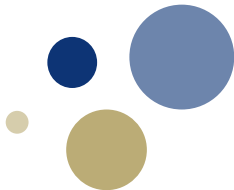




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Baryogenesis and the Two-Higgs Doublet Model

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Why do you exist?



We observe a matter dominated universe [1].

Unsolved mysteries

- Why does the universe contain more matter than antimatter?
- Why does the asymmetry have the specific amount that it does ($\eta \approx 10^{-9}$)?

Conditions for baryogenesis



Sakharov conditions [2]

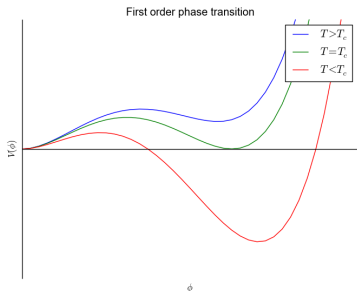
1. Baryon number is violated
2. Deviation from thermal equilibrium
3. CP violation

Phase transition in the universe

First or second order phase transition?

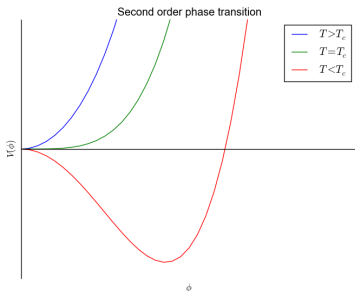
First order phase transition have a barrier:

- Deviation from thermal equilibrium



Second order phase transition does not have a barrier:

- Can happen continuously

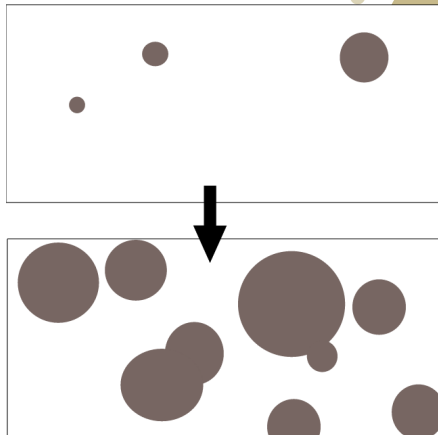


Bubble nucleation

Bubbles with a different vacuum nucleates.

Bubbles of a certain size will:

1. Expand
2. Collide
3. Coalesce and fill the whole universe.

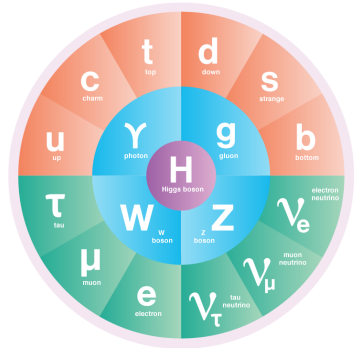


The Standard Model

$$\begin{aligned}
 \mathcal{L}_{SM} = & - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\
 & + i\bar{\psi} \not{D} \psi + \text{h.c.} \\
 & + \psi_i Y_{ij} \psi_j \phi + \text{h.c.} \\
 & + |D_\mu \phi|^2 - V(\phi)
 \end{aligned}$$

where

$$V(\phi) = -\frac{1}{2} m^2 \phi^2 + \frac{\lambda}{4} \phi^4$$



Is the Standard Model sufficient to explain baryogenesis?



1. Ok
2. Not quite
3. Not quite

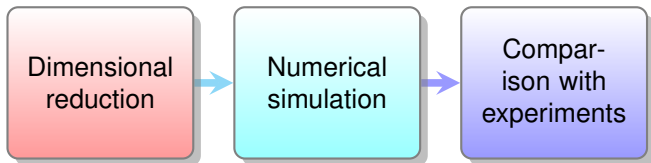
We need an extended model.

Examples: SM + singlet [3], Minimal Supersymmetric SM [4].

Two-Higgs Doublet Model

We add an extra Higgs doublet.

$$\begin{aligned} V(\phi_1, \phi_2) = & -\frac{1}{2}m_{11}^2\phi_1^\dagger\phi_1 - \frac{1}{2}m_{22}^2\phi_2^\dagger\phi_2 \\ & + \lambda_1(\phi_1^\dagger\phi_1)^2 + \lambda_2(\phi_2^\dagger\phi_2)^2 \\ & + \lambda_3(\phi_1^\dagger\phi_1)(\phi_2^\dagger\phi_2) + \lambda_4(\phi_1^\dagger\phi_2)(\phi_2^\dagger\phi_1) \\ & + \frac{\lambda_5}{2}\left[(\phi_1^\dagger\phi_2)^2 + (\phi_2^\dagger\phi_1)^2\right] \end{aligned}$$



Thermal field theory

We replace the integral over 4-momentum with a sum-integral.

$$\int \frac{d^4 k}{(2\pi)^4} \rightarrow T \sum_{n=-\infty}^{\infty} \int \frac{d^3 k}{(2\pi)^3} \equiv \int_K$$

where $K = (\omega_n, \mathbf{k})$ is the 4-momentum.

Lorentz symmetry is broken by the heat bath!

Particles acquire a thermal (Debye) mass.

Bosons

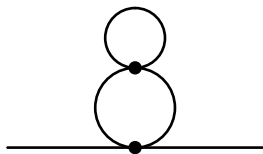
$$\omega_n = 2n\pi T, \quad n \in \mathbb{Z}$$

Fermions

$$\omega_n = (2n + 1)\pi T, \quad n \in \mathbb{Z}$$

Infrared divergences

A quantum correction to the scalar self-energy of order $\mathcal{O}(\lambda^2)$ is


$$= -\frac{\lambda^2}{4} \int_P \frac{1}{P^2} T \sum_n \int \frac{d^d k}{(2\pi)^d} \frac{1}{[\omega_n^2 + k^2 + m^2]^2}$$

where $d = 3 - 2\epsilon$.

For $m^2 = 0$ og $n = 0$ the diagram diverges when $k^2 \rightarrow 0$.

We must sum all diagrams of this form (daisy-diagrams), and they yield a finite result.

β -functions for Two-Higgs Doublet Model

The β -functions can be extracted from the counter terms.

$$\begin{aligned}\beta_{g'} &= \frac{g'^3}{6(4\pi)^2} \left[2 + \sum_f Y_f^2 \right], & \beta_g &= -\frac{g^3}{3(4\pi)^2} \left[21 - \sum_{\text{left}} \right] \\ \beta_{\lambda_1} &= \frac{1}{(4\pi)^2} \left[\frac{9}{4}g^4 + \frac{3}{4}g'^4 + \frac{3}{2}g^2g'^2 - \lambda_1(9g^2 + 3g'^2) + 12\lambda_1^2 + 4\lambda_2^2 \right. \\ &\quad \left. + 2\lambda_4^2 + 2\lambda_5^2 + 4\lambda_3\lambda_4 + 4\lambda_1 \sum_i h_i^2 \right] \\ \beta_{\lambda_2} &= \frac{1}{(4\pi)^2} \left[\frac{9}{4}g^4 + \frac{3}{4}g'^4 + \frac{3}{2}g^2g'^2 - \lambda_2(9g^2 + 3g'^2) + 12\lambda_2^2 + 4\lambda_3^2 + 2\lambda_4^2 \right. \\ &\quad \left. + 2\lambda_5^2 + 4\lambda_3\lambda_4 + 4\lambda_2 \sum_i h_i^2 - 4 \sum_i h_i^4 \right] \\ \beta_{\lambda_3} &= \frac{1}{(4\pi)^2} \left[\frac{9}{4}g^4 + \frac{3}{4}g'^4 + \frac{3}{2}g^2g'^2 - \lambda_3(9g^2 + 3g'^2) + 2(\lambda_1 + \lambda_2)(3\lambda_3 + \lambda_4) \right. \\ &\quad \left. + 4\lambda_3^2 + 2\lambda_4^2 + 2\lambda_5^2 + 4\lambda_3 \sum_i h_i^2 \right] \\ \beta_{\lambda_4} &= \frac{1}{(4\pi)^2} \left[\lambda_4(2\lambda_1 + 2\lambda_2 + 4\lambda_4 + 8\lambda_3 - 9g^2 - 3g'^2) + 8\lambda_5^2 + 4\lambda_4 \sum_i h_i^2 \right] \\ \beta_{\lambda_5} &= \frac{1}{(4\pi)^2} \left[\lambda_5(2\lambda_1 + 2\lambda_2 + 4\lambda_3 + 6\lambda_4 - 9g^2 - 3g'^2) + 4\lambda_5 \sum_i h_i^2 \right]\end{aligned}$$

Dimensional reduction

At high temperatures all fermions and almost all bosons (except the $n = 0$ mode) are (super)heavy [5]. We can integrate out the (super)heavy particles, and are left with an *effective theory*:

$$\mathcal{L}(\mathbf{A}, \phi, \psi) \rightarrow \mathcal{L}_{\text{eff}}(\mathbf{A}_0, \mathbf{A}, \phi) \rightarrow \mathcal{L}'_{\text{eff}}(\mathbf{A}, \phi).$$

We find parameters of the three-dimensional theory by matching, i.e. requiring the long distance (zero external momentum) behaviour of correlators to be the same as in the four-dimensional theory.

Numerical simulations are more tractable [6]: reduced number of dimensions, purely bosonic theory etc.

We can find static (time independent) observables, as the theory is purely spatial.

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