



Status and Perspectives in String Theory

Dieter Lüst, MPP and LMU-München

MPP: 100 Year Anniversary, October 12th, 2017



Outline:

String theory in a nutshell

The string theory landscape

Some general lessons about strings and quantum gravity

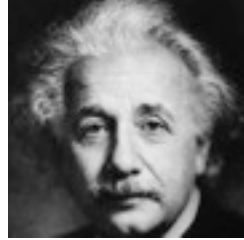
Tension between different theories:

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Electromagnetism \longleftrightarrow classical mechanics

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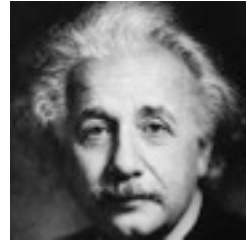
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Special and General Relativity

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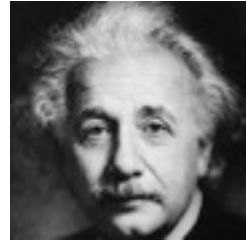


Special and General Relativity

Classical mechanics \longleftrightarrow Stability of atoms

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Special and General Relativity

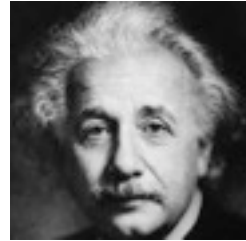
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Quantum Mechanics

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Electromagnetism ↔ classical mechanics



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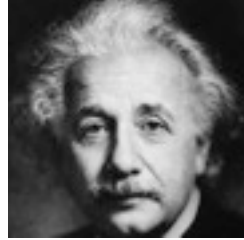


Quantum Mechanics

Einstein gravity ↔ Quantum mechanics

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Quantum Gravity ?

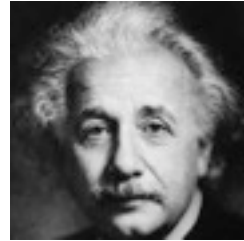
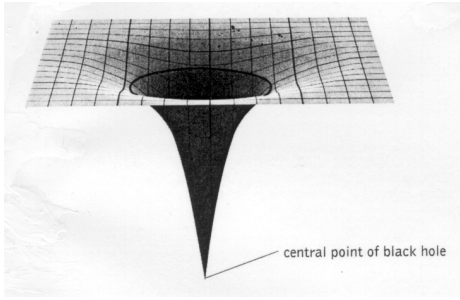
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Time dilatation,
black holes



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Classical mechanics



Stability of atoms



Quantum Mechanics

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Quantum mechanics



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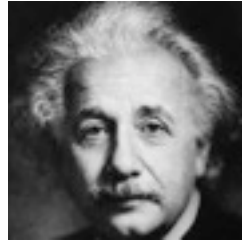
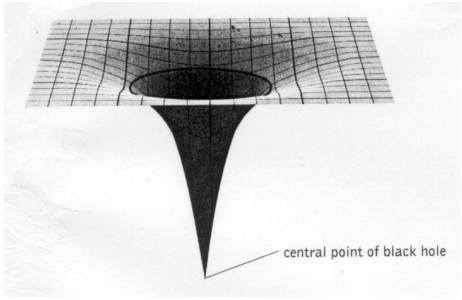
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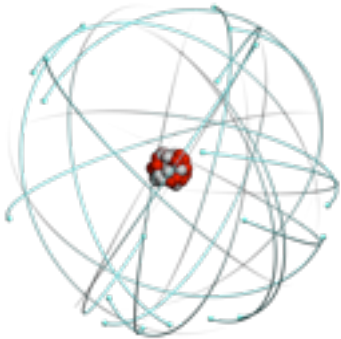


Special and General Relativity

Classical mechanics



Matter waves,
entanglement



Quantum Mechanics

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Quantum mechanics



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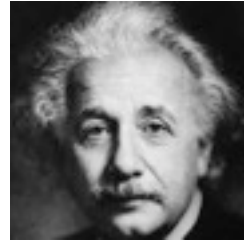
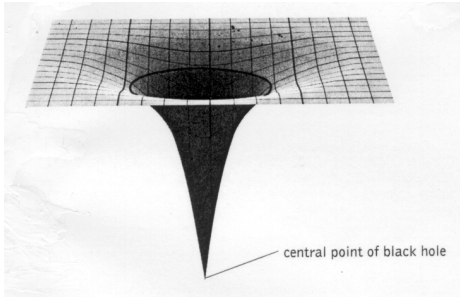
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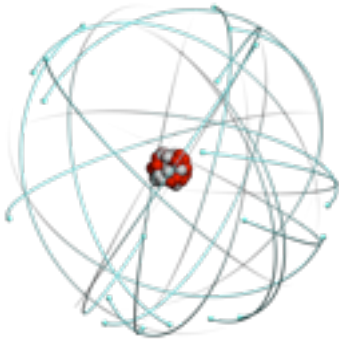


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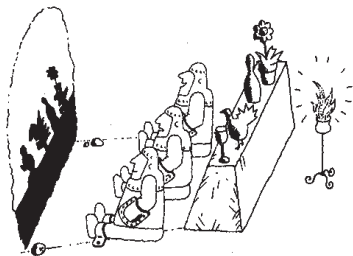
Quantum Mechanics

Einstein gravity



Quantum Gravity

Strings,
extra dimensions
emergent space-time ?

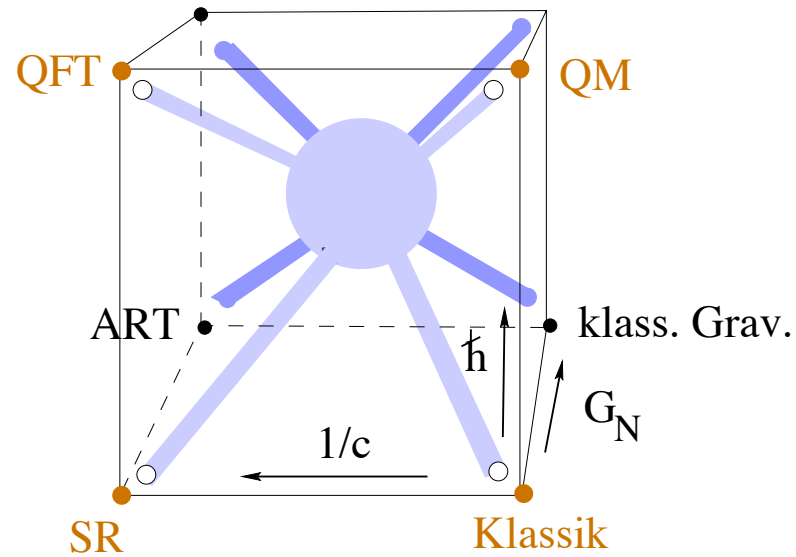


Quantum Gravity ?

Mass scale of Quantum Gravity:

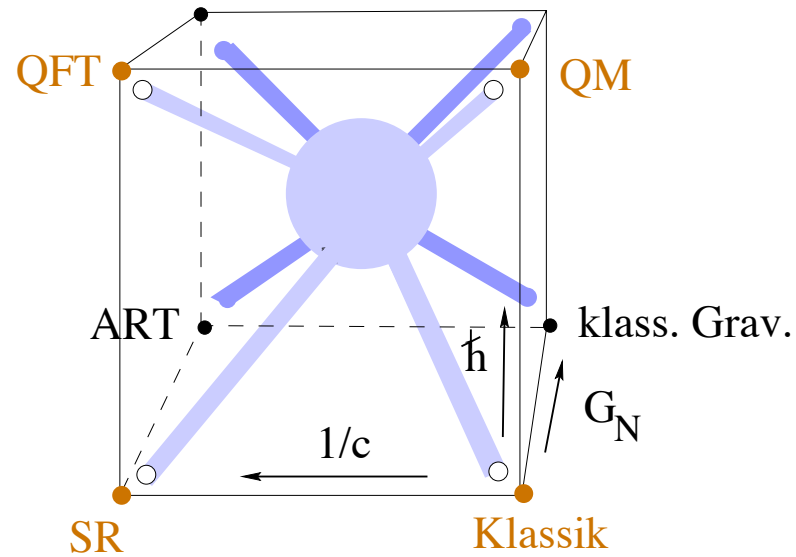
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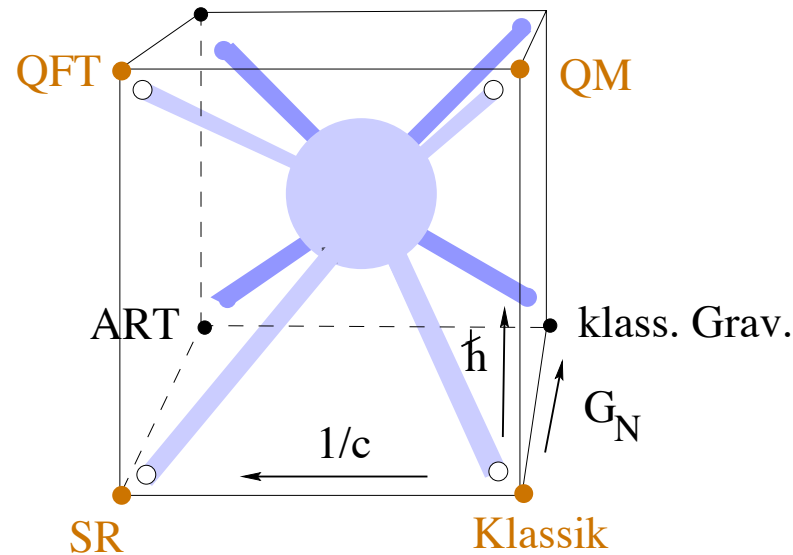
Strings/M-Theorie ?



$$M_{\text{Planck}} = \sqrt{\frac{\hbar c}{G_N}} \simeq 10^{19} \text{ GeV}, \quad L_{\text{Planck}} = \sqrt{\frac{\hbar G_N}{c^3}} \simeq 10^{-35} \text{ m}$$

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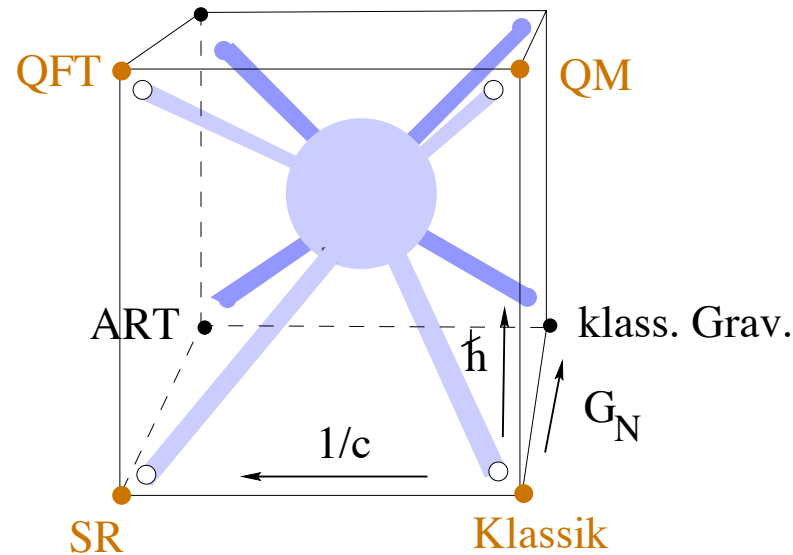
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These are extremely high energies and very short distances:

⇒ **Physics of the Big Bang!**

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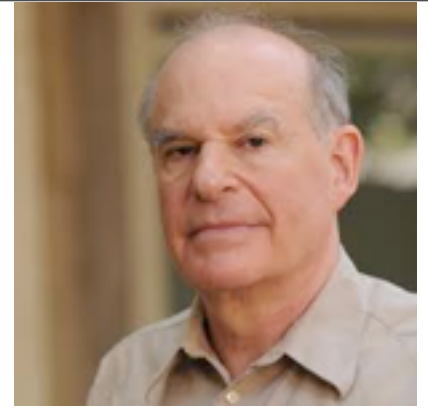
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⇒ **Physics of the Big Bang!**

At this scale gravity becomes strong !!

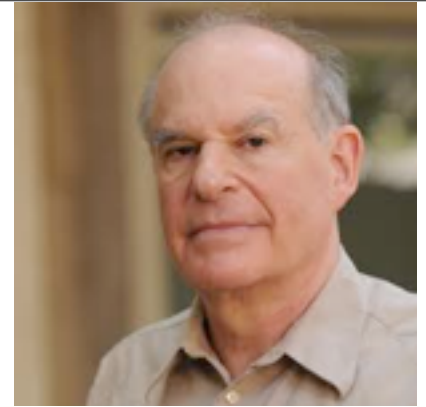
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(G.Veneziano, 1968; J. Scherk, J. Schwarz, 1974)



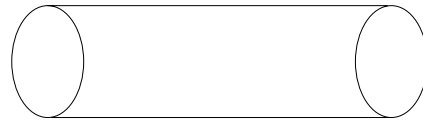
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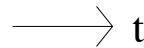


Theory of (1-dim) extended objects:

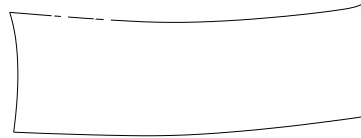
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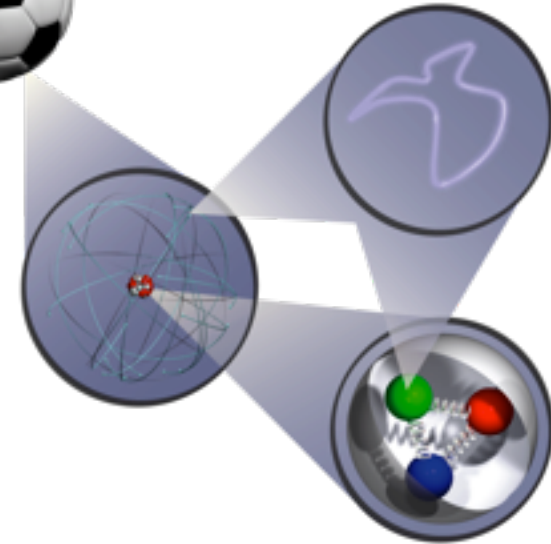
geschlossener
String



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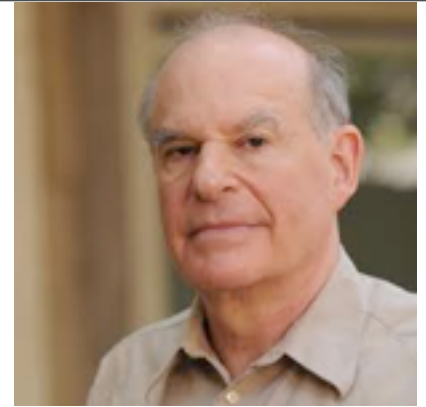


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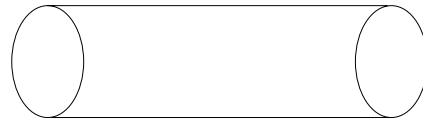
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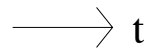


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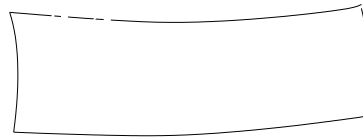
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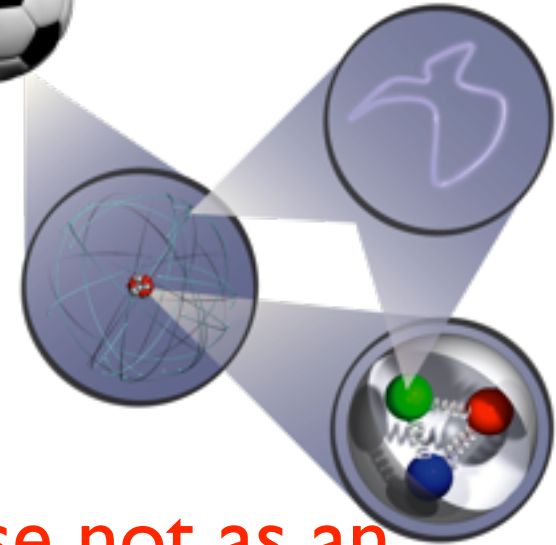
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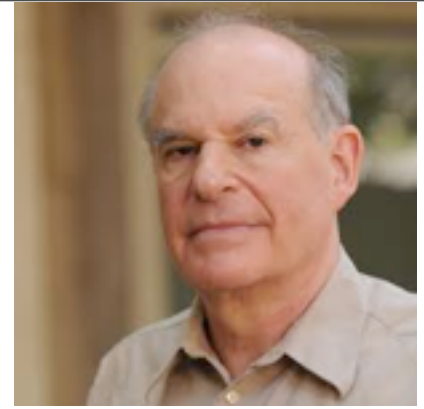
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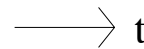


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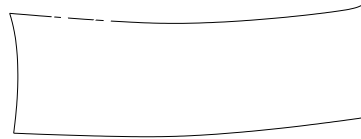
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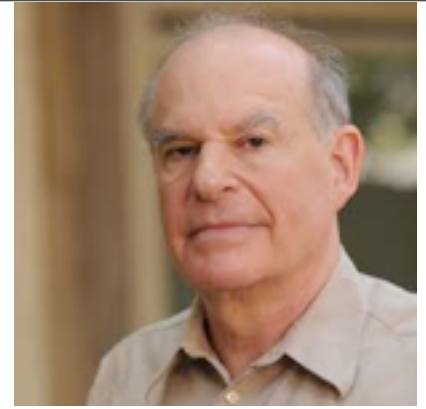


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- Massless spin-two graviton: $\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{Einstein}}$

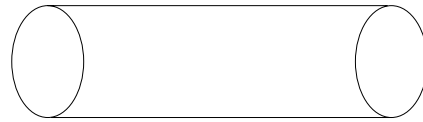
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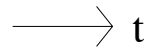


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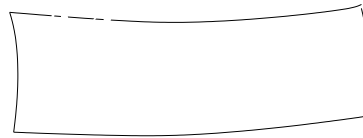
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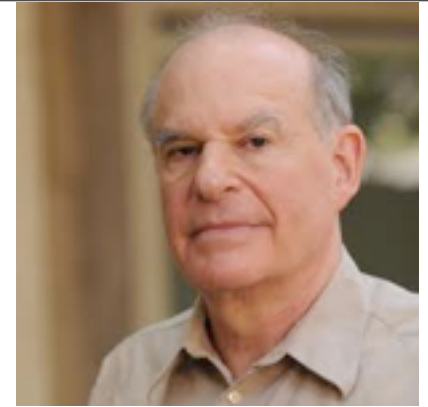


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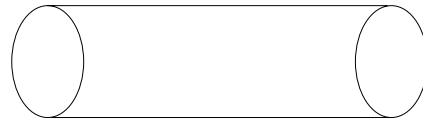
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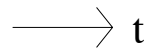


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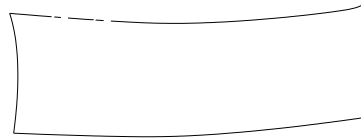
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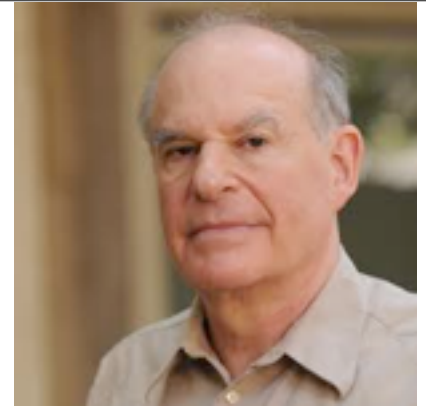
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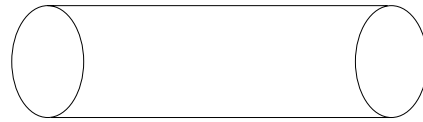
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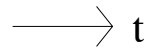


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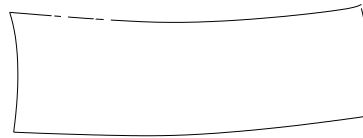
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- No UV divergences: consistent theory of quantum gravity

(S. Mandelstam, 1992)

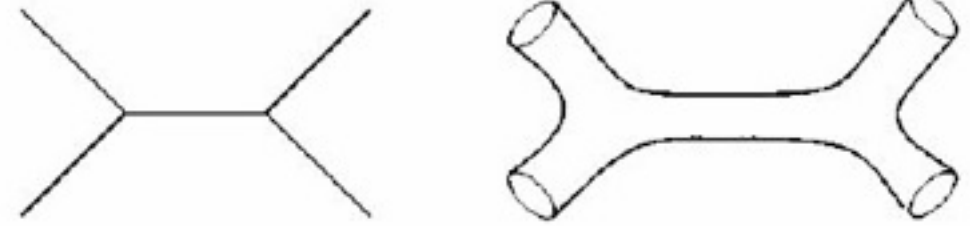


Fig.3: Particle scattering processes (left), string scattering processes (right).

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- Critical dimension: $D=10$

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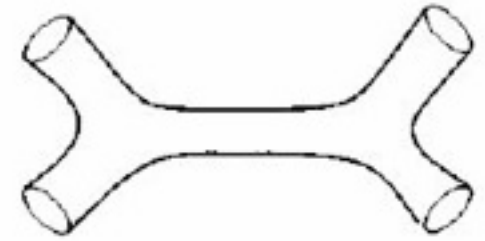
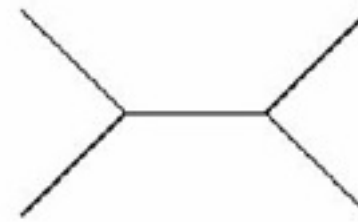


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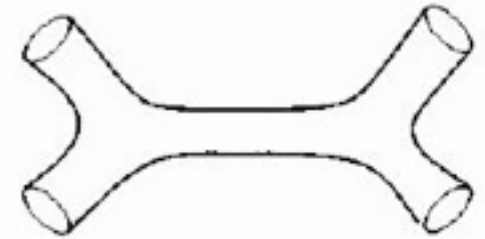
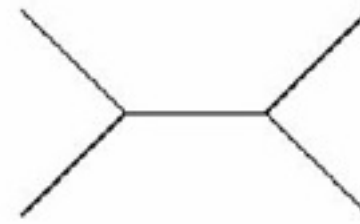


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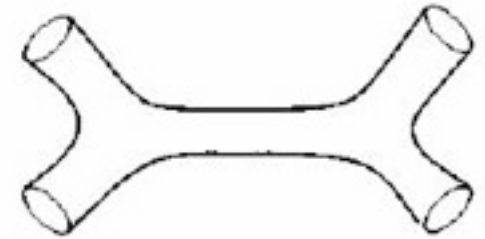
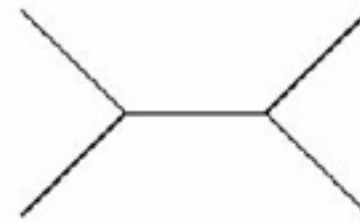


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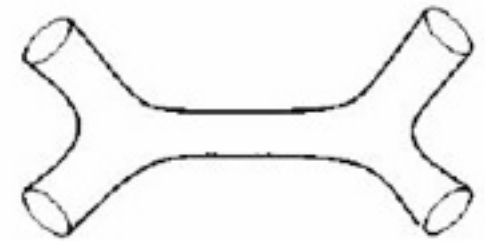
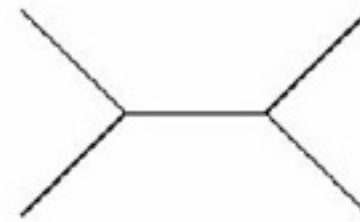


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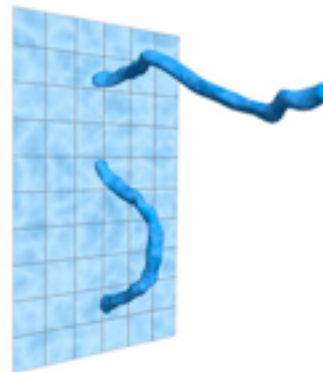
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Gauge fields and gauge interactions

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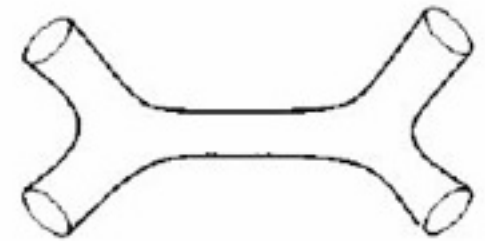
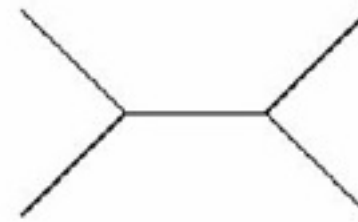


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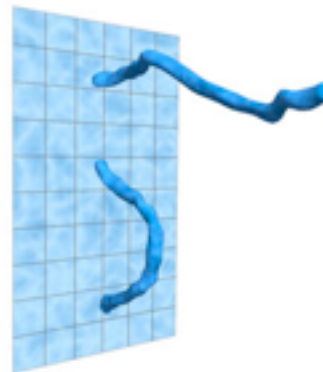
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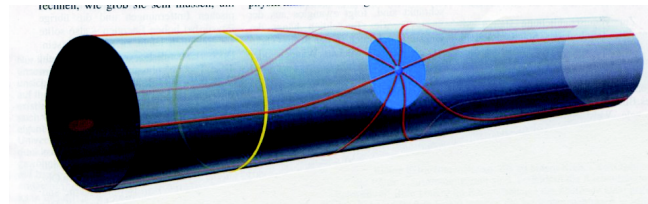
- Strongly coupled string theory \Rightarrow M-theory in $D=11$

(E.Witten, 1995)

Why do we observe only 4 space-time dimensions at low energies?

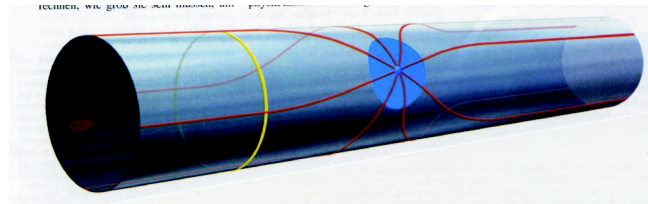
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Calabi-Yau space

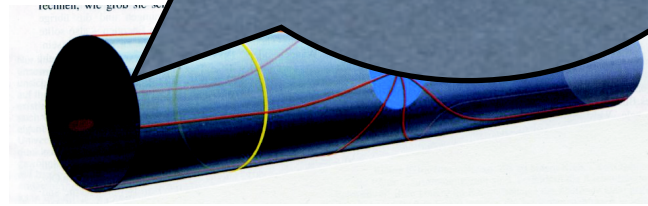
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Compactification: Changing the geometry of the extra dimensions

New scale R^\perp :
size of extra dimensions



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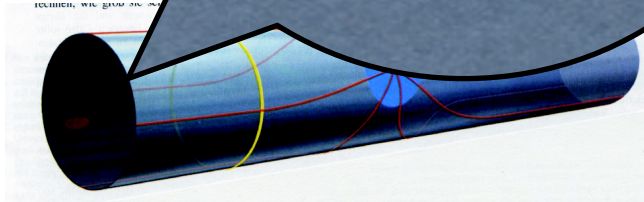
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Universal relation:

$$M_{\text{Planck}}^2 \simeq M_{\text{string}}^{2+d^\perp} (R^\perp)^{d^\perp}$$

In 1985/1986 many people believed that there are very few possible compactifications to four space-time dimensions.

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How many Calabi-Yau spaces exist?

How many string compactifications to four dimensions arise as consistent string solutions?

CHIRAL FOUR-DIMENSIONAL HETEROTIC STRINGS FROM SELF-DUAL LATTICES

W. LERCHE

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Received 24 November 1986

It is shown how our previous work on lattice constructions of ten-dimensional heterotic strings can be applied to four dimensions. The construction is based on an extension of Narain's lattices by including the bosonized world-sheet fermions and ghosts, and uses conformal field theory as its starting point. A natural embedding of all these theories in the bosonic string is automatically provided. Large numbers of chiral string theories with and without $N = 1$ supersymmetry can be constructed. Many features of their spectra have a simple interpretation in terms of properties of even self-dual lattices. In particular we find an intriguing relation between extended supersymmetry and exceptional groups.

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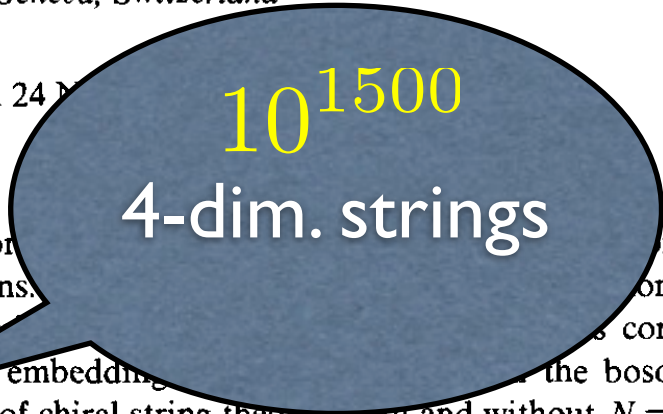
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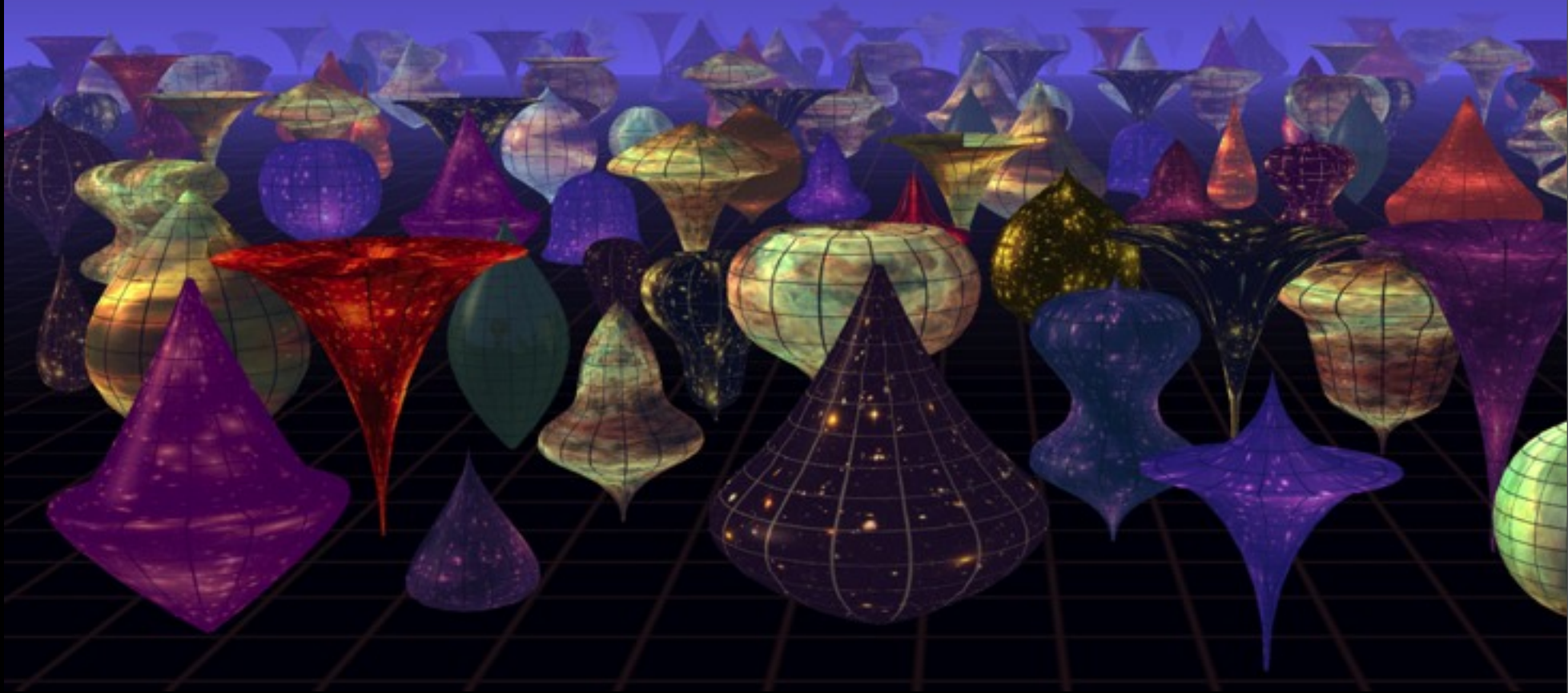
10^{1500}
4-dim. strings

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copy right: Kristin Riebe

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Many (10^{1500} ?) compactified solutions!

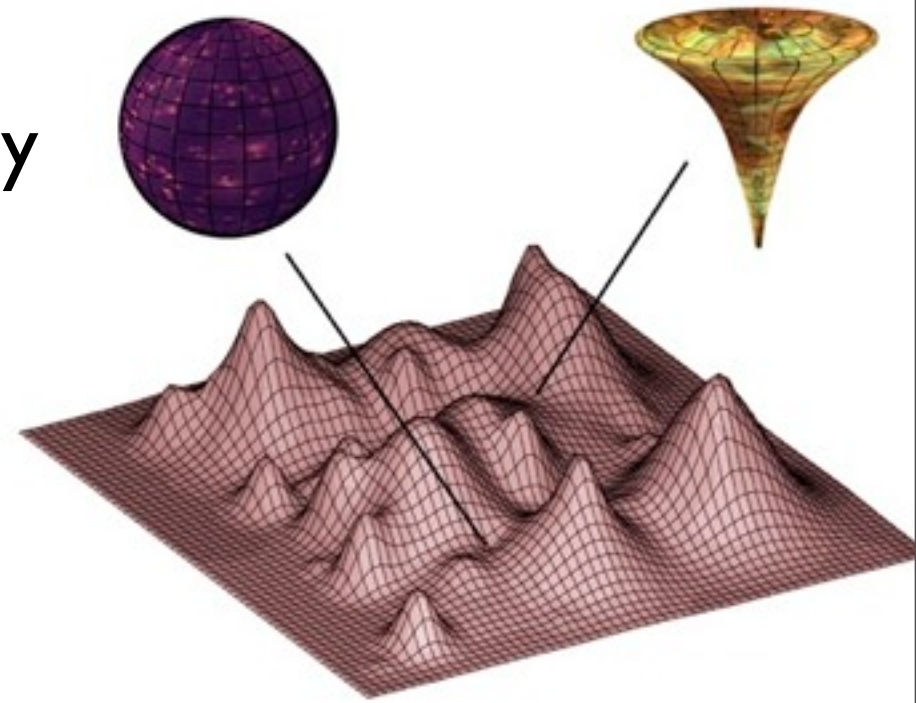
(Lerche, D.L., Schellekens, 1986; Bousso, Polchinski, 2000; Douglas, 2003;
Kachru, Kallosh, Linde, Trivedi)

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Landscape \Rightarrow Multiverse picture:

(A. Linde, L. Susskind, ...)

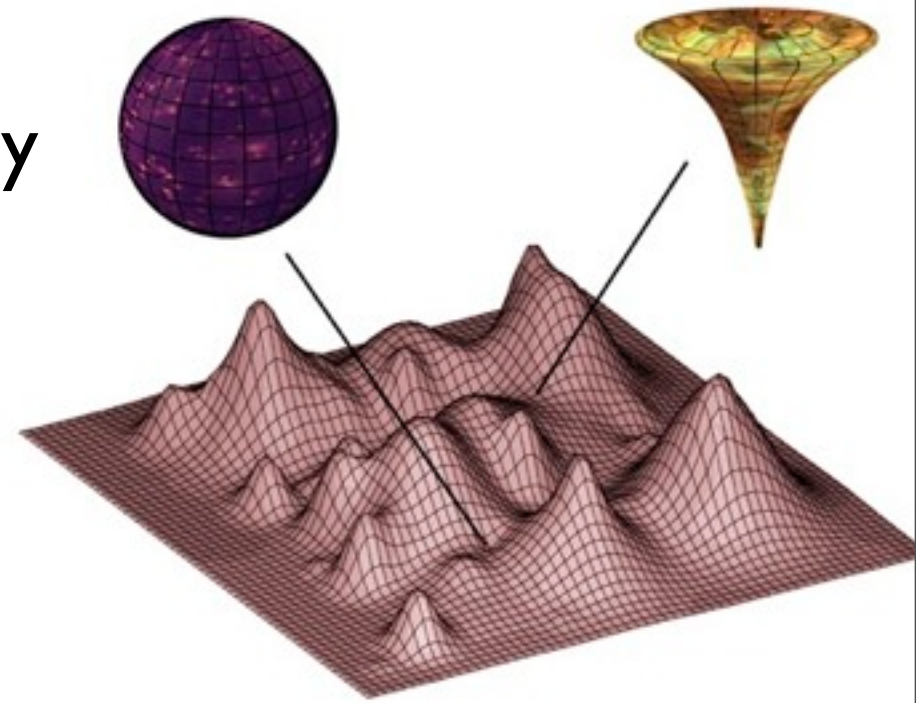
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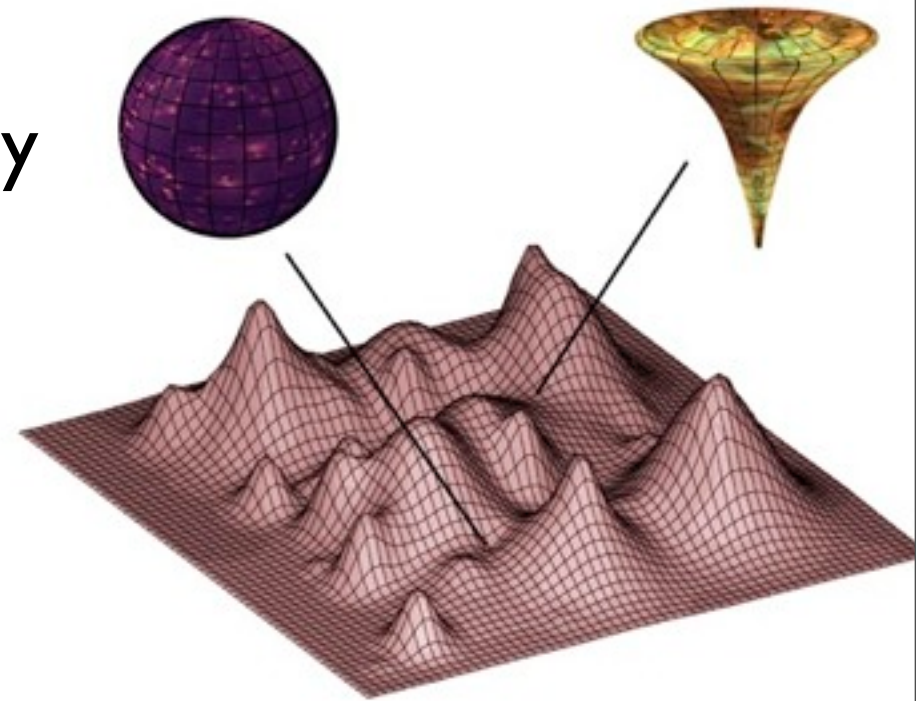


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Murray Gell-Mann:

“If we really live in a multiverse, Physics will have been reduced to an environmental science like Botany.”

(from M. Duff, 1987)



How to handle the huge string landscape?



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- F-theory GUTS, heterotic models, D-brane models
- **Supersymmetry breaking is still a big problem.**

- Explore all mathematically consistent possibilities.
 - ⇒ **Top down approach:**

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String statistics:

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D-brane models:

In total $3.4 \cdot 10^{28}$ susy D-brane models.

$5.7 \cdot 10^6$ of them possess MSSM like spectra!

Only one in 10^{-22} models gives rise
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New strategies to handle big data: Machine and deep learning

(Y. He, 2017)

- Look for generic features in the landscape in some classes of models:

Low string scale and large extra dimensions:

(N. Arkani-Hamed, S. Dimopoulos, G. Dvali, 1998)

- Gauge and matter fields are living on lower dimensional branes:



- Gravity lives in an higher-dimensional bulk

$$M_{\text{Planck}}^2 \simeq M_{\text{string}}^{2+d^\perp} (R^\perp)^{d^\perp}$$

If R^\perp is very large, M_{string} can be very low.

Generic feature of low string/gravity scale models:

Production of universal string Regge excitations at the LHC.

String Hunters' Companion: Model independent calculation
of production cross sections:

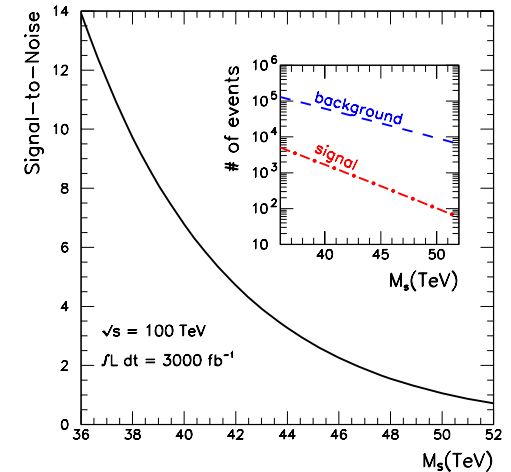
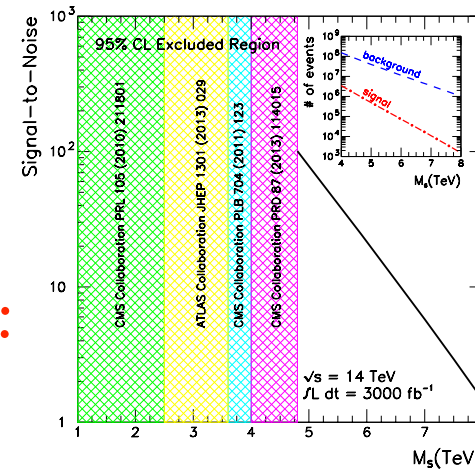
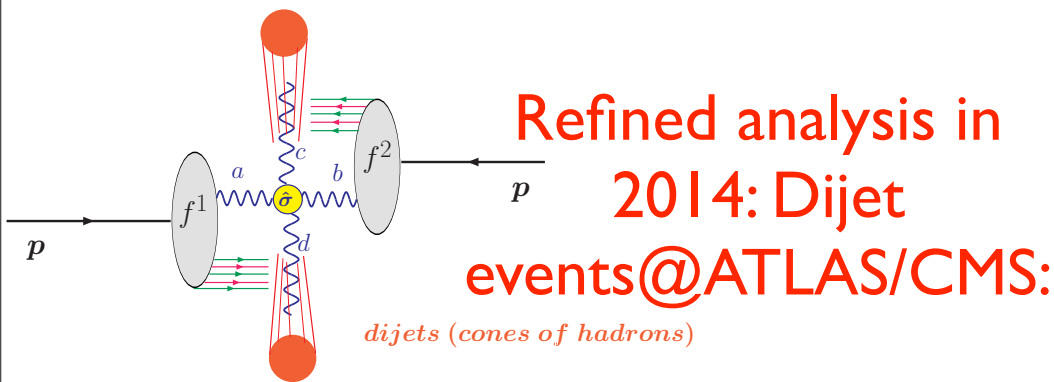
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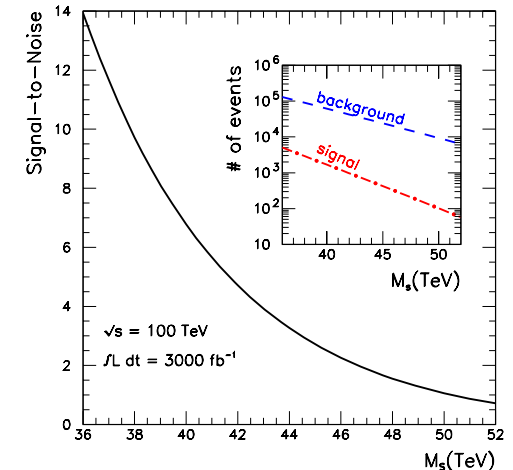
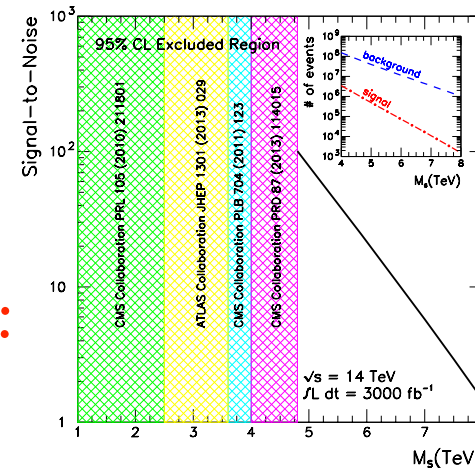
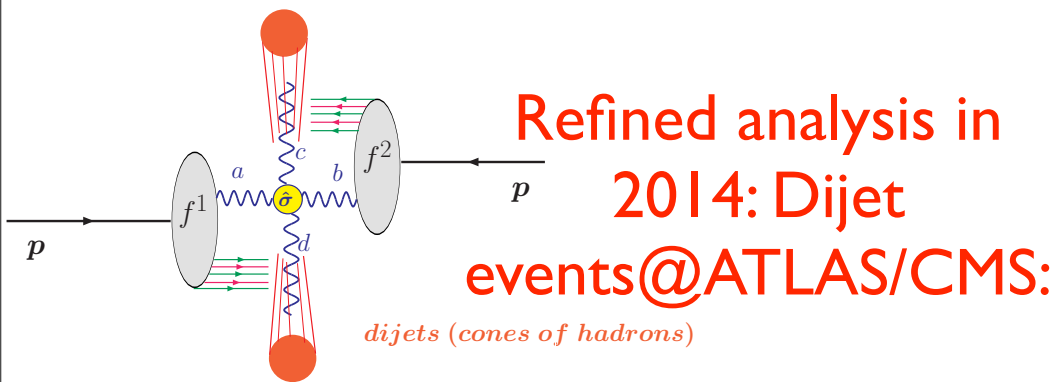


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LHC(13) (ATLAS/CMS): $M_{\text{string}} \geq 7 \text{ TeV}$

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- Structure of perturbative gravitational (and also gauge theory) scattering amplitudes
- Holography and the AdS/CFT correspondence
- Quantum properties of black holes
- The swamp land approach and the weak gravity conjecture

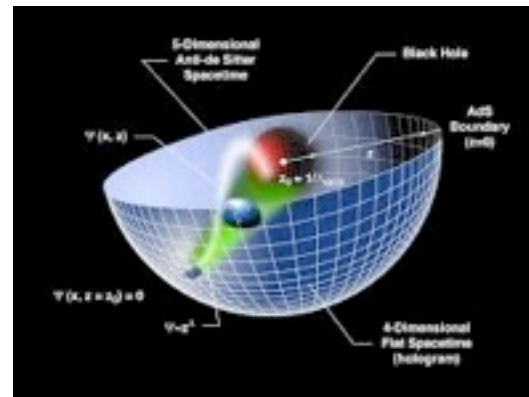
(i) Holography and the AdS/CFT correspondence:

(J. Maldacena, 1997; E. Witten, 1998)

(Quantum)
gravity in the
 $D+1$ bulk
theory

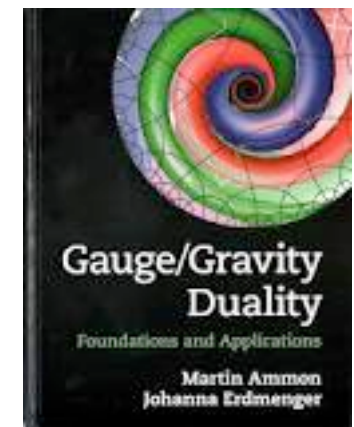
\Leftarrow Duality \Rightarrow

Quantum field theory
on the D -dimensional
holographic boundary
theory



Many applications for condensed matter physics: AdS/CMT:

Compute correlation functions in strongly coupled systems via dual quantities in classical gravity theory.



Reverse strategy:

Some new important insights for quantum gravity from the dual field theory side.

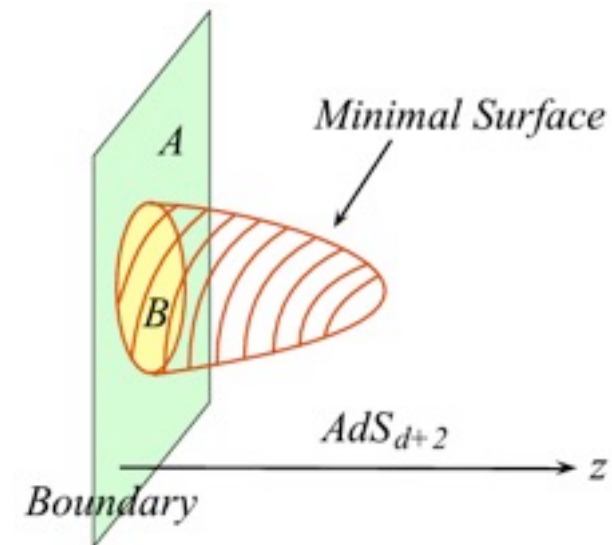
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- Holographic entanglement entropy (S. Ryu, T. Takayanagi, 2006)

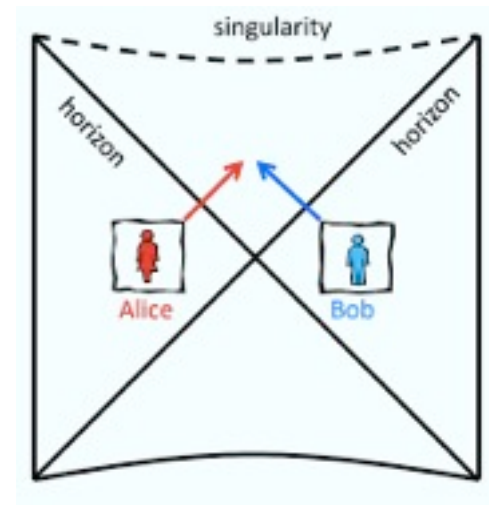
von Neumann entropy

$$S(\rho_A) = \frac{\text{Area}(\gamma_A)}{4G_N}$$



- Gravity and quantum information - emergent geometry

Bulk space-time can be described by entangled states in boundary quantum mechanics



Worm hole is an entangled state:

(J. Maldacena, L. Susskind, 2013)

$$|\Psi\rangle = \sum_n e^{-\beta E_n/2} |\bar{E}_n\rangle_L \times |E_n\rangle_R$$

(thermal field double)

(ii) Quantum properties of black holes

One of the most celebrated results of string theory as quantum gravity theory:

Microscopic derivation of black hole entropy (in 5D):

(A. Strominger, C.Vafa, 1996)

Black holes can be viewed as bound states of $3 \times N$ D-branes.



$$S_{BH} = \frac{\text{Area}}{4} = 2\pi \sqrt{\frac{N^3}{2}}$$

Entropy is due to open string degeneracy.

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- **Asymptotic BMS symmetries of classical black hole metric.**

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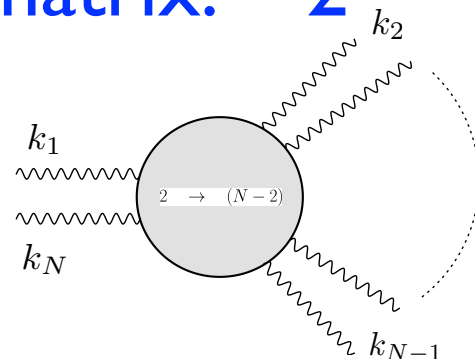
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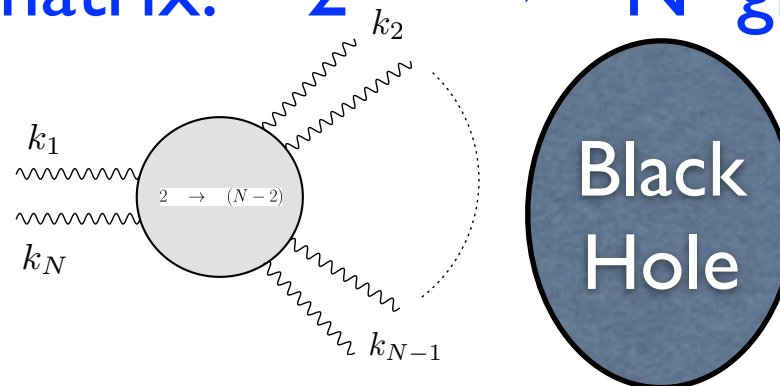
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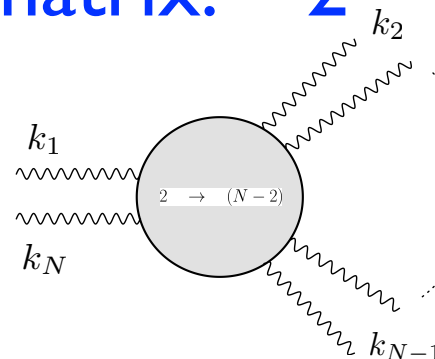
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Strings
at very high
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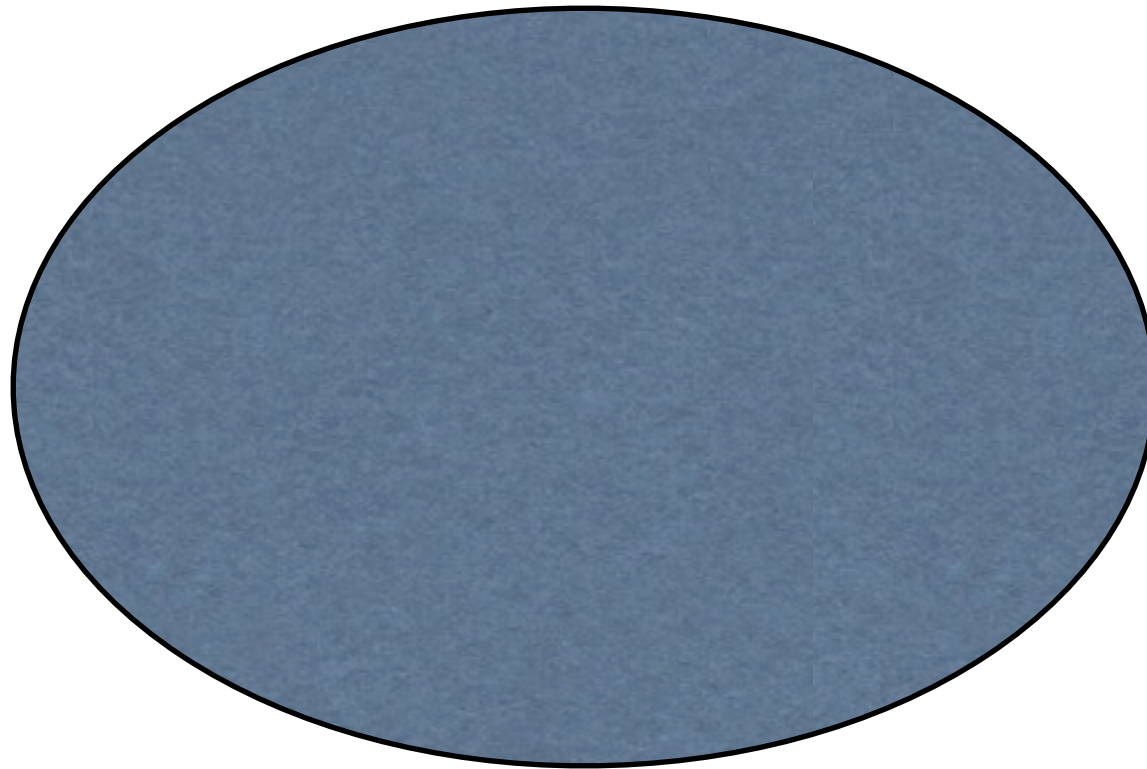
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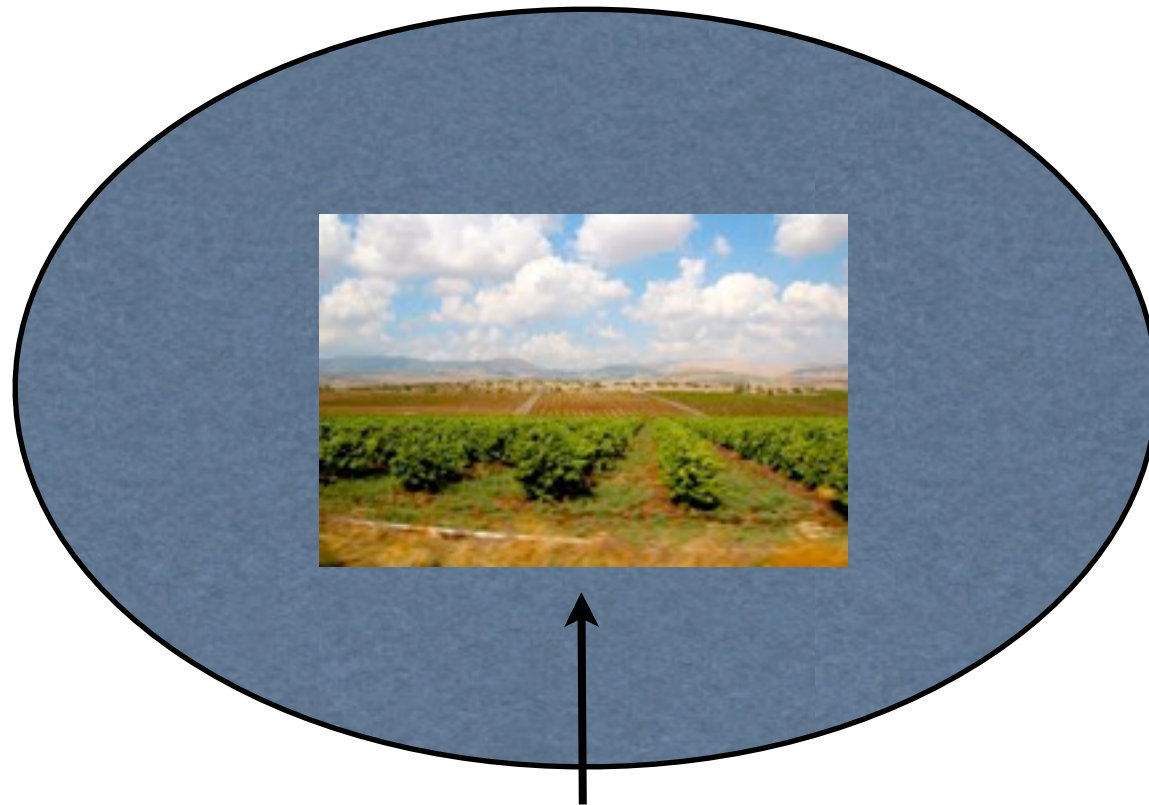
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Fruitful quantum field theories that can be consistently embedded in quantum gravity/string theory in the UV.

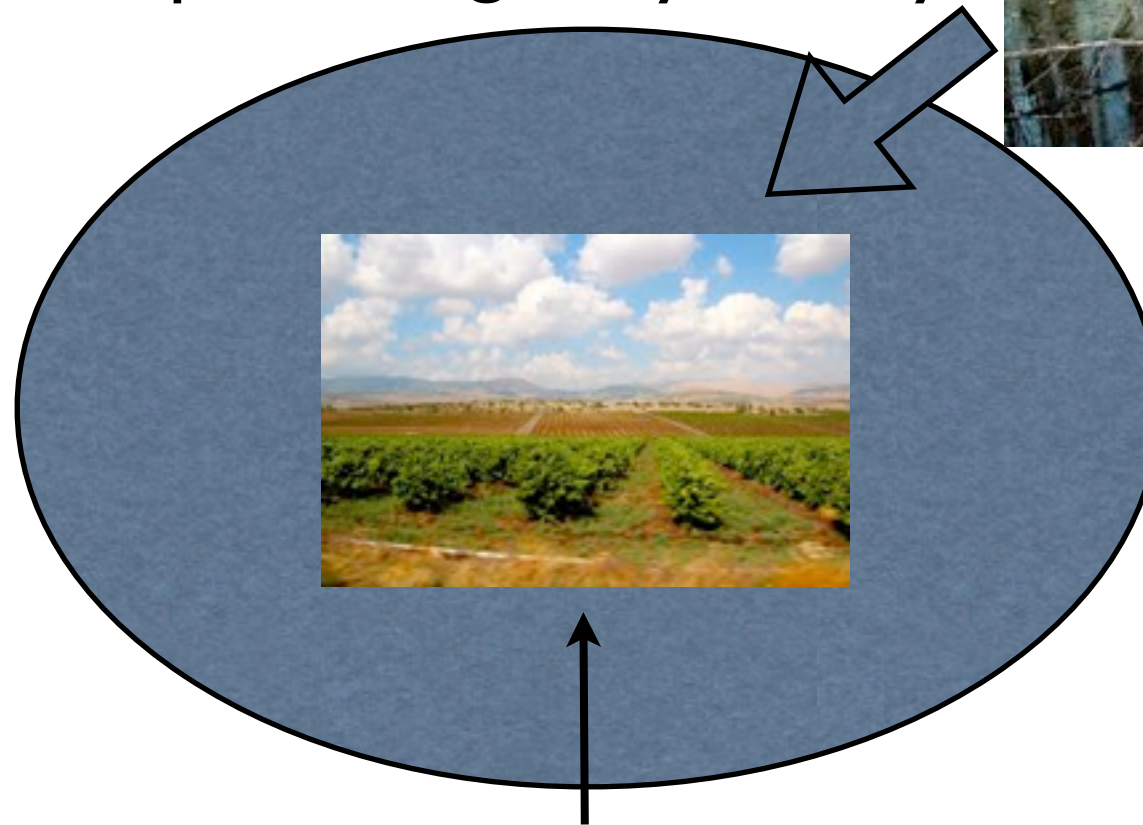
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Swampland of theories that are inconsistent as quantum gravity theory



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Axion WGC \Rightarrow Large field inflation: $\Phi_{\text{inflaton}} \leq M_{\text{Planck}}$

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What is happening if one crosses the UV cut-off scale ?

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New feature of quantum gravity: UV/IR mixing.

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- String production (massive strings are very large)
- T-duality (a small compact space is equivalent to a large compact space)
- Duality invariant generalization of Einstein gravity
⇒ Double Field Theory, Exceptional Field Theory

(D. Berman, H. Godazgar, O. Hohm, C. Hull, E. Malek, H. Nicolai, M. Perry, F. Rudolph, H. Samtleben, B. Zwiebach, ...)

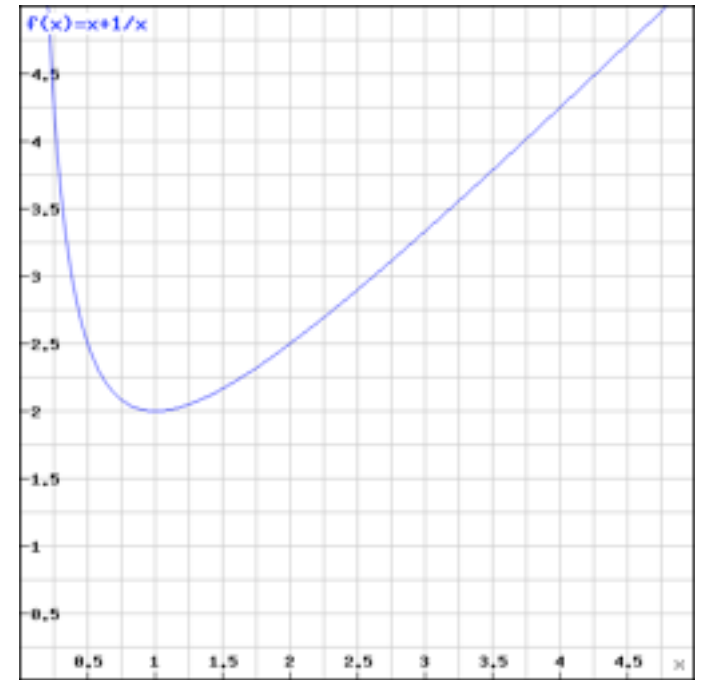
String scattering at super-string scale energies:

Refinement of Heisenberg relation: $\Delta x \geq \frac{\hbar}{\Delta p} + \alpha' \Delta p$

(A. Amati, M. Ciafaloni, G. Veneziano, 1987)

⇒ Smallest possible distance:

$$\Rightarrow \Delta x \geq \Delta x_{min} \simeq L_{string}$$



Δp

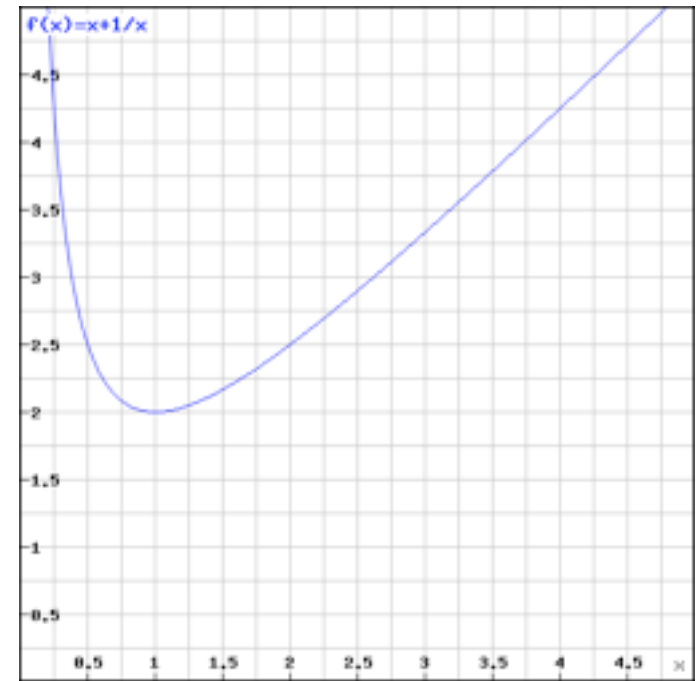
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Mathematically this can be described by non-commutative and non-associative geometry.

(D.L., A. Blumenhagen, E. Plauschinn, 2010)

Volume uncertainty: $\Delta V = \Delta X^1 \Delta X^2 \Delta X^3 \geq L_{string}^3$

(D. Mylonas, P. Schupp, R. Szabo, 2013; Chamseddine, Connes, Mukhanov, 2014)

Another problem in (quantum) gravity:

Hierarchy problem:

Why is the scale of gravity apparently so much higher that the mass scale of the Standard Model?

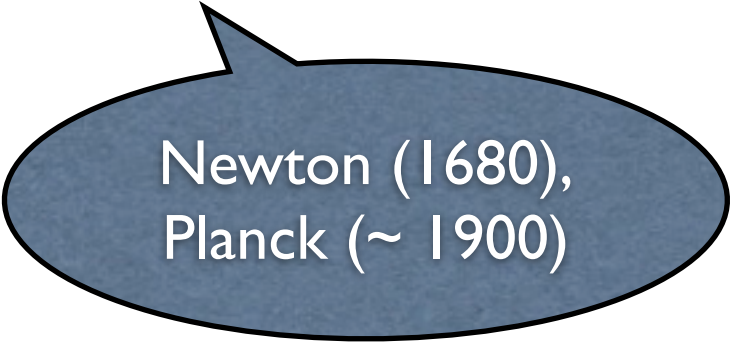
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Newton (1680),
Planck (~ 1900)



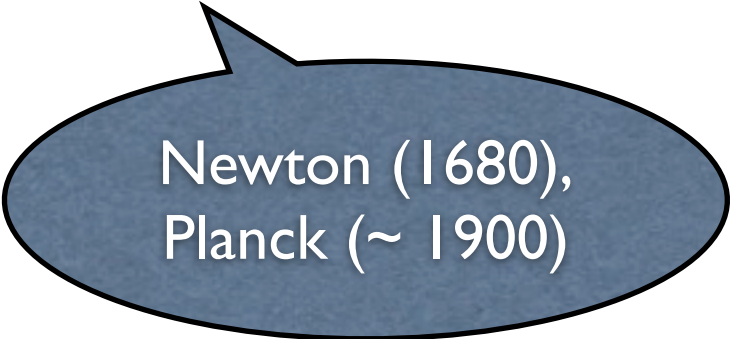
ATLAS, CMS (2012)

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Cosmological constant problem:

$$\Lambda_{\text{cosm}} \simeq 10^{-120} M_{\text{Planck}}^4$$

„Standard“ Wilsonian solutions to the hierarchy problem:

Lowering the UV cut-off: New physics at the weak scale.

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- String theory with a low fundamental scale of gravity:

$$M_{\text{string}} \simeq 1 - 10 \text{ TeV}$$

String resonances (excited quarks) at a few TeV

Possible new solution of the hierarchy problem based on quantum gravity and UV / IR mixing:

The **IR weak scale at low energies** is a consequence of **UV gravity at high energies**.

No new physics/particles will emerge at the weak scale.

Concrete proposal: **Refined weak gravity conjecture:**

(E. Palti, 2017)

Consider gravity plus U(1) plus a neutral scalar field ϕ , which is coupled to a charged particle h :

$$\mathcal{L} = \frac{M_p^2}{2} R - \frac{1}{4g^2} F^2 - |Dh|^2 - (\partial\phi)^2 - m^2 h^* h - m_\phi^2 \phi^2 - 2m\mu\phi h^* h + \dots$$

gauge coupling g scalar field coupling μ

UV cut off scale of the theory is still given as:

$$\Lambda_{UV} \leq g M_{\text{Planck}}$$

But UV quantum gravity now implies an IR mass scale :

(D.L., E. Palti, 2017)

$$m_h \leq \sqrt{g^2 - \mu^2} M_{\text{Planck}} \equiv \beta M_{\text{Planck}}$$

This particle could be e.g. the Higgs or a light fermion and its mass can be arbitrarily far from the UV cut-off scale.

Summary:

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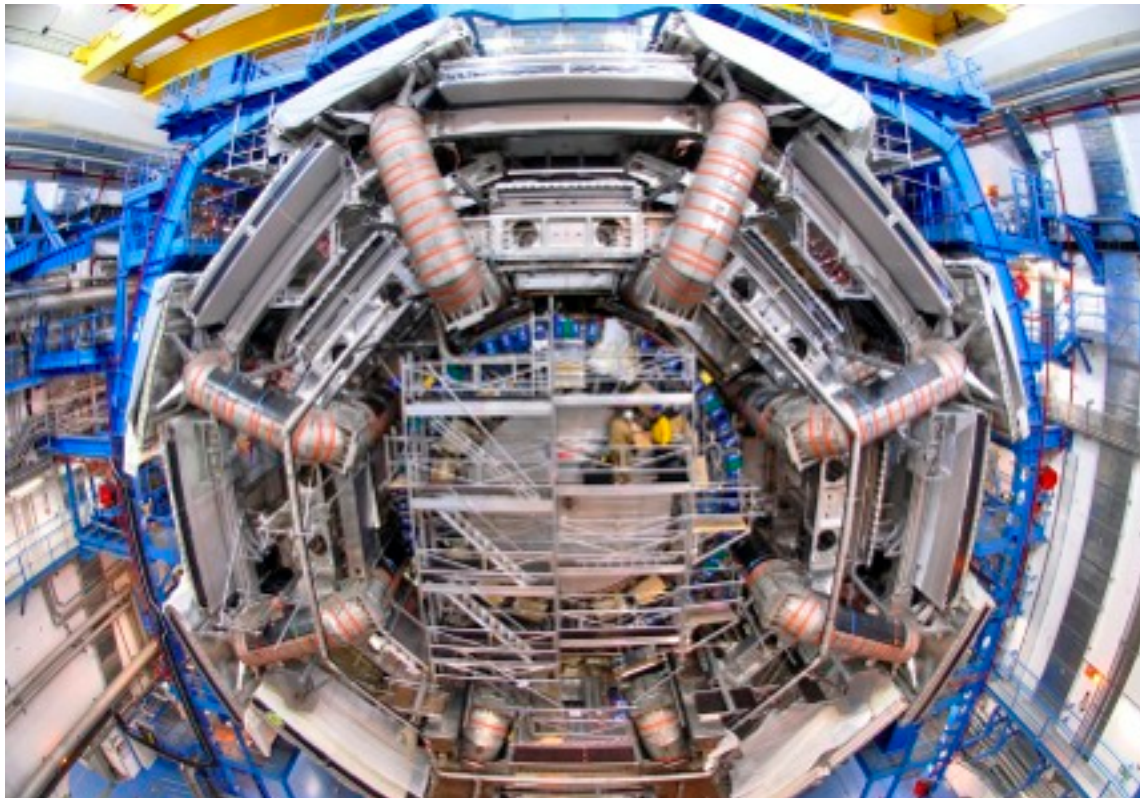
Perhaps some of these ideas can be tested in experiment:

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Particle physics:

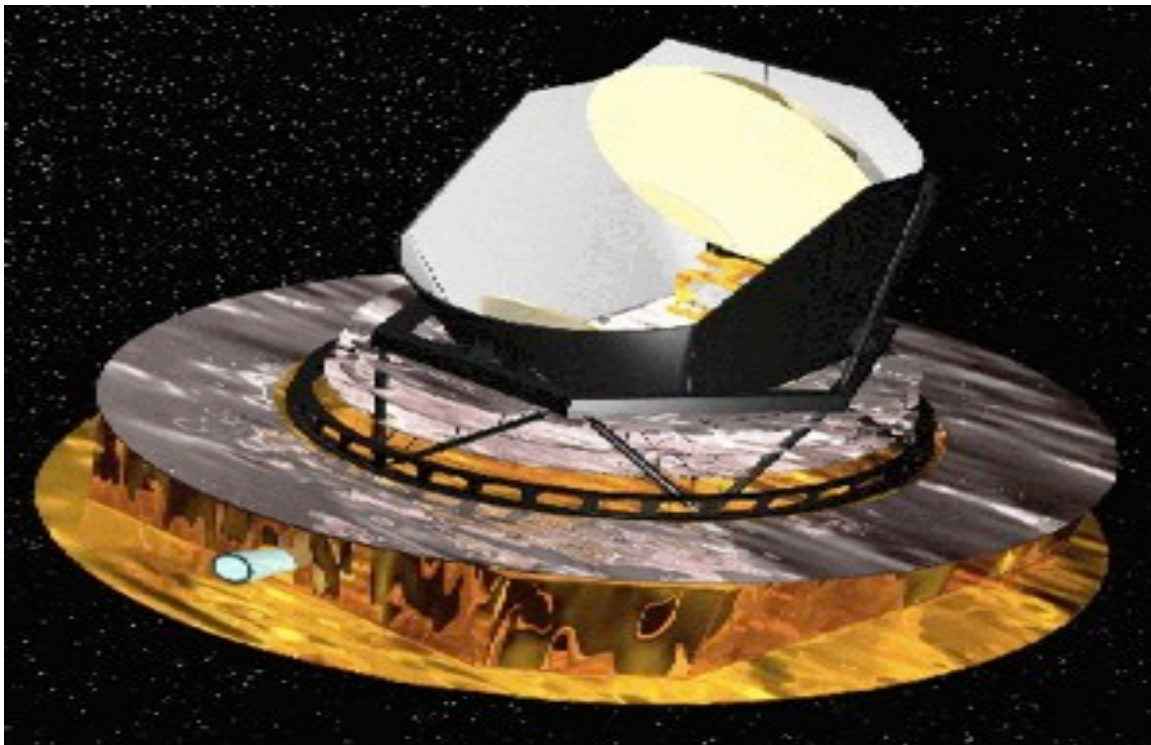


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Astrophysics/cosmology

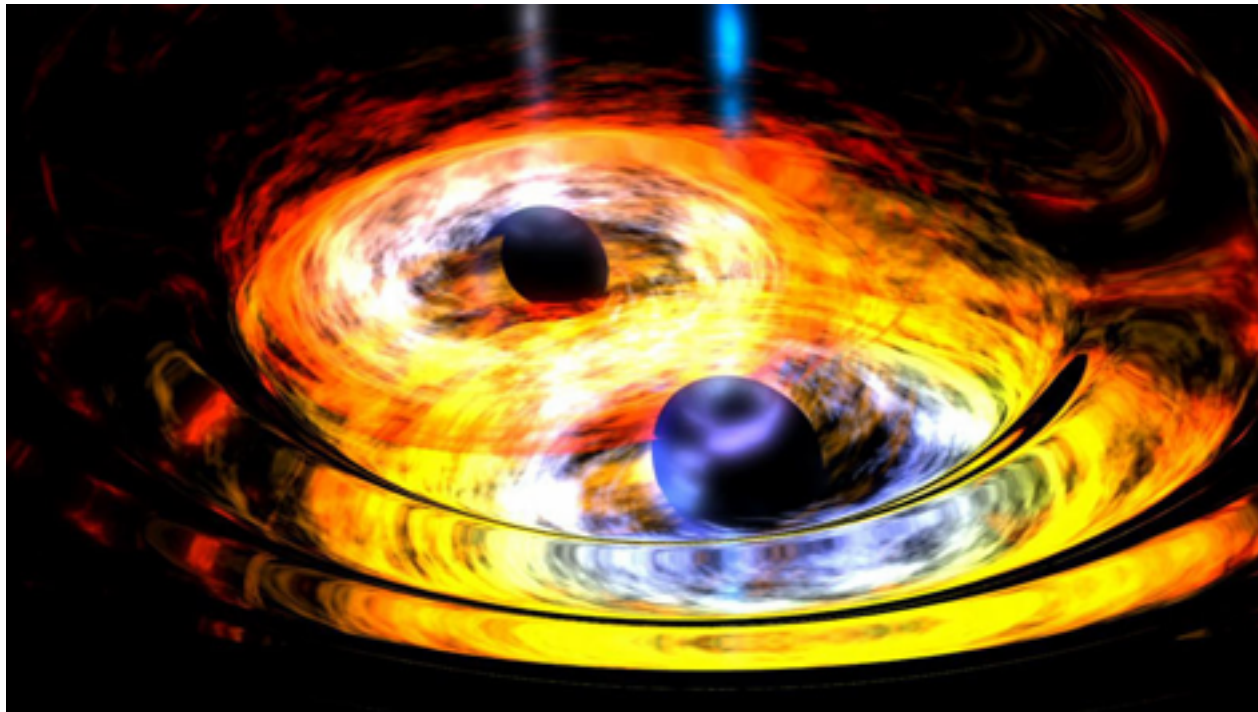


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Gravitational waves/black holes

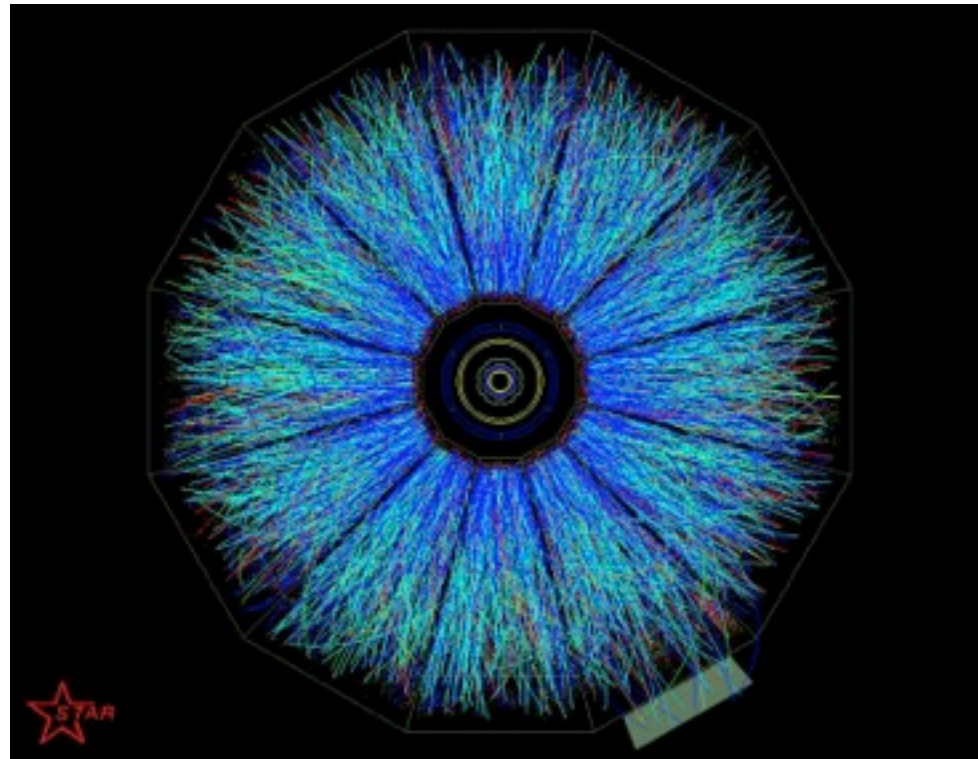


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Strongly coupled system (QCD, condensed matter)



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Perhaps some of these ideas can be tested in experiment:

Thank you!