

Jets + Bremsstrahlung = OPE + ?

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Ringberg Workshop on **New Physics**, **Flavors** and **Jets**

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What does all this mean?

- Fermi motion
- Bremsstrahlung
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- Discussions and conclusions

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- Is there an effective field theory for SCET in reality?

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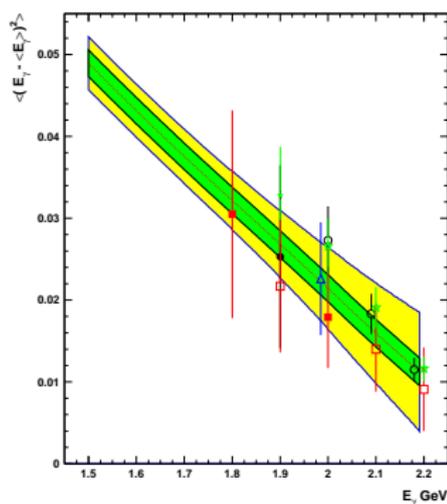
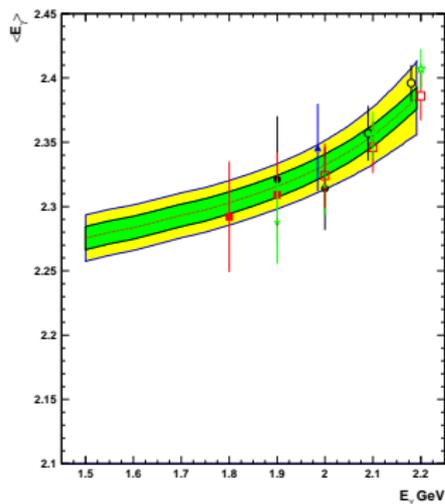
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‘Theoretical correlations’

Check in a different *environment*:

$b \rightarrow$  light  $q$  decays

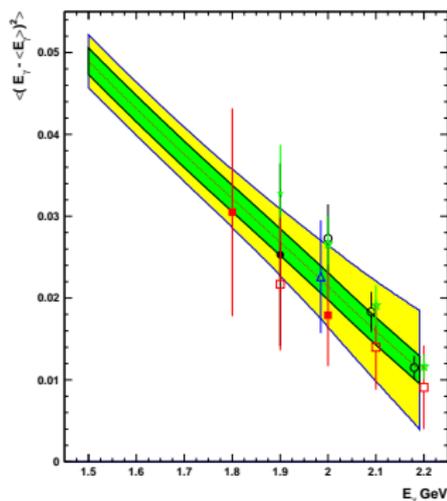
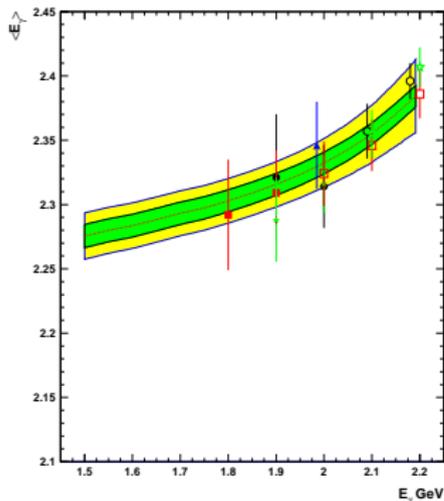
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$b \rightarrow$  light  $q$  decays have some subtleties which represent the problems for applying the OPE

These motivated the theoretical study underlying this talk

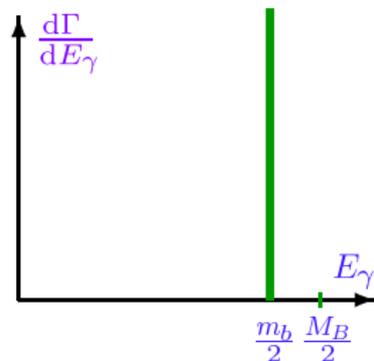
# Fermi motion

Inclusive distributions are close to partonic?

$$b \rightarrow s + \gamma:$$

$$E_\gamma \simeq \frac{m_b}{2} \quad \frac{d\Gamma}{dE_\gamma} \sim \delta\left(E_\gamma - \frac{m_b}{2}\right) ?$$

$$B \rightarrow X_s + \gamma: \quad E_\gamma = \frac{M_B^2 - M_X^2}{2M_B} \quad - \quad \text{There should be a distribution...}$$



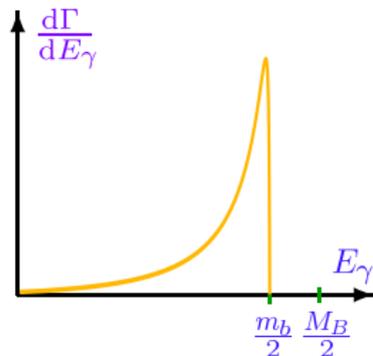
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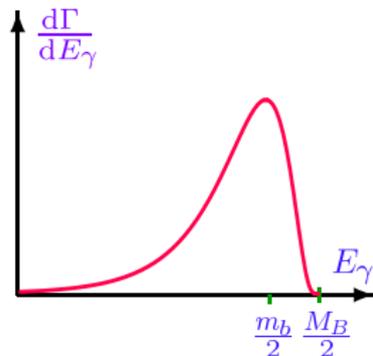
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- Intrinsic motion of the heavy quark

with  $\vec{p}_b \neq 0$   $E_\gamma = \frac{m_b}{2} + \frac{\vec{p}_b \vec{n}_\gamma}{2}$ : Doppler smearing

$b$ -quark shivers in  $B$ -hadron,

$$\vec{p}_b \sim \Lambda_{\text{QCD}}$$

$$\vec{v}_b \sim \frac{\Lambda_{\text{QCD}}}{m_b}$$

Phenomenological models incorporating Doppler smearing in the spectrum were called upon early on

Ali, Pietarinen 1979  
 $AC^2 M^2$  1982

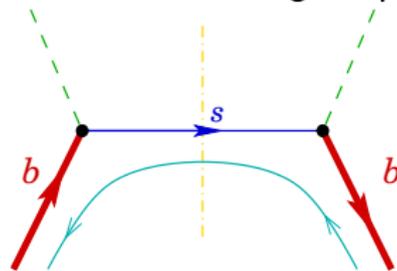
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In fact the *light-cone* momentum component  $\pi_0 - \vec{\pi} \vec{n} = \pi_\mu n^\mu$  enters,  $\vec{n}^2 = 1$

Comes from the *s*-quark propagator in the forward scattering amplitude

$$\bar{b} \dots \frac{1}{m_b^2 - (p_b + \pi_b - q)^2 - \frac{i}{2} \sigma G} \dots b ,$$





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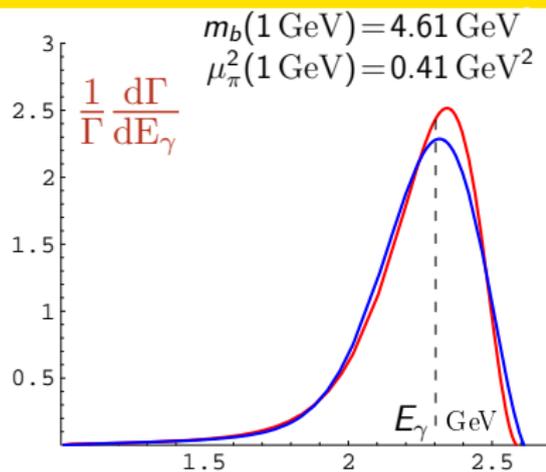
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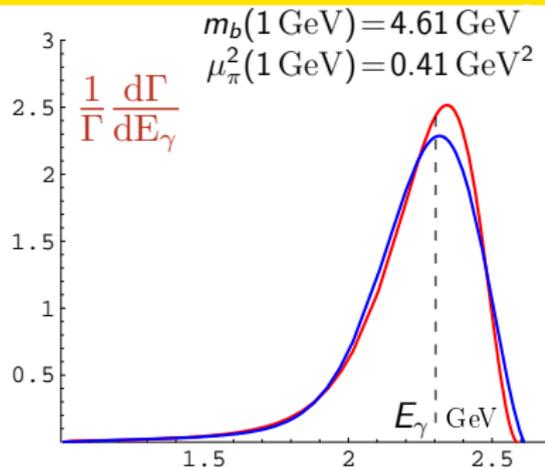
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treating nonperturbative effects and the OPE



## OPE:

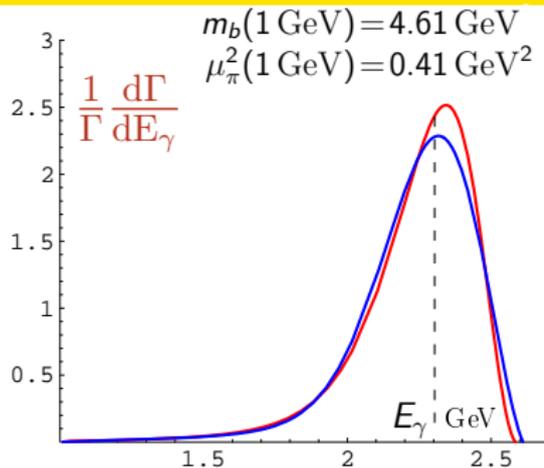
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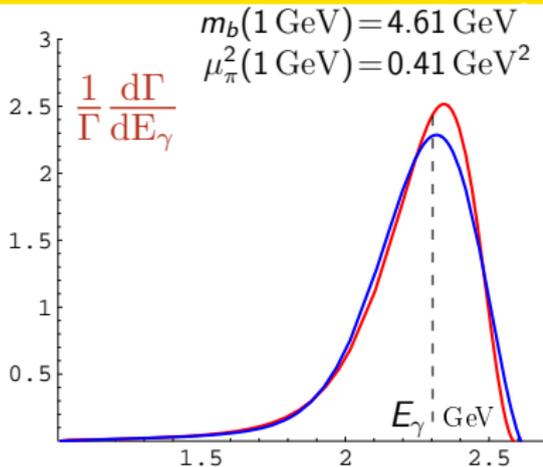


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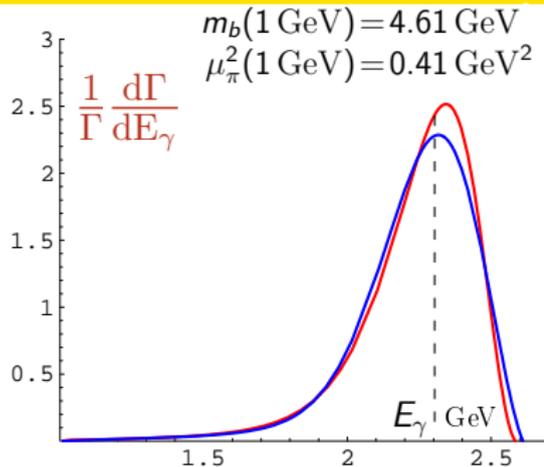
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 $|\vec{k}| > \mu$  - short distances; then  $\alpha_s$  enters at  $Q^2 > \mu^2$

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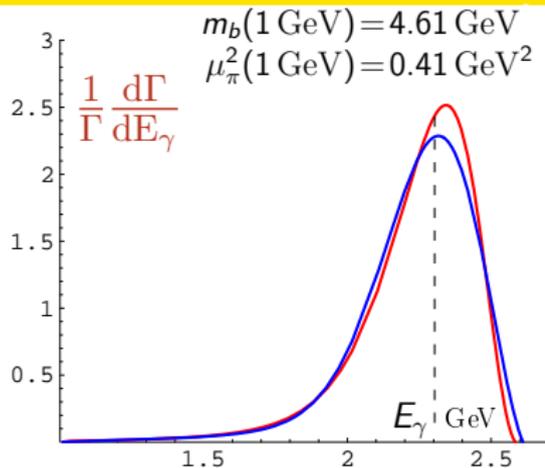
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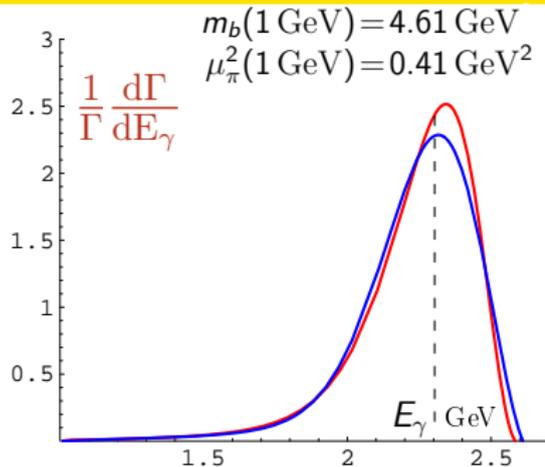
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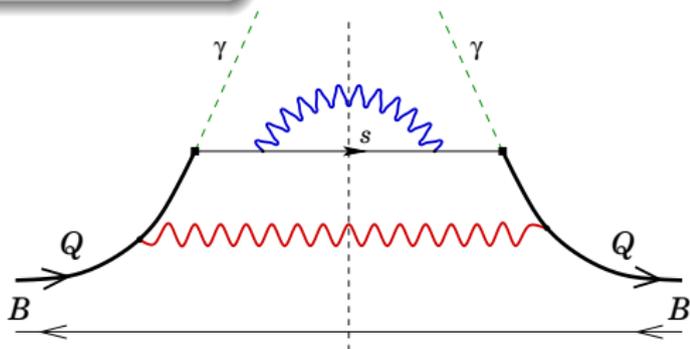
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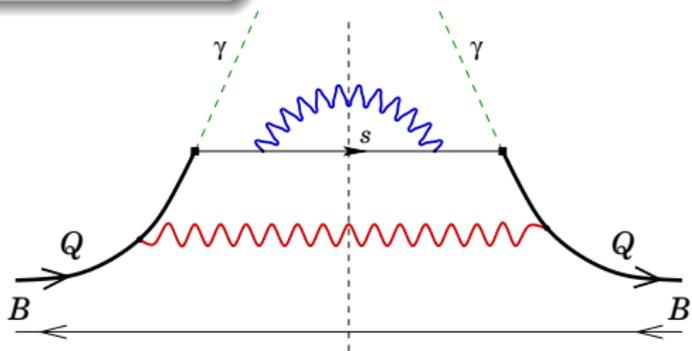
OPE expands in the fields' momenta  $\implies$  have to assume that all  
 the components of the momentum are literally small for the n/p fields.



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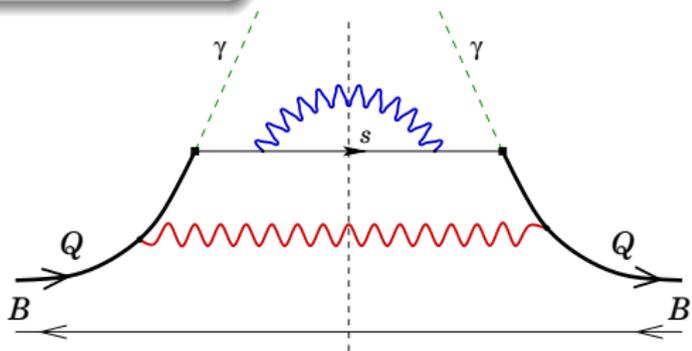
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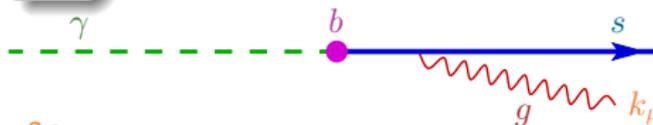
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$\frac{d\Gamma_{\text{pert}}}{dE}$  accounts for all other modes

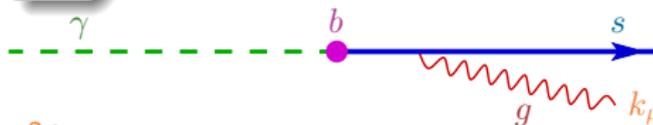
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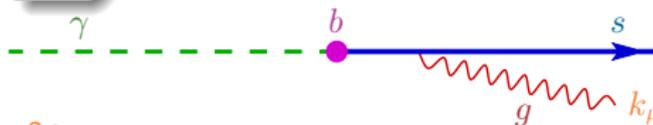


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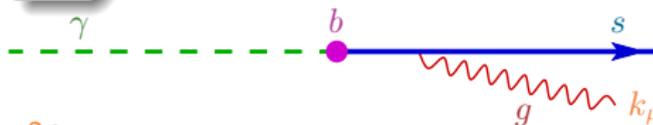
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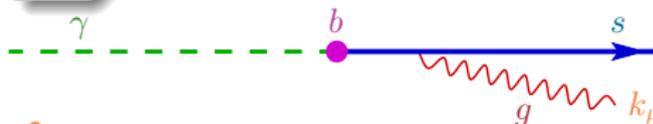
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Moments of  $E_\gamma$  in the Wilsonian OPE ?

No virtual contributions to cancel...

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Additional nonperturbative terms on top of the OPE?

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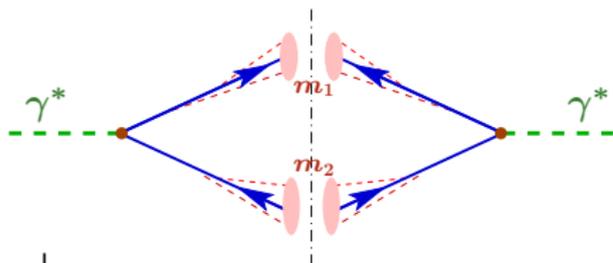
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Extra effects here would lead to certain paradoxes

Consider  $\sigma(e^+e^- \rightarrow \text{hadrons})$  from this perspective

no OPE for jets

Look at it now as a decay of the photon with the mass  $\sqrt{s}$  into two jets:



$$\Gamma \sim 1 - \frac{m_1^2 + m_2^2}{s} + \frac{(m_1^2 + m_2^2)^2}{s^2} + \dots$$

If there were  $\delta m_{\text{jet}}^2$ , we would normally have the phase-space-related  $1/s$ ,  $1/s^2$ , ... corrections in the total rate, not only in jet distributions where no Euclidean OPE applies.

However, for the *total cross section* we know from the OPE that

$$\delta_{\text{IR}} R(s) \sim \frac{1}{s^3} + \alpha_s \frac{1}{s^2} + \dots$$

$$\text{since } D(Q^2) \sim 1 + (1 + \alpha_s)/Q^4 + \dots$$

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This is a consequence of analyticity and unitarity for the polarization operator of the currents

$$\Pi(q^2)(-q^2\delta_{\mu\nu} + q_\mu q_\nu) = \int d^4x e^{iqx} \langle 0 | iT \{ J_\mu(x) J_\nu(0) | 0 \rangle =$$



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If we know the asymptotics of  $D(Q^2)$  in deep Euclid, we know power corrections in  $R(s)$  at large  $s$ :



Total probability depends only on 'usual' nonperturbative corrections  
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No, integer moments of jet mass<sup>2</sup> are not affected!

# Kinematics and the OPE

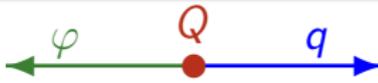
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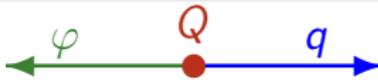
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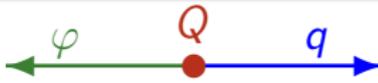
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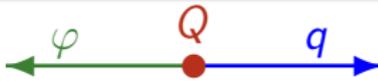
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Usual nonperturbative (OPE) window is  $M_X^2 \lesssim \mu_{\text{hadr}} m_b$

larger  $M_X^2 \gg \mu_{\text{hadr}} m_b$  come from hard bremsstrahlung

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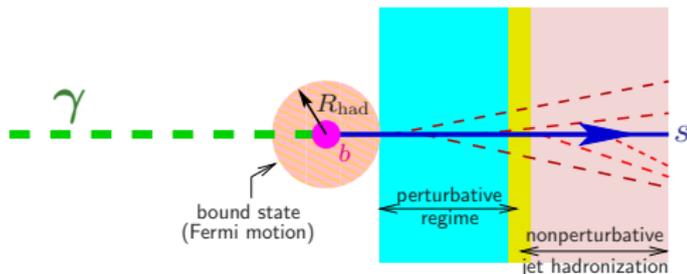
Physics: this is not the bound-state dynamics!

Travel distance

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much larger than  $\mu_{\text{hadr}}^{-1}$ 

The jet hadronization process is space- and time- separated from the bound state once  $\omega \gg \mu_{\text{hadr}}$

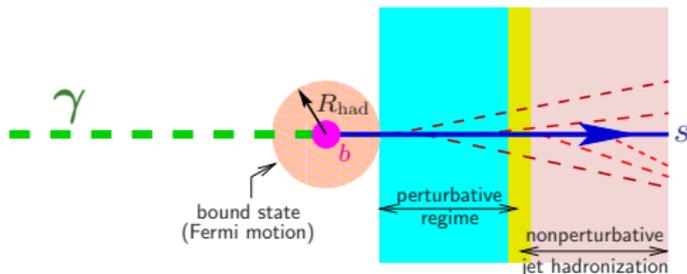


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In the **total width** virtual corrections cancel these radiation effects no matter what  $\alpha_s$  coupling is

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How to show this?

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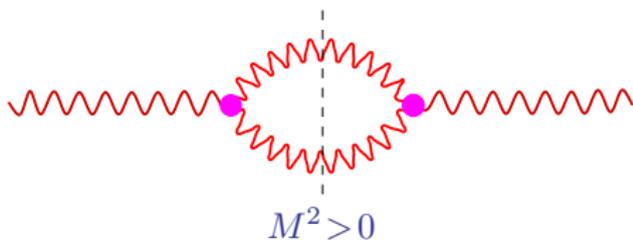
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illustrate in the framework of

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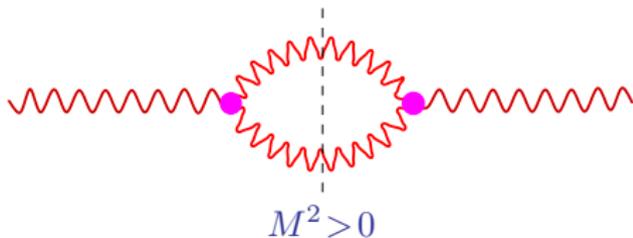
$$\frac{\alpha_s^\epsilon(Q^2)}{Q^2} = \pi \int \frac{d\lambda^2}{\lambda^2} \frac{\rho(\lambda^2)}{\lambda^2 + Q^2},$$

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Instead of  $\frac{\alpha_s}{k^2}$  we then put inside the diagrams  $\int d^4k \frac{\alpha_s(k^2)}{k^2} \dots$ :

$$\pi \int \frac{d\lambda^2}{\lambda^2} \rho(\lambda^2) \underbrace{\int \frac{d^4k}{(2\pi)^4} \frac{1}{k^2 + \lambda^2} \dots}$$

the would be loop correction with the massive gluon

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in jet physics the nonperturbative effects are given by the log-moments of  $\delta\rho$

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but the domain  $k_{\perp}^2 \lesssim \mu_{\text{had}}^2$  yields  
a nonperturbative piece  $\sim \mu_{\text{had}}^{2n}$

Do explicitly:

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Product of independent integrals over  $k_\perp^2$  and  $\lambda^2$ ! No  $\alpha_s(k_\perp^2)$ ,  
minimal scale  $Q^2$  is determined by the kinematics, it is  $\mu$  at worst,

$$\text{for } \int \frac{d\lambda^2}{\lambda^2} \rho(\lambda^2) = \lim_{Q^2 \rightarrow \infty} \delta\alpha_s^\epsilon(Q^2) = 0$$

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# Where the miracle hides?

Calculate the  $M_X^2$ -spectrum itself:

$$\begin{aligned}\frac{d\Gamma^{\text{pert}}}{dM_X^2} &= C_F \int \frac{d\omega}{\omega} \vartheta(\omega - \mu) \int \frac{d\lambda^2}{\lambda^2} \rho(\lambda^2) \int \frac{dk_{\perp}^2}{k_{\perp}^2 + \lambda^2} \delta(M_X^2 - (k_{\perp}^2 + \lambda^2) \frac{m_b}{2\omega}) \\ &= \frac{C_F}{M_X^2} \int \frac{d\omega}{\omega} \vartheta(\omega - \mu) \int \frac{d\lambda^2}{\lambda^2} \rho(\lambda^2) \vartheta(M_X^2 - \frac{m_b}{2\omega} \lambda^2)\end{aligned}$$

instead of

$$\frac{C_F}{M_X^2} \int \frac{d\omega}{\omega} \vartheta(\omega - \mu) \underbrace{\int \frac{d\lambda^2}{\lambda^2} \rho(\lambda^2) \frac{\frac{2\omega}{m_b} M_X^2}{\lambda^2 + \frac{2\omega}{m_b} M_X^2}}_{\frac{1}{\pi} \alpha_s^{\epsilon} \left( \frac{2\omega}{m_b} M_X^2 \right)}$$

The radiation is now driven by a *different* effective coupling  $\tilde{\alpha}_s(k_{\perp}^2)$ :

$$\delta\tilde{\alpha}_s(Q^2) = \pi \int_0^{Q^2} \frac{d\lambda^2}{\lambda^2} \rho(\lambda^2) \quad \text{vs.} \quad \delta\alpha_s^{\epsilon}(Q^2) = \pi \int_0^{\infty} \frac{d\lambda^2}{\lambda^2} \frac{Q^2}{\lambda^2 + Q^2} \rho(\lambda^2)$$

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The difference is significant: Integer moments of  $\delta\tilde{\alpha}_s(Q^2)$  all vanish, while those of  $\delta\alpha_s^\epsilon(Q^2)$  are 'positive'

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Effective  $\alpha_s$  becomes observable-dependent since using it implies we forcibly interpret the process in the massless one-gluon language

## Physics behind:

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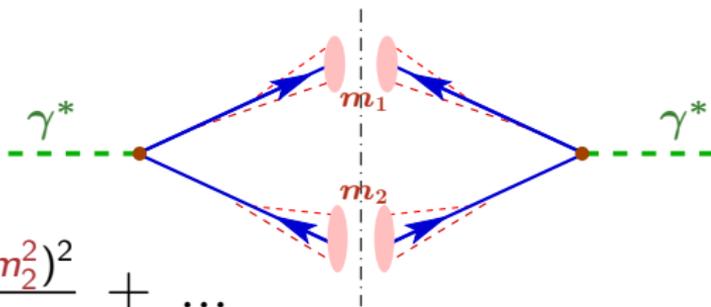
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KLN implies that 'rescattering' does not affect the total probabilities and other truly local observables. That is what we get

Go back to  $\sigma(e^+e^- \rightarrow \text{hadrons})$

$$\delta_{\text{IR}} R(s) \sim \frac{1}{s^3} + \alpha_s \frac{1}{s^2}$$

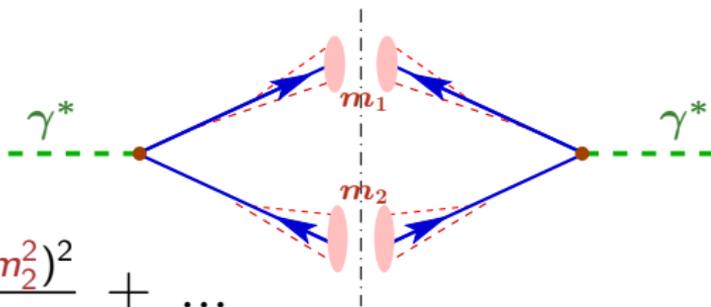


$$\Gamma \sim 1 - \frac{m_1^2 + m_2^2}{s} + \frac{(m_1^2 + m_2^2)^2}{s^2} + \dots$$

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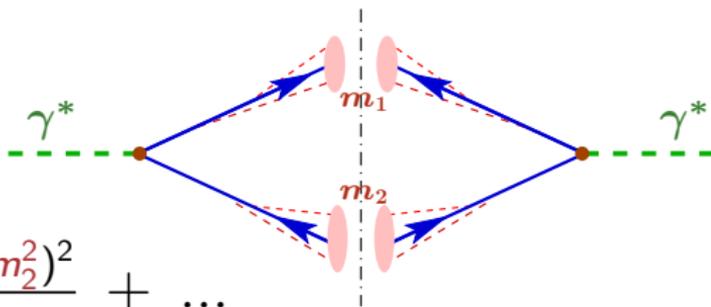
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We got this for inclusive jet moments as well

## My conclusions

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The Wilsonian OPE nonetheless is not affected. The usual moments are given by the local heavy quark expectation values, plus the perturbative corrections originating only from the short-distance domain

$\mu$ -dependence of the moments is given by  $\alpha_s^\epsilon(\mu)$

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*These are not bogus!*



Quite general reasons suggest it may be problematic to construct a real effective field theory like HQET for the light-like jet physics. Its Lagrangian may hardly be anything but the original QCD one: the notion of “fast jet” is meaningful only in the context of having a spectator **around** (or another jet, initial color source, ...). For a given quark alone one can always take a boost to view it as having ‘normal’ energy or momentum

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**Jet bremsstrahlung** adds a more specific doubt, it has to do with what one expects from an effective field theory. Like  $m_Q \rightarrow \infty$  in HQET, we need to take here the limit of jet energy to infinity. This may be meaningful provided increasing  $m_Q$  resulted in a pure perturbative convolution. It would then have been accounted for by perturbative ‘renormalization’ of the initial jet (cf. HQET). However, increasing the physical high mass scale — say, passing from  $m_Q$  to  $2m_Q$  in  $b \rightarrow s + \gamma$  — brings in new emissions of strongly coupled collinear gluons with  $k_\perp \lesssim \Lambda_{\text{QCD}}$  but energy between  $m_Q/2$  and  $m_Q$ , that are **nonperturbative**. Their effect generally is not suppressed, rather logarithmically enhanced. If one has to involve an **unsuppressed nonperturbative** matching, an effective theory may lose its primary textbook motivation

The standard expression used for the resummed distribution in the Mellin space

$$\Delta_N = \int_0^1 dz z^{N-1} \Delta(z)$$

has the form

$$\ln \Delta_N(Q^2) = \frac{C_F}{\pi} \int_0^1 dx \frac{x^N - 1}{1-x} \left[ \int_{Q^2(1-x)^2}^{Q^2(1-x)} \frac{dk_{\perp}^2}{k_{\perp}^2} \alpha_s^{\epsilon}(k_{\perp}^2) - B \alpha_s^{\epsilon}(Q^2(1-x)) - D \alpha_s^{\epsilon}(Q^2(1-x)^2) \right]$$

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While correctly describing the contribution of multiple emissions with different momenta,  $\alpha_s(k_{\perp}^2)$  does not count precisely the contributions of a particular gluon “virtuality”  $\mu^2$  in the emission: it yields  $\ln\left(\frac{Q^2}{\mu^2} + \frac{1}{1-x}\right)$  instead of  $\ln \frac{Q^2}{\mu^2} + \ln \frac{1}{1-x}$ ; a **different** effective coupling  $\tilde{\alpha}_s$  must be there. This is reshuffled into  $B$  and  $D$  through explicit integration using the logarithmic running of  $\alpha_s$  with the scale, to any particular N<sup>k</sup>LO, but potentially introduces large higher-order corrections

This is related to running of the coupling, and therefore is complementary to the Catani-Mangano-Nason-Trentadue analysis of resummation

I think it is advantageous to pass here to the proper effective coupling. This usually significantly reduces higher-order corrections and yields accurate numerical estimates already with much simpler quasi-LO calculations, at least for sufficiently inclusive distributions