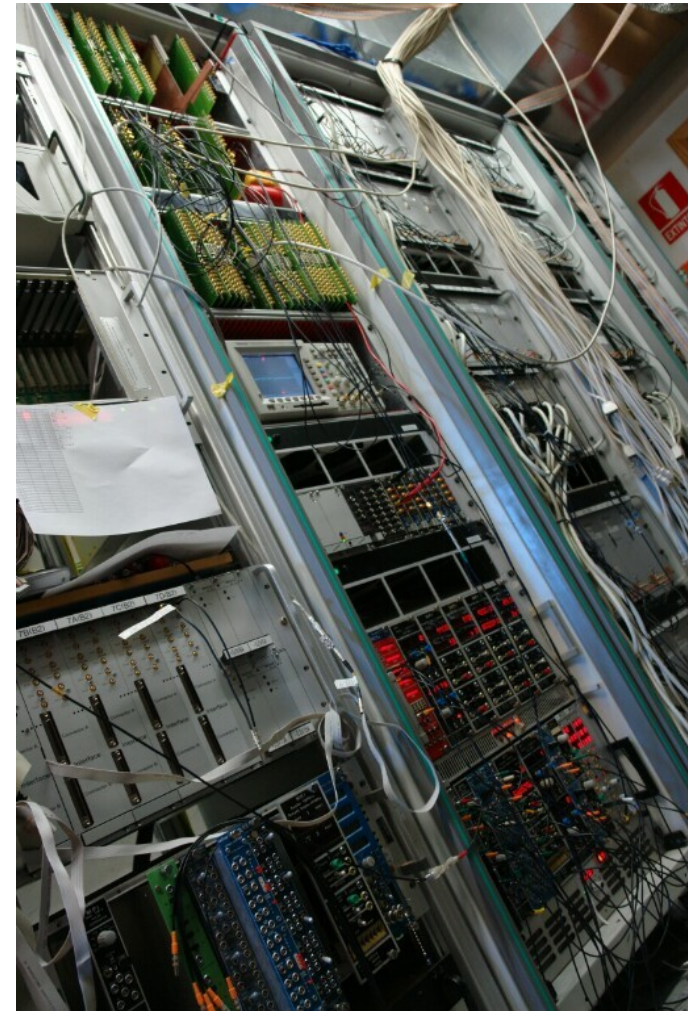
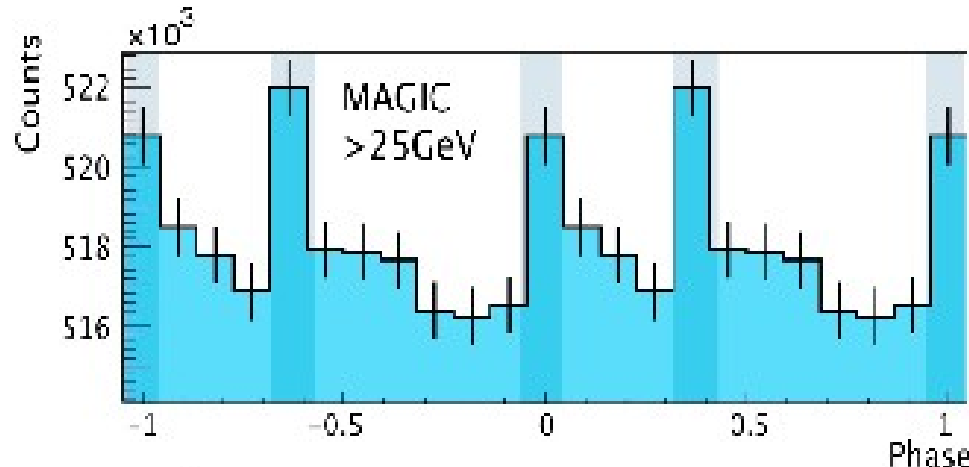
Parameter S21 ("Gain")

MAGIC-I Sum-Trigger

- Start (Development): April 2007
- Fabrication and test: June – August 2007
- Installation in La Palma: September 2007
- Taking first data: October 2007

Reaching the lowest energy threshold (~25 GeV)
ever achieved by any Cherenkov telescope up to
date

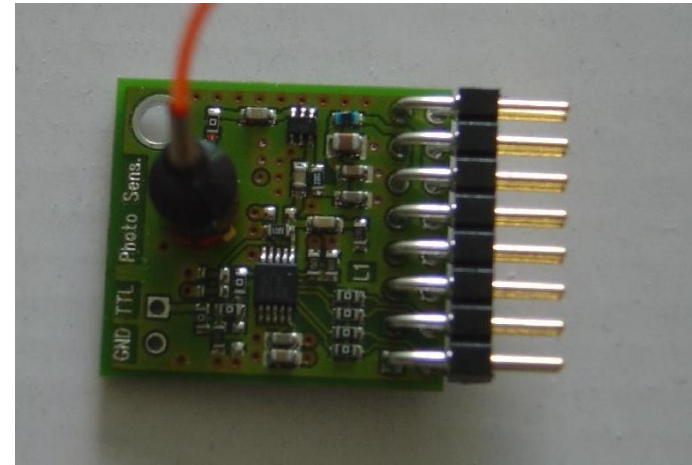
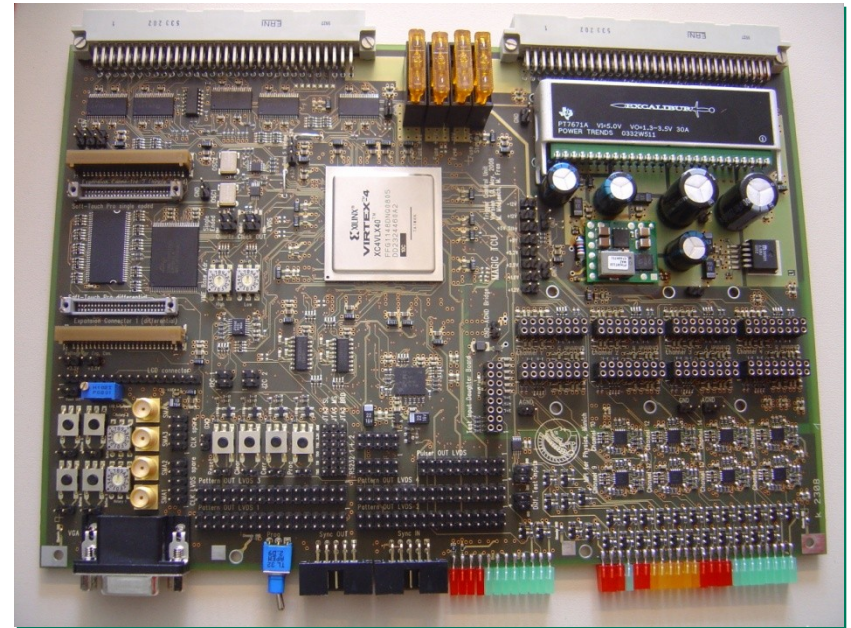
Light curve of Crab pulsar (2008):



Additional MAGIC-II Development

Control system for calibration laser,
test pulser, ...

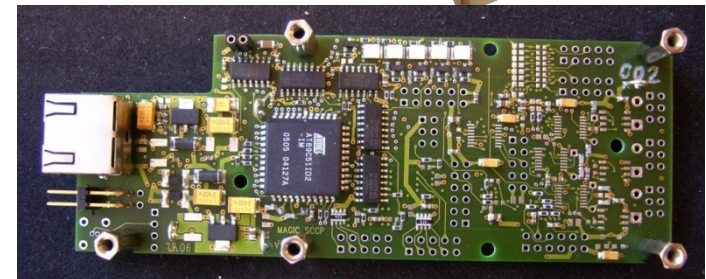
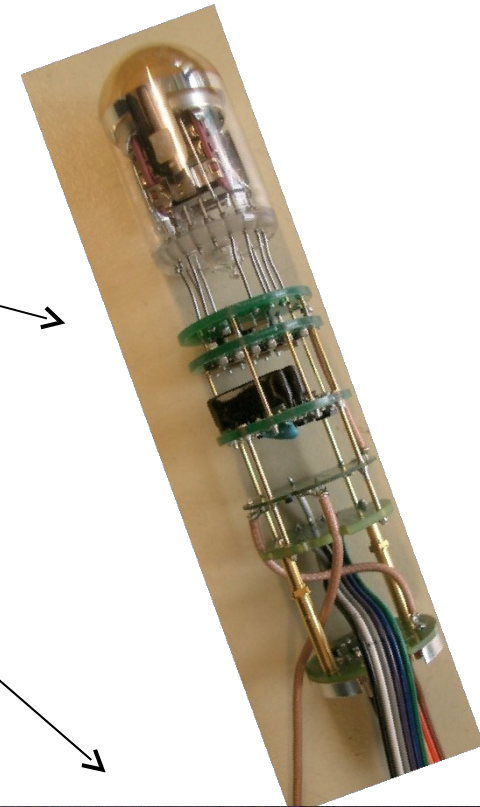
- FPGA: Xilinx Virtex-4 4VLX40FF1148
- 8 sockets for pulser daughter boards
- Pulse frequency: 0.023 Hz .. 50 MHz
- Pulse width: 10 ns .. 42.9 s
- Leading edge of the 16 pulses adjustable in steps of 11 ps
- 1 socket for pulse input daughter board with max. 2 input channels
- 2 x RS232 input, one on default front panel
- Connector to attach a 4x20 LCD
- Mezzanine board socket with 36 single ended signals and 19 differential
- Option for external clock for pulsers
- Option to cascade several boards
- VGA output
- board for optical link



Production for MAGIC-II in 2008

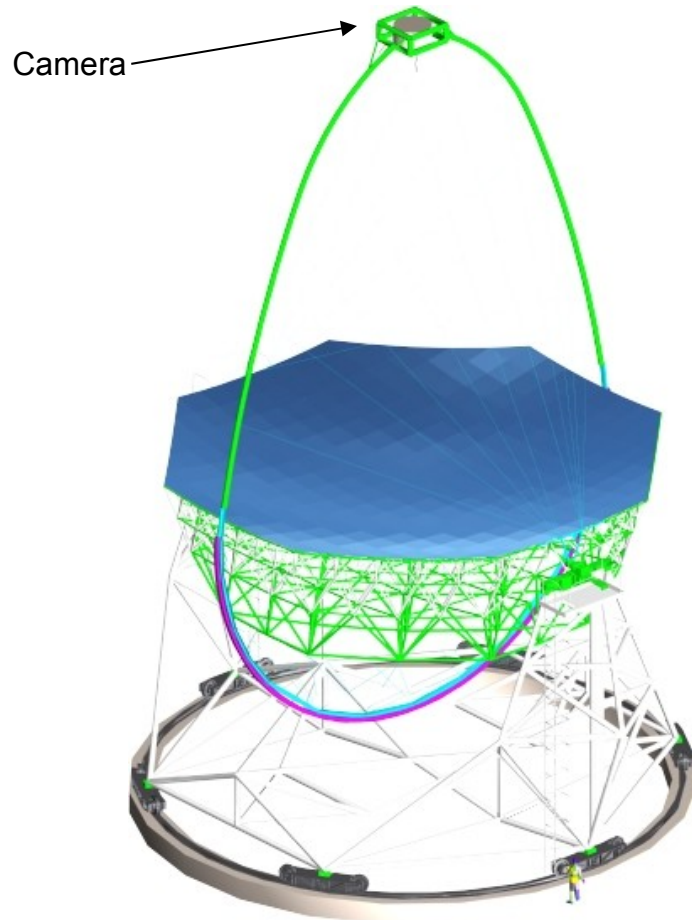
- Production of pixel-boards
- Assembling the pixels
- Production of control boards (SCCP)
- Production of test pulser boards
- Test of all parts and the assembled pixels
- Assembling the clusters
- Testing the clusters (function , flatfielding, ...)
- 169 clusters (+spares) are ready in October
- Shipped to La Palma in November
- Installing the clusters in only two weeks
- After installation the camera is “switched on” in only one week

The installation is completed 3-4 months before the plan!



MAGIC-II Overview

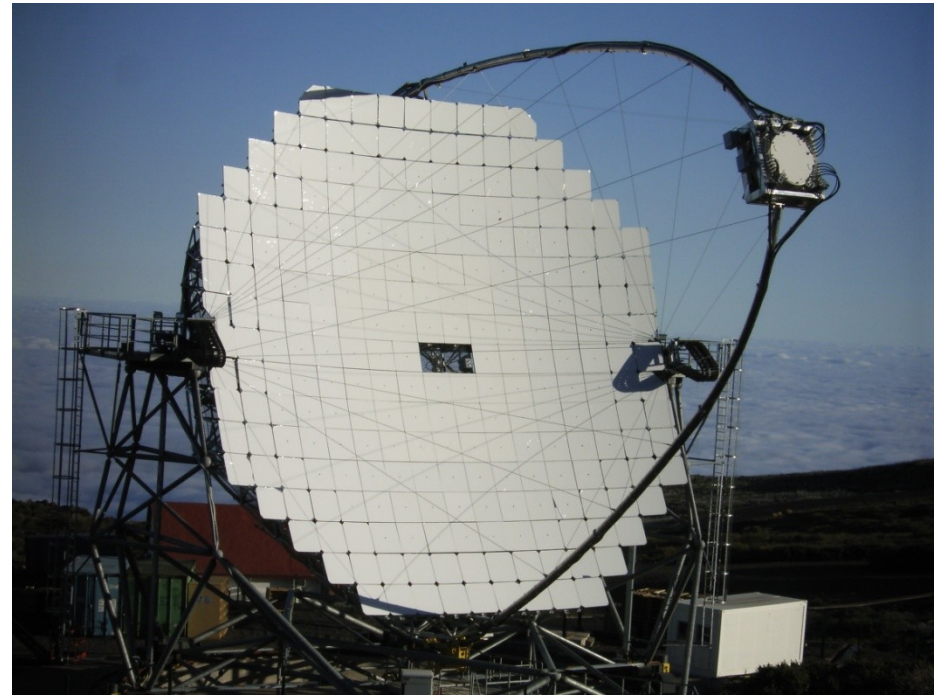
Situation 2007:



Main task:

- Development of camera electronics
 - Signal transmission system
 - Camera control system
 - Test signal generation
- Power distribution

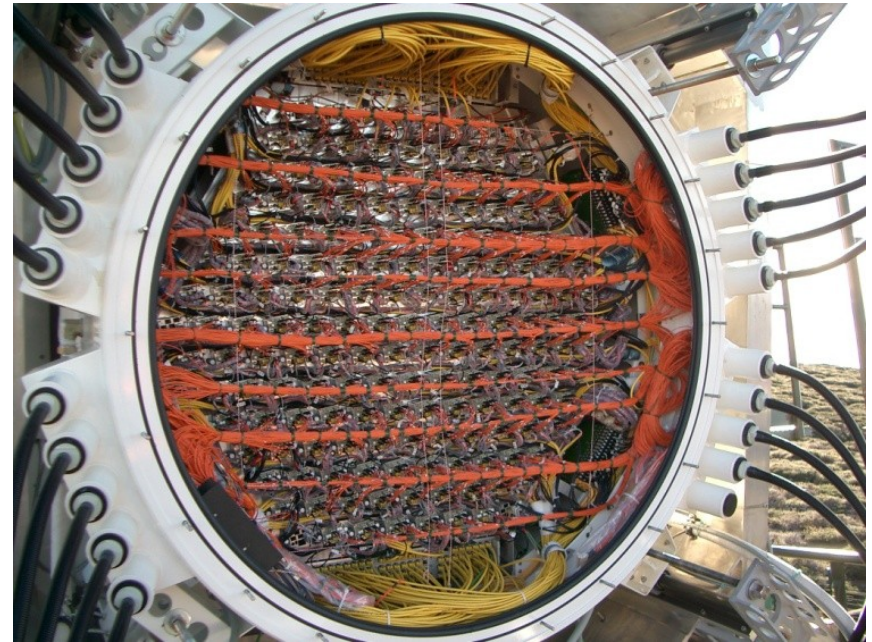
November 2008:



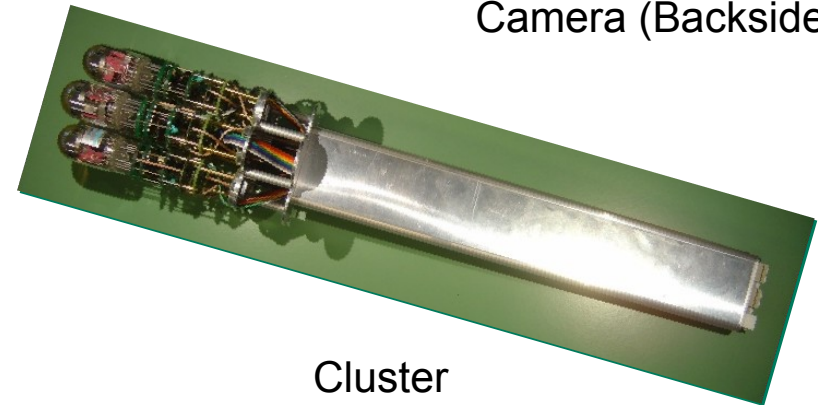
MAGIC-II Camera

- 1039 pixels
(photomultipliers + signal transm.)
- 7 pixels are grouped into a cluster
- Each cluster has its own test pulse generation and control system

Camera (Frontside)

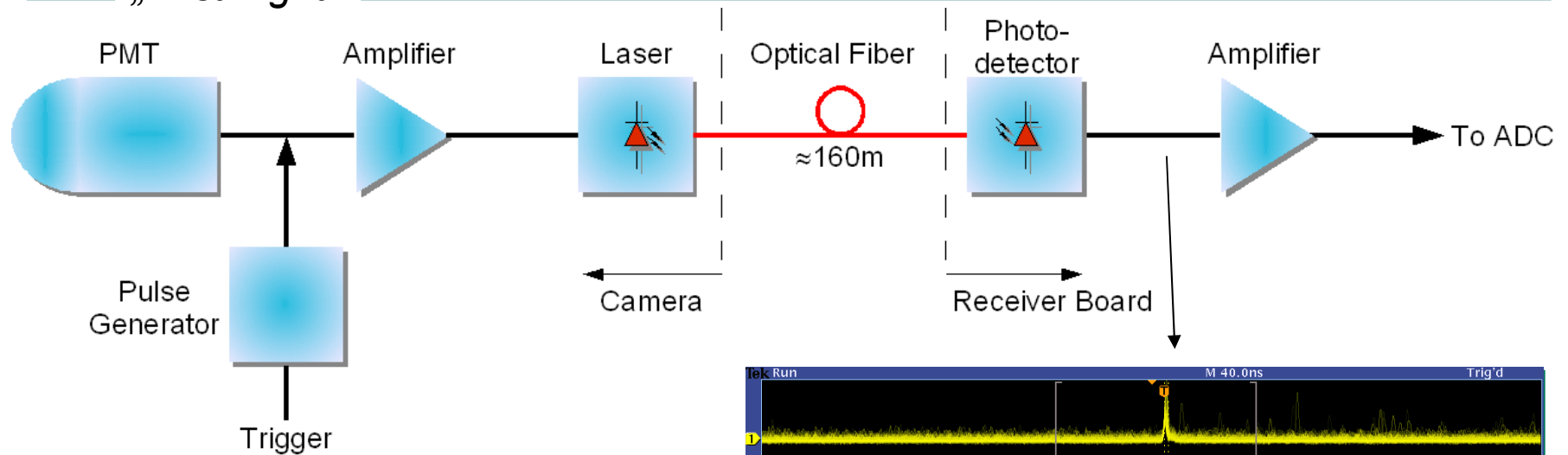


Camera (Backside)



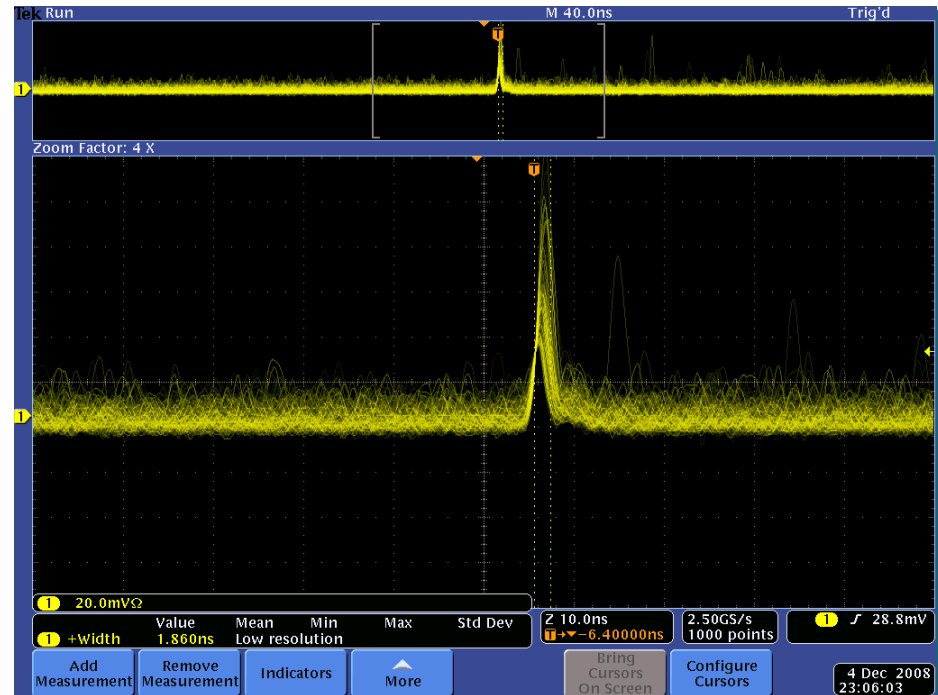
Cluster

„First Light“



4. December 2008:

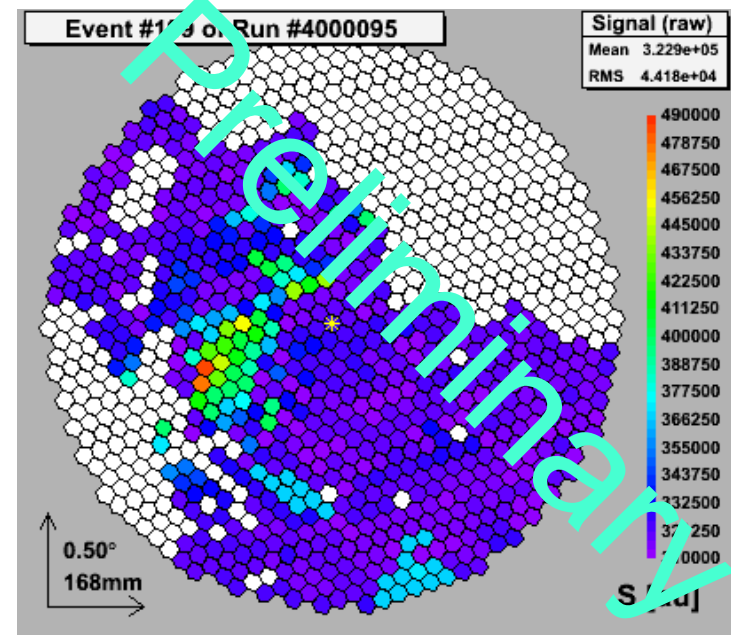
- First look to the night sky background
- Typical pulsewidth: 1.8ns
- Small coupling from VME-access at some outer ring pixels
(Will be solved in spring 2009)



First Measurement

Excerpt from an e-mail received from La Palma (Juan Cortina , last week):

- Checked DCs for dark time (~ 1 uA, fine) and we saw stars drifting when pointing at zenith
- Introduced default HV settings based on measurements at MPI.
They produce relatively flat anode DCs
- We have taken the first showers with MAGIC-II!



Thanks to all people making possible these nice results, especially the HEC-I crew and the MAGIC crew for their encouraged work to be ready in time.

*Thank you very much
for your attention*

