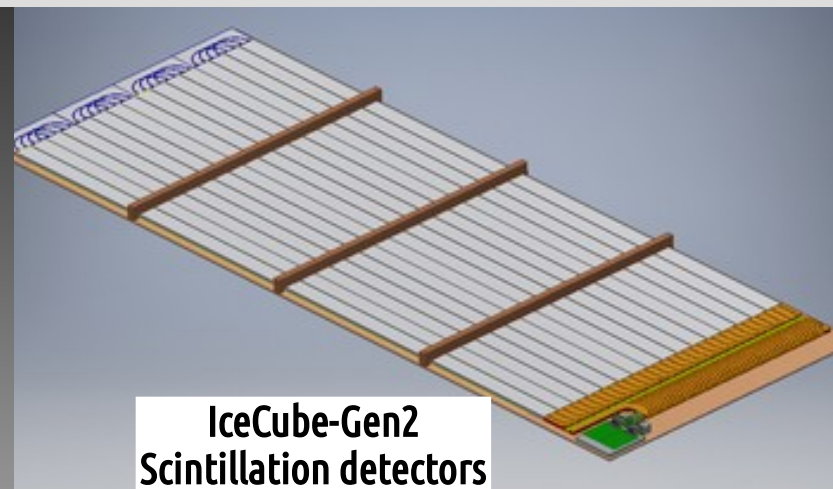
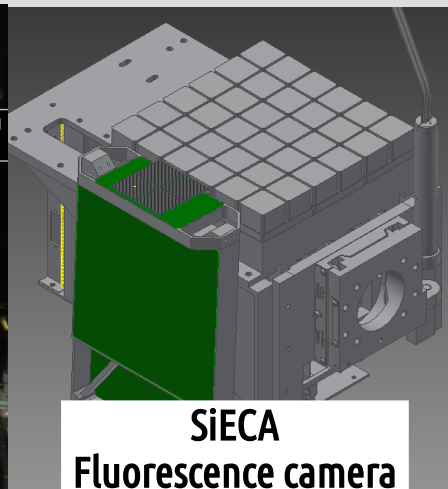
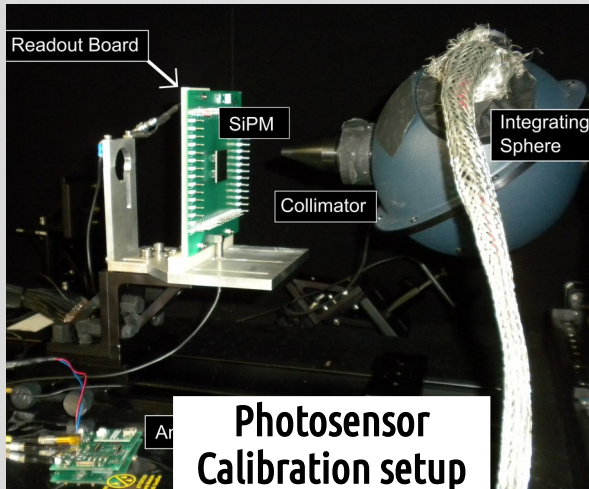


# SiPMs for astroparticle physics applications: Two specific examples

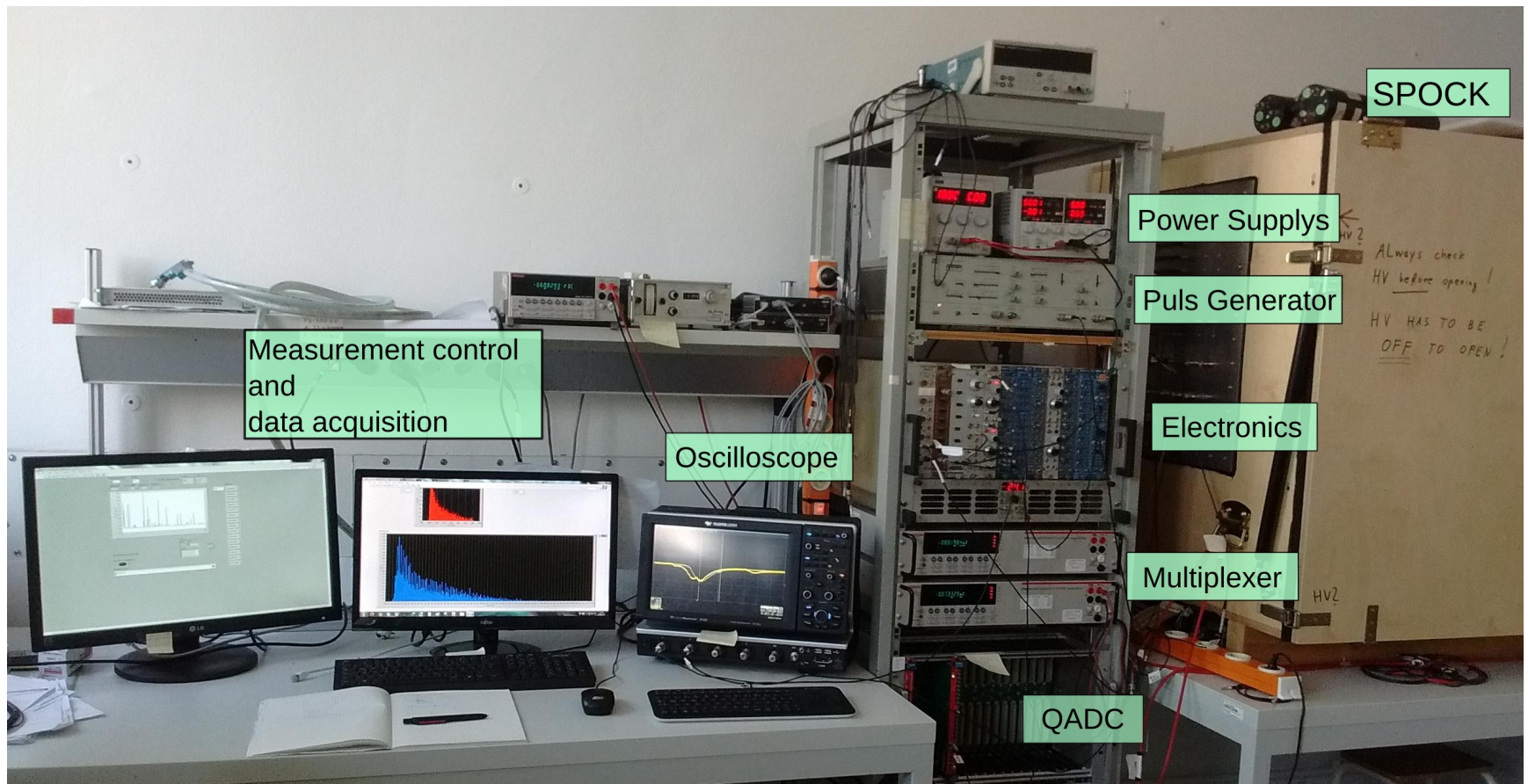
Thomas Huber (DESY / KIT-IKP), on behalf

Max Renschler | Institut für Kernphysik (IKP) | KIT

Institut für Kernphysik (IKP)  
Karlsruhe Institut für Technologie



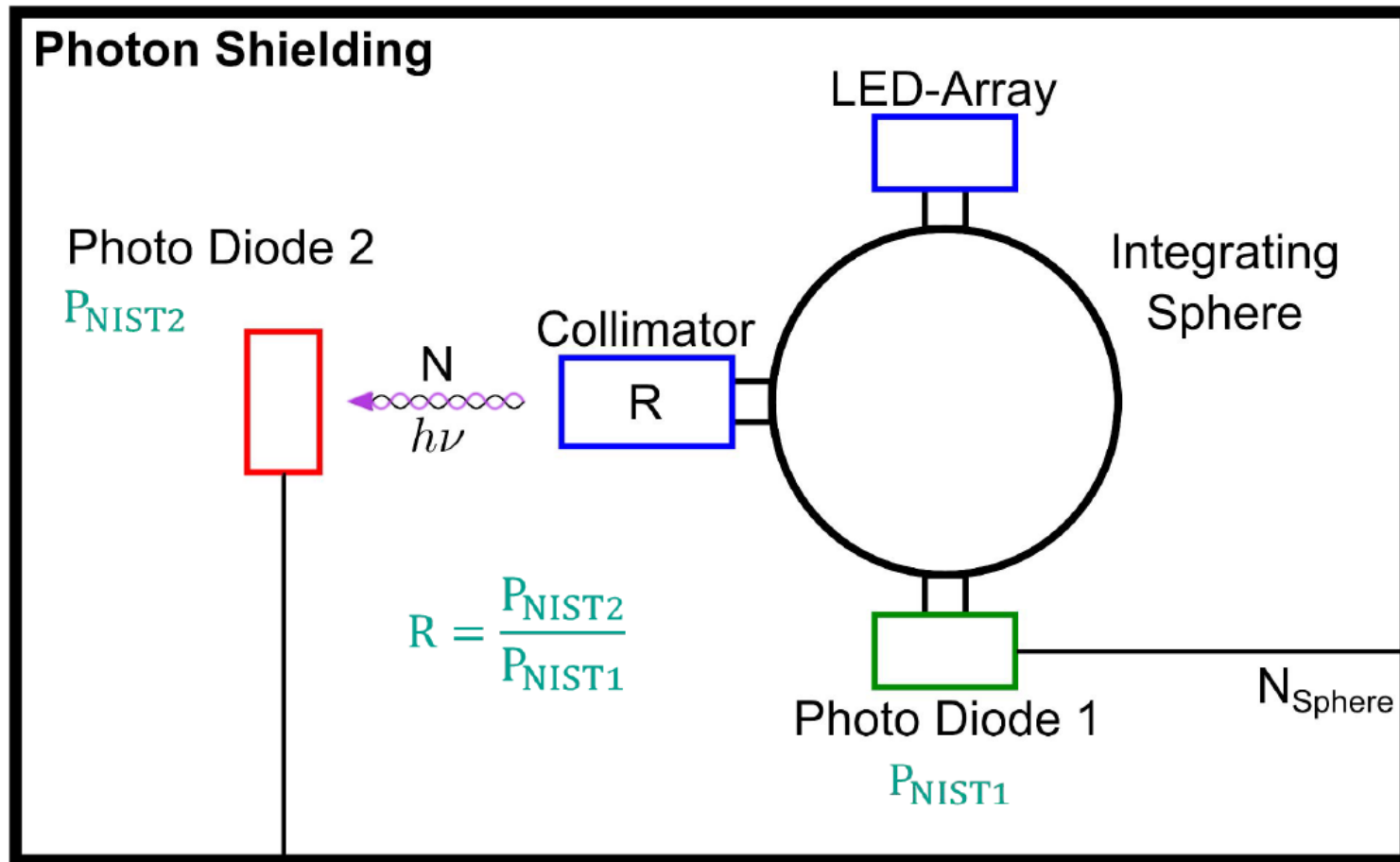
# Characterising measurement setup (SPOCK)



SPOCK: Single PhOton Calibration stand at KIT

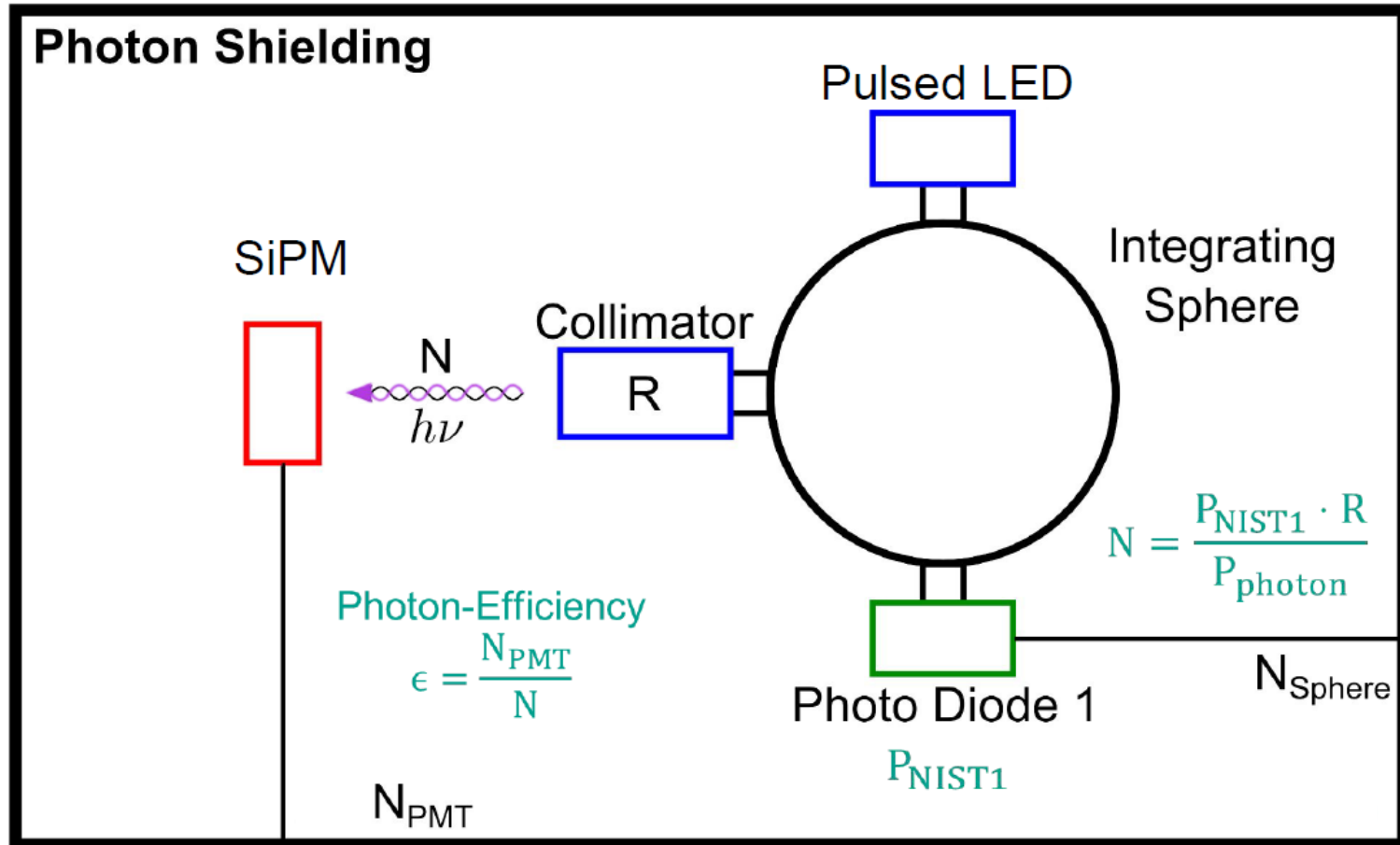
# Simplified sketch of measurement method

1 st step: Calibrating the measurement setup:

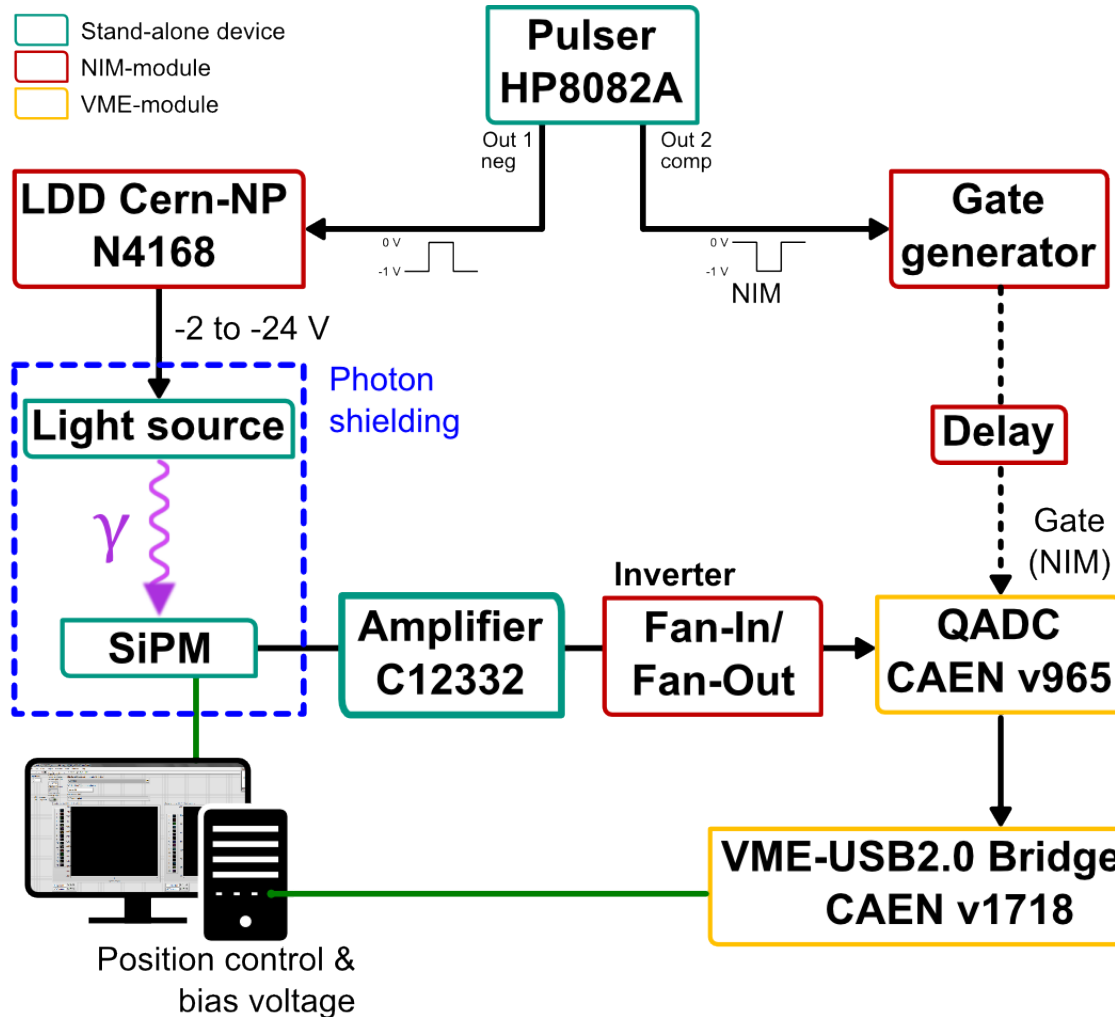


# Simplified sketch of measurement method

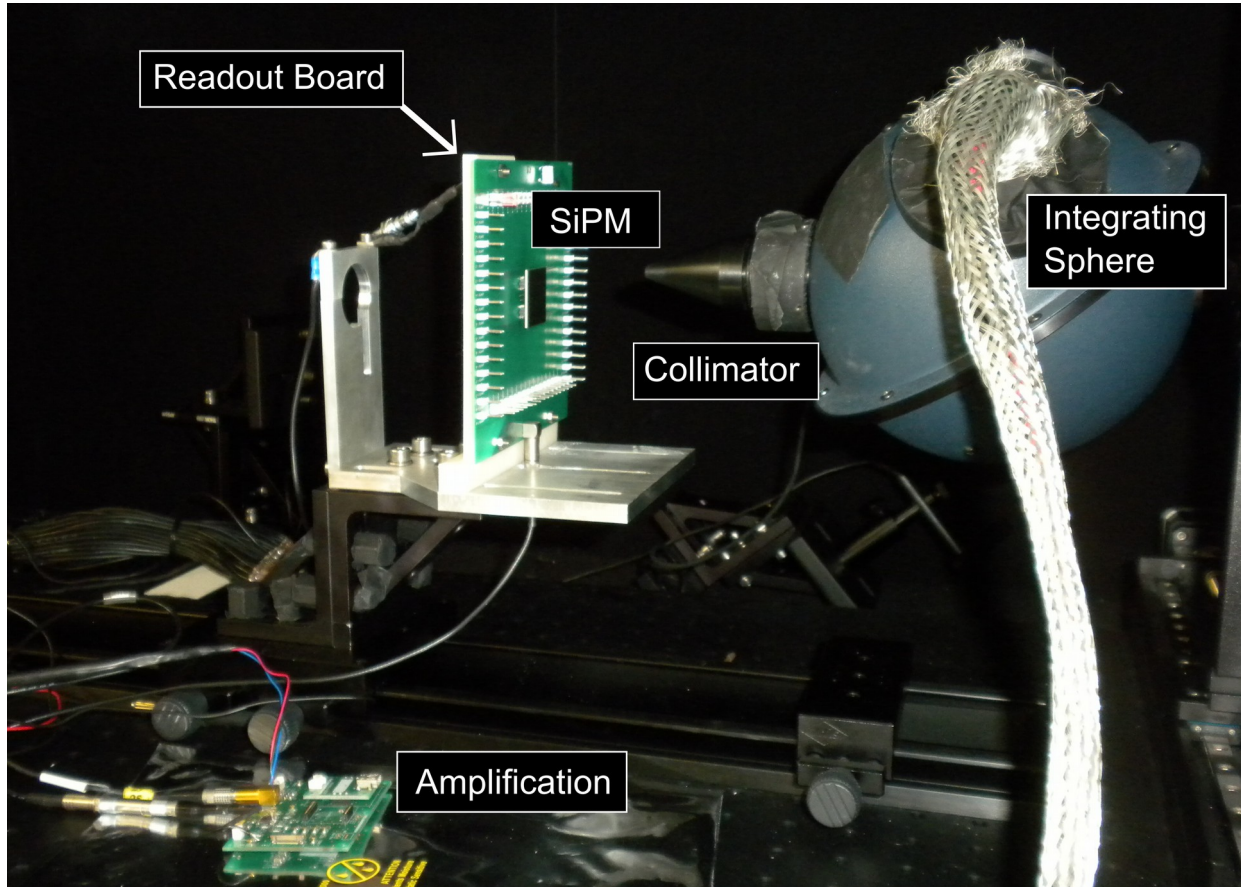
2 nd step: Charaterize the photo sensor



# Characterizing SiPMs



# Inside of SPOCK



Collimator Attenuation

- $\sim 10^6$  ( $\rightarrow$  Single photon pulses)

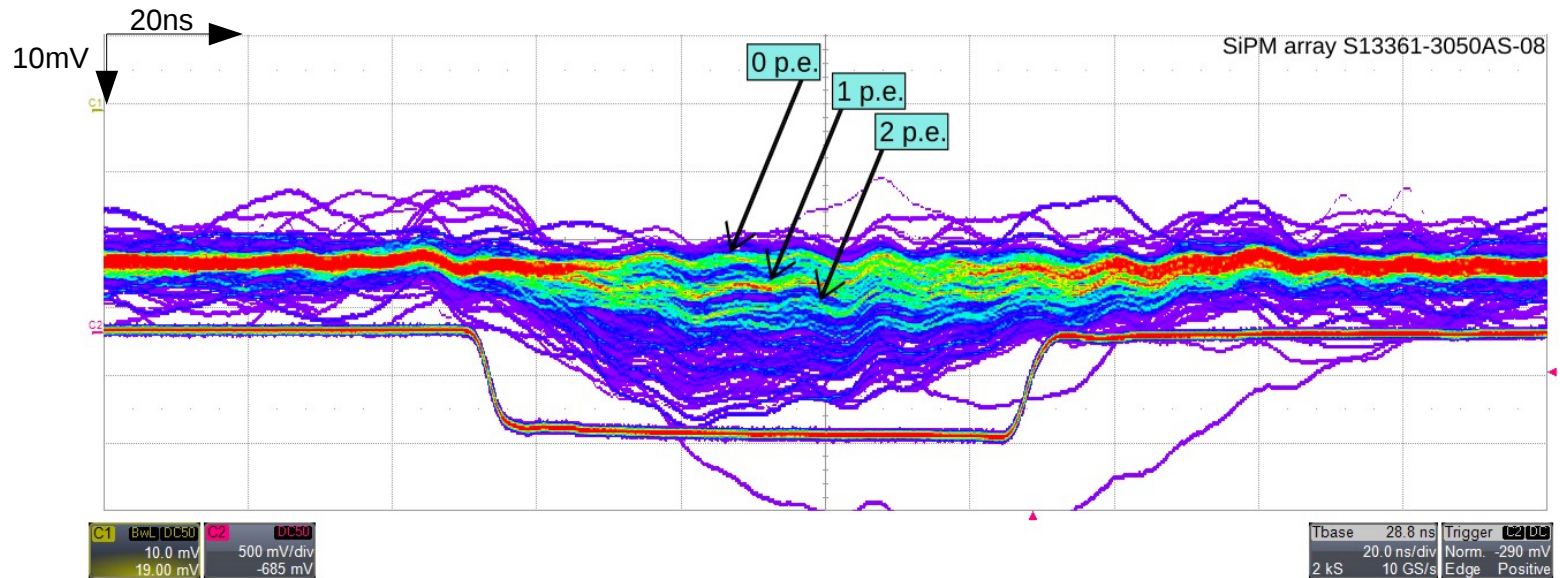
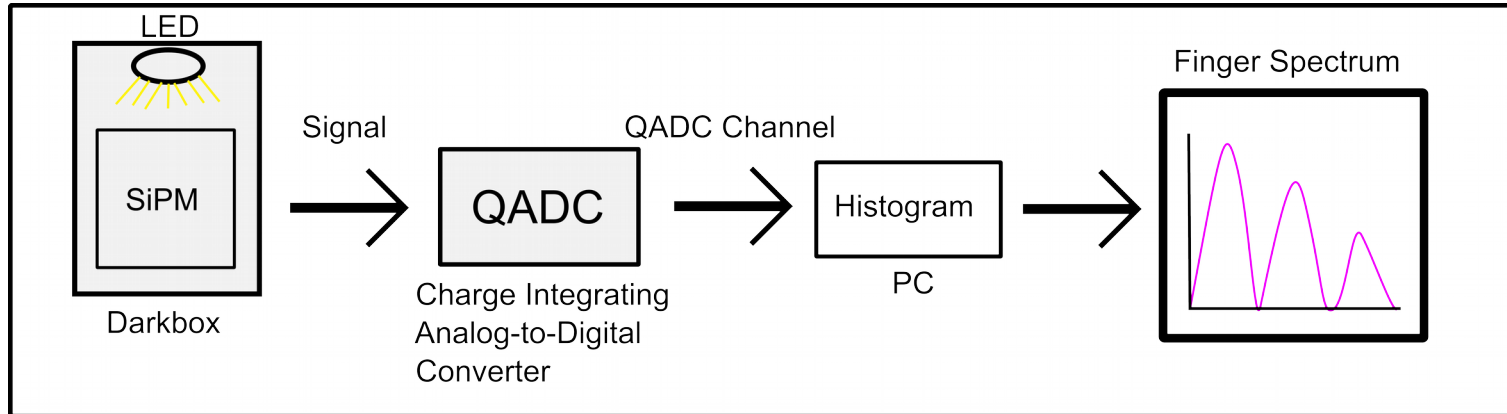
Available Wavelengths:

- 423nm
- 395nm
- 376nm
- 371nm

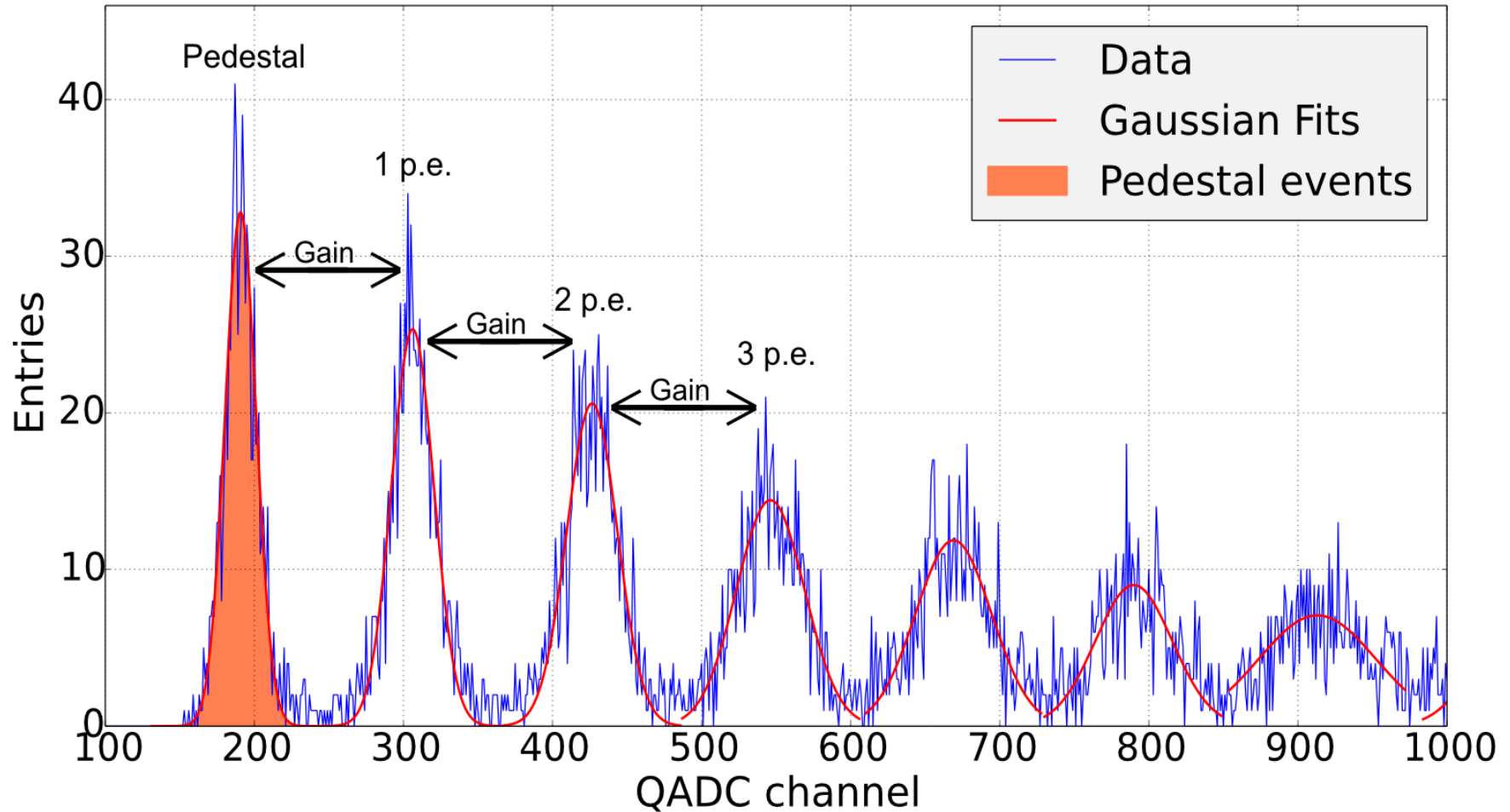
- Linear stage to automated measurement of more SiPM channels

- Only room Temperature (yet)

# Towards charge spectra

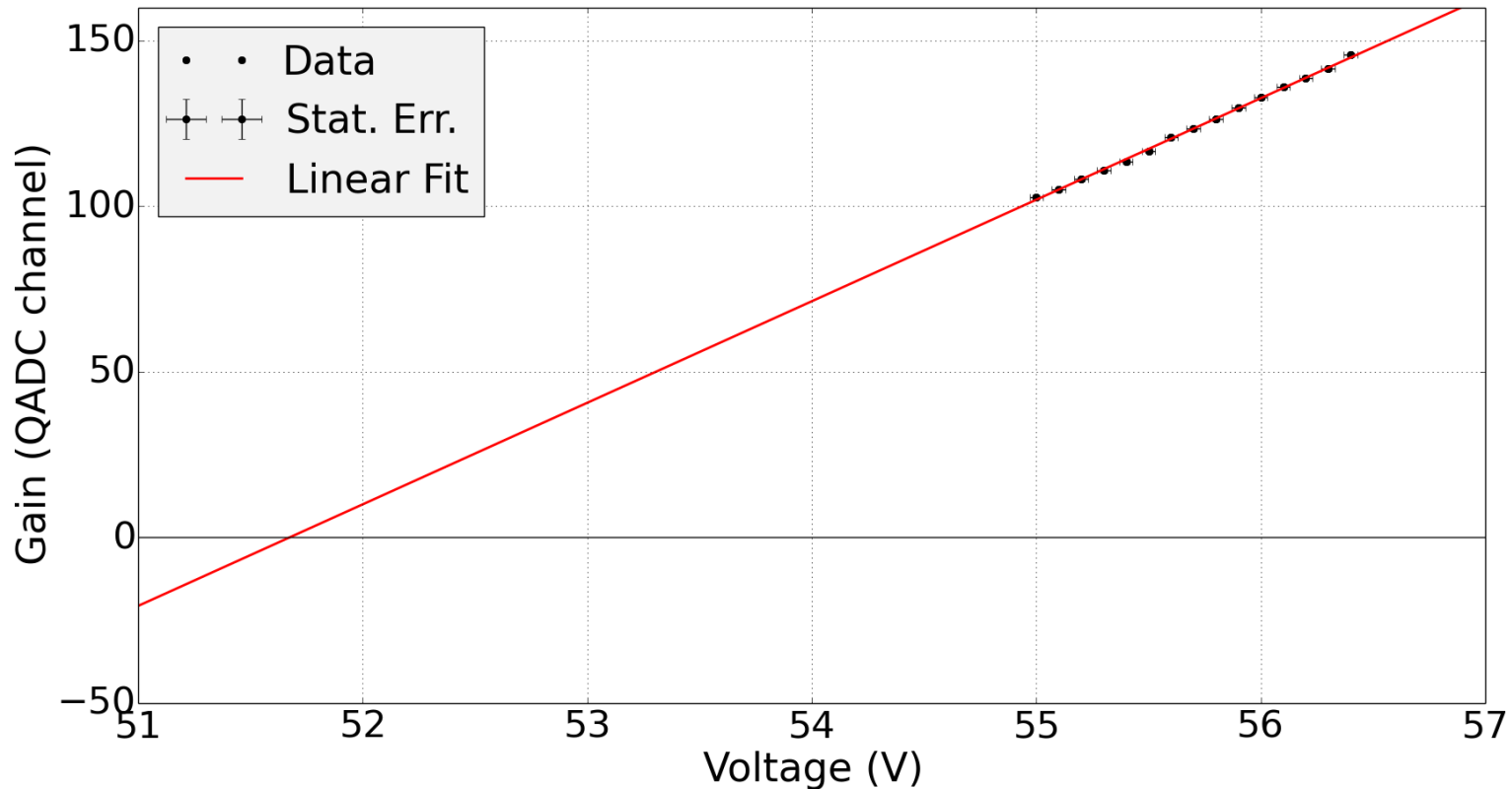


# Characterizing SiPMs – Finger spectrum



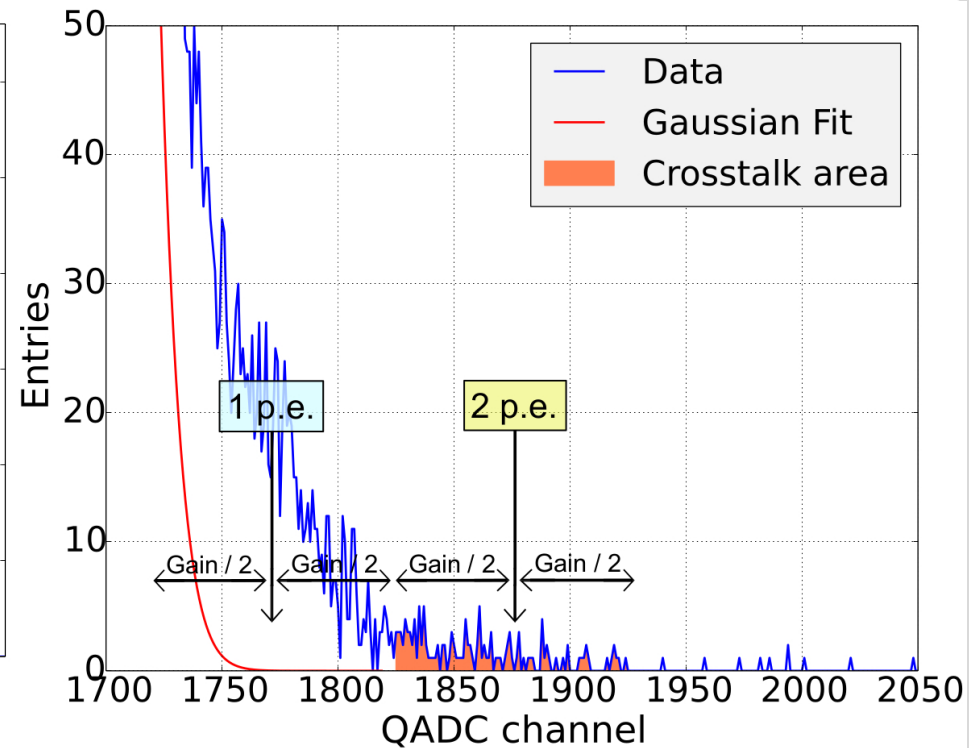
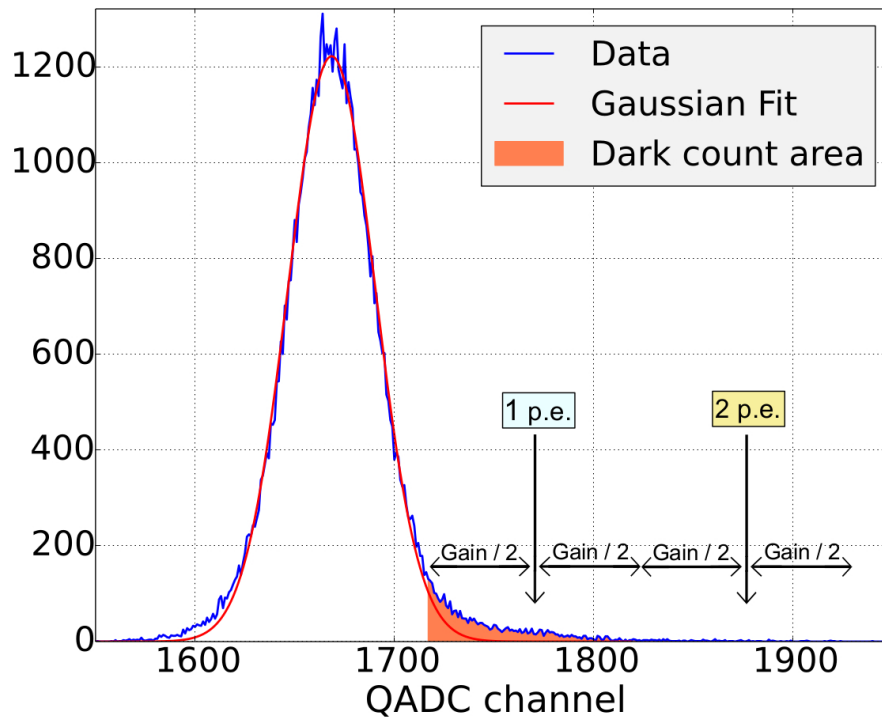


# Characterizing SiPMs – Breakdown Voltage



Gain-measurements at different over voltages  
→ Extrapolating gives breakdown voltage estimate

# Characterizing SiPMs – Dark Spectrum



(p.e. peak positions as result of gain measurement before)

# Characterizing SiPMs – Dark Count Rate and Crosstalk Probability

## Dark Count Rate

All dark events after the pedestal are assumed to be dark count events.

→ Dark Count Rate  $\xi$  :

$$\xi = \frac{N_{p.e.>0}}{t_{gate} \cdot N_{tot}}$$

## Crosstalk Probability

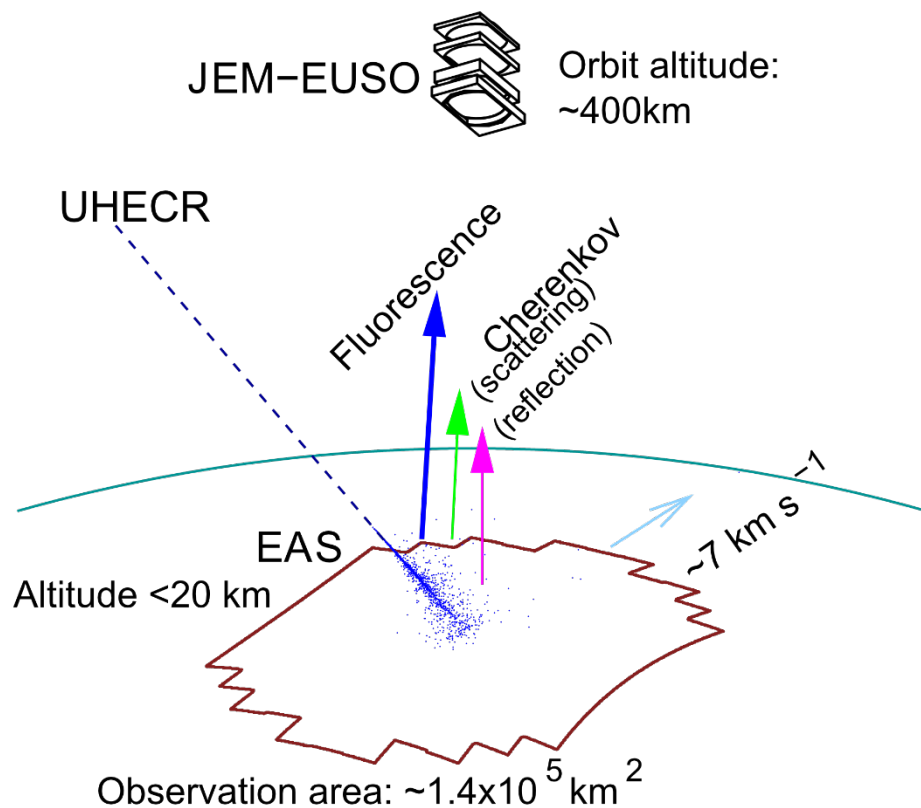
All dark events in the second p.e. region are assumed to be crosstalk events generated by dark counts.

→ Crosstalk Probability  $\epsilon$  :

$$\epsilon = \frac{N_{p.e.=2}}{N_{p.e.>0}}$$



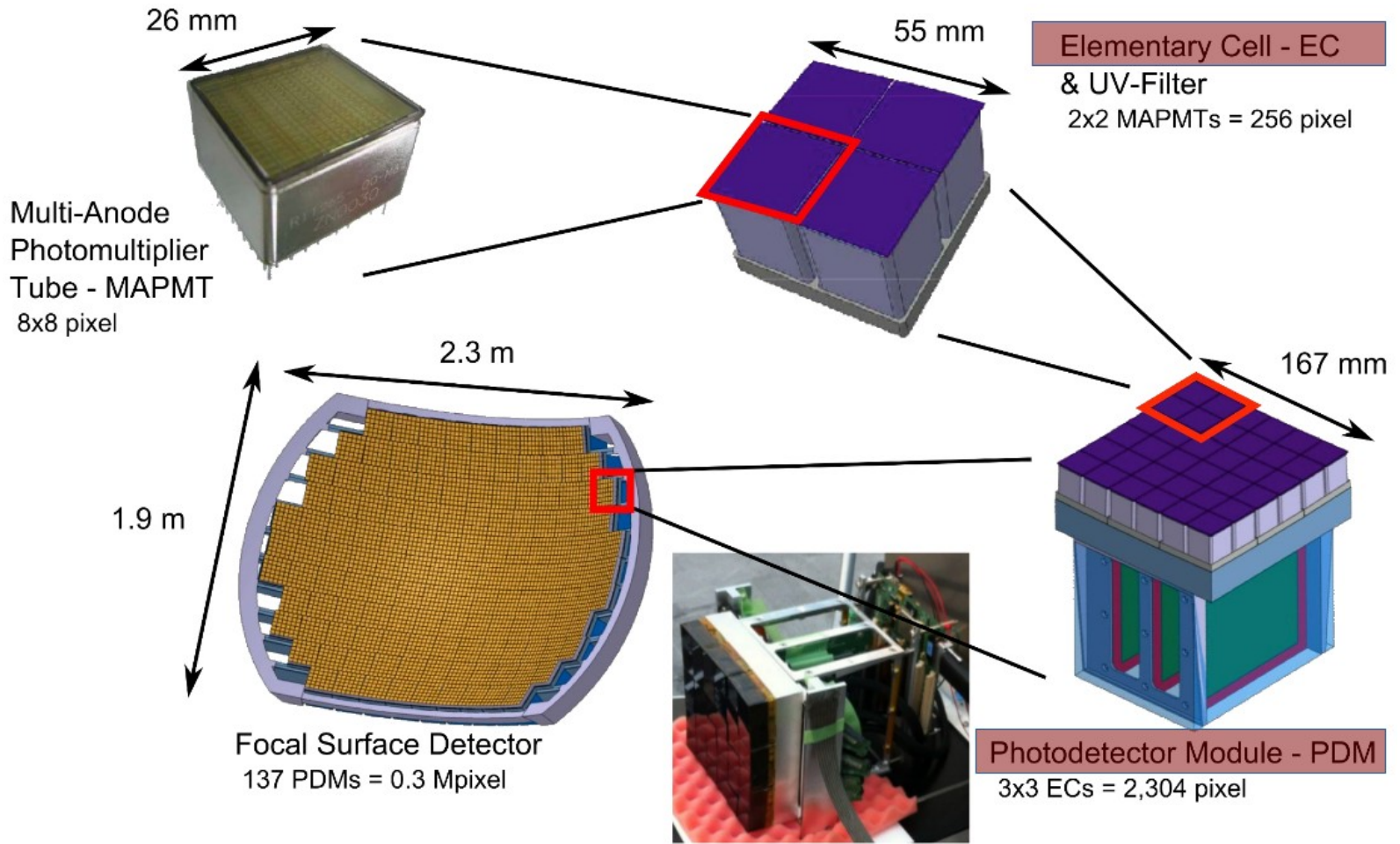
## The JEM-EUSO program



- Detection of Ultra High Energy Cosmic Rays by measuring induced Fluorescence light and scattered Cherenkov light
- 1.9m x 2.3m focal surface out of 4932 PMTs with 64 channel each PMT (standard design)
- Pathfinder Mission: EUSO – Super Pressure Balloon



# The JEM-EUSO program



# SiECA

## Silicon EC Add-on for EUSO-SPB.

- Fluorescence camera with 265 SiPM channels
- Placed next to an EUSO photo detection module (PDM)  
→ (Next slides)

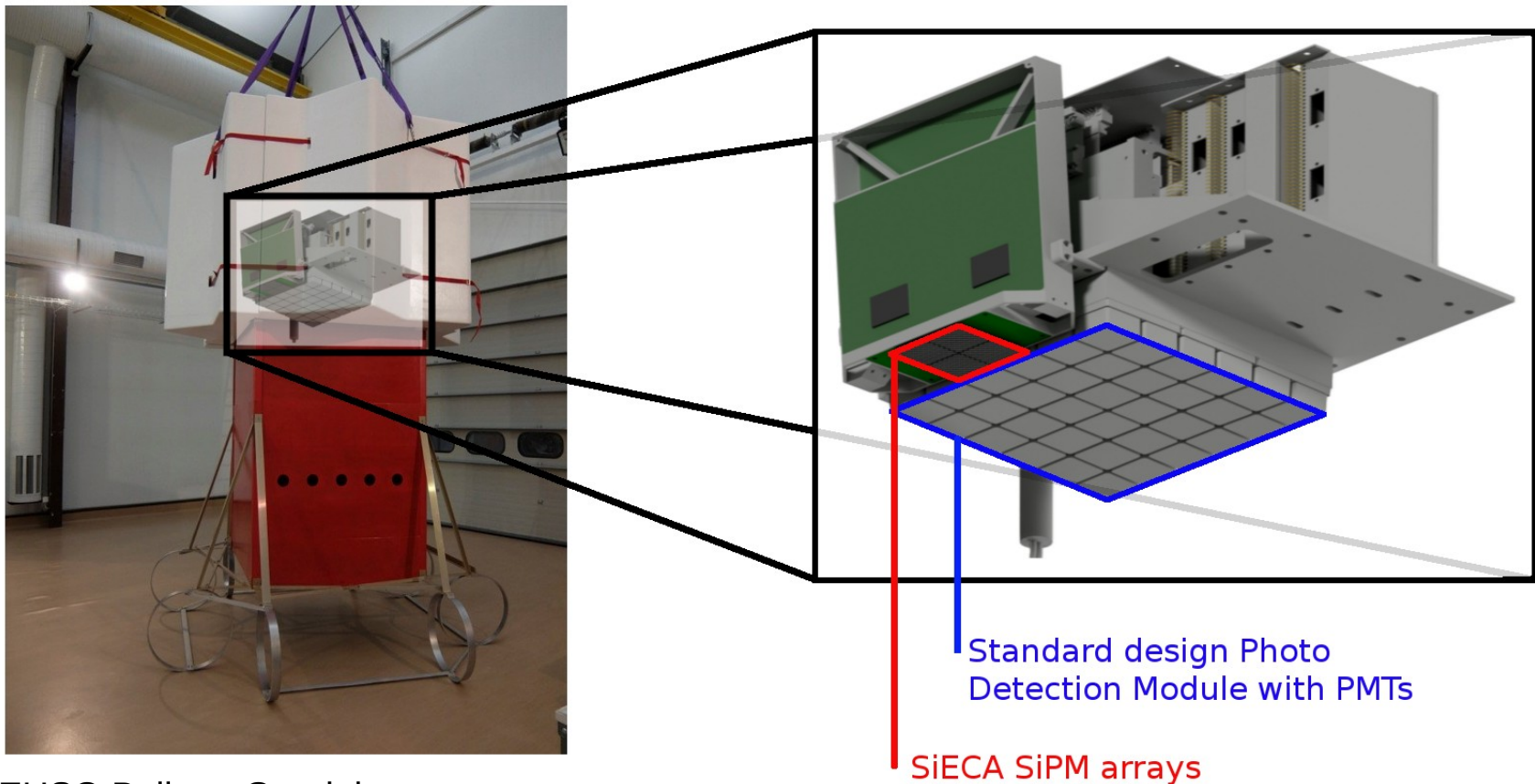


( Used: 4x 64 channel SiPM array  
S13361-3050AS)



EUSO-SPB.

# SiECA integration in Super pressure balloon gondola



EUSO Balloon Gondola

\*\*New Gondola has been built for EUSO-SPB\*\*

# Measured 64 channel SiPM arrays

## S12642 (old series)

- 64 Channels
- 50 $\mu$ m pitch
- Specifications by Hamamatsu:
  - Breakdown Voltage:  $65 \pm 10$  V
  - Dark Count: 2 - 3 M counts/s
  - Gain:  $1.25 \times 10^6$
  - Photo Detection Efficiency: 35%
- **Epoxy resin**
  - Down to  $\lambda = 320$ nm

## S13361

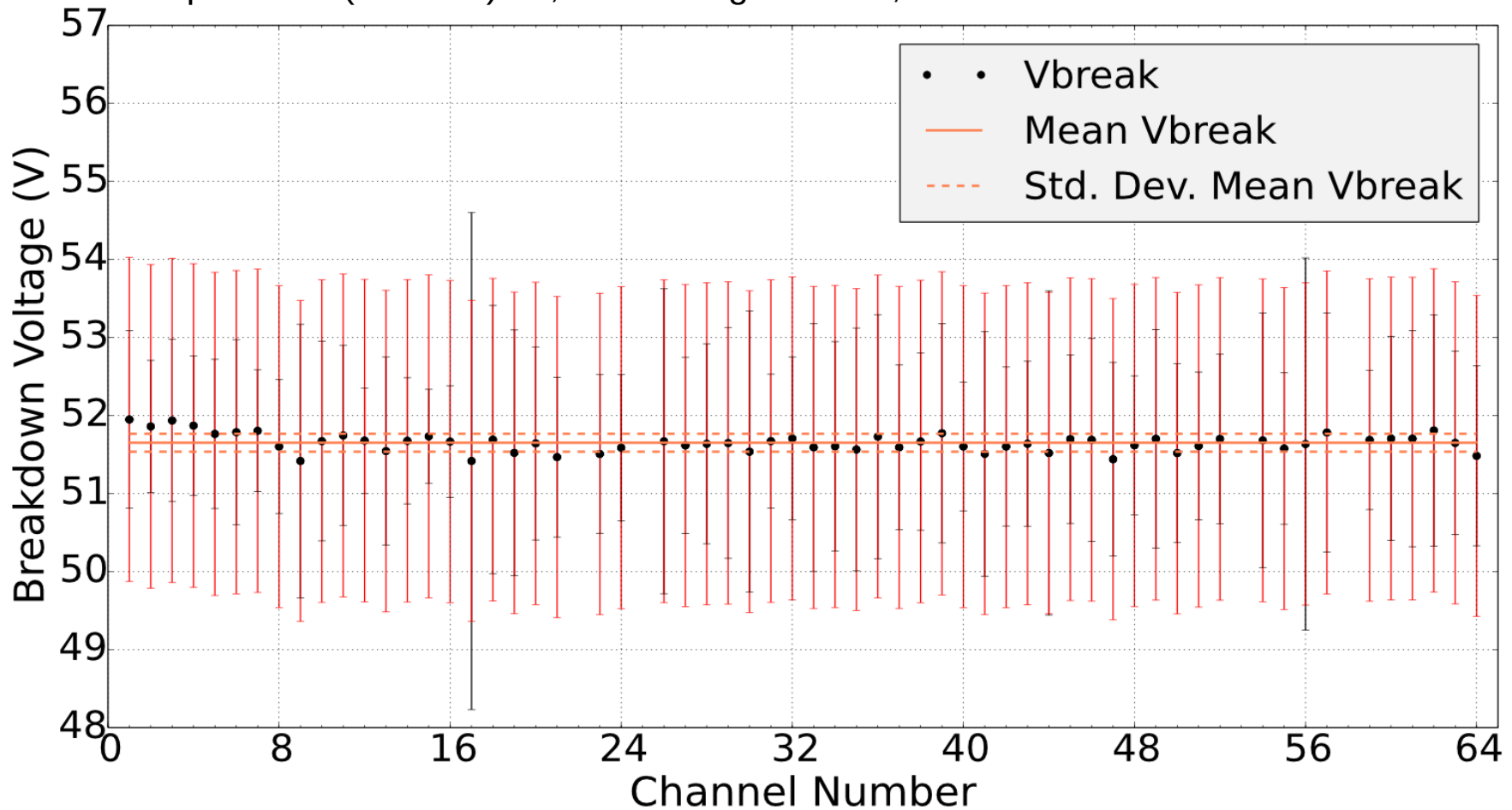
- 64 Channels
- 50 $\mu$ m pitch
- Specifications by Hamamatsu:
  - Breakdown Voltage:  $53 \pm 5$ V
  - Dark Count: 0.5 – 1.5 M counts/s
  - Gain:  $1.7 \times 10^6$
  - Photo Detection Efficiency: 40%
  - Reduced Crosstalk
- **Silicone resin**
  - Down to  $\lambda = 270$ nm
  - chosen for SiECA





# S13361 – Breakdown Voltage

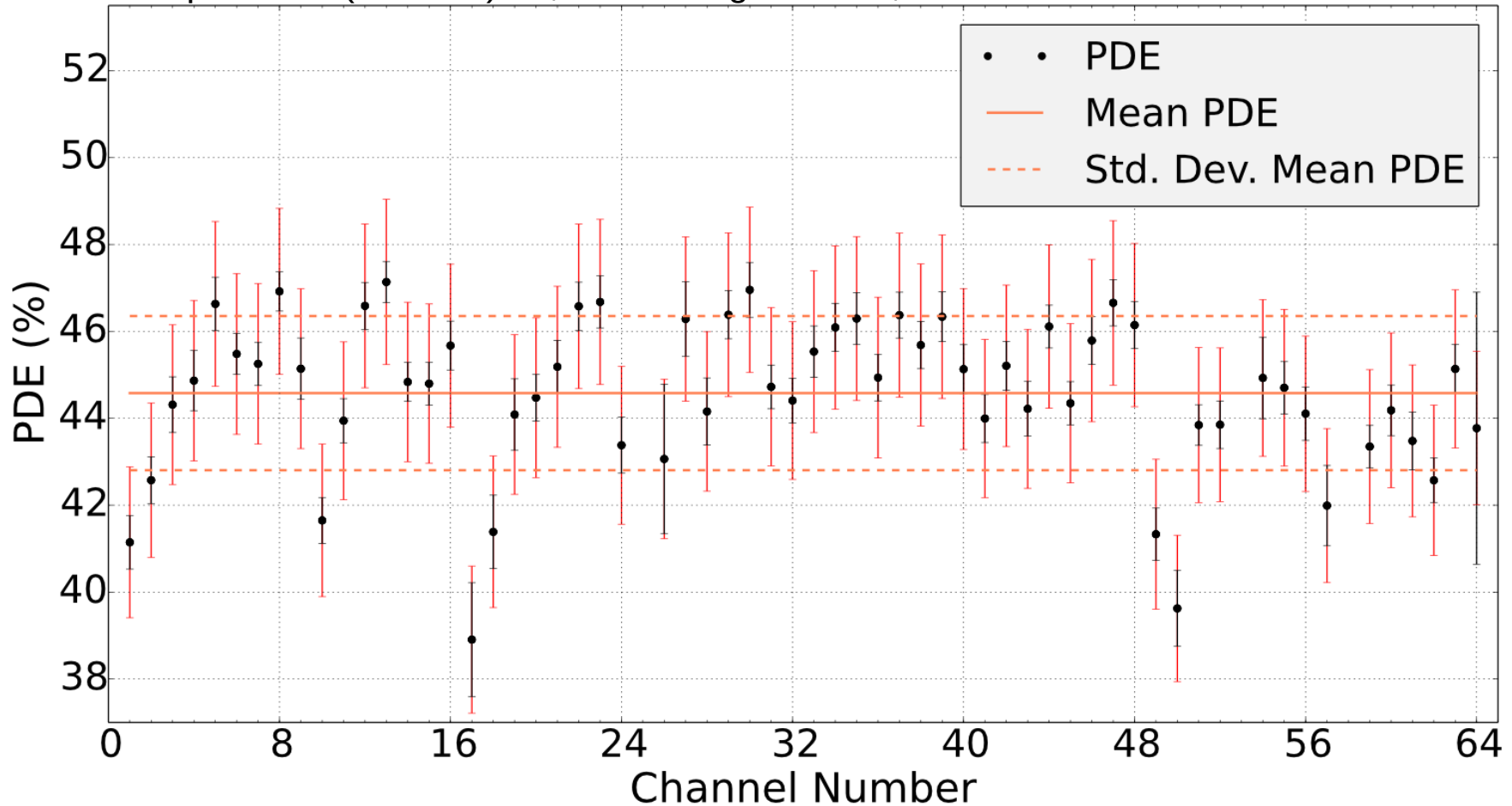
Temperature:  $(19.5 \pm 1)^\circ\text{C}$  , Bias voltage: 55.2 V,  $\lambda = 423\text{nm}$



Mean breakdown voltage:  $(51.65 \pm 0.12)$  V

# S13361 – Photo Detection Efficiency

Temperature:  $(19.5 \pm 1)^\circ\text{C}$  , Bias voltage: 55.2 V,  $\lambda = 423\text{nm}$



Mean PDE :  $(44.60 \pm 1.78) \%$

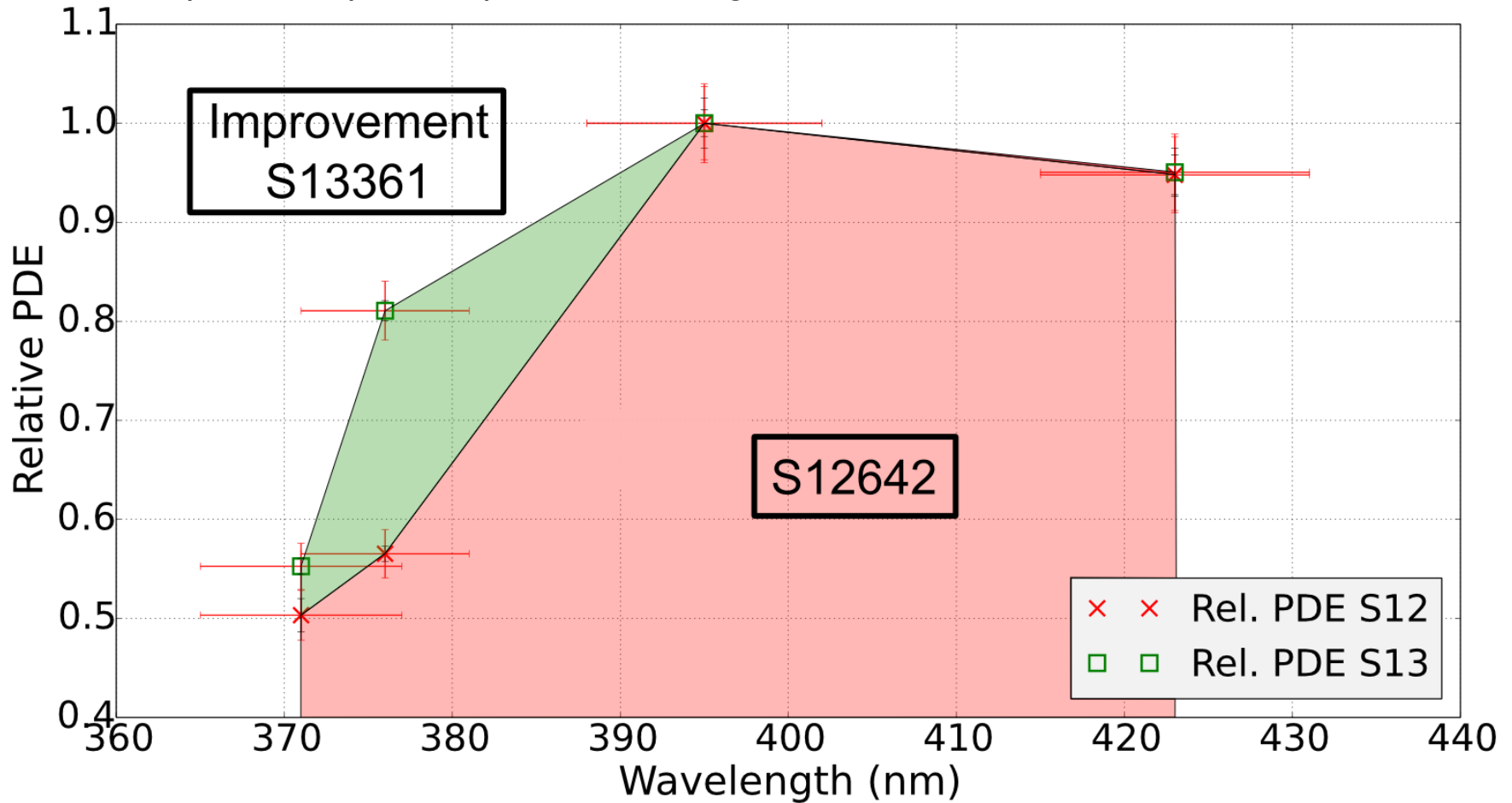
# Characterising results and Comparison Hamamatsu S12642 ↔ S13361

	S12642	S13361	Comp.
Vbreak (V)	$64.62 \pm 0.1$	$51.65 \pm 0.12$	20%
Gain (E+06)	$1.65 \pm 0.04$	$2.12 \pm 0.07$	28 %
PDE (%)	$35.96 \pm 1.09$	$44.60 \pm 1.78$	25 %
Dark Counts (MHz)	$1.29 \pm 0.14$	$0.68 \pm 0.11$	43 %
Crosstalk Prob. (%)	$11.17 \pm 1.27$	$3.9 \pm 0.66$	65 %

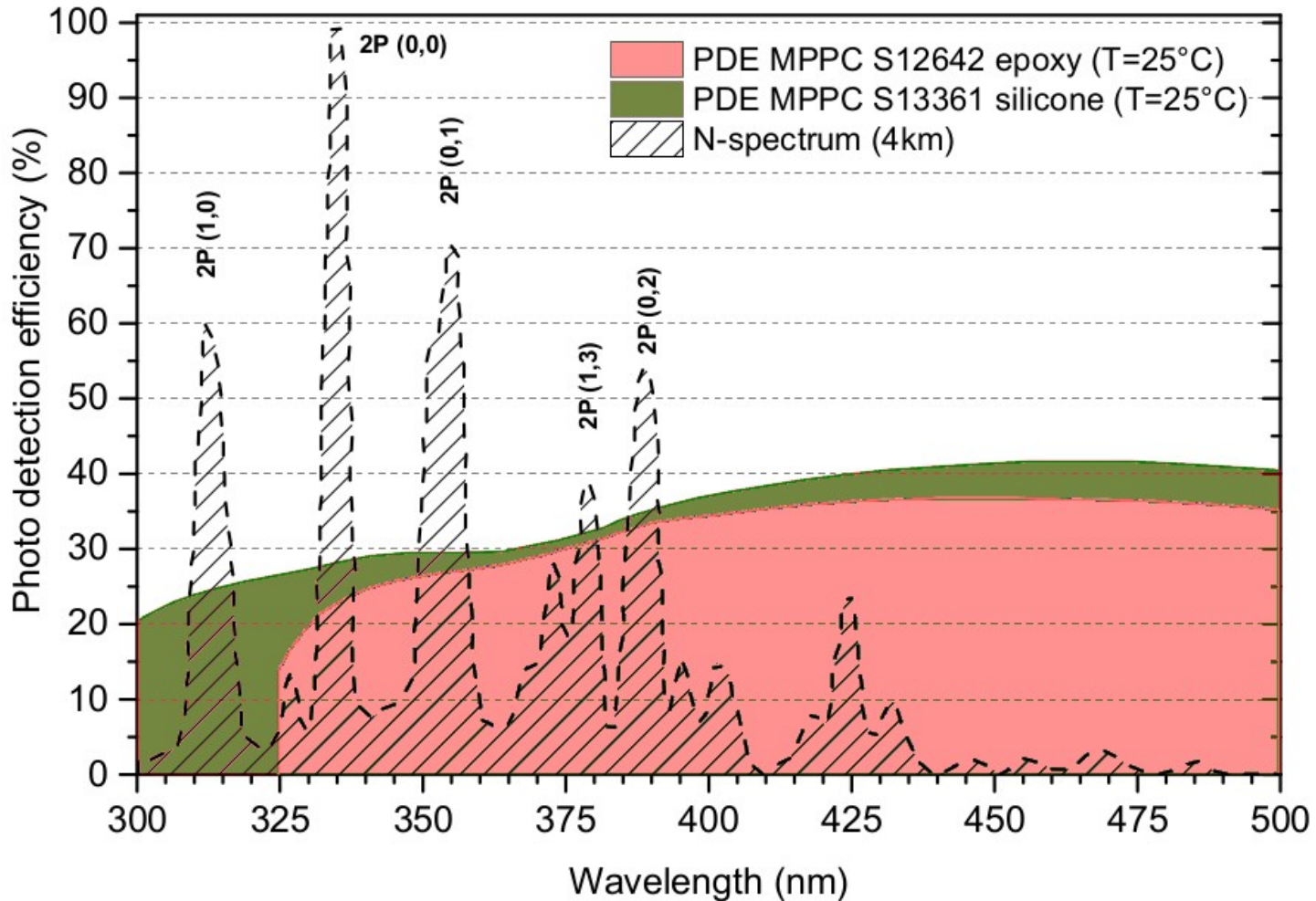
- Temperature :  $(19.5 \pm 1)^\circ\text{C}$
- Overvoltage:
  - S13361: 3.58 V (+0.58 V than recommended)
  - S12642: 2.95 V (+0.55 V than recommended)
- Wavelength: 423nm

# Wavelength Sensitivity

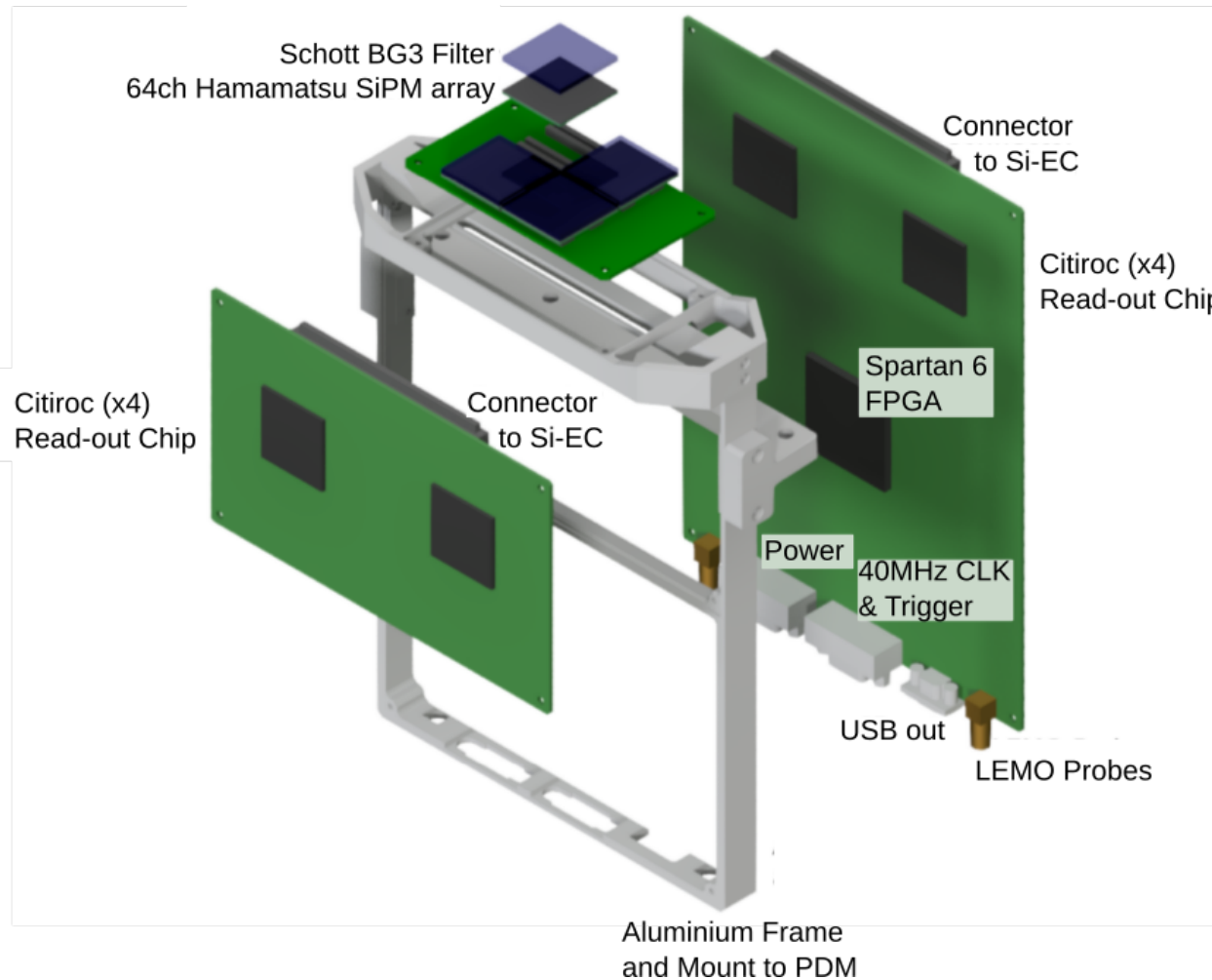
Temperature:  $(19.5 \pm 1)^\circ\text{C}$  , Bias voltage: 55.2 V



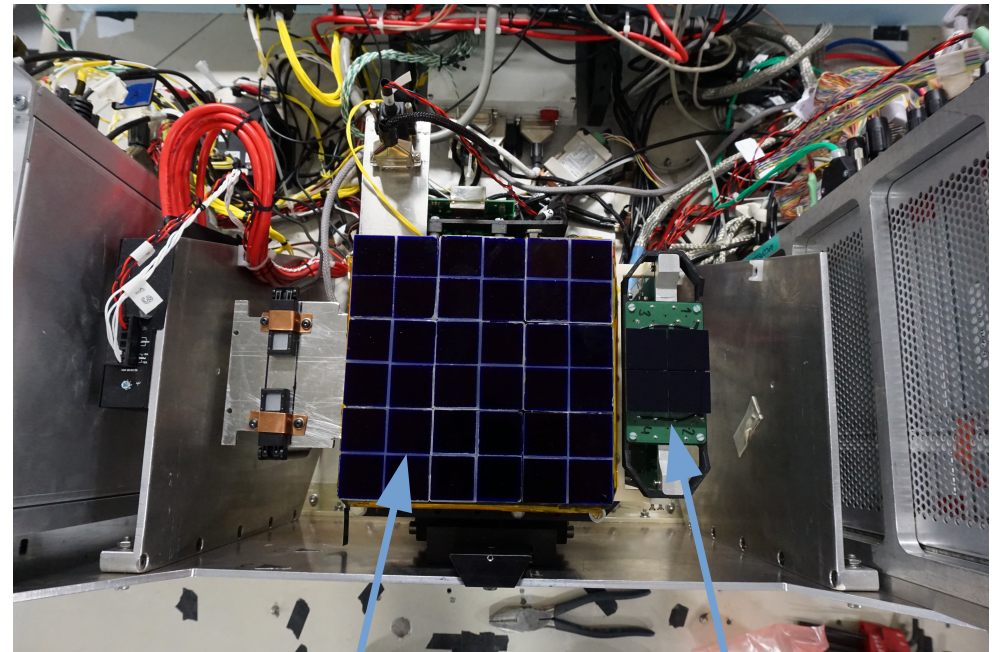
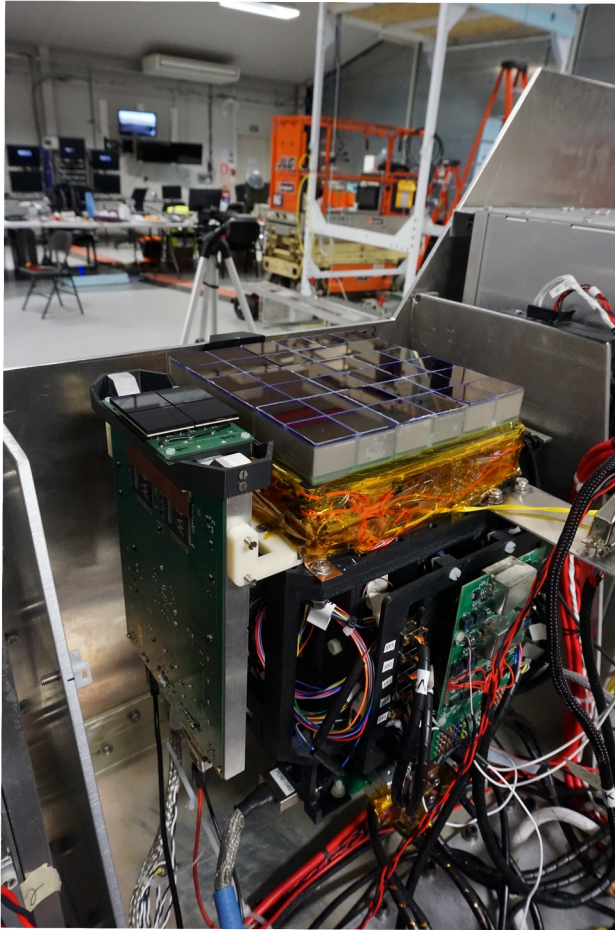
# Predicted effect of different SiPM resins



# SiECA in detail



# Euso-SPB with SiECA: ready to start



PDM  
(MAPMT  
camera)

(SiECA  
SiPM  
camera)

# First super pressure balloon flight 2015



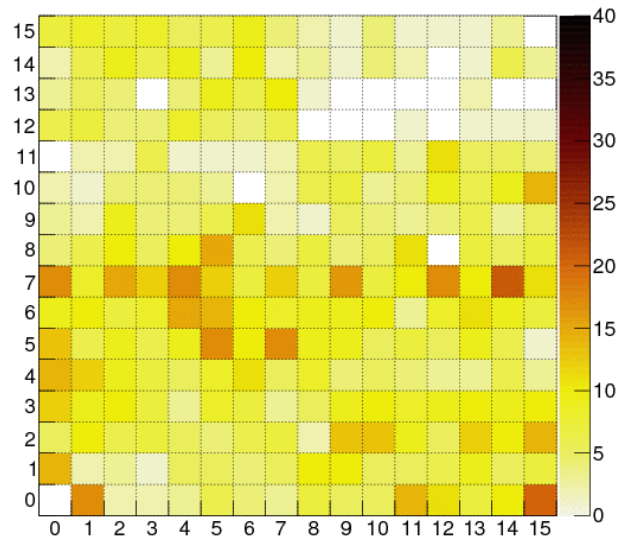
NASA's first Super Pressure Balloon flight, March 2015, Wanaka, NZ:

- Flight duration 32 days



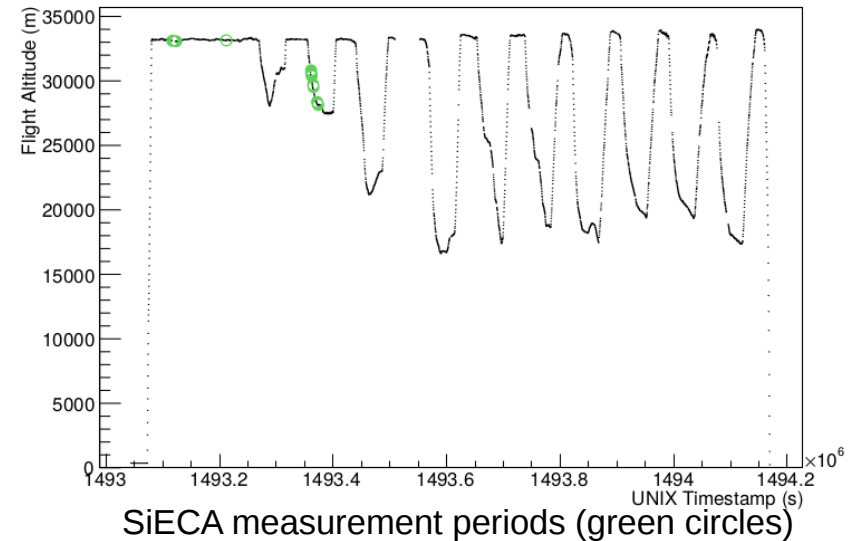
# SiECA Results

Total Counts per Event per Channel

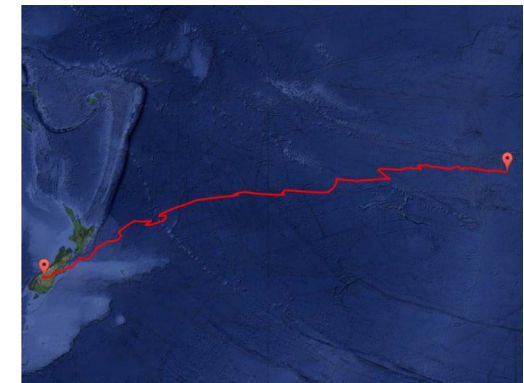


SiECA background measurement

SiECA Measurement Periods



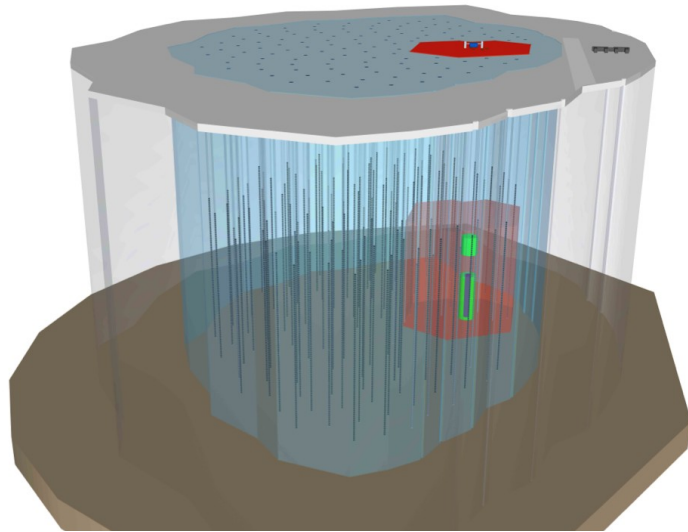
- Successful measurements with SiECA during flight time
- Super pressure balloon went down after 12 days due to leak in balloon



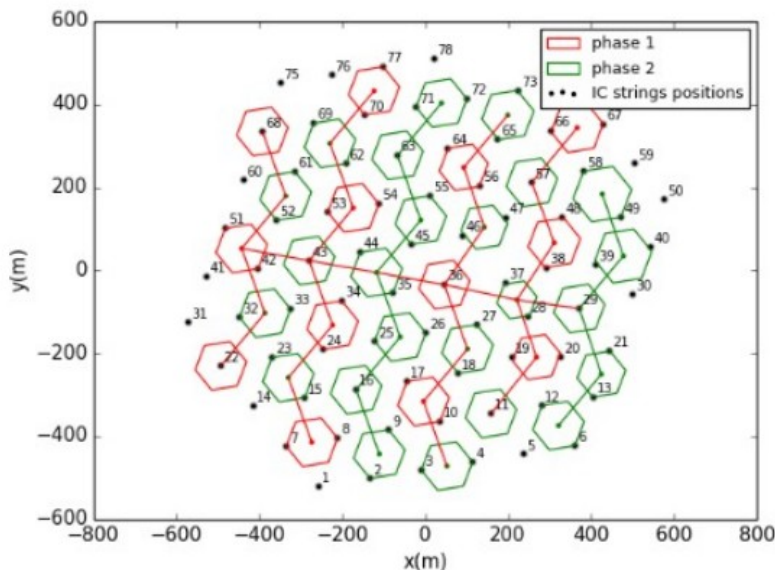
SPB flight path

# IceCube-Gen2 scintillation detectors

(1 channel SiPM S13361-6025PE)



- Extended surface array consisting of scintillation detector panels with SiPM read-out
- Veto for in-ice detector and improved measurement of cosmic ray air showers
- Prototype study for IceCube-Gen2 upgrade

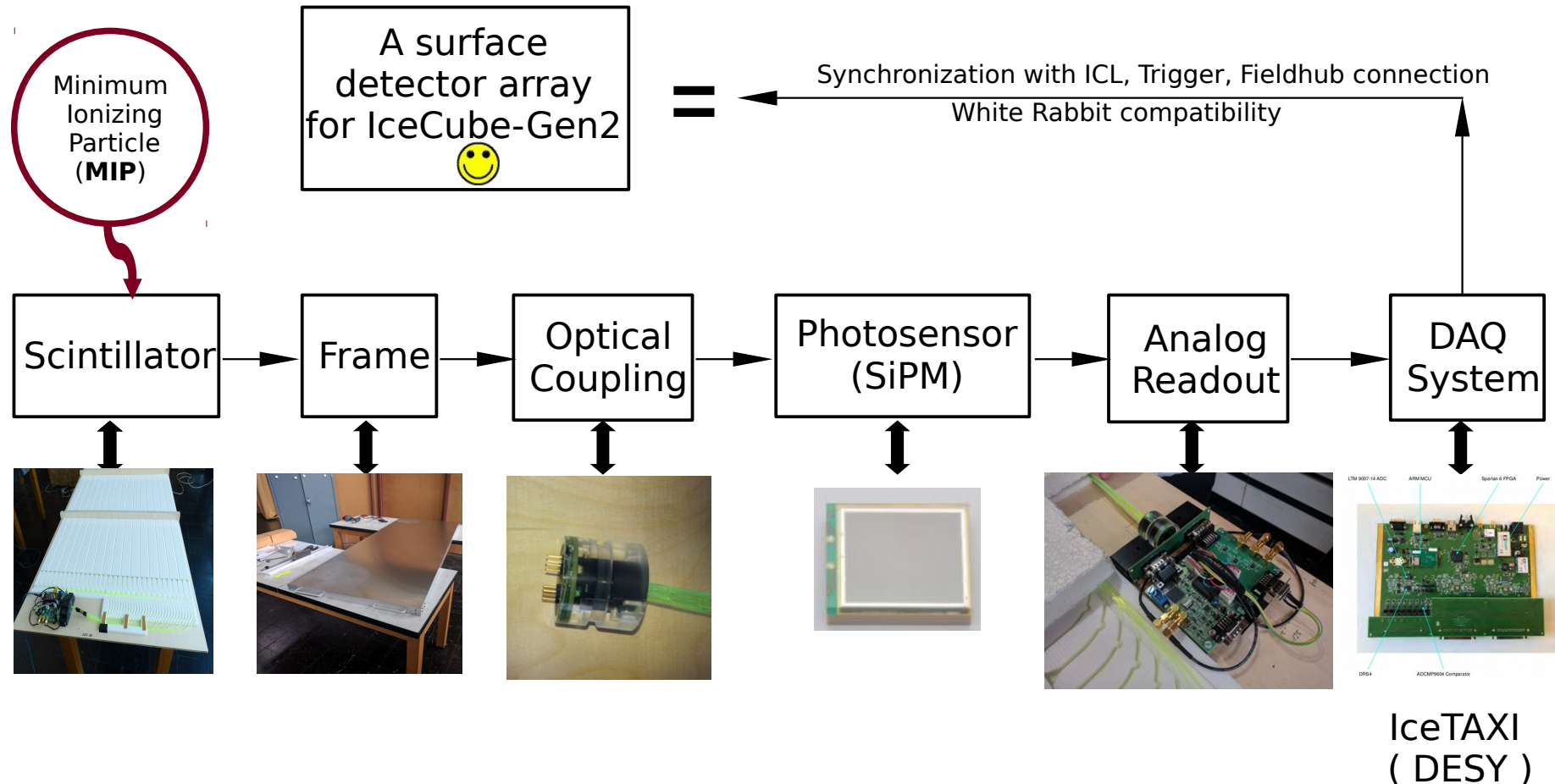


- Seven scintillation panels in one station
- One TAXI-DAQ for every station
  - SiPM signal sampling with DRS4 sampling chips @ 1GSamples/s

# IceCube-Gen2 scintillation detectors



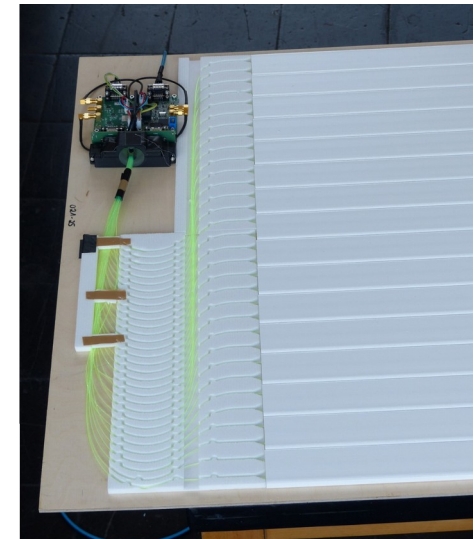
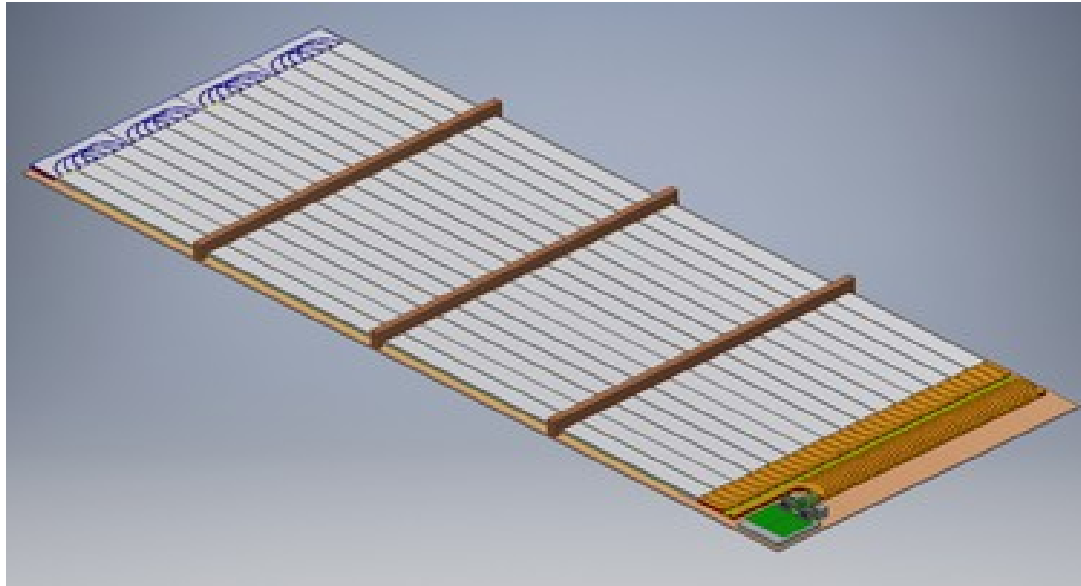
ICECUBE  
GEN2



# Detector design – bars and fibers



ICECUBE  
GEN2

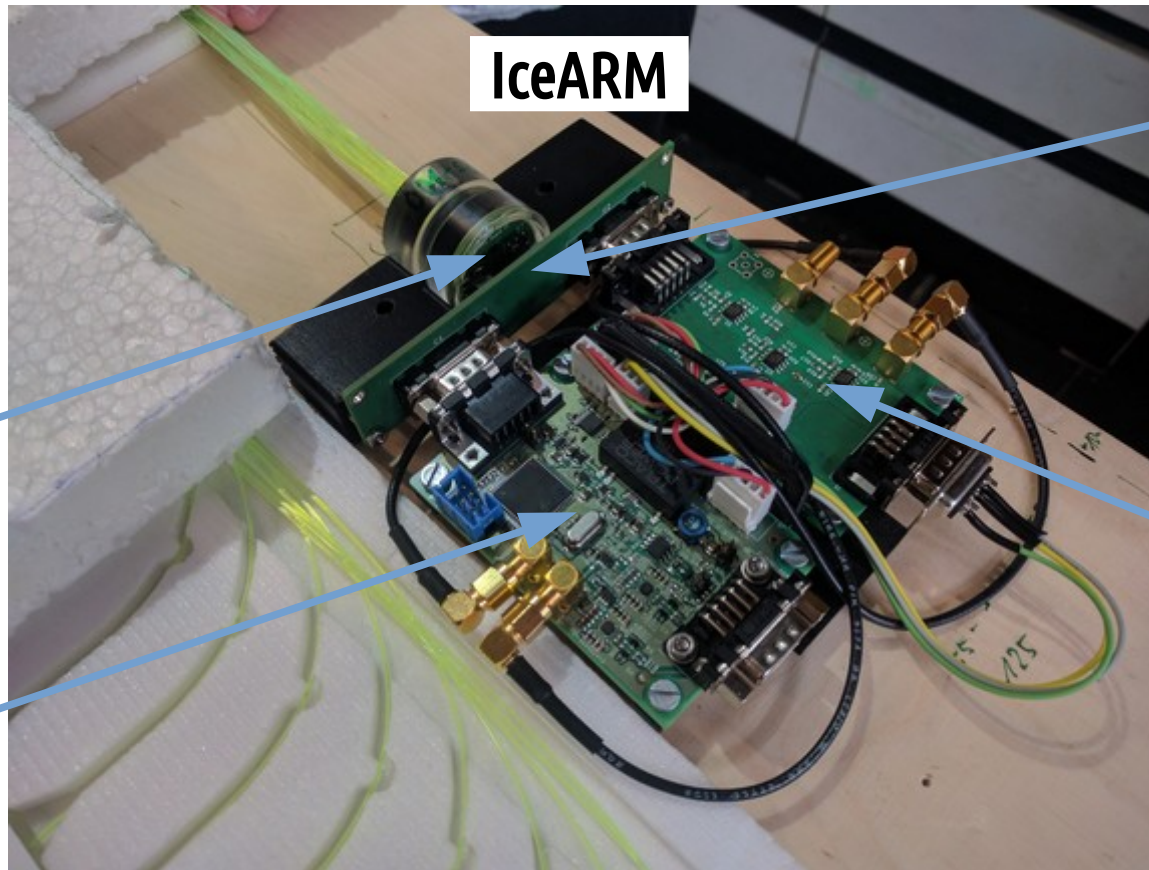


- 16 Fermilab scintillator bars, 2 optical fiber ports each bar
- 32 Kuraray wavelength shifting fiber ends coupled to SiPM
- MIP-Sensitive Area: 1.5 m<sup>2</sup>
  
- Dimensions: 2.2m x 0.8m x 0.05m

# Analog readout



ICECUBE  
GEN2



Cookie  
board

General  
purpose board  
(DESY)

**IceARM**

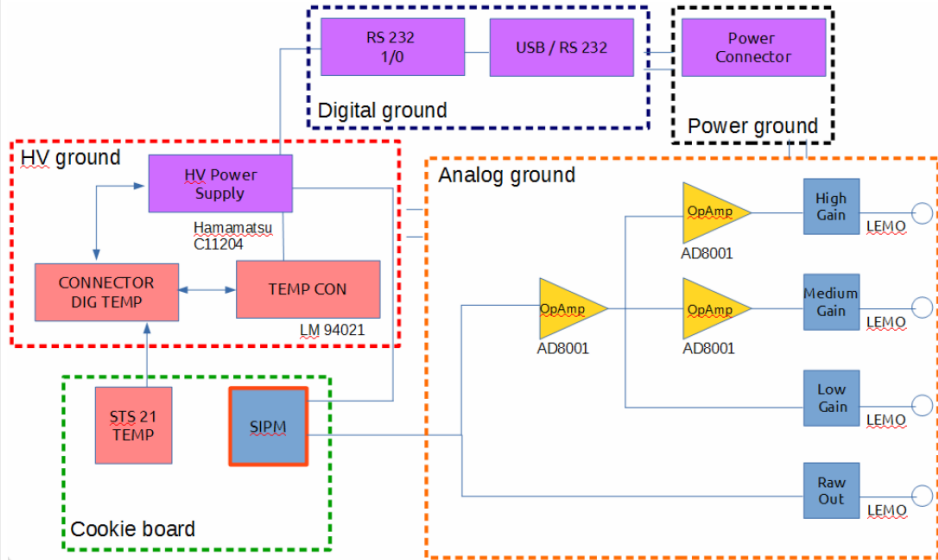
Adapter  
board

Readout  
board

# Analog readout

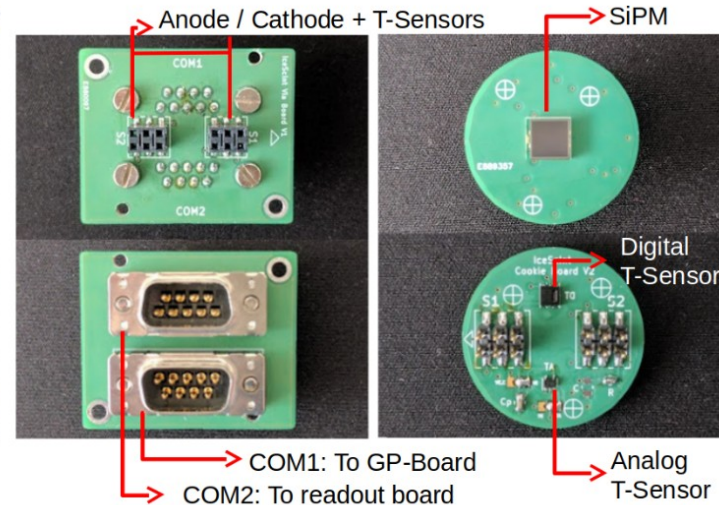
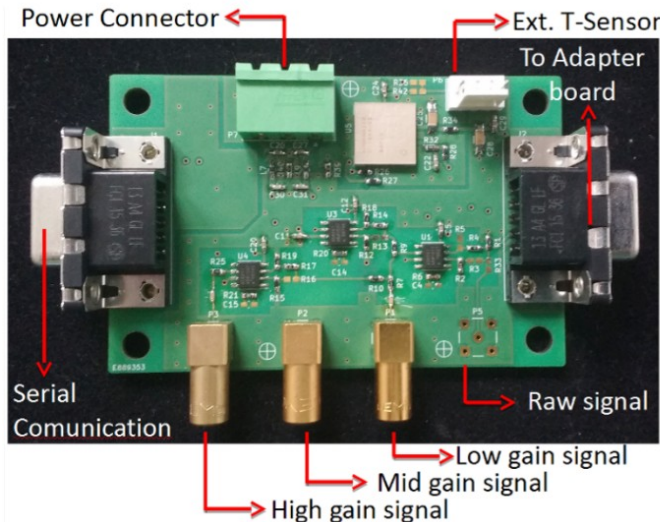


ICECUBE  
GEN2



- SiPM amplification and voltage control
- Automatic temperature compensation implemented in microcontroller
- High gain (x17) and low gain (x1.7) (after 70m cable)

Simplified ARM block diagram with amplifier chain for lab tests



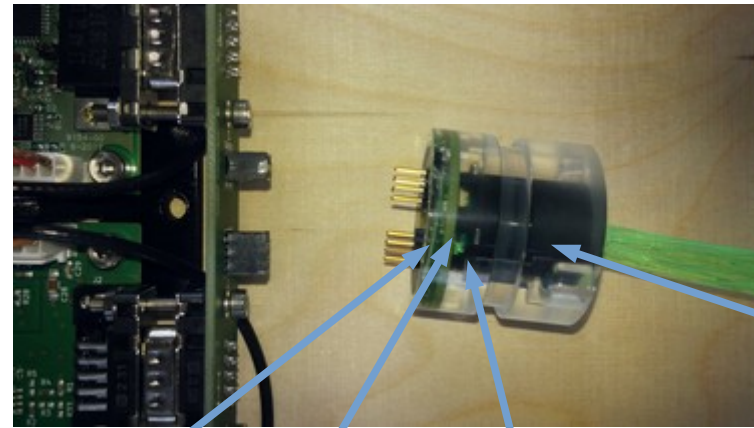
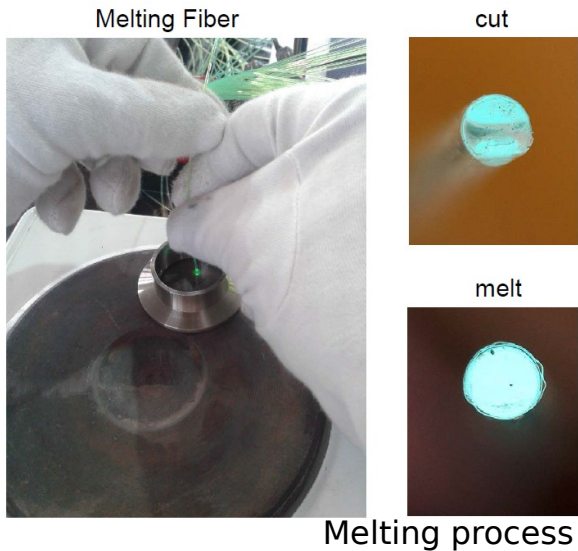
# Optical coupling



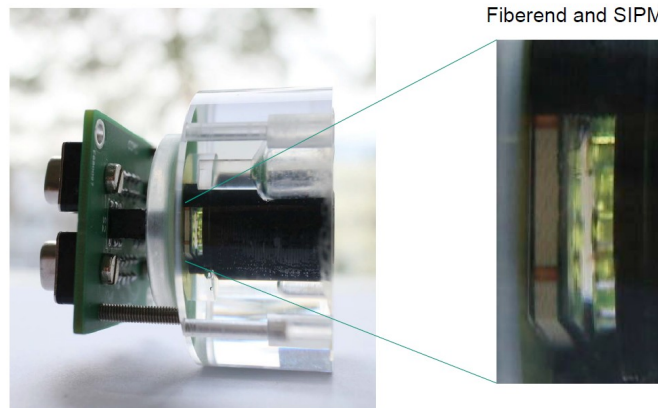
ICECUBE  
GEN2



Optical Fibers → “Cookie” → Cut and glued optical fibers → SiPM →  
→ Cookie board → Adapter board → Readout electronic



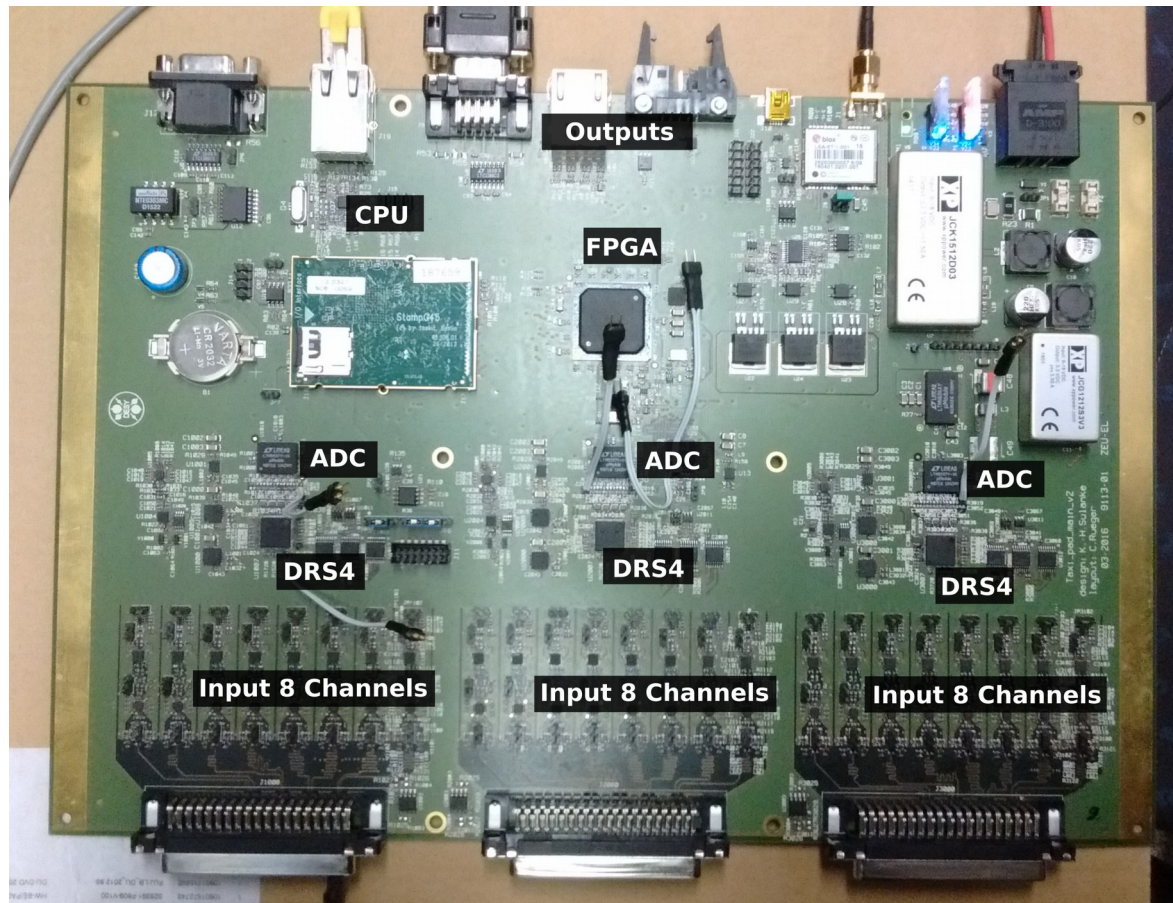
3D printed socket  
("round" to  
"squared")



1-2mm gap, filled with optical glue, to ensure triggering every APD cell to increase the dynamic range



# DAQ: IceTAXI board (DESY)



( Not actual version )

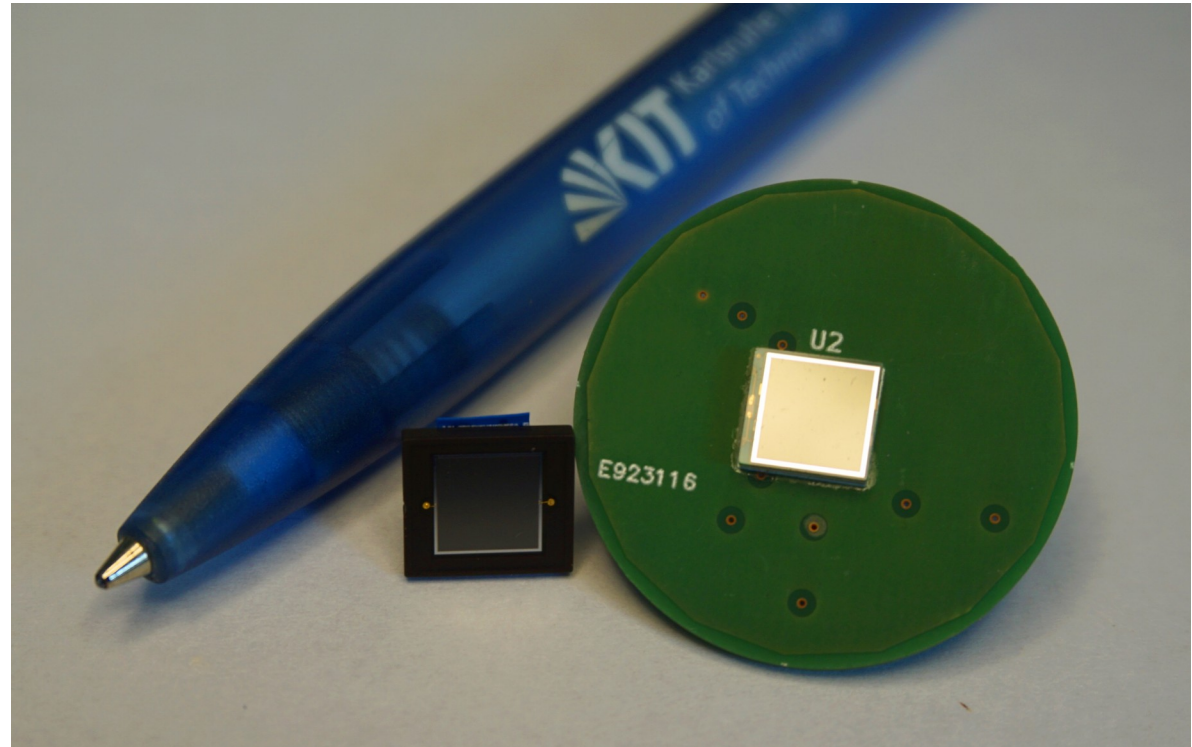
- Three DRS4 sampling chips with 8 input channels each chip
- Adjustable sampling rate 0.7GHz – 5GHz
- Triggered by signal-over-threshold
- DRS4 sampling cells are read out through slow ADCs (33MHz)

# Chosen: Silicon Photomultiplier S13360

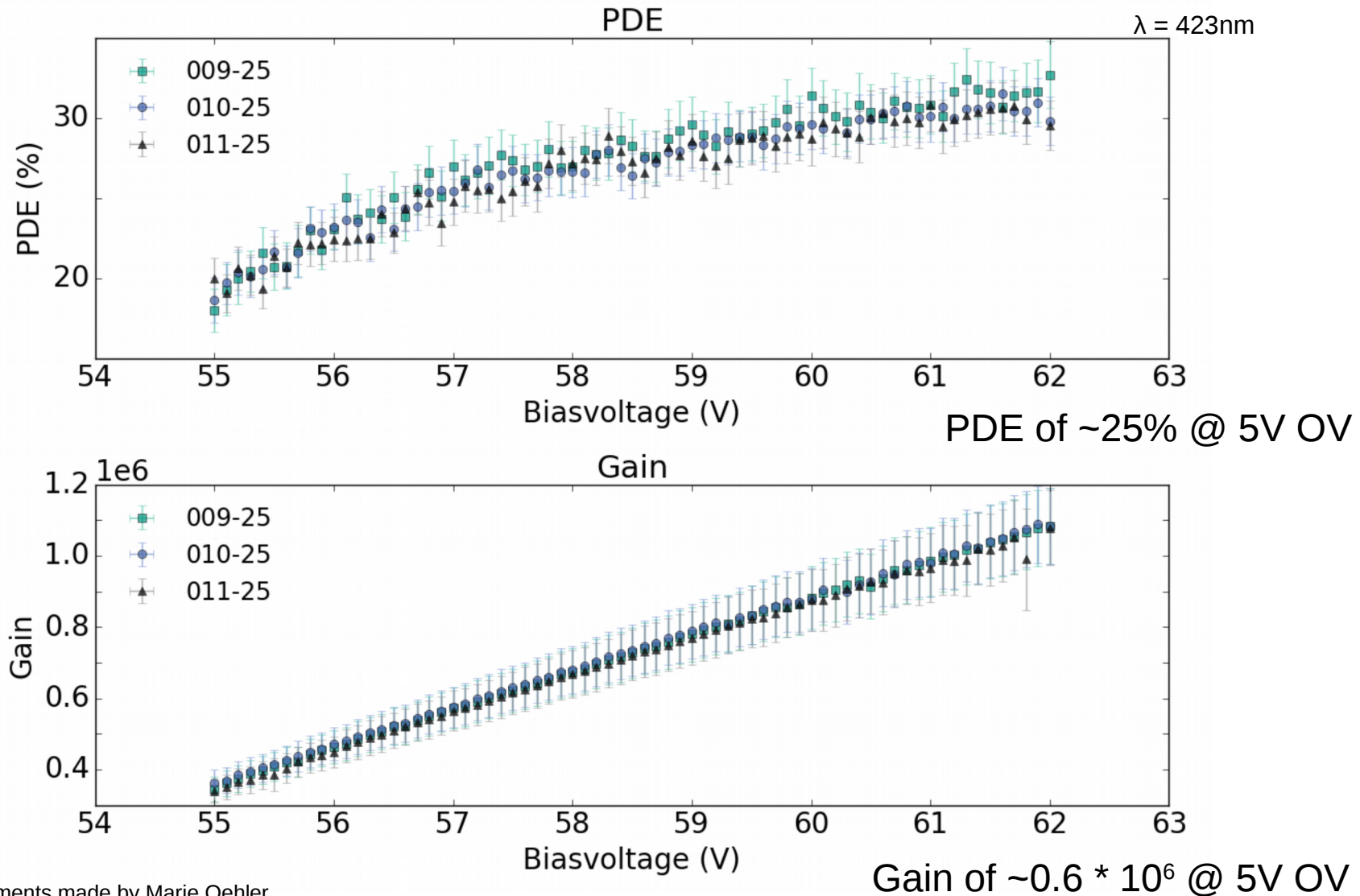
Hamamatsu S13360-6025

- 25 $\mu$ m pitch
- 6mm x 6mm active area
- Recommended overvoltage = 5V

Chosen for the IceCube-Gen2 scintillator prototype study, shipped for this season to the South Pole for deployment

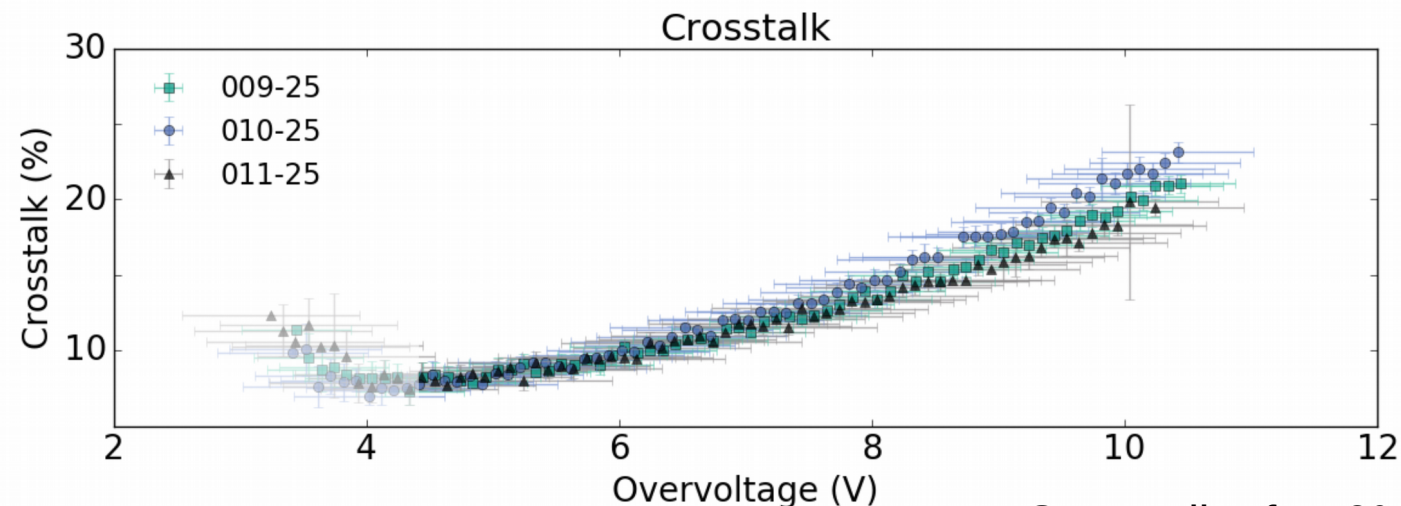


# Results SiPM Characterization

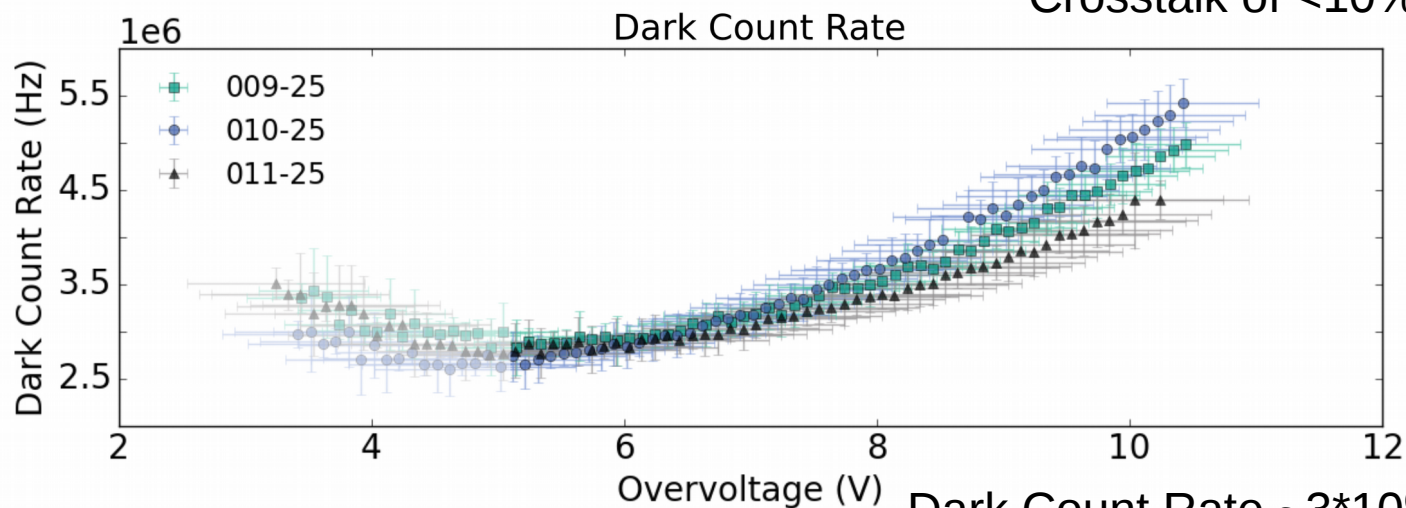


Measurements made by Marie Oehler

# Results SiPM Characterization



Crosstalk of  $<10\%$  @ 5V OV



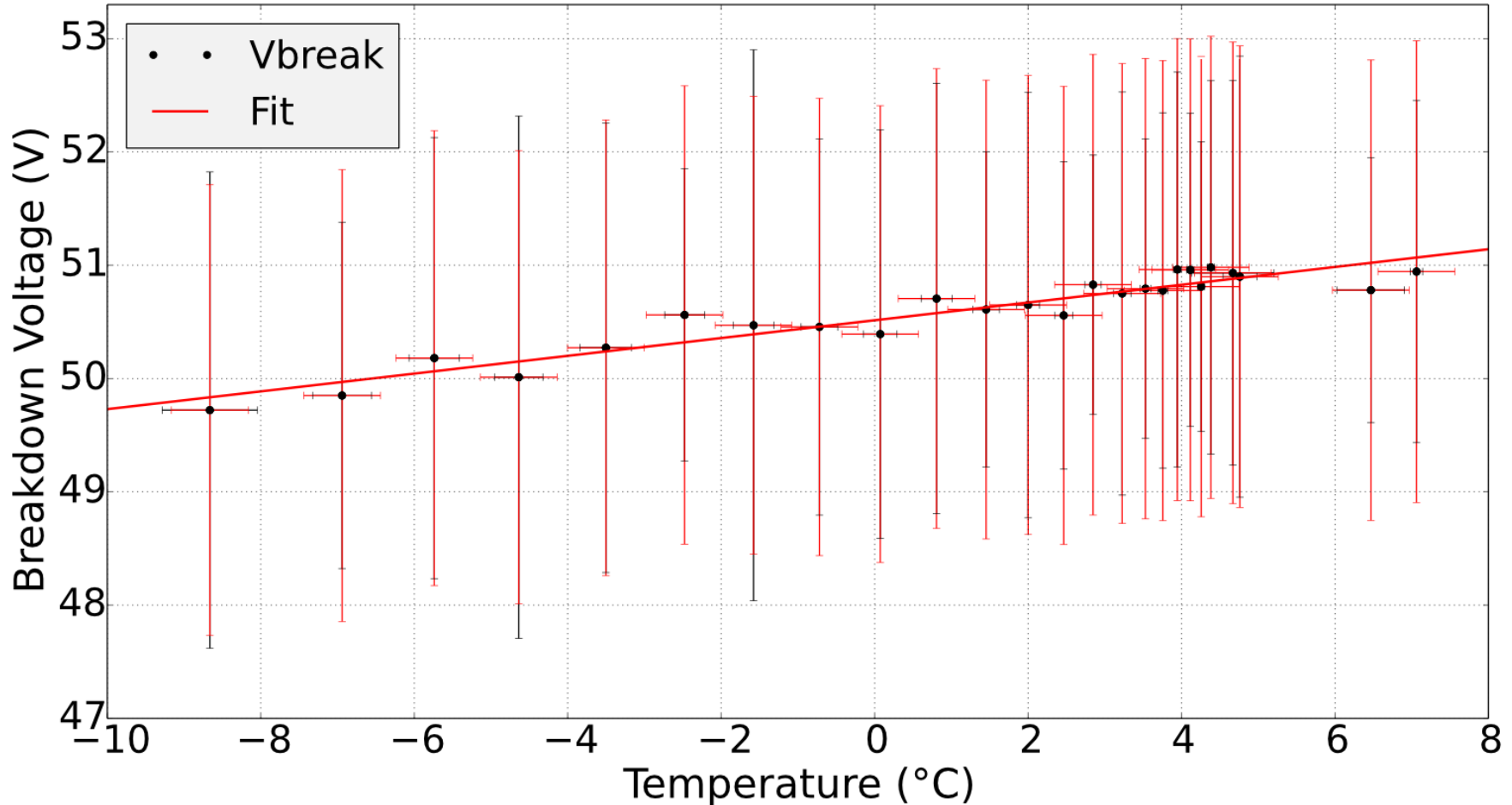
Dark Count Rate  $\sim 3 \times 10^6$  @ 5V OV

Measurements made by Marie Oehler

# Temperature dependent SiPM Characterization

# Temperature dependent breakdown voltage

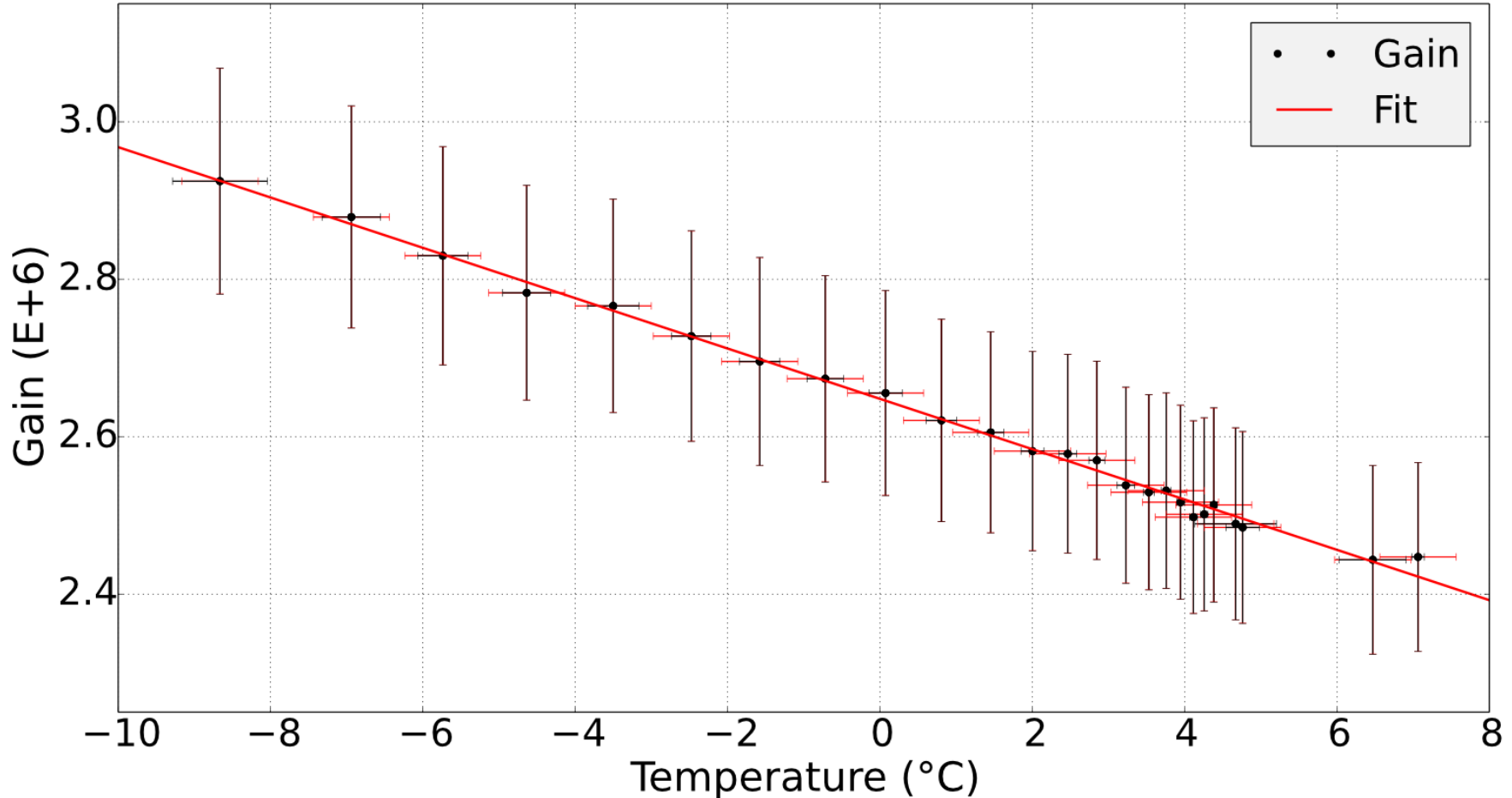
S13361 / 1 channel, constant bias voltage: 55.2 V,  $\lambda = 423\text{nm}$



$$\text{Fit: } V_{break}(T) = 0.078 \frac{\text{V}}{^{\circ}\text{C}} * T + 50.514 \text{ V}$$

# Temperature dependent gain

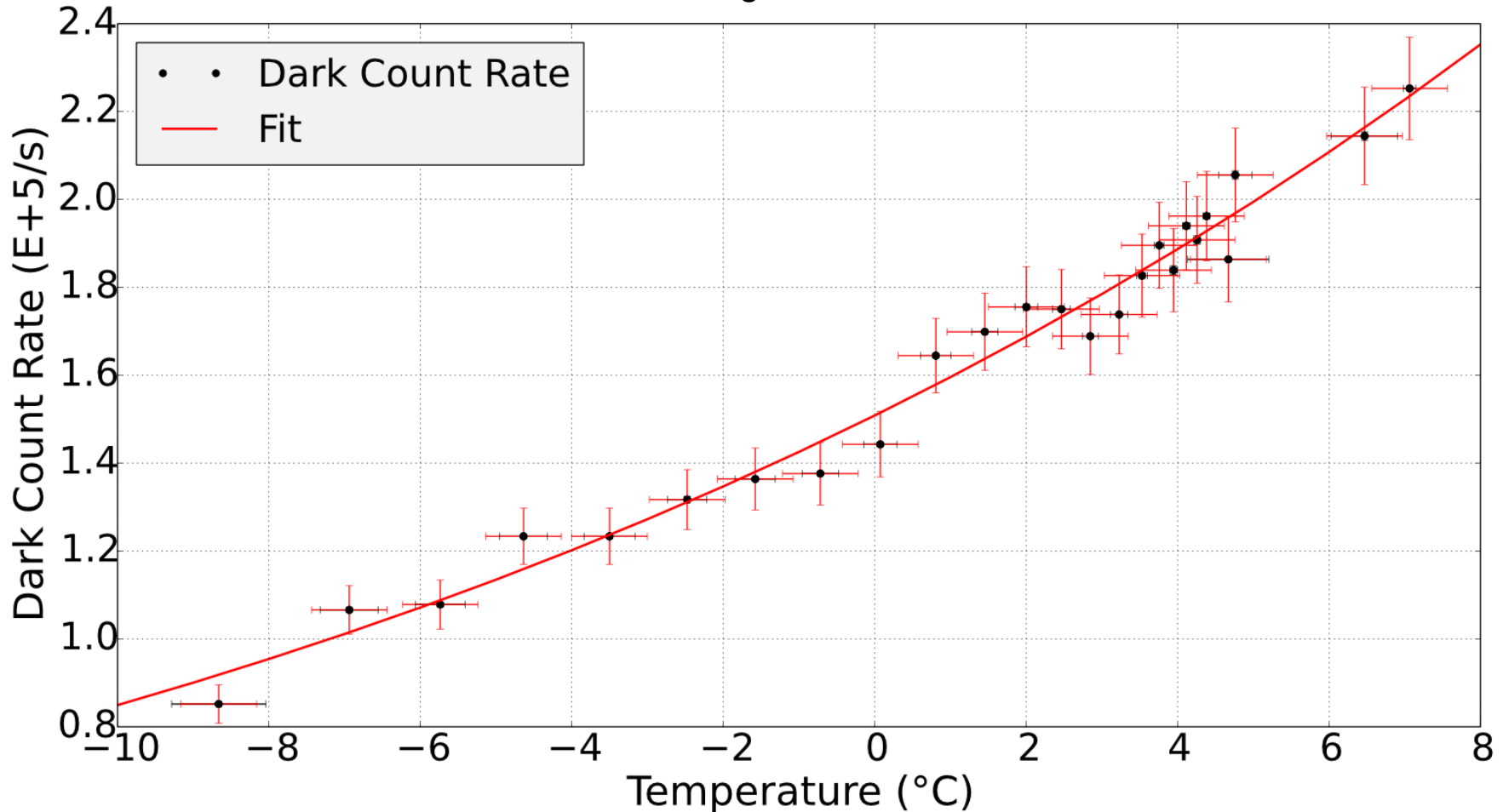
S13361 / 1 channel, constant bias voltage: 55.2 V,  $\lambda = 423\text{nm}$



$$\text{Fit: } G(T) = -0.03 \frac{10^6}{^\circ\text{C}} * T + 2.65 * 10^6$$

# Temperature dependent dark count rate

S13361 / 1 channel, constant bias voltage: 55.2 V,  $\lambda = 423\text{nm}$

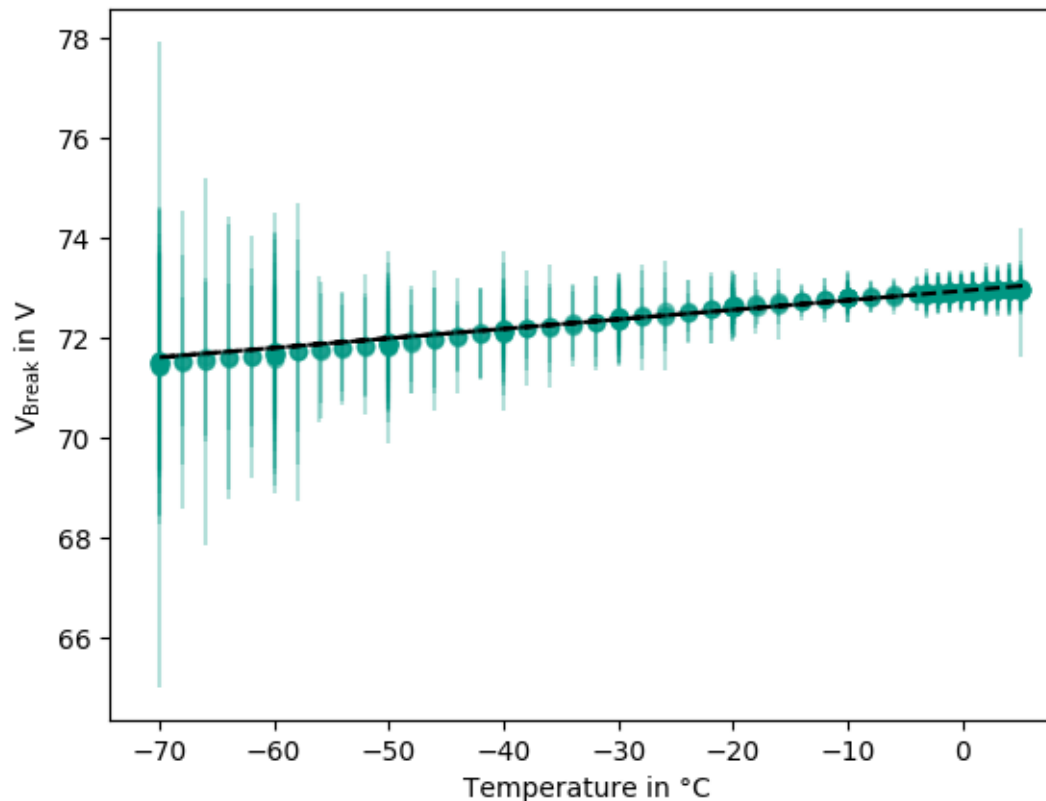


Fit:  $\xi(T) = 4.55 * 10^{-33} \frac{\text{Hz}}{\text{K}} * T^{15.4}$



# Temperature dependent measurements with S13360-6025

- First Measurements with SiPM S13360-6025 with analog read-out placed in temperature chamber (-70°C – 120°C )
- Measurements of the “breakdown voltage” by recording U-I-curves in darkness
- Additional measurements planned



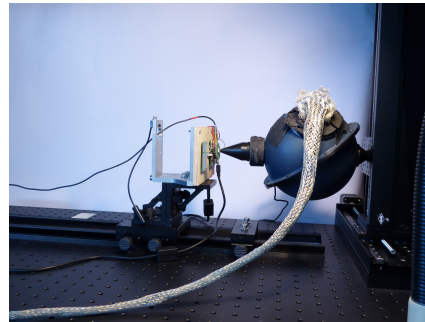
Slope: 19 mV/K

# Conclusion and Outlook

## SPOCK photosensor calibration setup

SiPM characterisation measurement stand SPOCK at KIT able to make a full SiPM characterization ( One channel SiPMs and SiPM arrays)

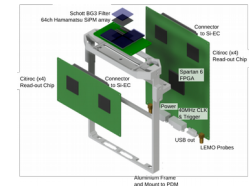
- Ongoing:
- Measuring the uniformity of single channel SiPMs
  - Characterization and comparison of SiPMs from different companies



## SiECA fluorescence camera

SiECA V 2.0 in development

Plan:  
Place it next to the PMTs of an Auger fluorescence telescope



## IceCube-Gen 2 scintillator surface extension

-7 detectors and IceTAXI DAQ on their way to the south pole. Deployment this season.

-7 detectors in Madison for low temperature test with a large cooling chamber

-7 detectors at KIT for cross calibration

-1 detector at DESY for signal processing development with IceTAXI



## SiPM Temperature tests

“Cold SPOCK” in a temperature chamber in development:

SiPM characterization to temperatures down to  $-70^{\circ}\text{C}$



# Backup

# Characterizing SiPMs - PDE

Photo Detection Efficiency: 
$$PDE = \frac{N_{pe}}{N}$$

$N_{pe}$  : Number of detected photoelectrons per pulse

$N$  : Number of incident photons per pulse

Number of incident photons per pulse:

$$N = \frac{P \cdot R \cdot R_{geom}}{E_{photon} \cdot f_{pulse}}$$

Number of detected photoelectrons per pulse:

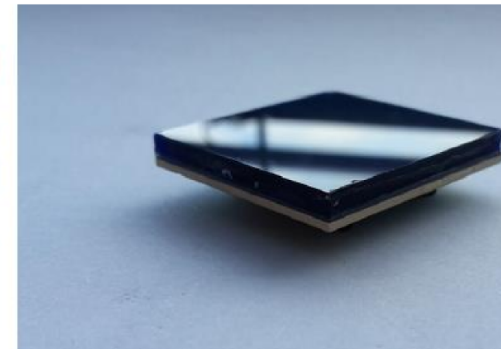
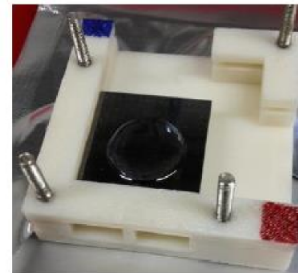
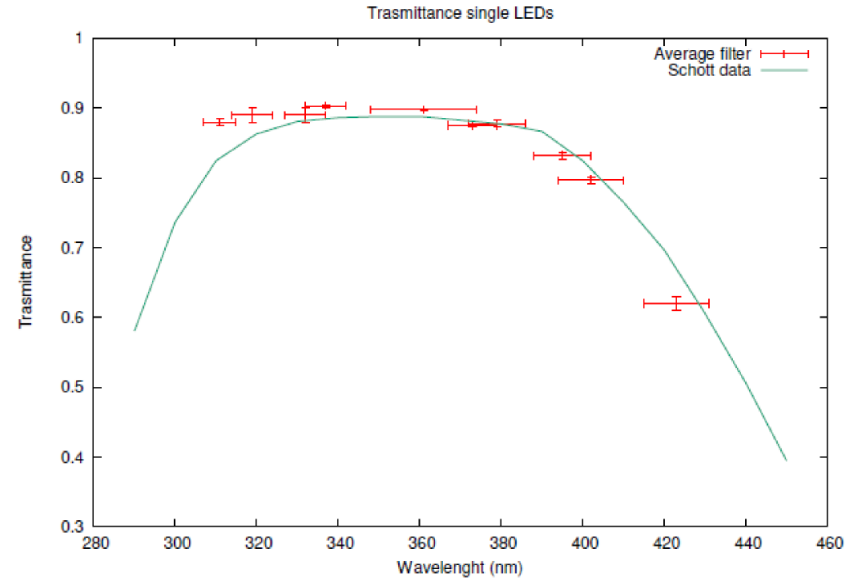
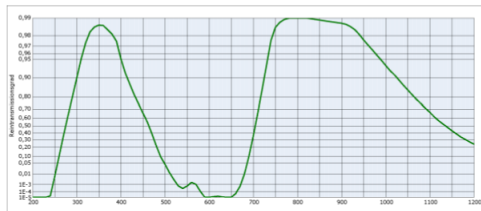
$$N_{pe} = \ln \left( \frac{N_{tot}}{N_{ped}} \right) - \ln \left( \frac{N_{tot}^{dark}}{N_{ped}^{dark}} \right)$$

*Characterisation Studies of Silicon Photomultipliers; ArXiv:1003.6071v2*

# Schott BG3 filter glass SiPM covering



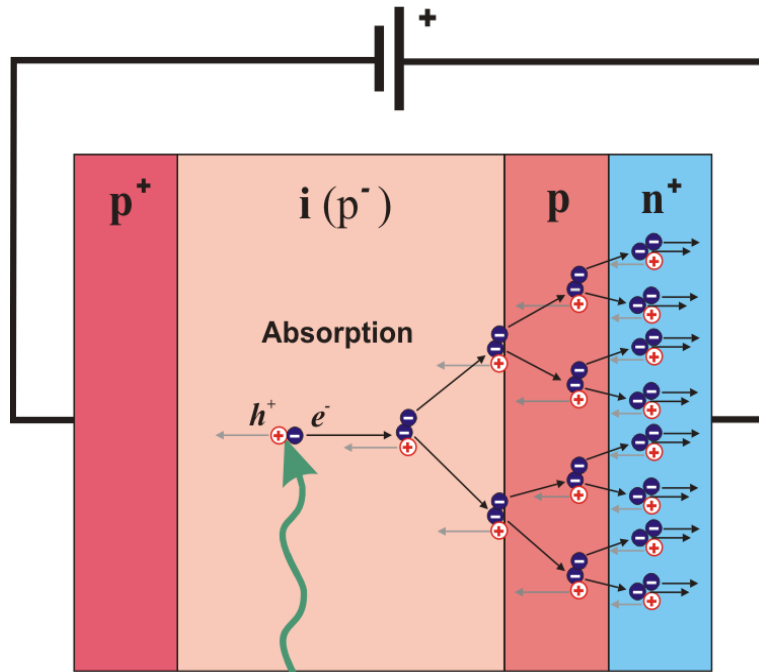
Schott BG3 29.32mm x 2mm thick  
(unframed glass circle only)



**Alberto Bertone, Univ Torino, Erasmus@KIT**

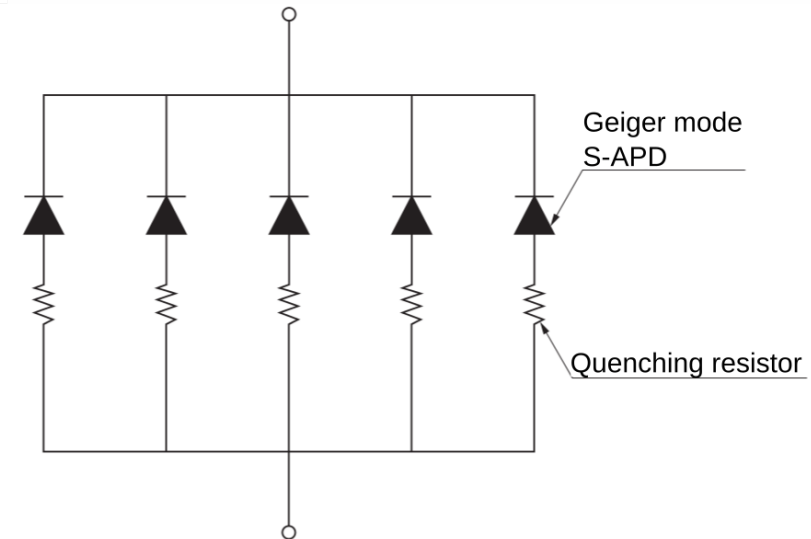
# Silicon Photomultiplier

## Avalanche Photo Diode













- Photon enters the APD and produces Electron/Hole pair
- Electrons are accelerated to n<sup>+</sup>-layer and produce avalanche

## Silicon Photomultiplier



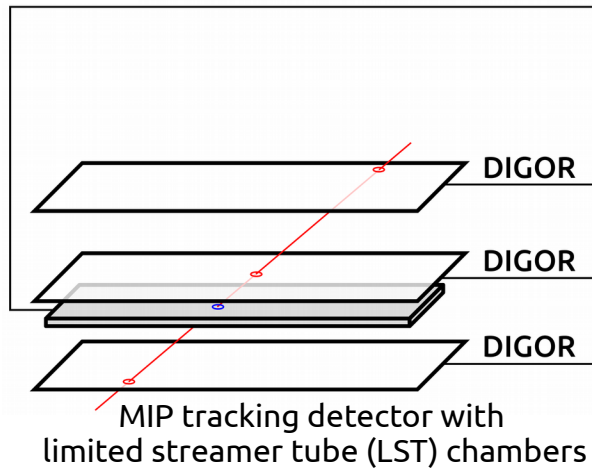
- Array of parallel connected APDs form a SiPM

# Comparison PMT ↔ SiPM

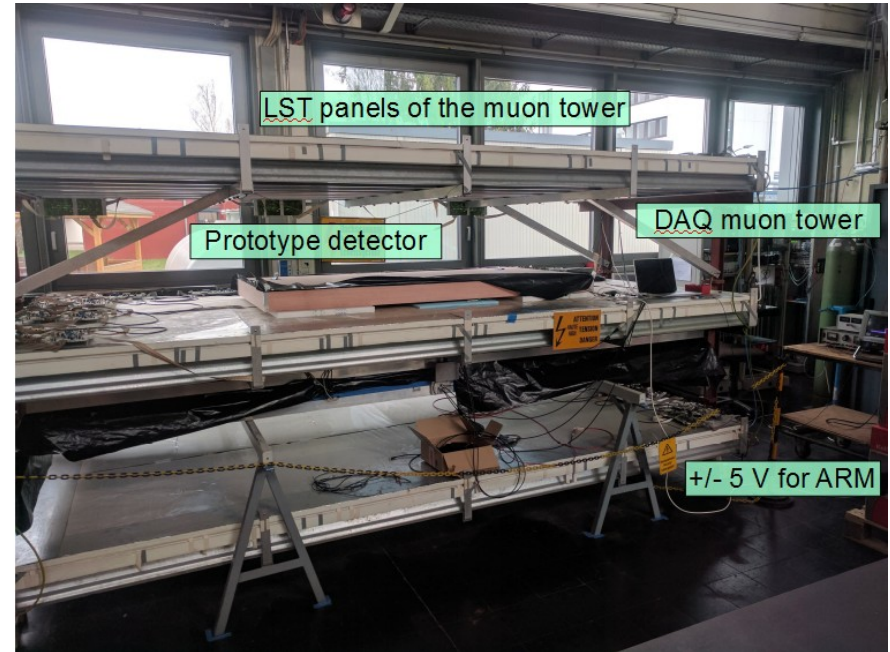
	PMT	SiPM
PDE	20-40%	20-60%
Gain	$10^6$	$10^6$
Dark noise rate	~Hz 	~MHz 
Behaviour in magnetic fields		
Operation Voltage	1000+ V 	50-70 V 
Temperature sensitivity		
Robustness and compactness		

# Muon Tower

## Prototype detector



Spatial resolution possible!  
(1cm x 1cm, later)

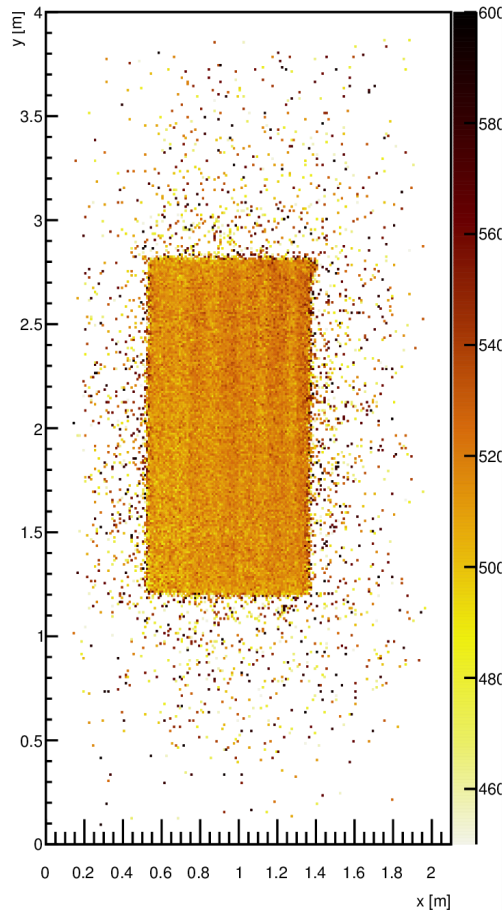


Muon tower with the prototype detector

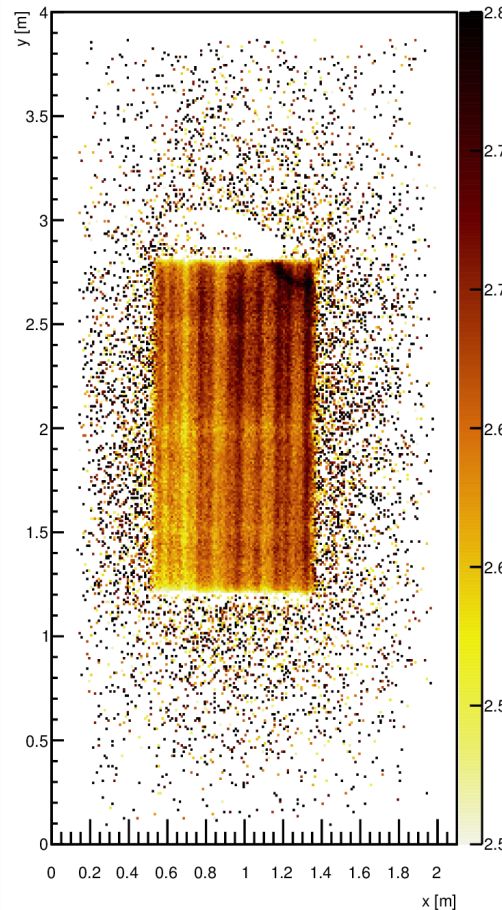


# Uniformity

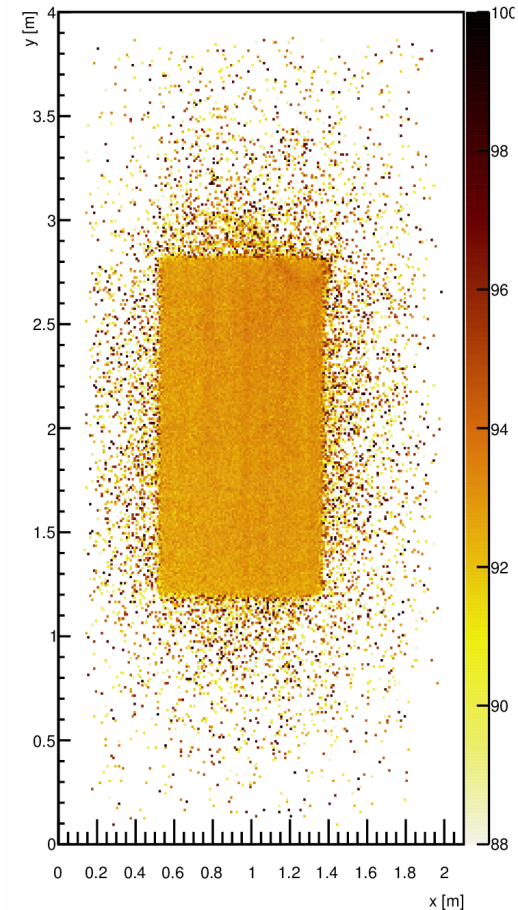
<VE charge>



<lg(VE charge)>

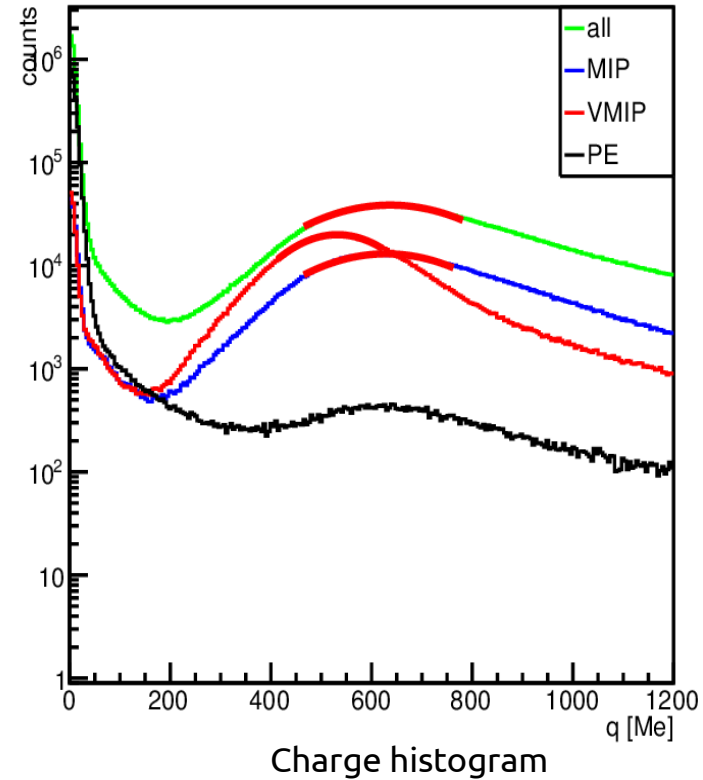
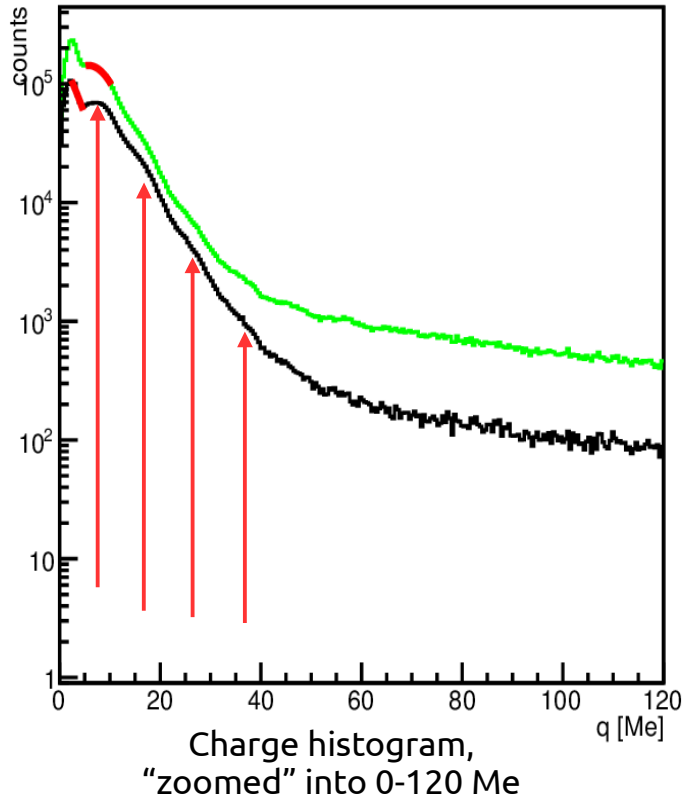


<pulse width>



Left: Average pulse charge depending on the particle detection position  
Center: In logarithmic scale. Single optical fibers are visible!  
Right: Average pulse width

# MIPs



**PE:  $25/3 = 8.3$  Me**  
**(mid gain 5x  $\rightarrow$  1.6 Me low gain)**

**MIP/PE:  $630/8.3 = 75$  photons**

# Temperature dependent measurements with S13361-3050

- SiPM array S13361-3050AS cooled down in isolating cooling box together with heat reservoirs
- Measurements made in SPOCK as the system returned to ambient room temperature
- Measurement of one channel out of the 64 channels in the array