

The Max-Planck Advanced Study Group

within the Center of Free Electron Laser Science

***High Speed, large format
imagers for X-rays from
50 eV up to 30 keV***



***for experiments in astrophysics
at X-ray FELs and in industry***

MPE-HLL projects

1. X-ray astronomy

- o XMM since 2000 'til now
- o eROSITA 2012
- o BepiColombo 2014
- o ASTRO-H 2015
- o IXO 2022 + x

2. Center for Free Electron Laser Science

- o FLASH 2008
- o BESSY 2008
- o LCLS/SCSS 2009/2010
- o XFEL 2013

3. X-ray Fluorescence Physics

- o Spirit and Opportunity since 2004 'til now
- o EXOMars 2013
- o SIDDHARTA 2009
- o HICAM and DRAGO 2010
- o CETRA 2014
- o Plasmadiagnostics 2014
- o X-ray microscopy 2012

4. Optical applications

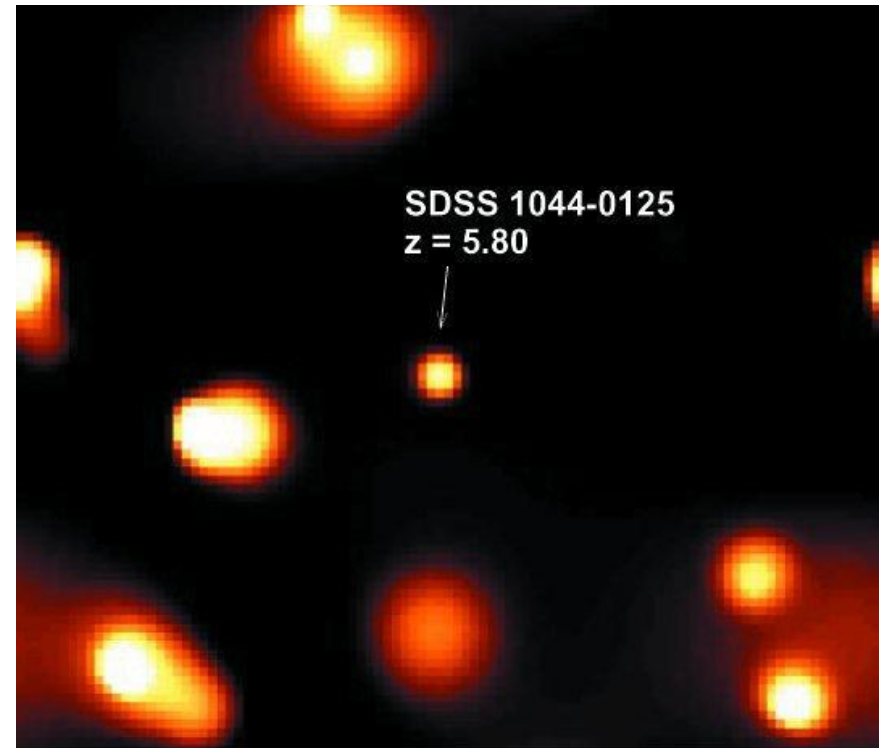
- o LBT, wave front sensing 2012
- o Avalanche pnCCDs 2011
- o RNDR - DePFETs 2015
- o IBC (BIB) detectors 2013
- o Gatable DePFETs 2015

Outline

- X-ray CCDs for
 - eROSITA
 - BESSY, FLASH
 - LCLS, SOSS, FERMI
- DePFETs for
 - BepiColombo
 - IXO
 - XFEL
- Future plans

XMM Summary

- Working since launch (10. Dez. 1999) without any problem.
- The energy resolution @ the Al_K line (1.5keV) decreased since launch from 98 eV to 99 eV (FWHM).
- Since launch the operating conditions have never been changed.
- Up to now more than **15.000** observations were made with XMM – Newton. **80 %** with the pnCCD as prime instrument
- Up to now, > **2.600** refereed astrophysics publications have been made
- pnCCD paper was cited **1.000** times in refereed journals



QSO SDSS 1044-0125

European Space Agency 

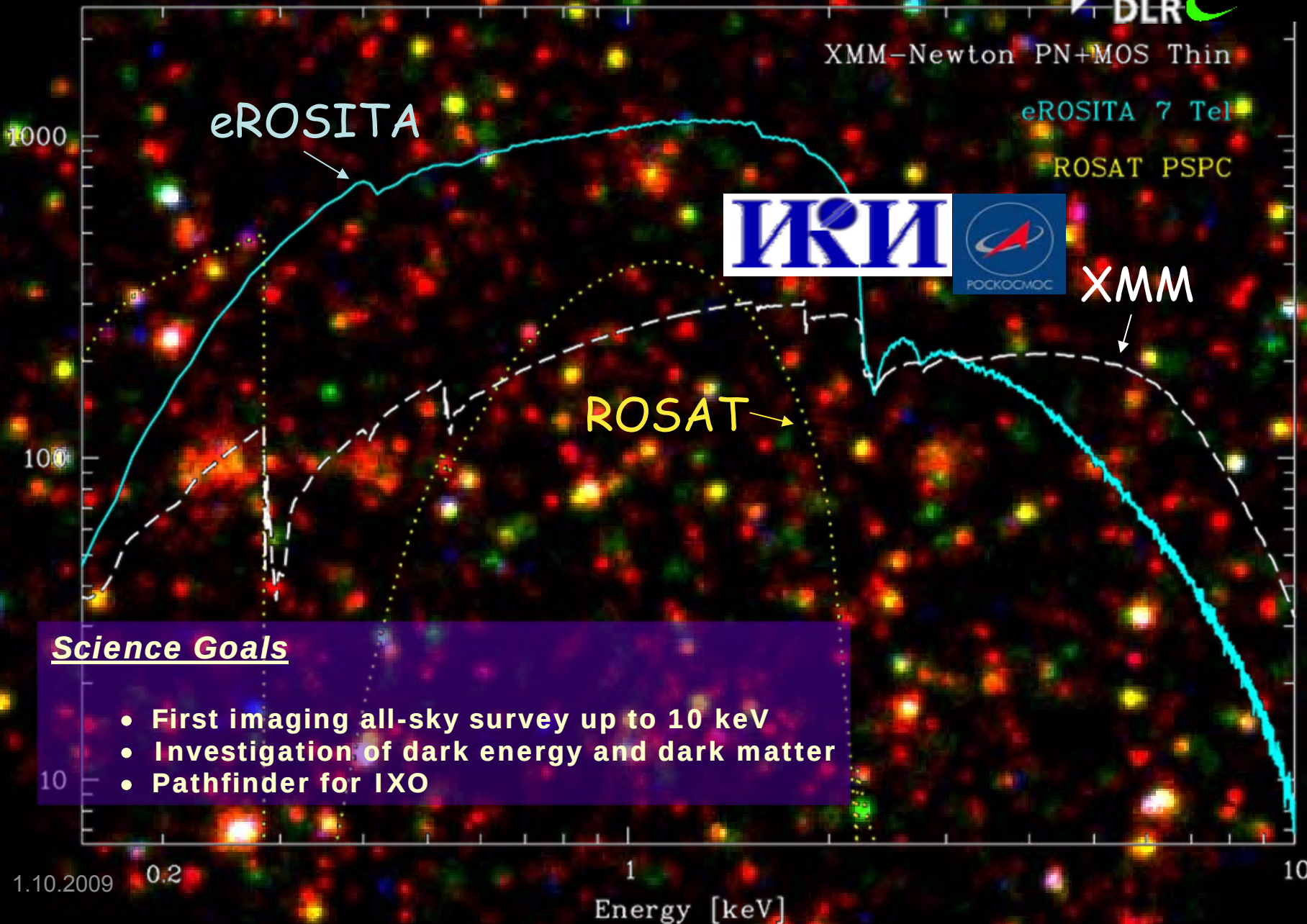
Very high redshift QSO observed by XMM EPIC

eROSITA

Dark Universe Consortium



Grasp: effective area * FOV [cm**2 deg**2]



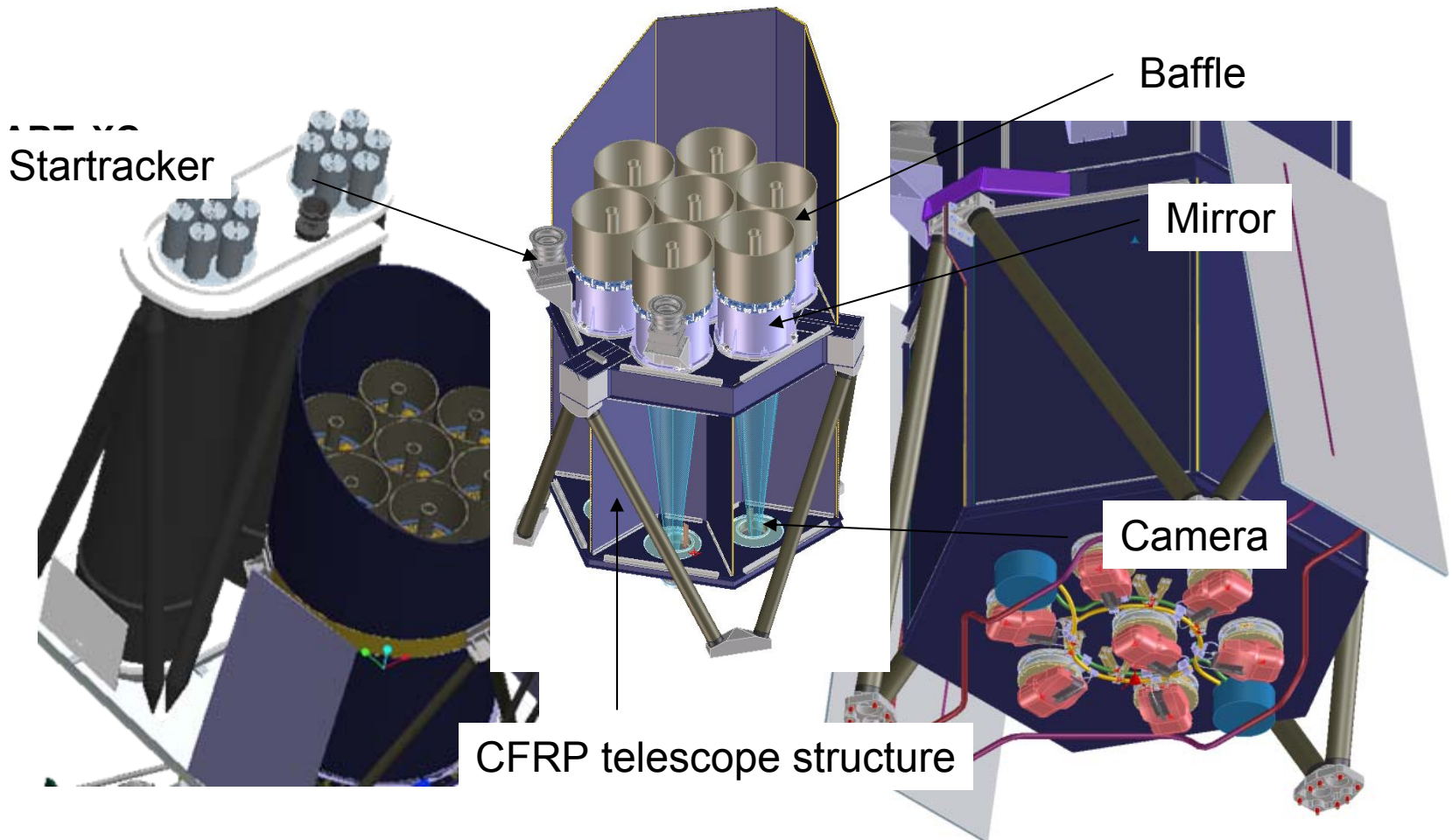
Science Goals

- First imaging all-sky survey up to 10 keV
- Investigation of dark energy and dark matter
- Pathfinder for IXO

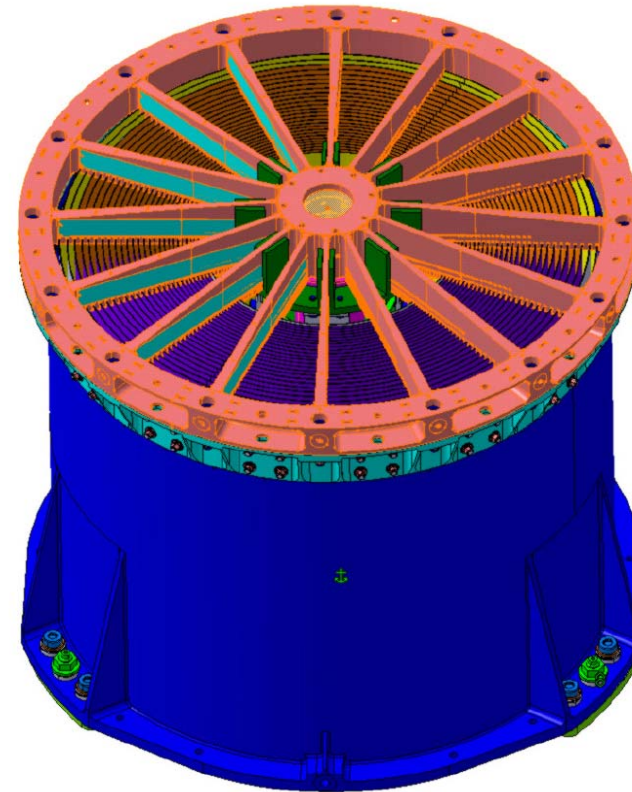
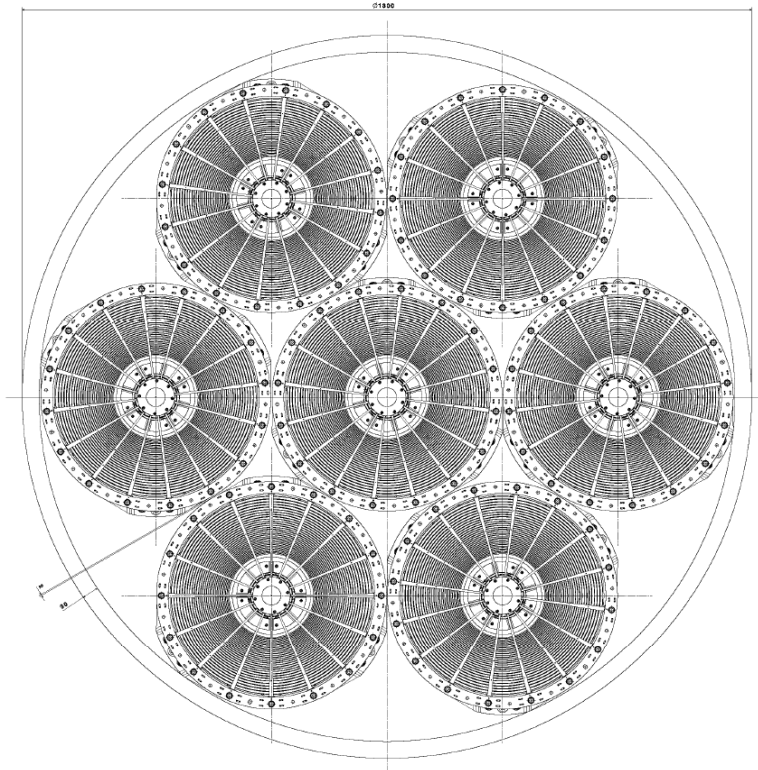
1.10.2009

Telescope Elements

- eROSITA: 7 „identical **mirror systems** and **pnCCD cameras**
- ART-XC: 2 X-ray optic Array of 7 telescopes **eras**

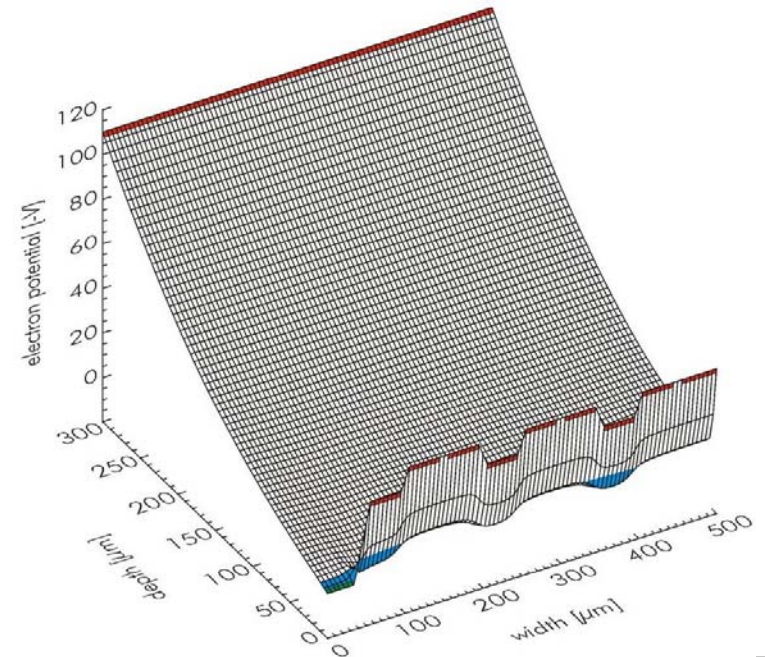
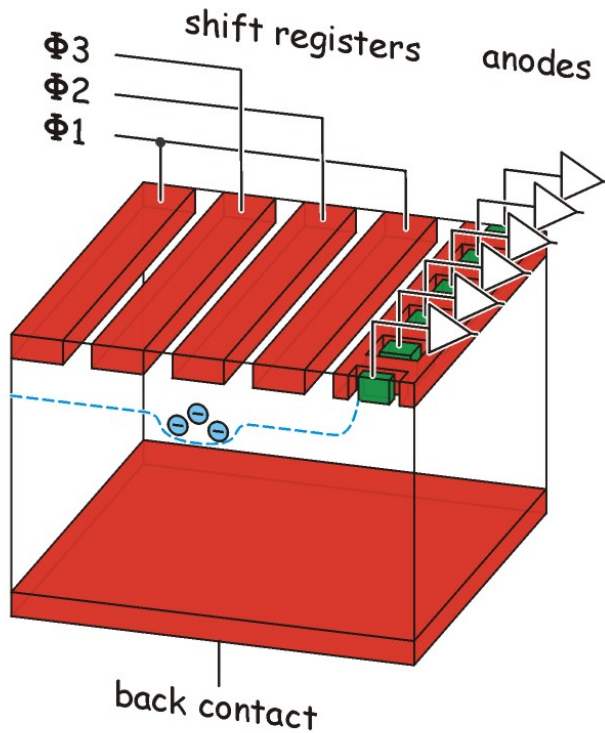


Mirror System

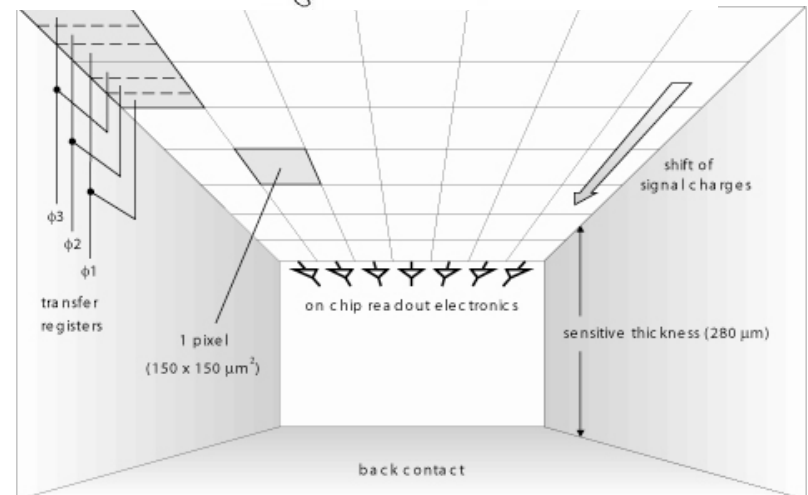


- **7 Mirror Modules, 54 shells each, 360mm Ø, f=1.600mm**
- **inner 27 shells to be replicated from old ABRIXAS mandrels**
- **3 outer mandrels already manufactured (#1, #2, #27)**

CCD basics



- full depletion (50 μm to 500 μm)
- back side illumination
- radiation hardness
- high readout speed
- pixel sizes from 36 μm to 650 μm
- charge handling: more than 10^6 e⁻/pixel
- high quantum efficiency



How many charges can be stored in one pixel ?

pixel volume:

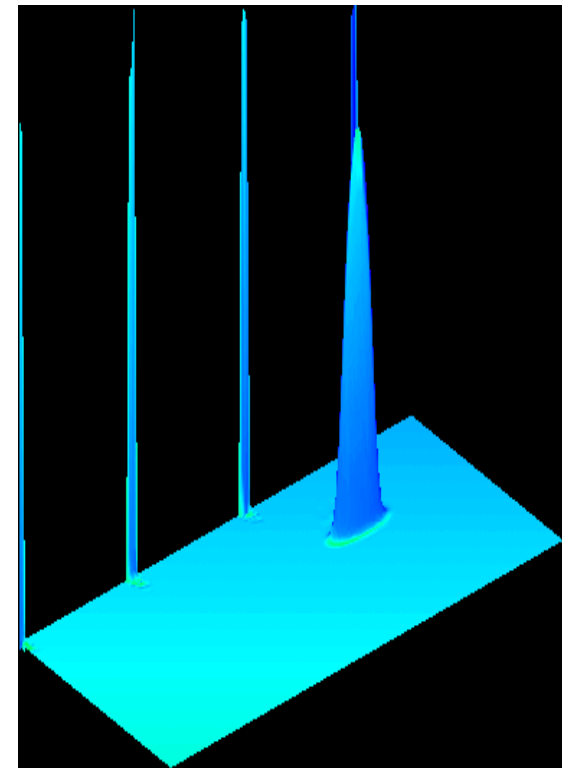
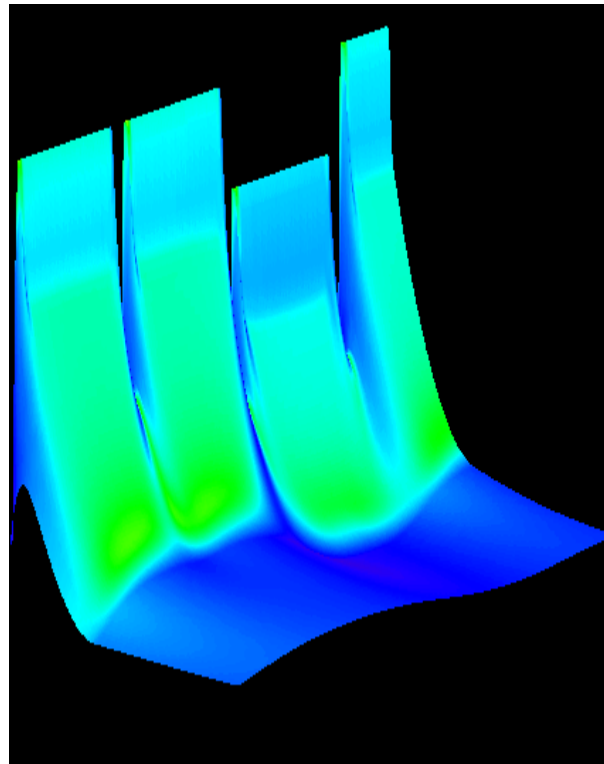
$$20 \times 40 \times 12 \text{ } \mu\text{m}^3 \approx 1 \times 10^4 \mu\text{m}^3$$

Doping: $10^2 \text{ P per } \mu\text{m}^3$

$$\text{CHC} = 3-8 \times 10^5 \text{ per pixel}$$

can be increased by
external voltages

can be increased by doping



The charge spread function

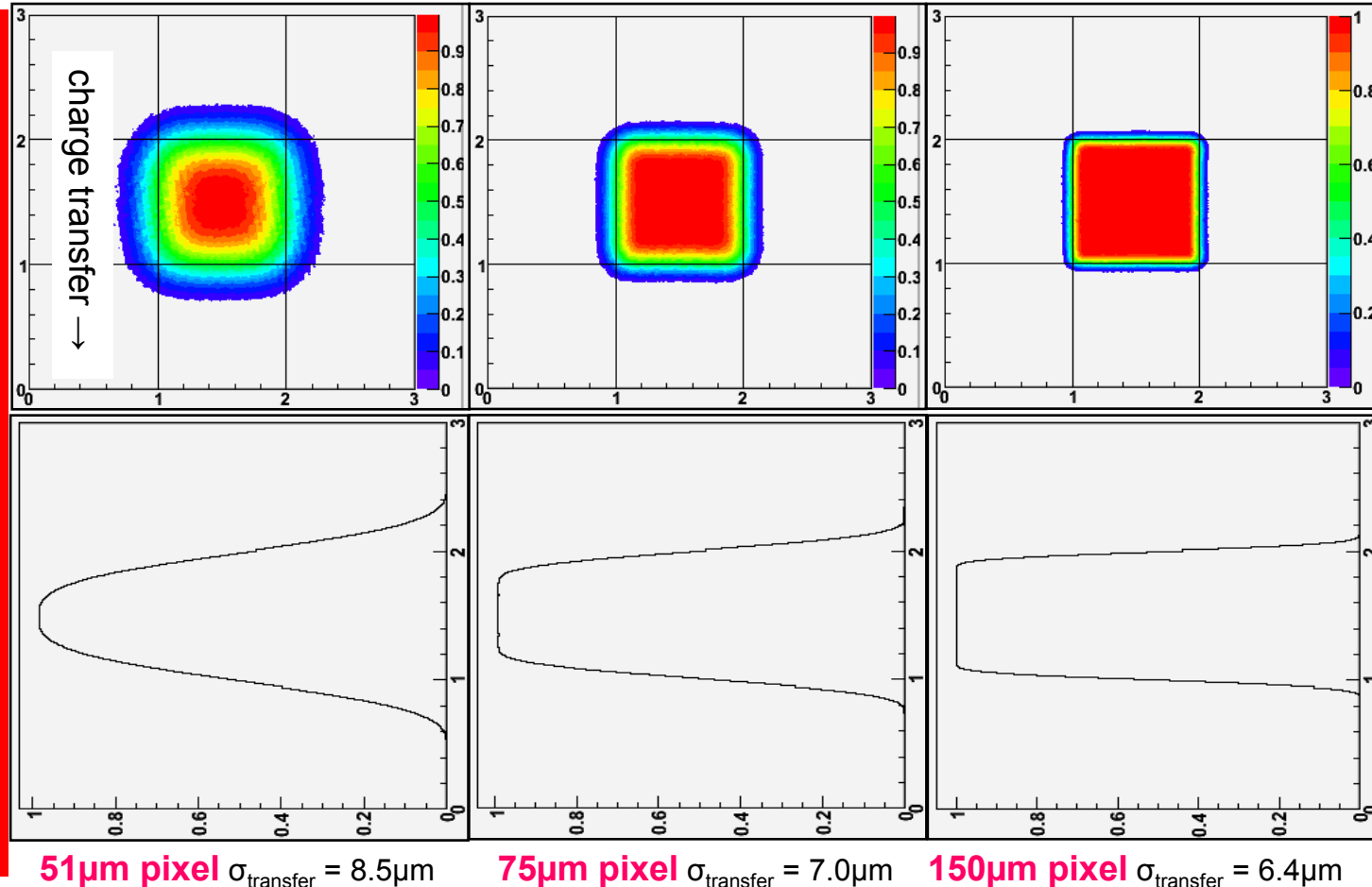
Injected charge:
4.5 keV, i.e.
1.250 electrons



At 1000 X-rays
of 12 keV the
of electrons
is 3.600.000

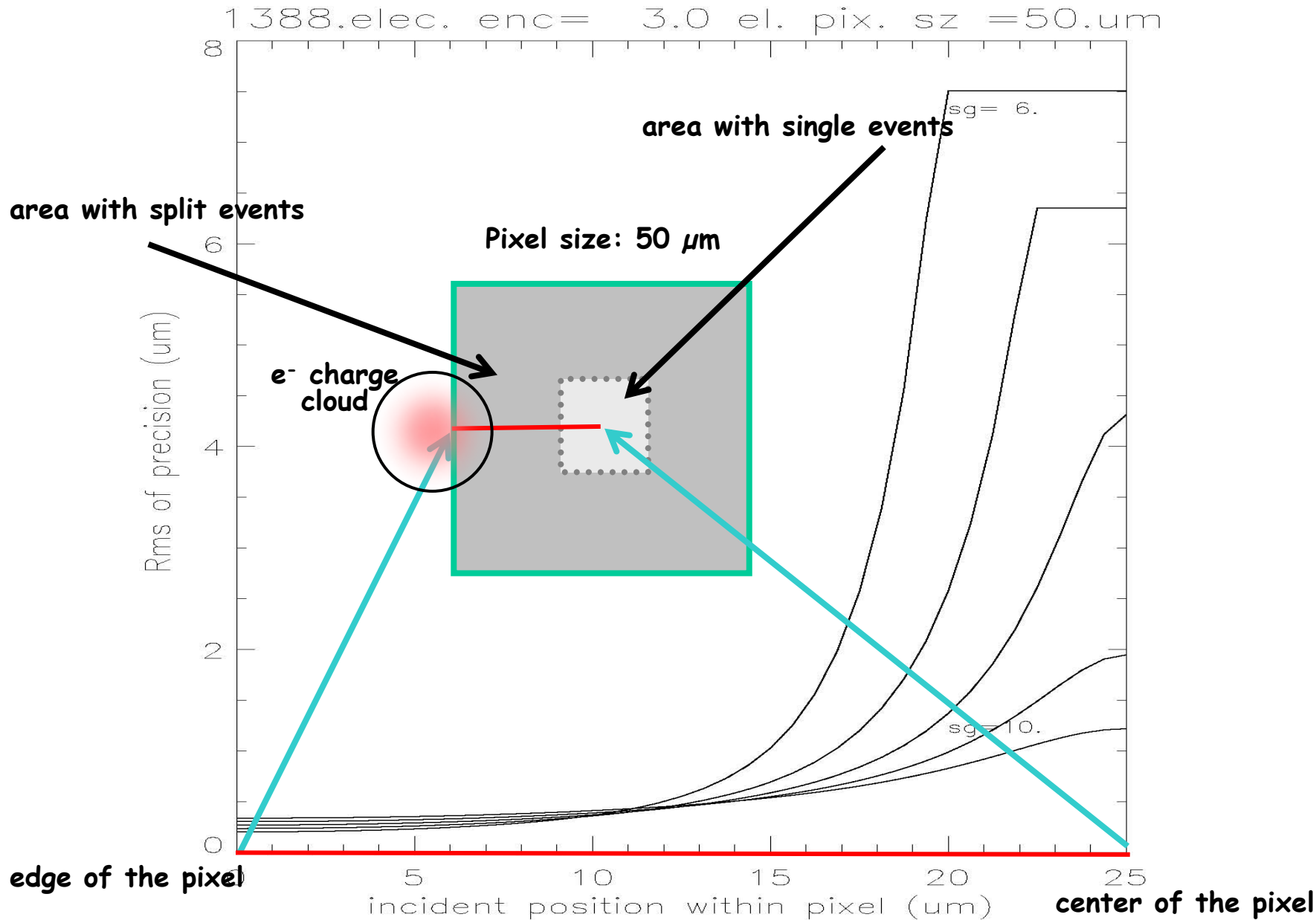
The charge spread
depends on:

- charge collection time
- pixel size
- number of charges
- operating temp.

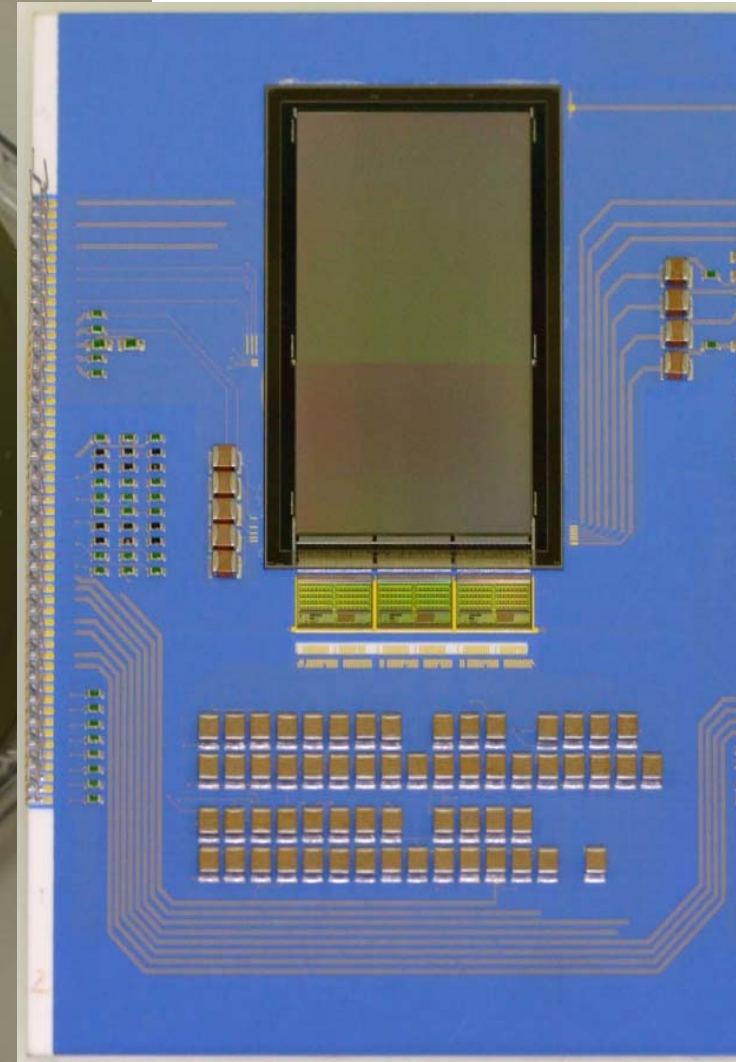
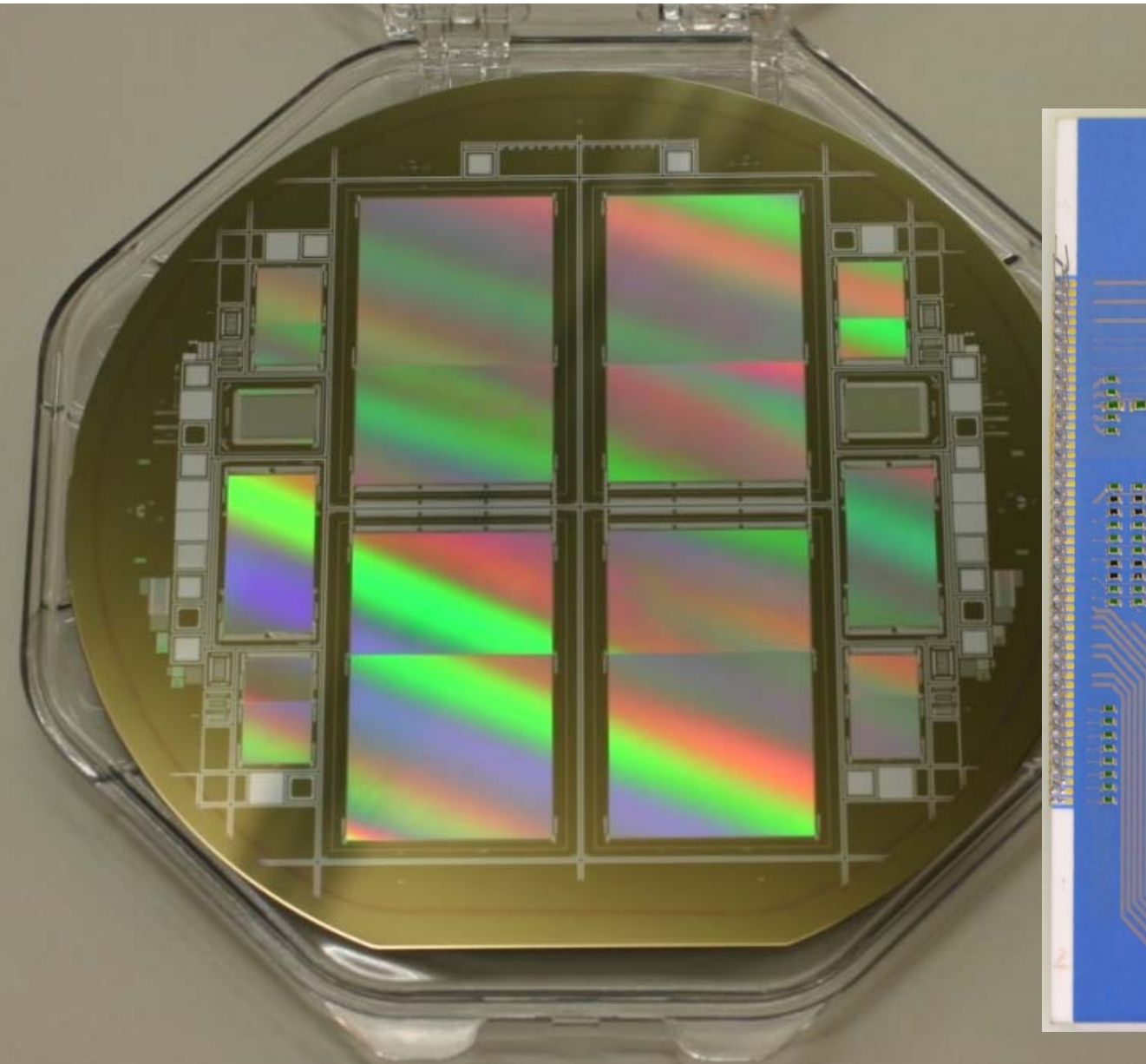


For more than 1.500.000 signal charges a pixel size of less than 150 μm is not adequate

Single photon counting position precision



II. eROSITA detector design



eROSITA camera concept

eROSITA (extended **RO**entgen **S**urvey with an **I**maging **T**elescope **A**rray) **X-ray telescope**

aboard **Spectrum-RG** satellite

(launch: 2012, L2 orbit)

7 identical eROSITA cameras: imaging + spectroscopy

0.3 - 10 keV

(identical cameras for ART-XC: 2-11 keV)

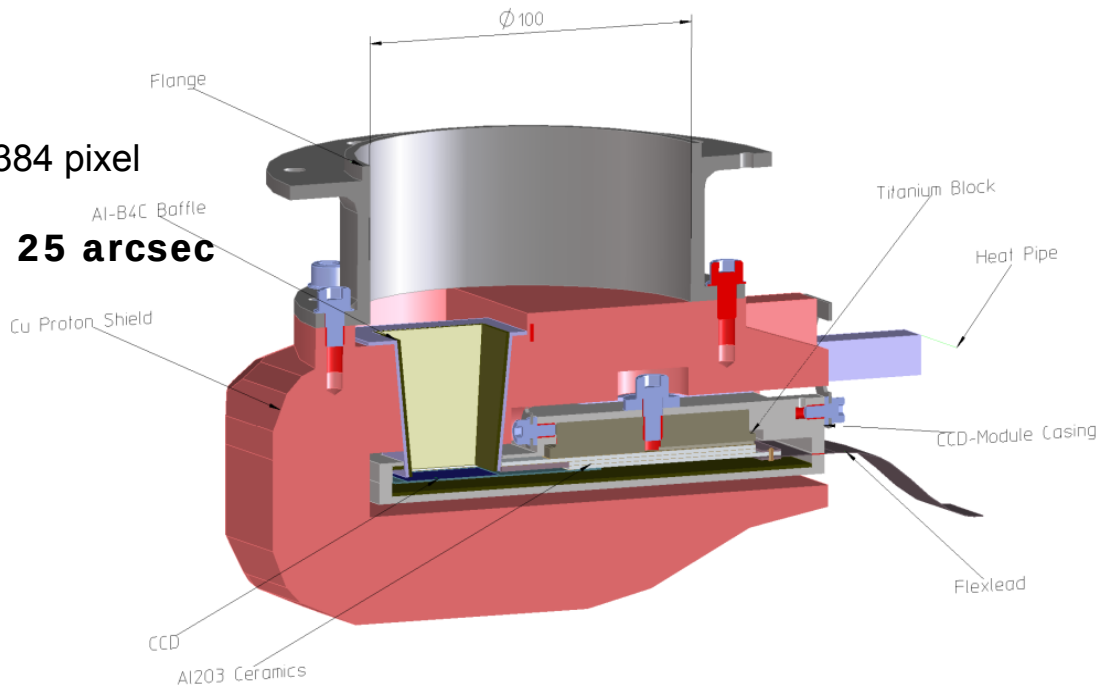
FoV: **1.0 degree** = 28.8 mm x 28.8 mm
= 75 x 75 μm^2 size $\rightarrow$ 384 x 384 pixel

expected angular resolution:

25 arcsec

one pixel of 75 x 75 μm^2 : 9 arcsec

**position resolution of the
camera for 1 - 10 keV: 4 arcsec**



First measurements with eROSITA CCDs: Intensity image

- **Performance measurement:**

eROSITA CCD **C12-03-42**

(384 x 384)

frame transfer mode

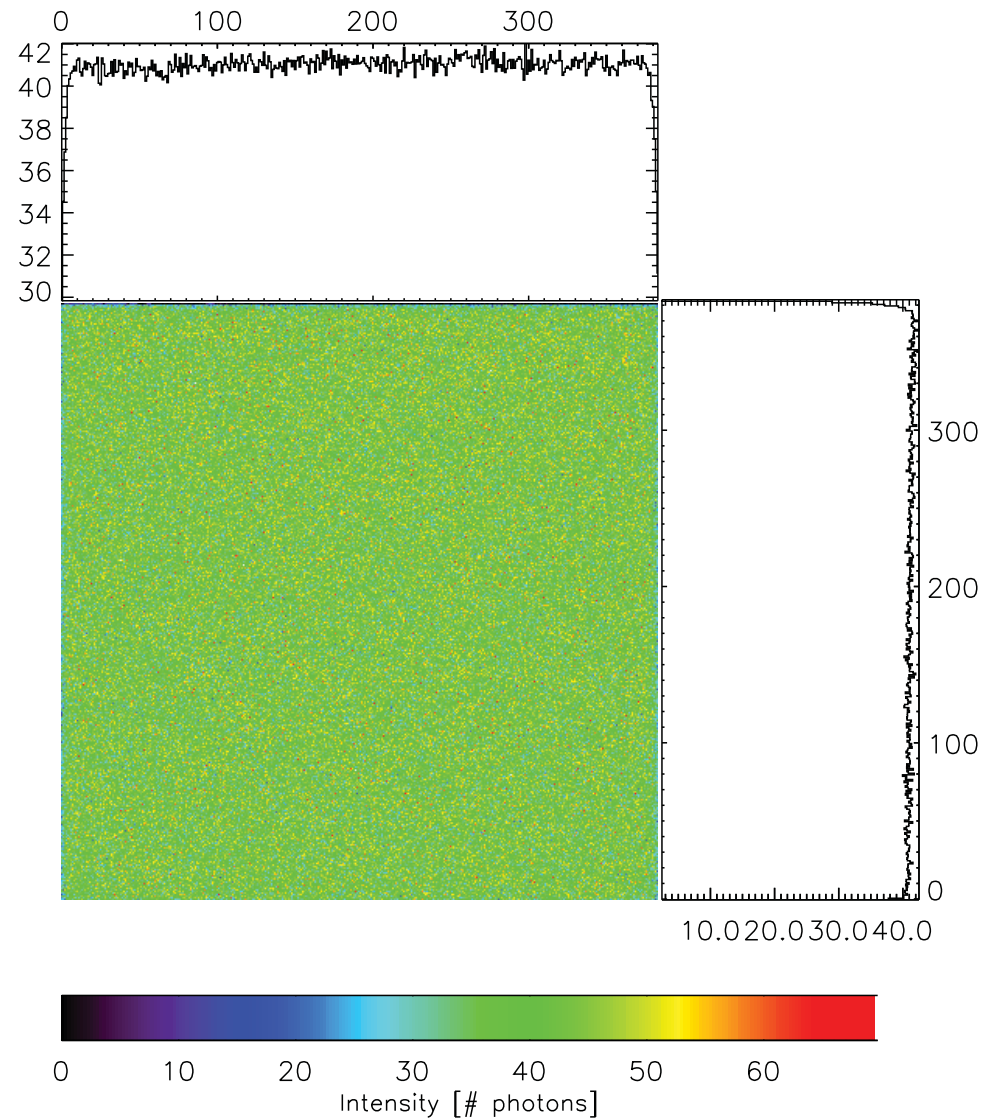
T = -80°C

20 images / s

Results:

No pixel defects / 14 cm²

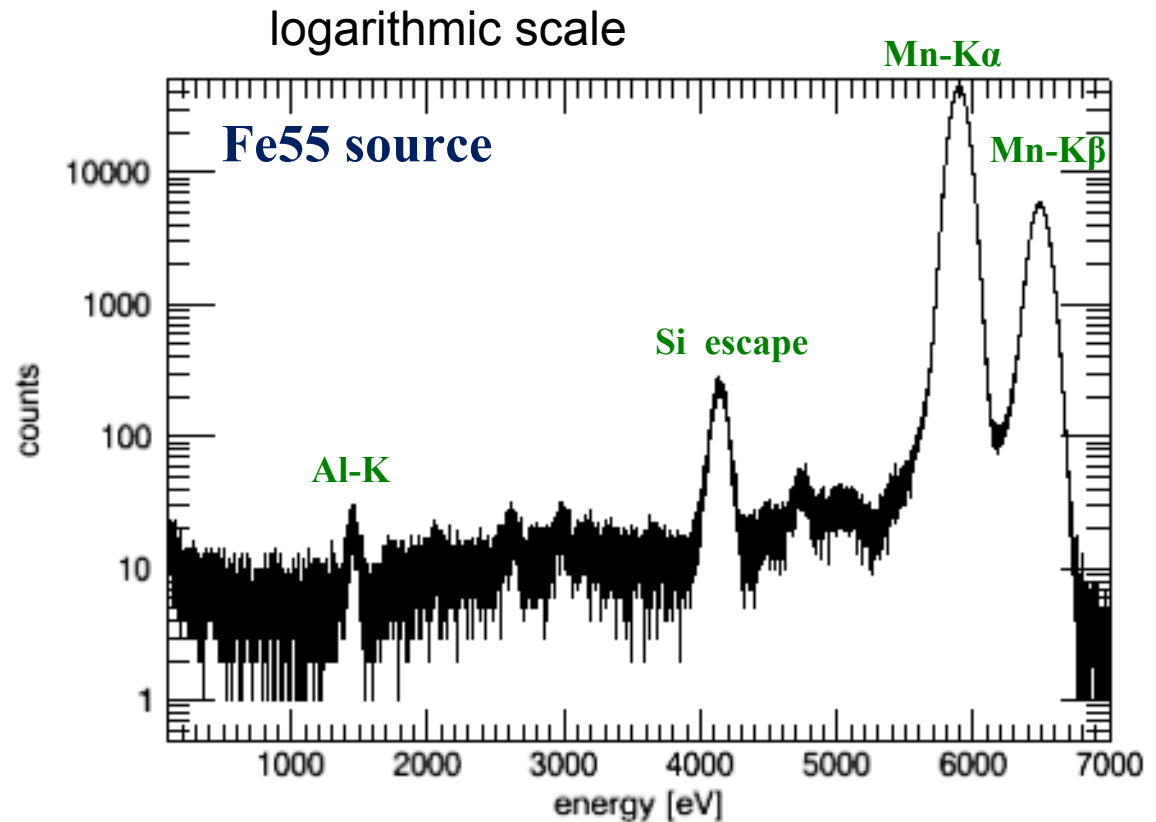
Noise: **2.4 e⁻ ENC**



First measurements with eROSITA CCDs: Spectroscopic images

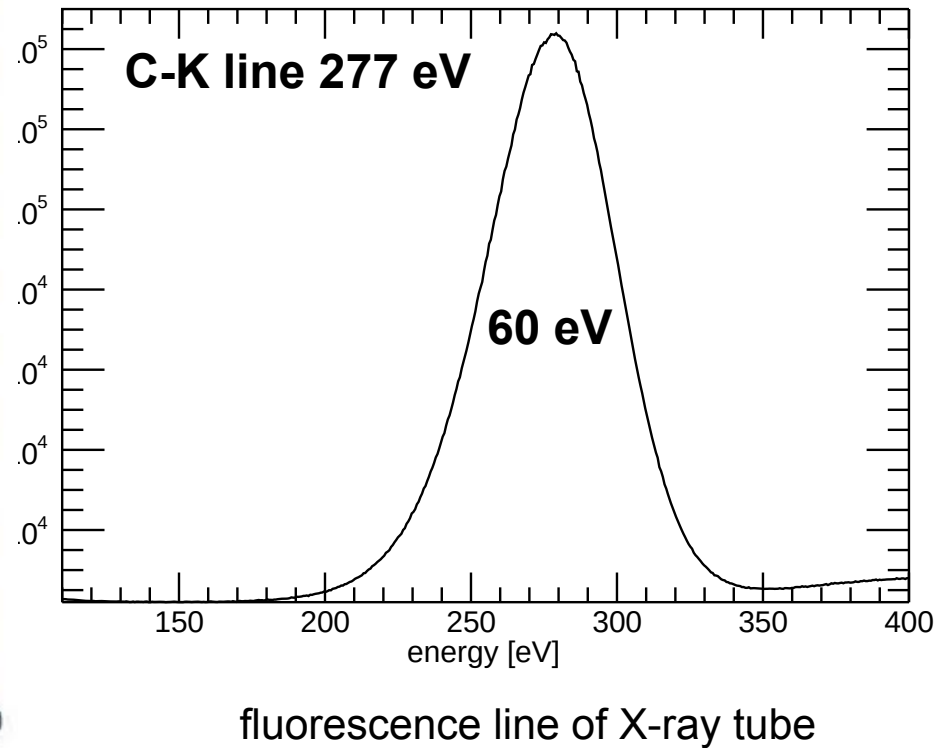
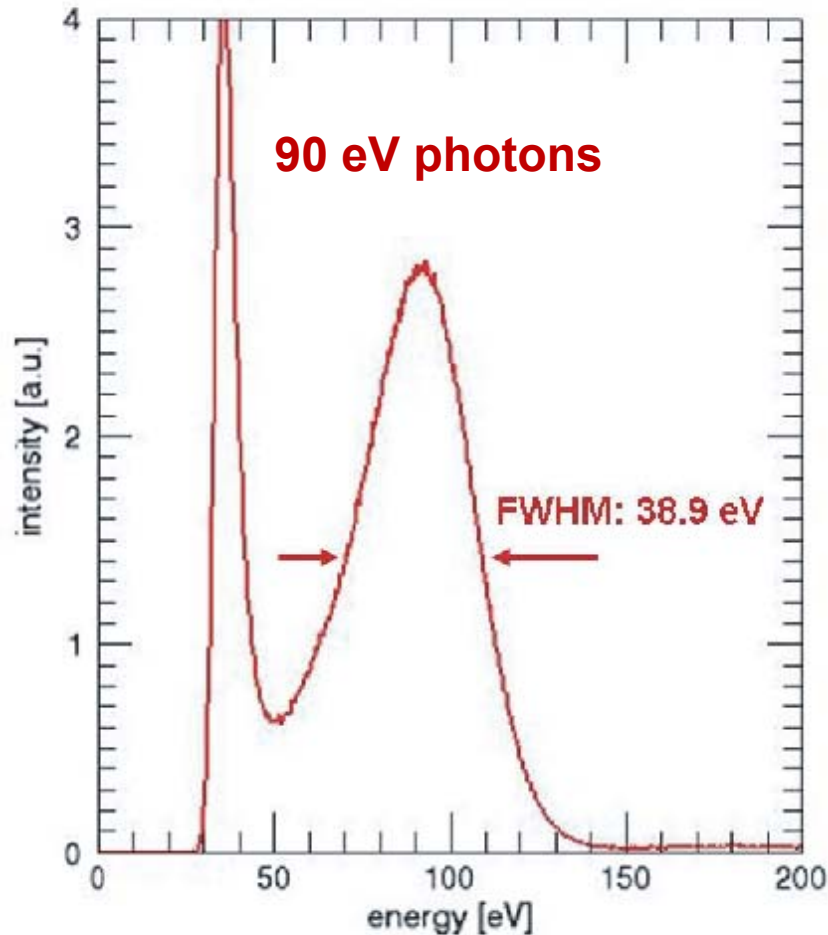
- Performance measurement:

Peak to valley: 7700 : 1
(photon entrance window
equipped with light filter)



First measurements with eROSITA CCDs:

Low energy performance



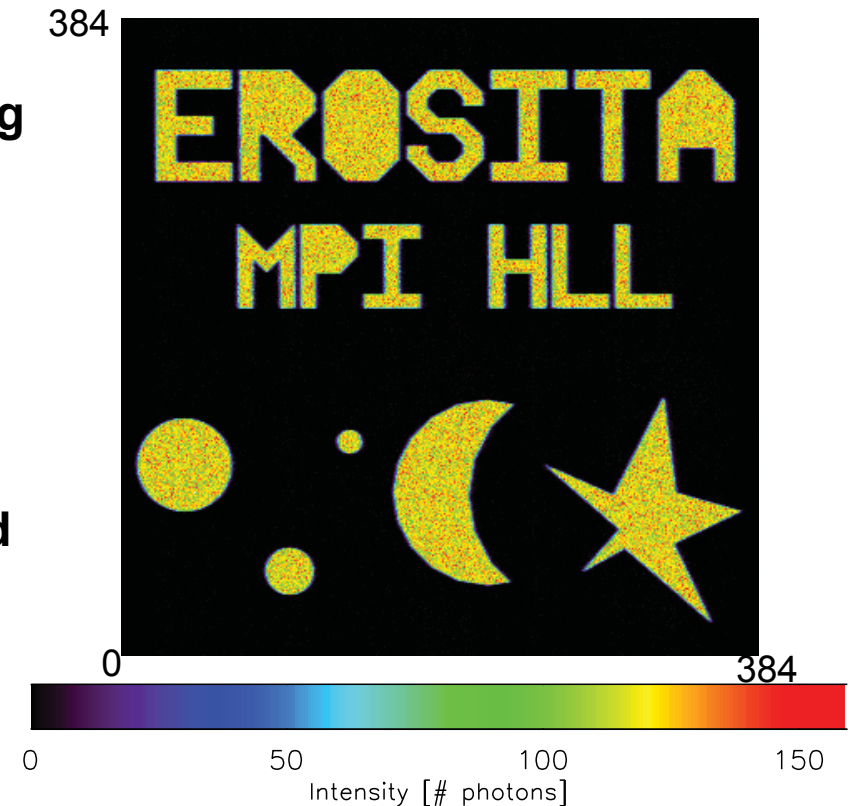
eROSITA is on track for launch in 2012

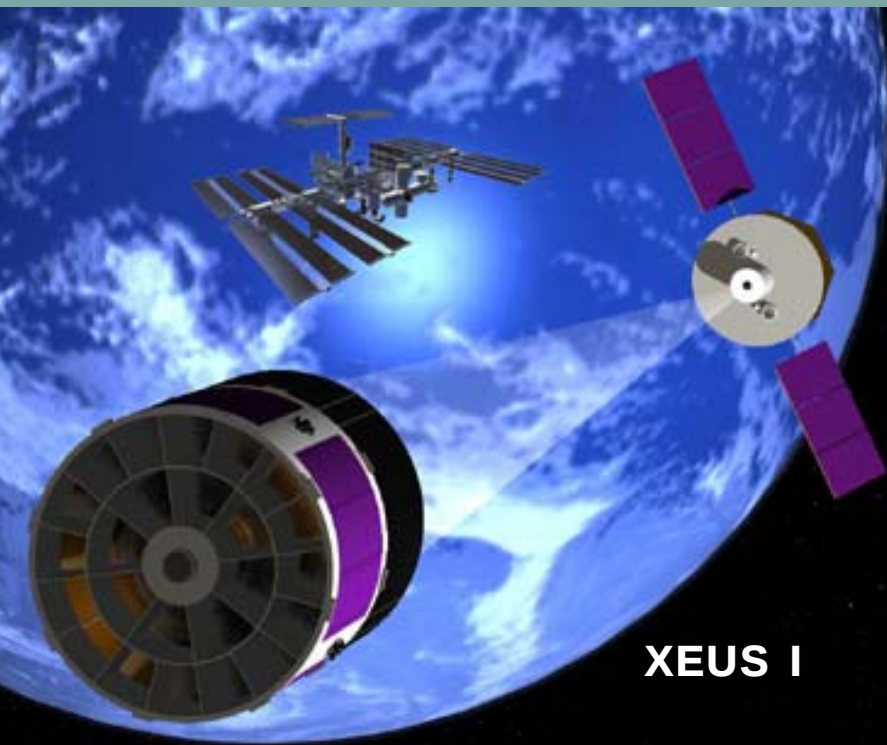
Status of camera components is according to schedule

- detailed design + testing of prototype
- PDR on Oct. 21/22 2009

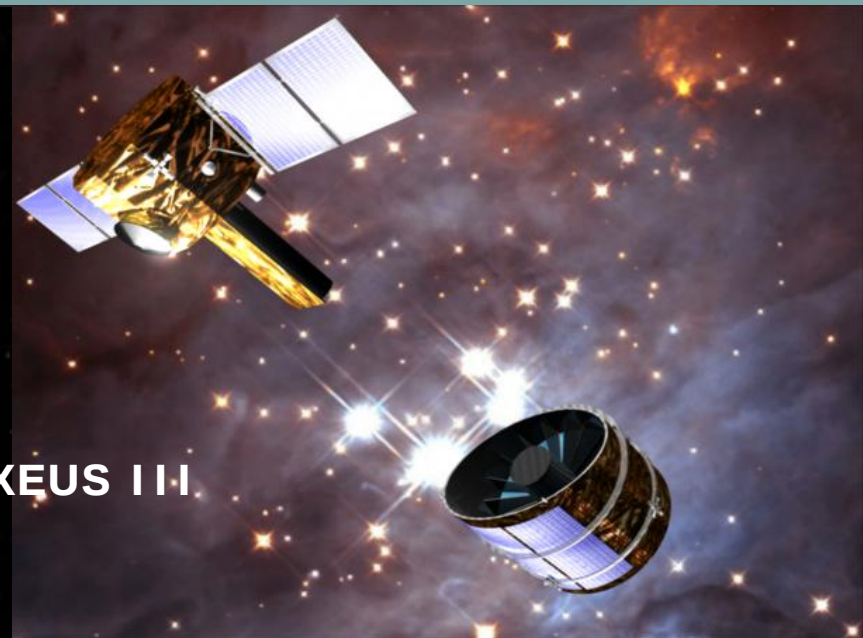
eROSITA flight CCDs produced and tested

- design + production process verified
- performance excellent
- more tests needed





XEUS I



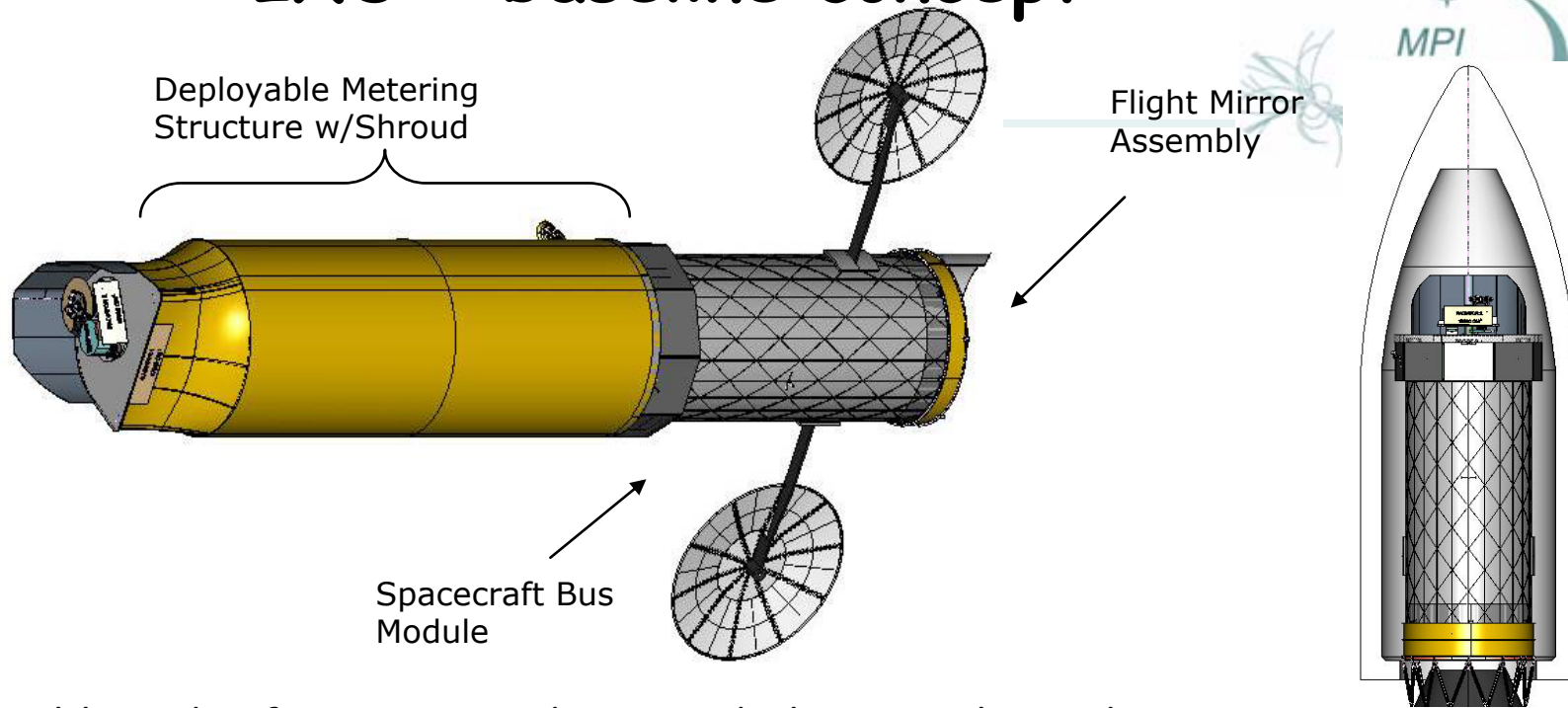
XEUS III

- **History of XEUS (IXO)**
- "constitutional meeting @ LU in ´96
- ESA mission under study 2002
- ESA cosmic vision process 2007
- ESA, JAXA, NASA study 2008
- assessment study, Phase A study:
end 2009
- definition phase study, Phase B1:
end 2011



XEUS II

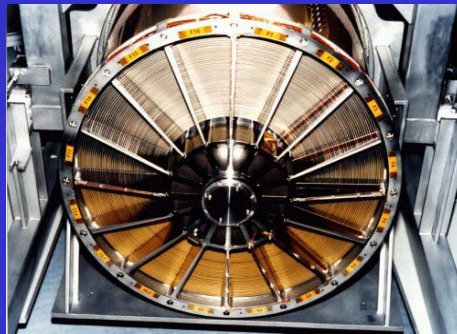
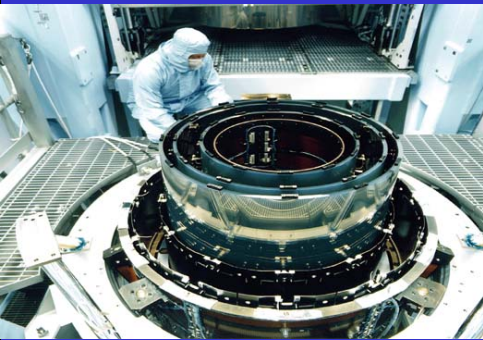
IXO - baseline concept



- Focal length of 20-25m with extendible optical bench
- ◆ Concept compatible with Ariane 5 and Atlas V 551
- ◆ Core instruments to include:
 - Wide Field Imager including Hard X-ray Imager
 - X-ray Micro-calorimeter / Narrow Field Imager
 - X-ray Grating Spectrometer
 - high-time resolution instrument, a polarimeter and a hard X-ray imager are under study

Atlas V 551
Medium
Composite
Fairing

XEUS – Optics Development

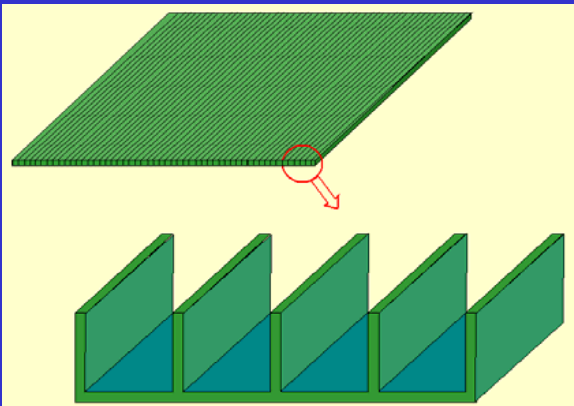


Chandra
0.5"
 18500 kg m^{-2}
 $A_{\text{eff}} @ 1 \text{ keV}$

XMM-Newton
15"
 2300 kg m^{-2}
 $A_{\text{eff}} @ 1 \text{ keV}$

Si-HPO
5"
 200 kg m^{-2}
 $A_{\text{eff}} @ 1 \text{ keV}$

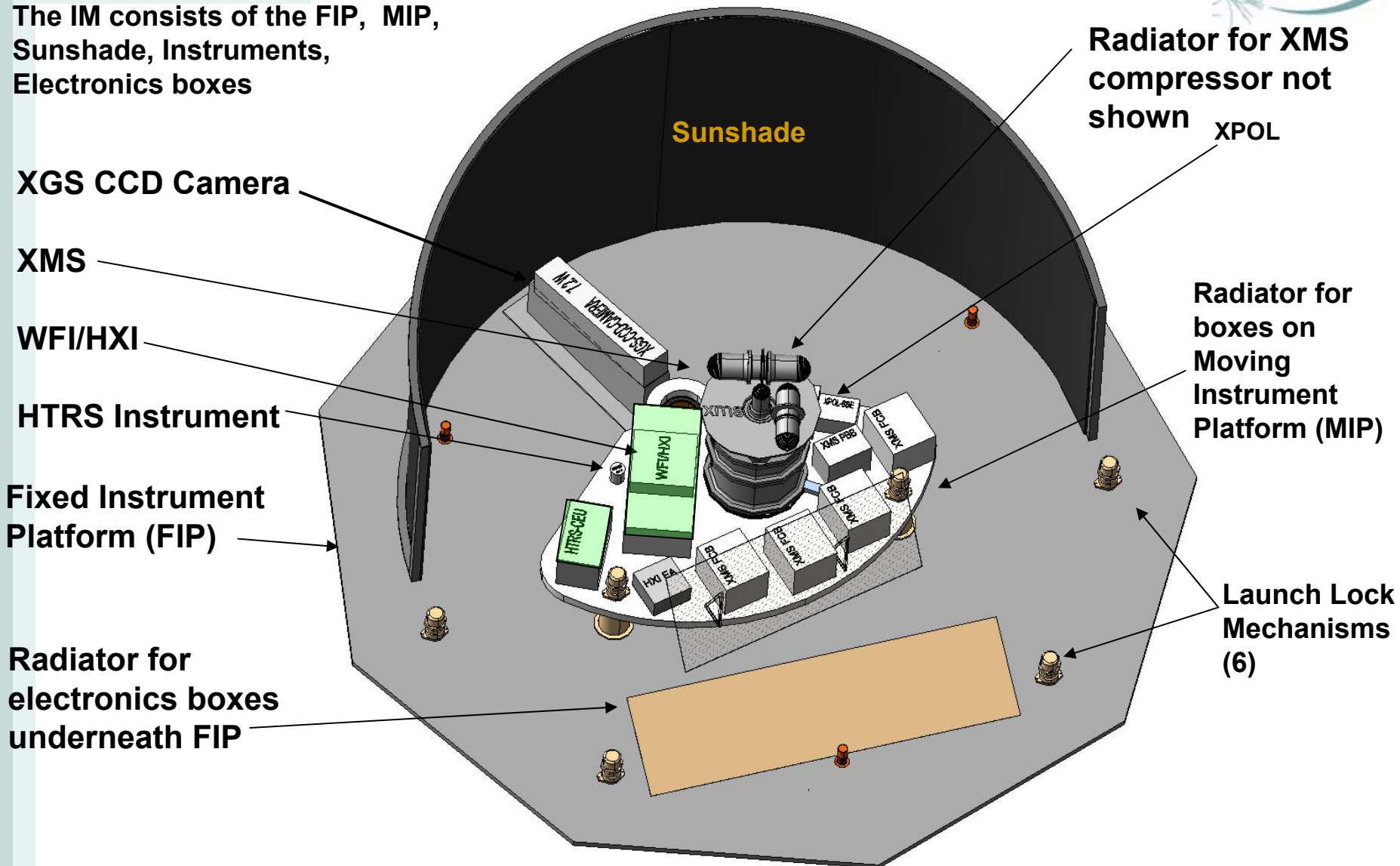
Glass-MPO
30"
 25 kg m^{-2}
 $A_{\text{eff}} @ 1 \text{ keV}$



XEUS Requirement

Instrument Module

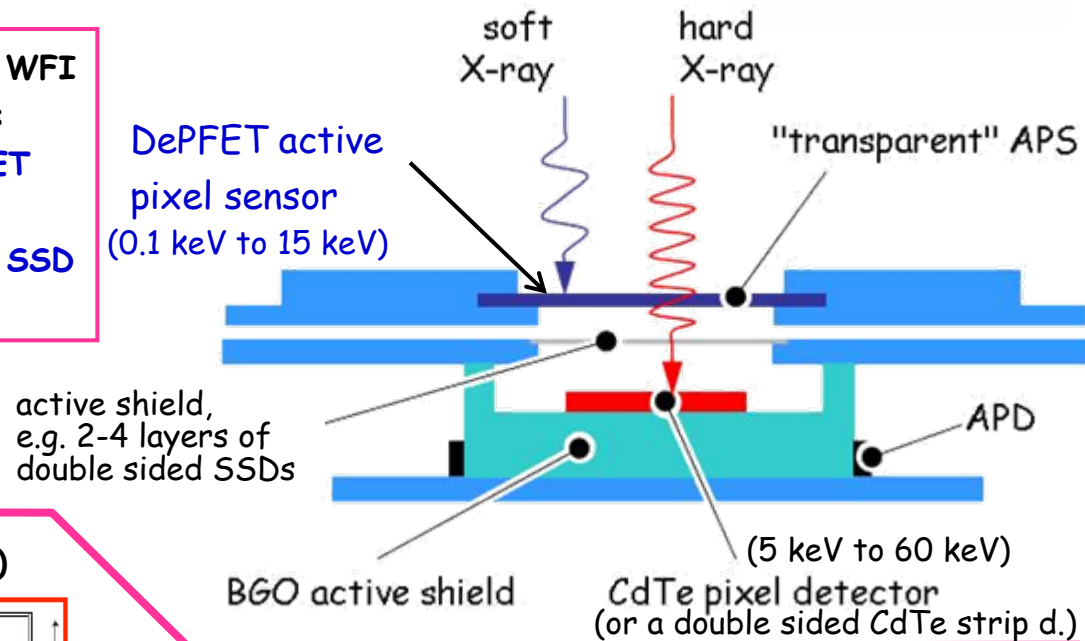
The IM consists of the FIP, MIP, Sunshade, Instruments, Electronics boxes



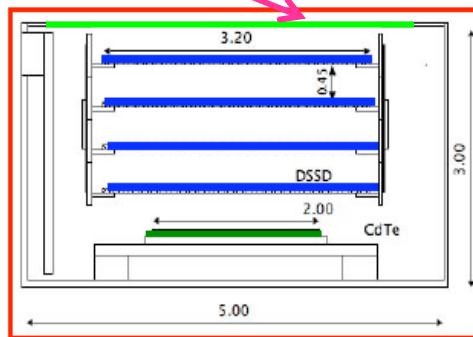
WFI and HXI on XEUS

one possibility of accommodation the DePFET WFI and the HXI within a single camera housing:

For X-rays above typ. 20 keV the DePFET becomes transparent and the X-rays will eventually interact with the double sided SSD or the CdTe detector



wide field imager (DePFET)



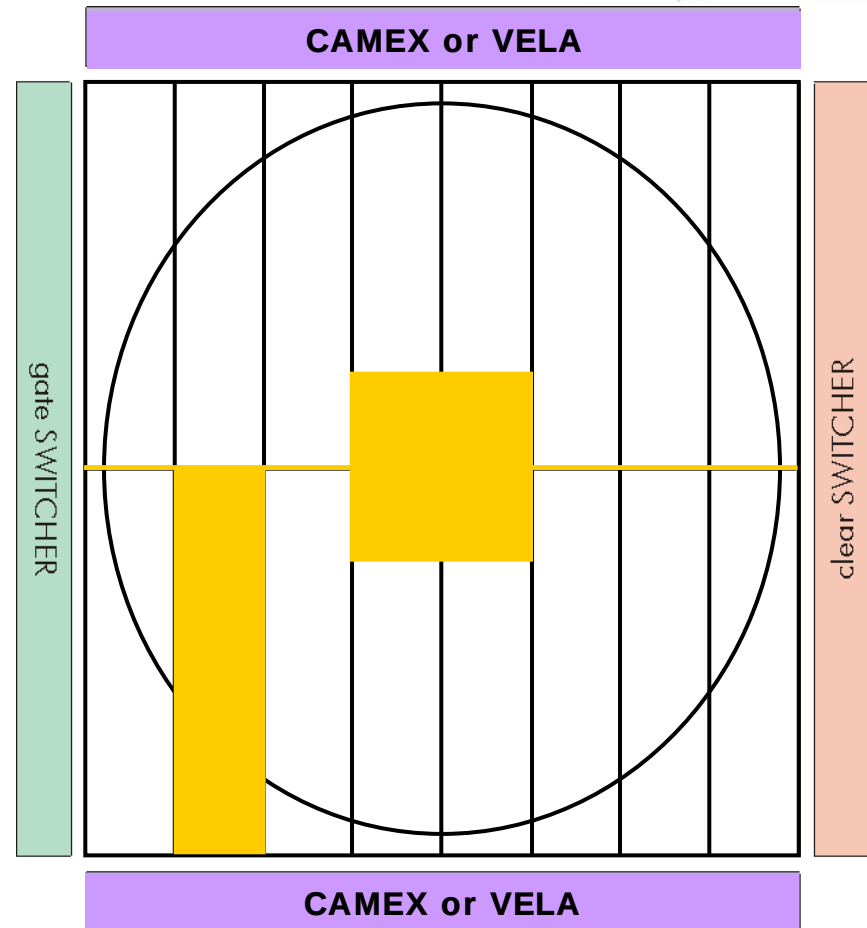
Test set-up for the combined double sided SSD and the CdTe detector. The distance between the WFI and the CdTe detector of the HXI is approximately 2 cm.

○ Double Side Silicon Strip (4 layers)

○ Double Side CdTe Strip (1 layer)

DEPFETs for the XEUS WFI

1. Flexible operating modes
2. low power dissipation (less than 2 W in 100 cm², DePFETs only)
3. Fano limited energy resolution from 0.5 keV to 30 keV
4. Spatial resolution better than 20 μm @ 100 μm pixel size
5. Homogeneous radiation entrance window
6. Intrinsic radiation hardness, no charge transfer needed
7. ENC was lowered to 0.2 e⁻ rms with RNDR
8. Thin optical ``Blocking Filter`` can be directly integrated
9. Operation at ``warm temperatures``, e.g. - 40 ° C

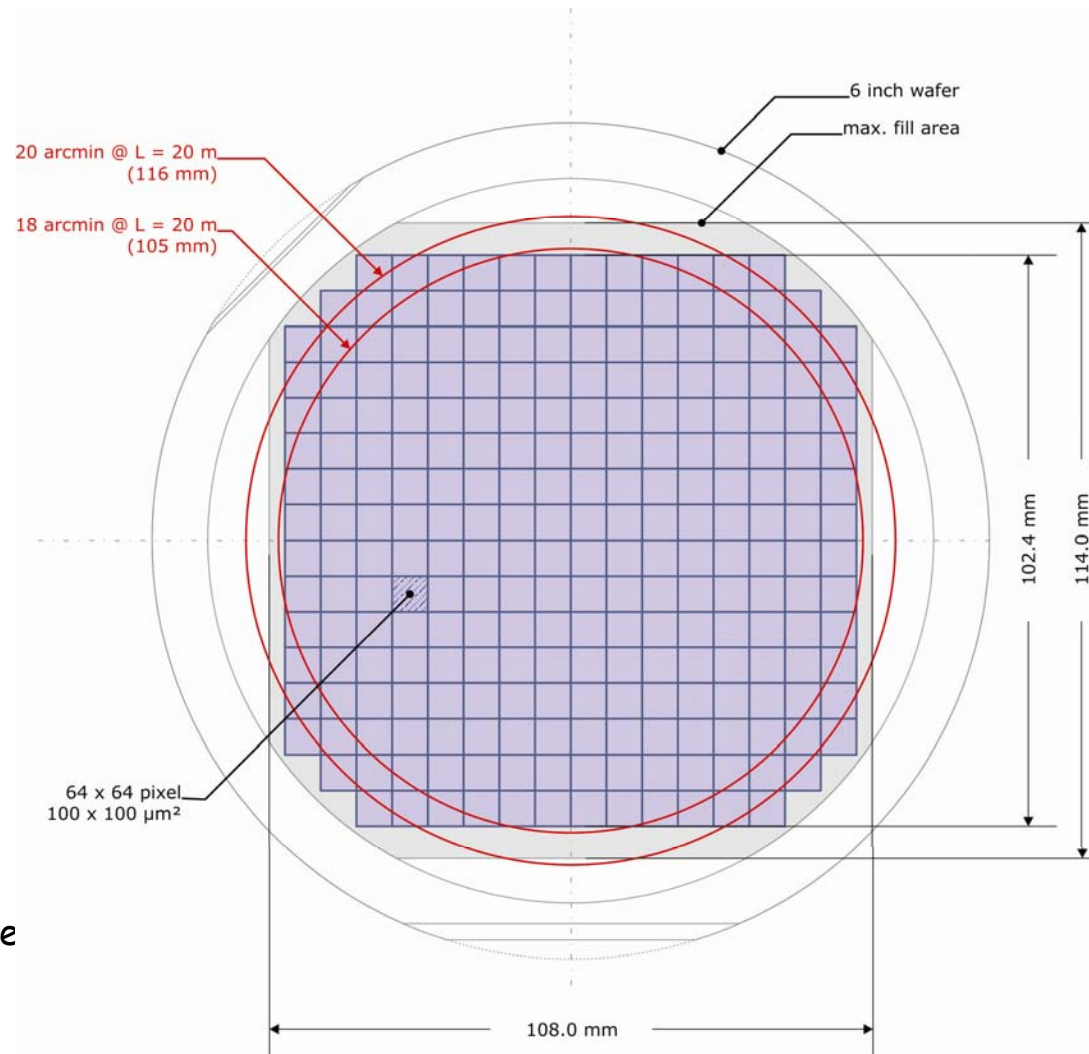


Wide Field Imager

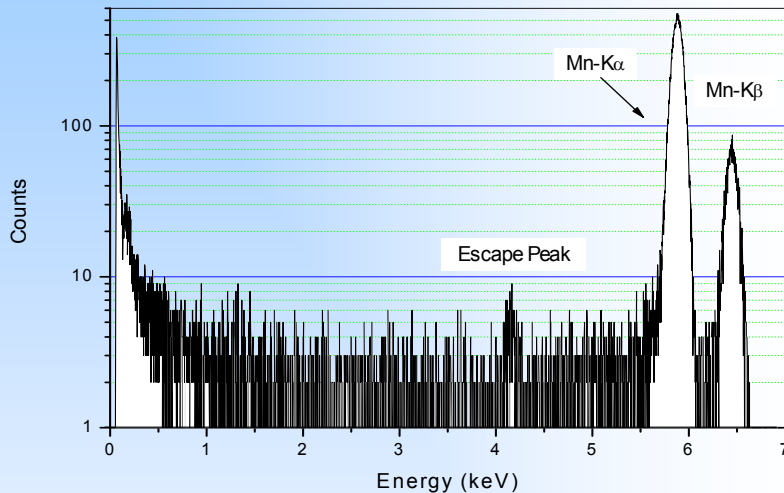


◆ tentative wafer layout

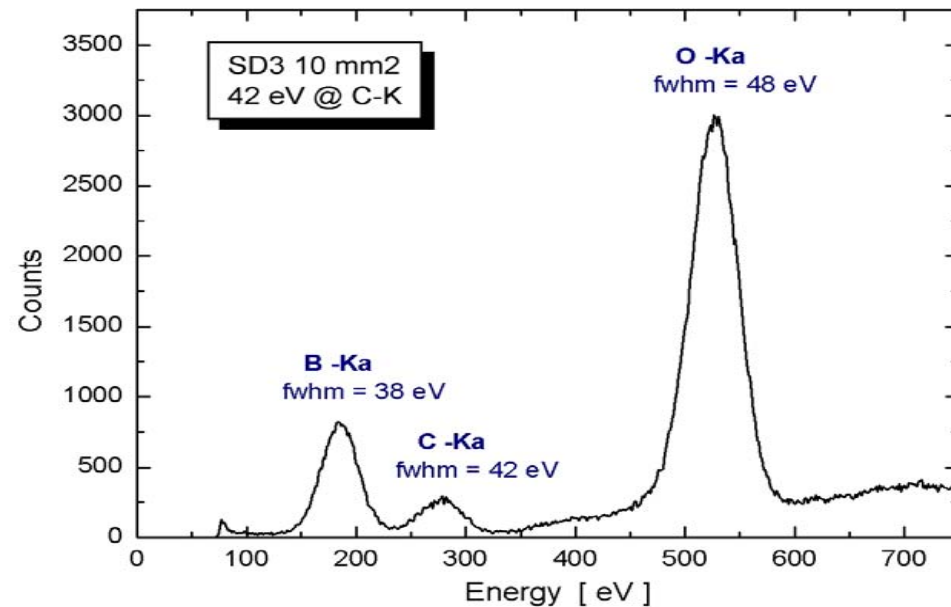
- **pixel size** $100\ \mu\text{m}$ $\square$
 $1\ \text{arcsec} \approx 97\ \mu\text{m} @ 20\ \text{m}$
 $\approx 1/5$ of PSF (5 arcmin)
- **format**
 $102.4 \times 102.4\ \text{mm}^2$
 1024×1024 pixels
with 'rounded' corners
- **field of view**
 $18\ \text{arcmin} \approx 105\ \text{mm} @ 20\ \text{m}$
 $20\ \text{arcmin} \approx 116\ \text{mm} @ 20\ \text{m}$
- no. of pixels 999.414
- FOV fraction outside sensor are
 $1.6\ \% @ 18\ \text{arcmin}$
 $9\ \% @ 20\ \text{arcmin}$



**"Backside" illumination:
Source on top of entrance window**



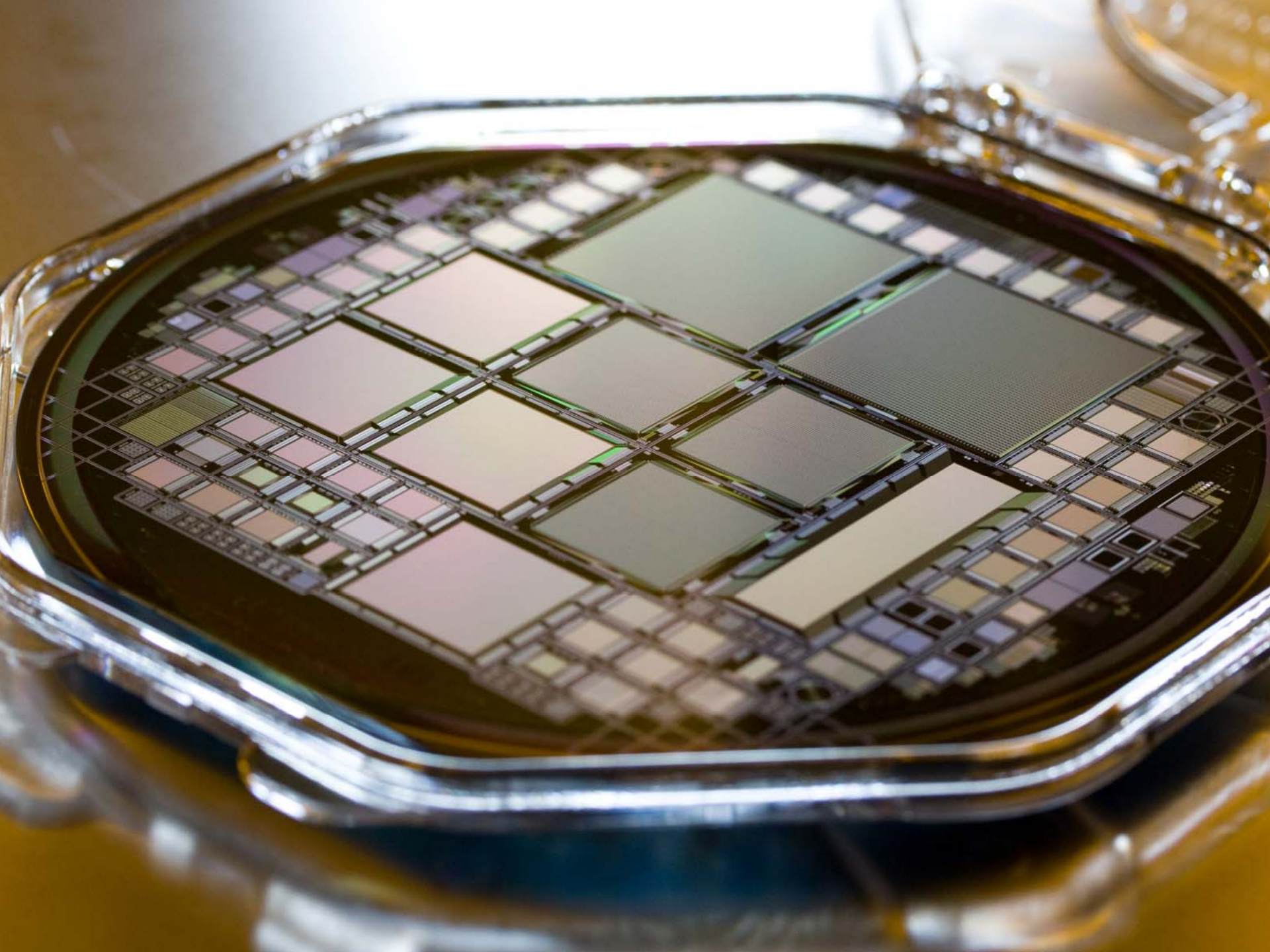
- timing
 $2 \mu\text{sec/row} \leftrightarrow 32 \mu\text{sec}/32 \times 512$
sensor
- room temperature
220 eV FWHM @ 5.9 keV (singles)
- moderate cooling -40 °C
127 eV FWHM @ 5.9 keV (singles)
132 eV FWHM @ 5.9 keV (all events)
- extrinsic speed & resolution limitations



◆ yield & homogeneity

- defect pixels
2 in 45 devices ($> 10^6$ pixels)
pixel yield > 0.99999
- dispersions

offset	$< 2 \%$	(of Mn-K α)
gain	$< 5 \%$	
noise	$< 10 \%$	



“Collaborating” Institutions for the WFI:

1. *Max-Planck-Institut für extraterrestrische Physik*

- Institut für Astronomie + Astrophysik, Tübingen***
- Universität Erlangen***
- Universität Darmstadt***

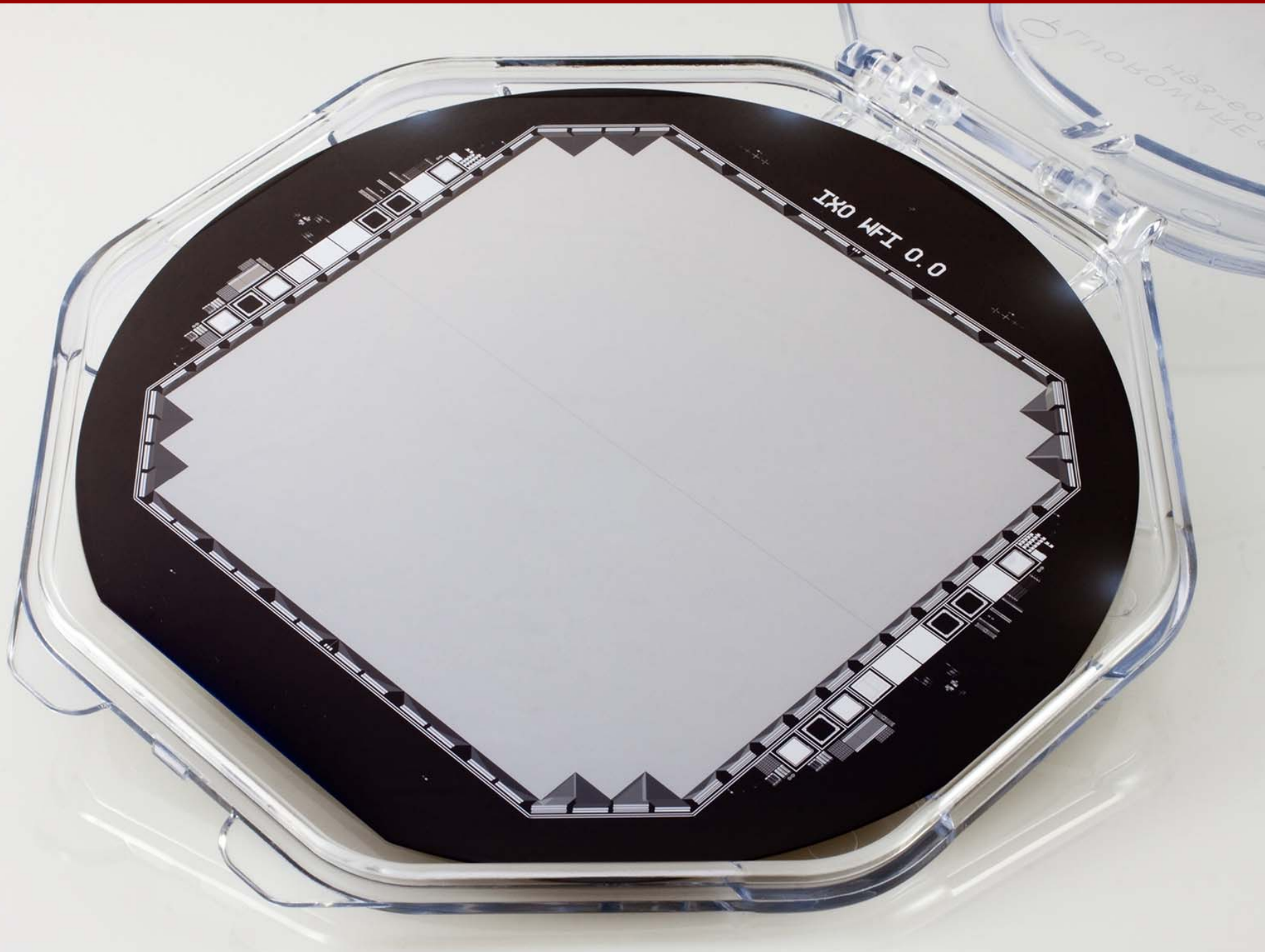
2. *University of Leicester, UK*

3. *Politecnico di Milano, Italy*

4. *University of Osaka, Japan*

5. *Center for Astrophysics, Cambridge, MA, Penn State University and MIT*

6. *PNSensor, Germany*



Large-Format, High-Speed, X-ray pnCCDs Combined with Electron and Ion Imaging Spectrometers in a Multipurpose Chamber for Experiments at 4th Generation Light Sources

Lothar Strüder ^{a,c,e,h}, Sascha Epp ^a, Daniel Rolles ^a, Robert Hartmann ^{b,e}, Peter Holl ^{b,e}, Gerhard Lutz ^{b,e}, Heike Soltau ^{b,e}, Rouven Eckart ^{b,e}, Christian Reich ^{b,e}, Klaus Heinzinger ^{b,e}, Christian Thamm ^{b,e}, Artem Rudenko ^a, Faton Krasniqi ^a, Kai-Uwe Kühnel ⁱ, Christian Bauer ⁱ, Claus-Dieter Schröter ⁱ, Danilo Miessner ^{c,e}, Matteo Porro ^{c,e}, Olaf Hälker ^{c,e}, Norbert Meidinger ^{c,e}, Nils Kimmel ^{c,e}, Robert Andritschke ^{c,e}, Florian Schopper ^{c,e}, Georg Weidenspointner ^{c,e}, Alexander Ziegler ^{c,e}, Sven Herrmann ^{c,e}, Ullrich Pietsch ^h, Albert Walenta ^h, Wolfram Leitenberger ^h, Christoph Bostedt ^f, Thomas Möller ^f, Daniela Rupp ^f, Marcus Adolph ^f, Heinz Graafsma ^g, Helmut Hirsemann ^g, Klaus Gärtner ^k, Rainer Richter ^{d,e}, Daniel Pietschner ^{c,e}, Lutz Foucar ^a, Robert L. Shoeman ⁱ, Ilme Schlichting ^{a,l}, Joachim Ullrich ^{a,i}

^a Max Planck Advanced Study Group, Center for Free Electron Laser Science (CFEL) at DESY, Notkestr. 85, D-22607 Hamburg, Germany

^b PNSensor GmbH, Römerstraße 28, D-80803 München, Germany

^c Max-Planck-Institut für extraterrestrische Physik, Giessenbachstraße, D-85741 Garching, Germany

^d Max-Planck-Institut für Physik, Föhringer Ring 6, D-80805 München, Germany

^e MPI Halbleiterlabor, Otto-Hahn-Ring 6, D-81739 München, Germany

^f Technische Universität Berlin, Institut für Optik und Atomare Physik, Hardenbergstraße 36, D-10623 Berlin, Germany

^g DESY, Notkestr.85, D-22607 Hamburg, Germany

^h University of Siegen, Emmy-Noether Campus, Walter Flex Str. 3, D-57068 Siegen, Germany

ⁱ Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, D-69117 Heidelberg, Germany

^k Weierstrass Institut, Berlin, Mohrenstr. 39, D-10117 Berlin, Germany

^l Max-Planck-Institut für medizinische Forschung, Jahnstr. 29, D-69120 Heidelberg, Germany



Free Electron Lasers



...across the classical disciplines
...across institutions

CFEL: 350 people
UHH + DESY + MPG

Center for Free-Electron Laser Science (CFEL)

DESY

Coherent
Imaging

Chapman

DESY

Experiment
Group

DESY

Theory
Group

MPG

Advanced
Study
Group

Ullrich

MPG/UNI

Condensed
Matter
Dynamics

Cavalleri

MPG/UNI

Experiment
Group

Miller

Univ. HH

Advanced
Study
Group

Wurth

Kernphysik

extraterr. Physik

med. Forschung

biophysik. Chem.

Fritz-Haber Inst.

kompl. Systeme

Junior
Research
Groups

Group 1

Group 2

Group 3

Max-Planck 'Forschungsgruppe'
at HH University

CFEL
Management
Board

The Max-Planck Advanced Study Group

Topic 1:
FLASH Experiments
(Ullrich, Moshhammer)

Topic 2:
LCLS/XFEL Exp.
across the sections
(Schlichting, Techert)

Topic 3
Detectors
(Strüder)

Topic 4
Theory
(Rost)

Core ASG based at the CFEL site

Scientist 1

Scientists 2 & 3

Scientist 4 & 5

SWNG*

supported by 2 technicians and ~ 4 PhD students

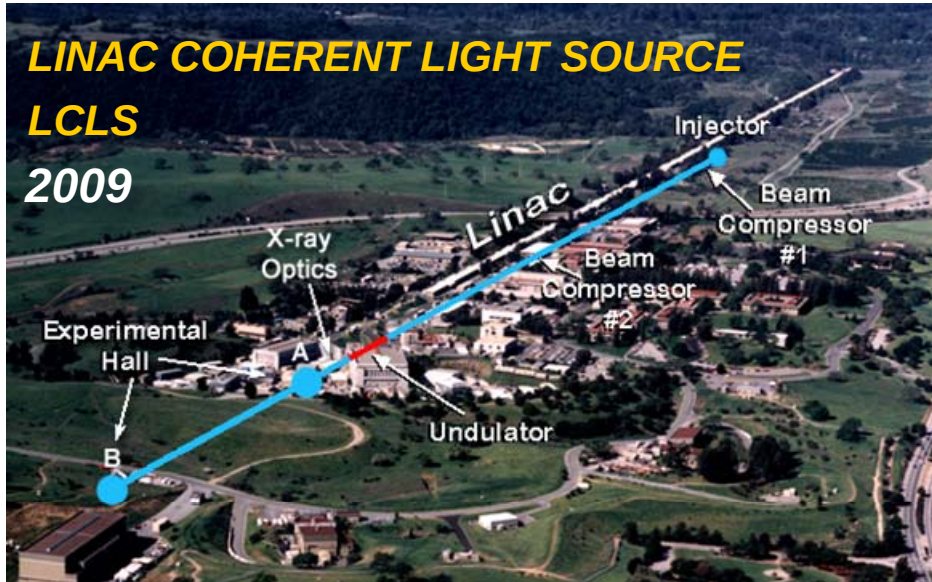
- Atoms & molecules
- Molecular Ions
- Cluster physics
- Ultra-cold plasmas
- Pump-Probe exp.
- Ion Spectroscopy (EBIT)

- Structure investigation of larger molecules and soft matter
- Alignment of molecules in the gas phase
- Fragment-imaging
- Ultra-cold gases
- Ion Spectroscopy (EBIT)

- **Pixel detectors**
 - * **pnCCDs**
 - * **DePFETS**
 - * **SDDs**
- **large formats**
up to 4 Mpix
- **MHz readout for XFEL**
- **low noise**
- **high dynamic range**
- **radiation hardness**

- Basics of light-matter interactions
- Cluster dynamics & bio-molecules
- Imaging, dynamics of fragmentation

Hard X-ray SASE Free Electron Lasers



FLASH: 5 Hz, 10 Hz
and 5 MHz

LCLS: 120 Hz

SCSS: 60 Hz

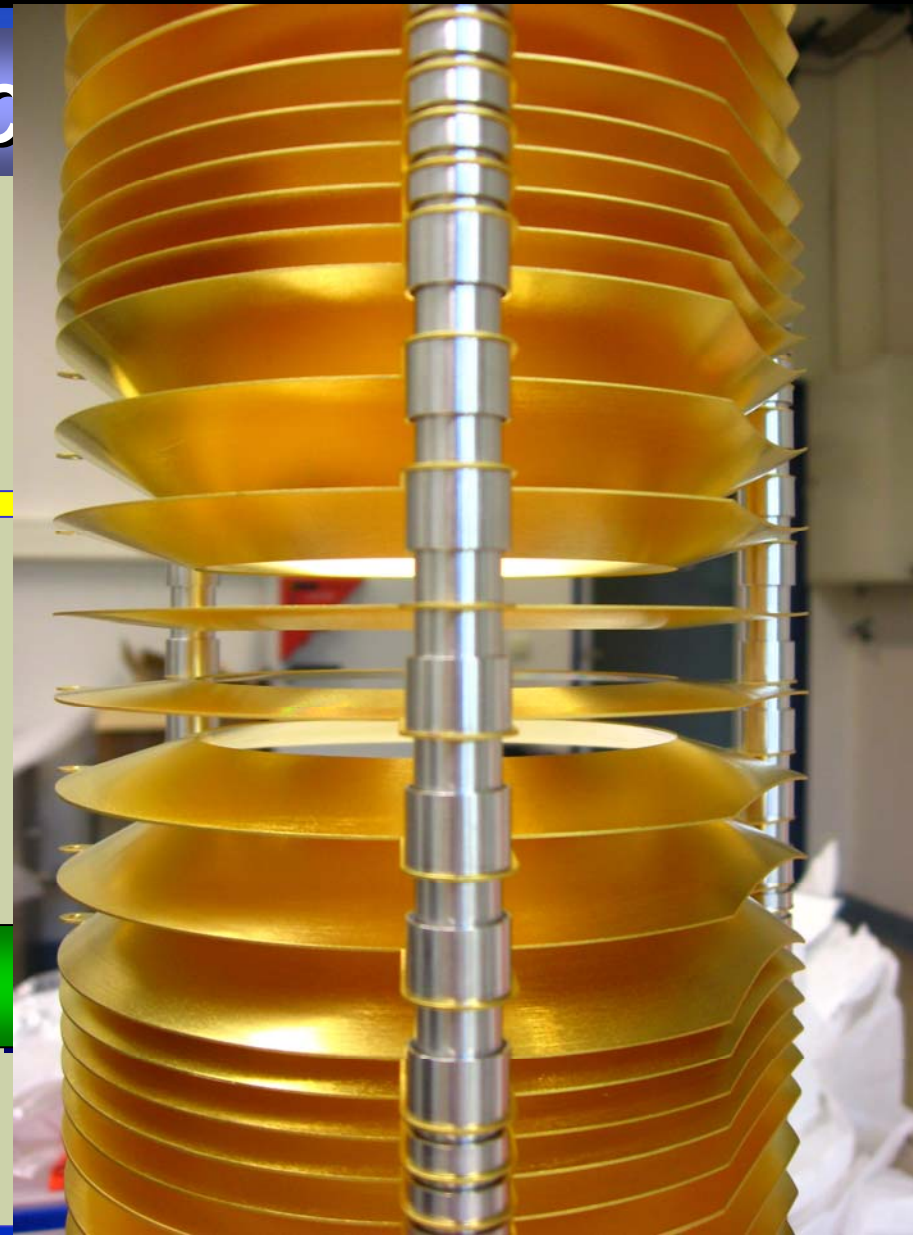
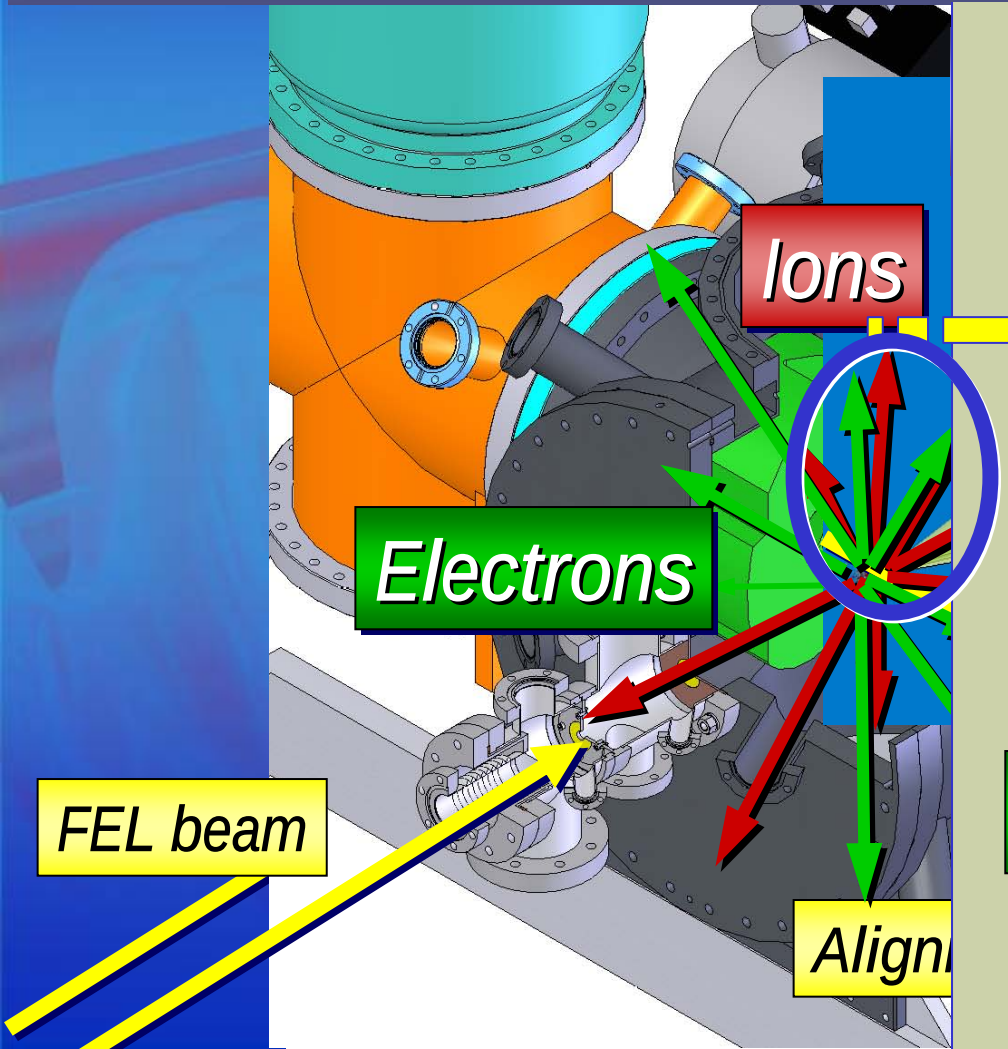
XFEL: 5 Hz, 10 Hz
and 5 MHz





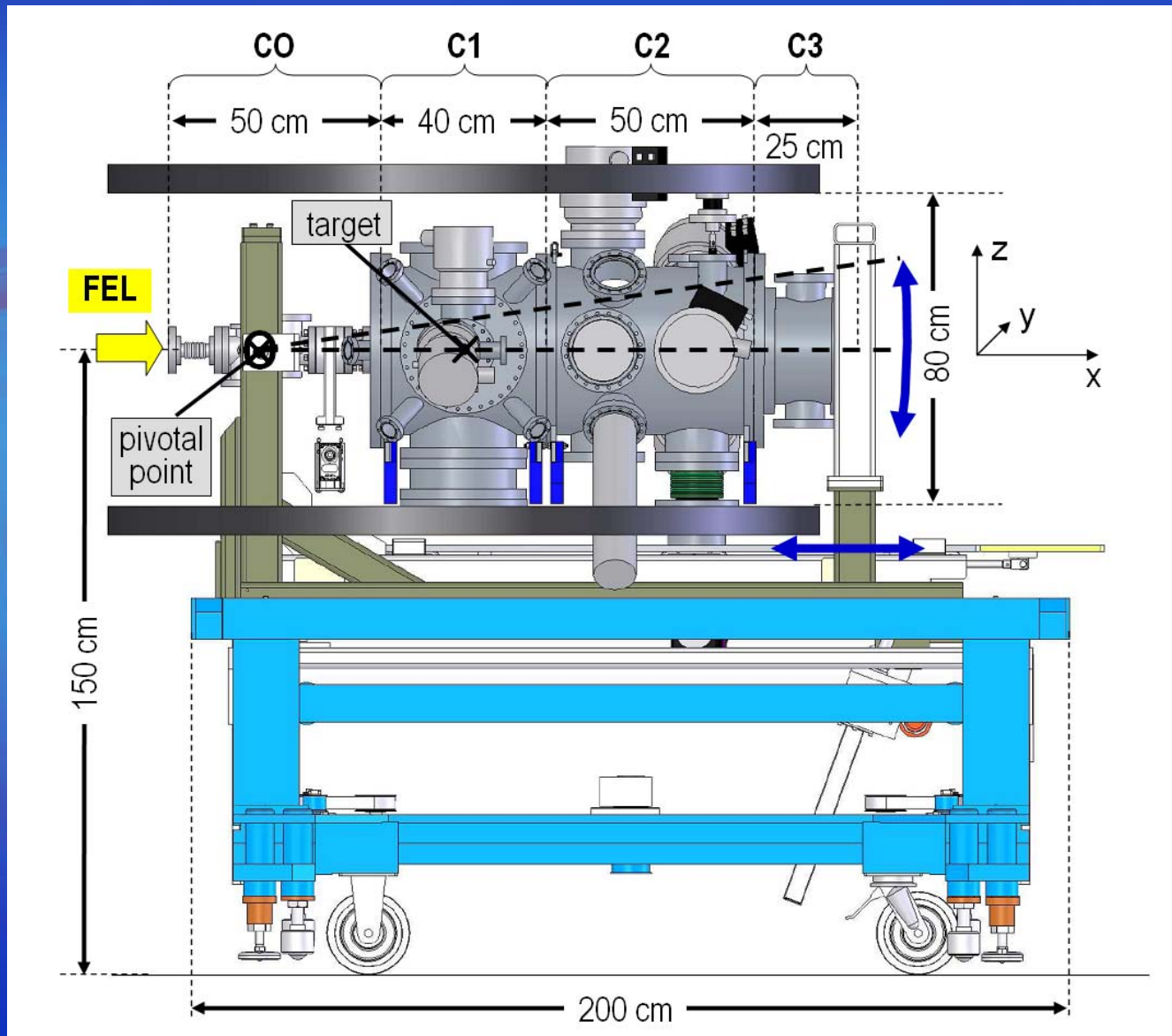
CFEL-ASG MultiPurpose Chamber

Integrated Reaction Micro



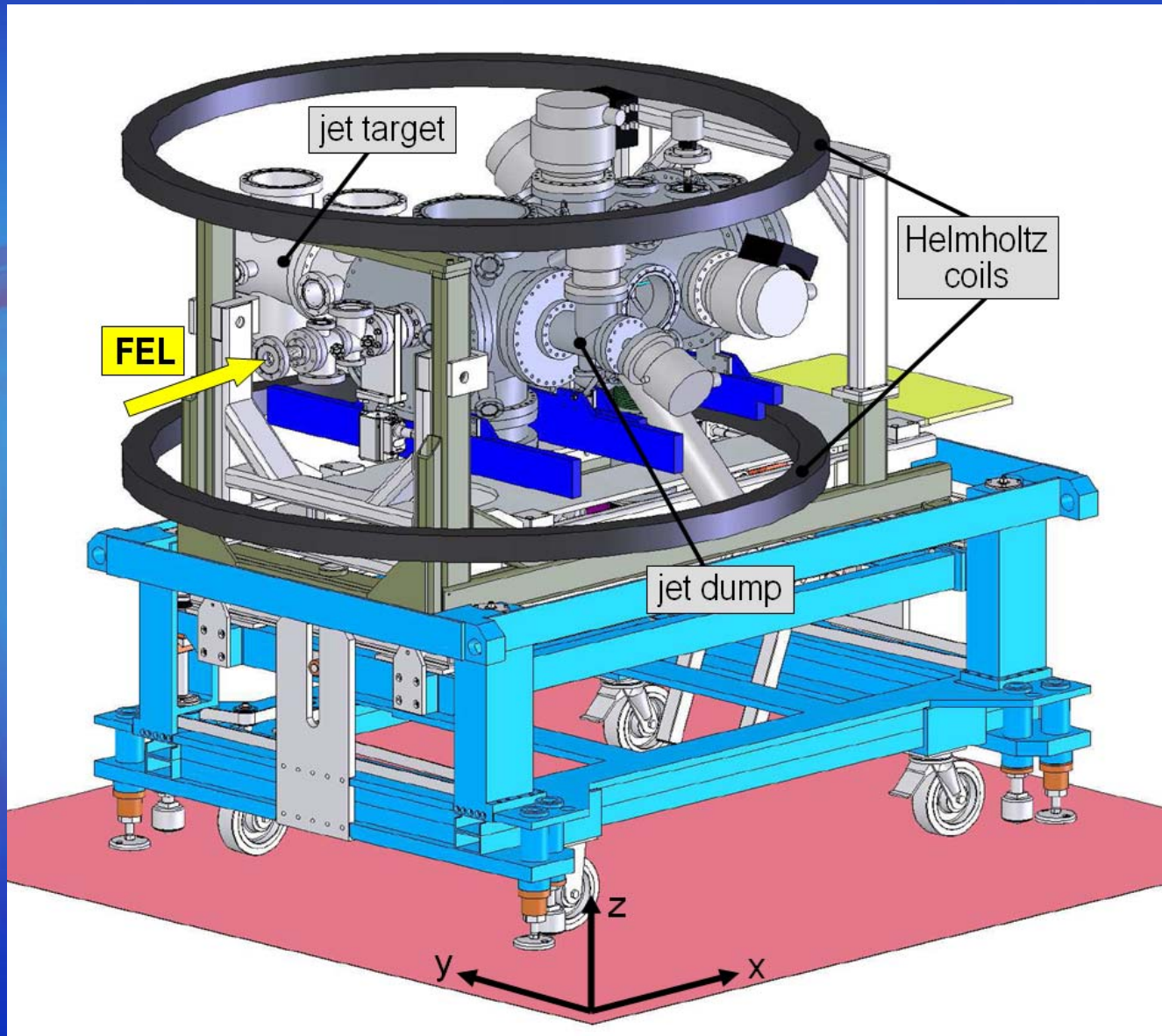


CFEL-ASG MultiPurpose Chamber



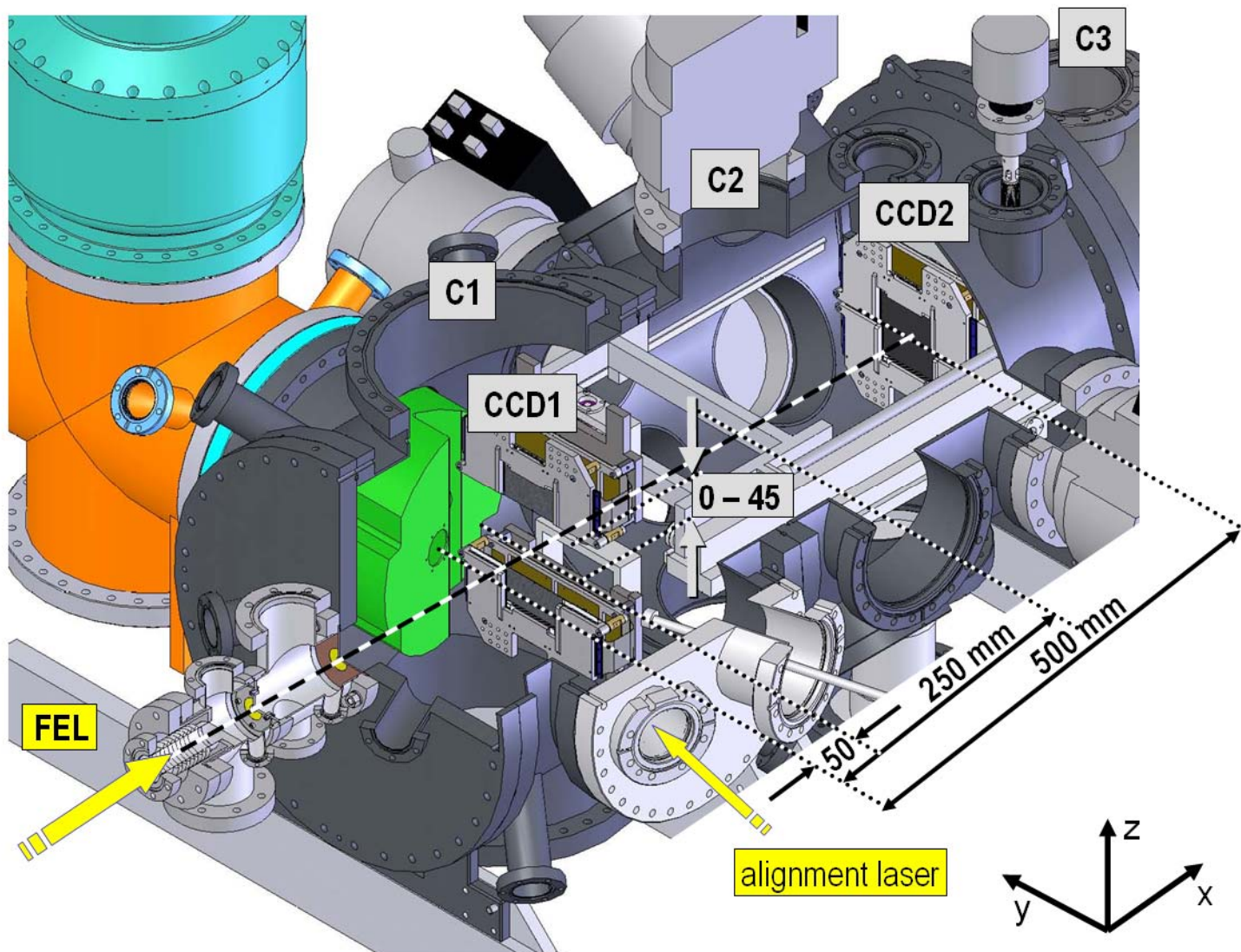


CFEL-ASG MultiPurpose Chamber





CFEL-ASG MultiPurpose Chamber



Requirements of the FLASH, LCLS and XFEL

Integrating Area Detectors

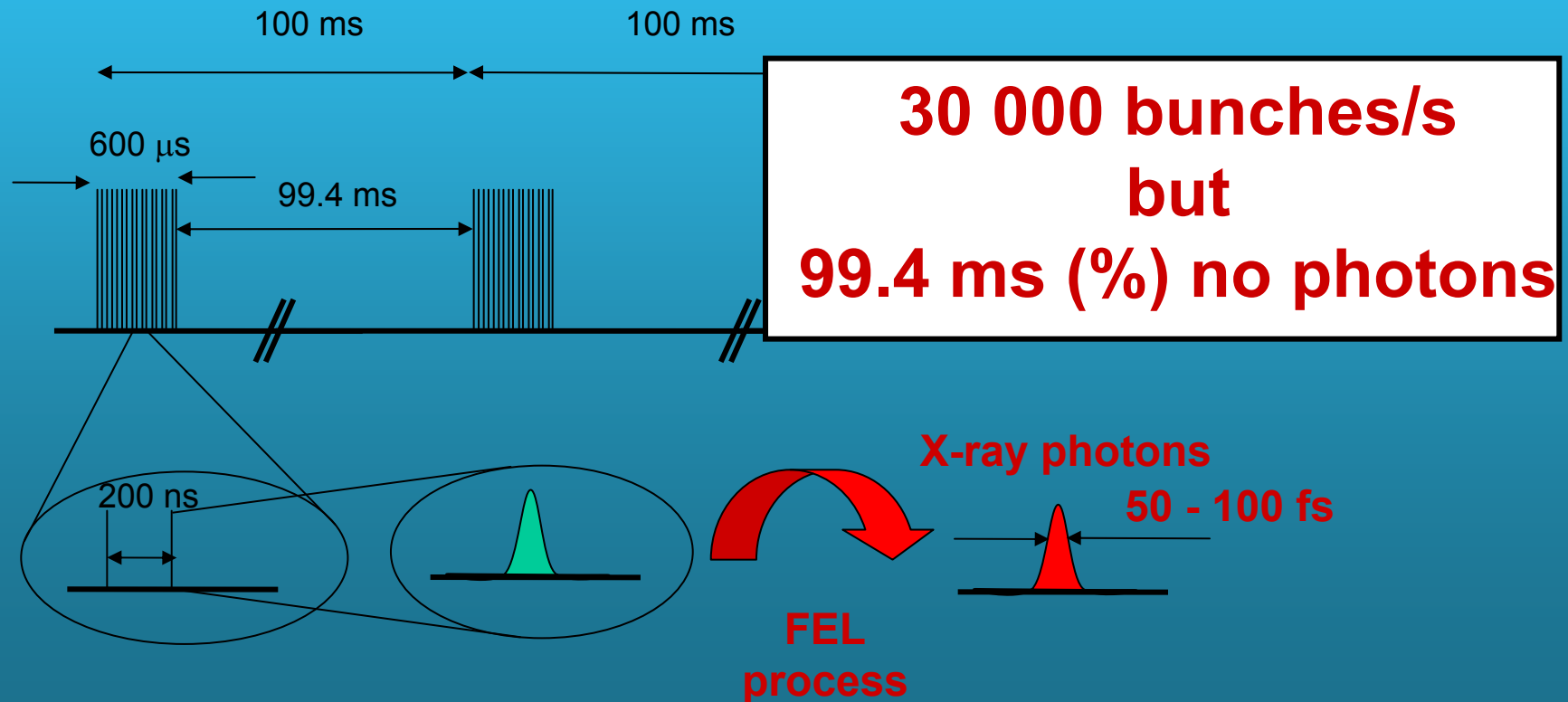


	FLASH, LCLS + XFEL	pnCCD and DePFET system
single photon resolution	yes	yes
energy range	$0.05 < E < 24$ (keV)	$0.05 < E < 25$ [keV]
pixel size (μm)	100	75 (200)
sig.rate/pixel/bunch	10^3 (10^5)	$10^3 - 10^4$
quantum efficiency	> 0.8	> 0.8 from 0.3 to 12 keV
number of pixels	512×512 (min.)	1024×1024 and 2048×2048
frame rate/repetition rate	10 Hz - 120 Hz	up to 250 Hz with pnCCD
XFEL burst mode	5 MHz (3.000 bunches)	> 5 MHz (3.000 bunches) with DePFET system
Readout noise	$< 150 e^-$ (rms)	$< 10 e^-$ (rms) ($2 e^-$ possible)
cooling	possible	- 20°C optimum room temperature possible
vacuum compatibility	yes	yes
preprocessing	no (yes) ?	possible upon request

What is the challenge for Detectors @ XFEL

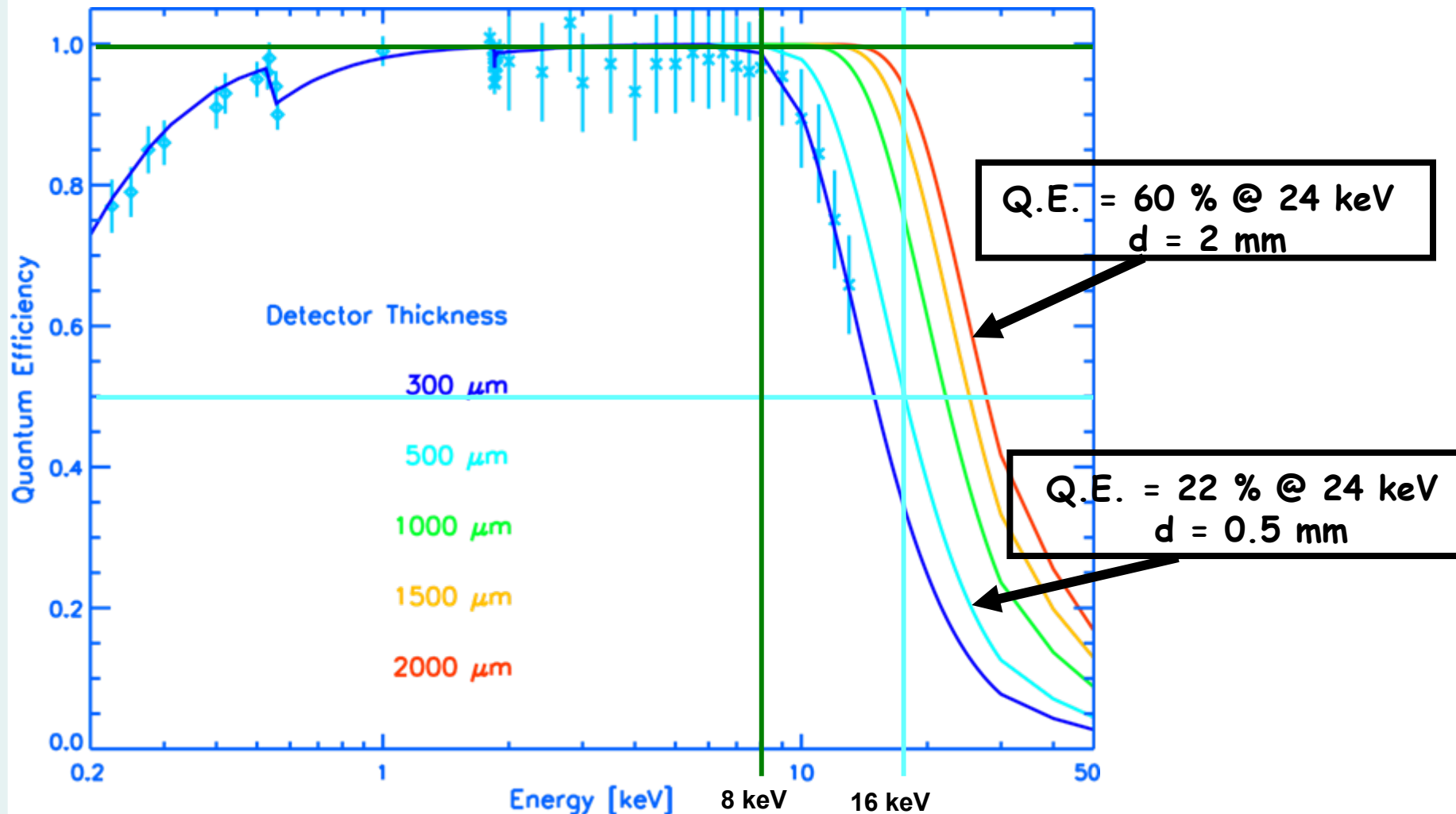
Time structure: difference with "others"

Electron bunch trains; up to 3000 bunches in 600 μs , repeated 10 times per second.
Producing 100 fsec X-ray pulses (up to 30 000 bunches per second).

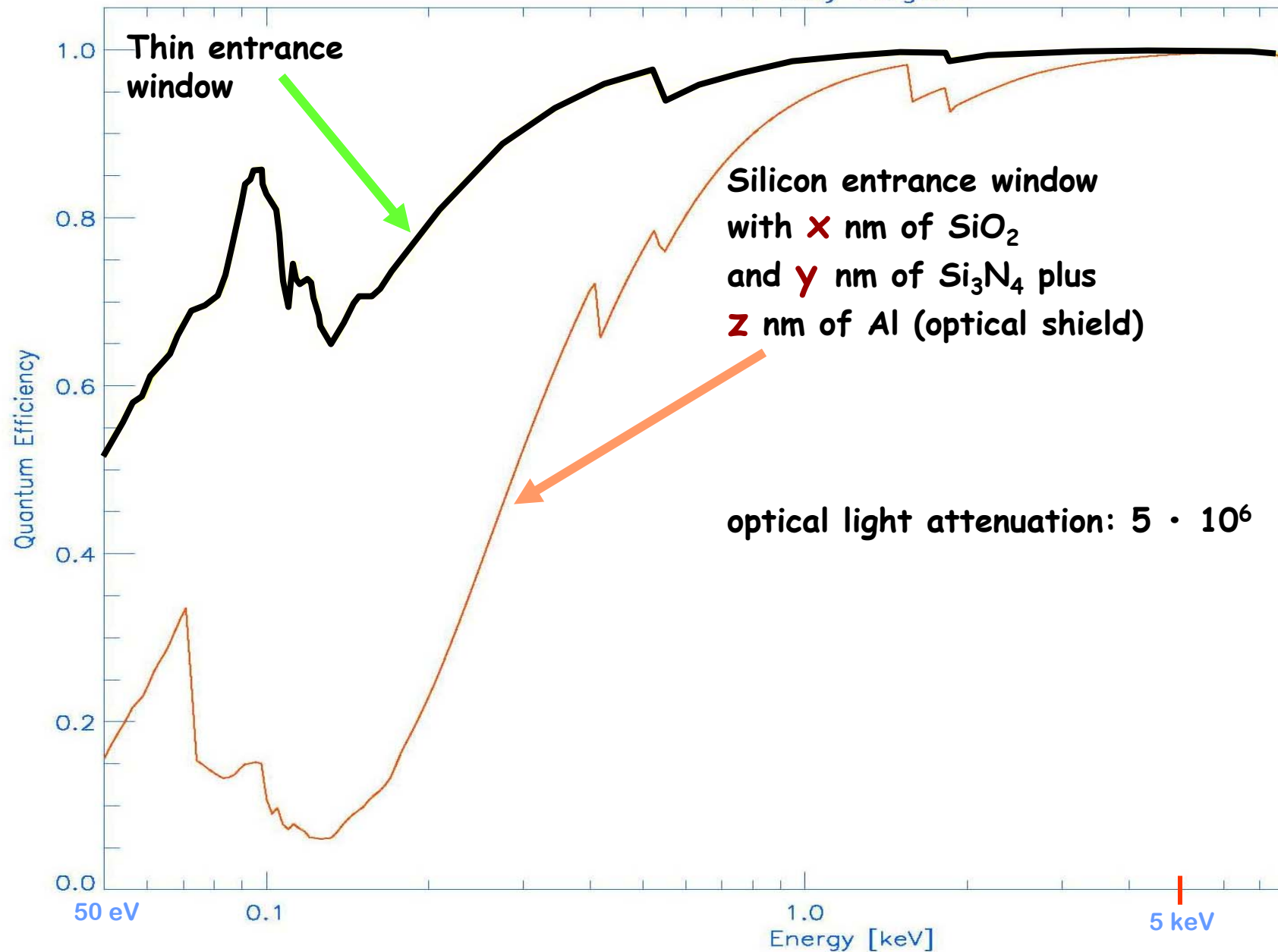


What is limiting the quantum efficiency ?

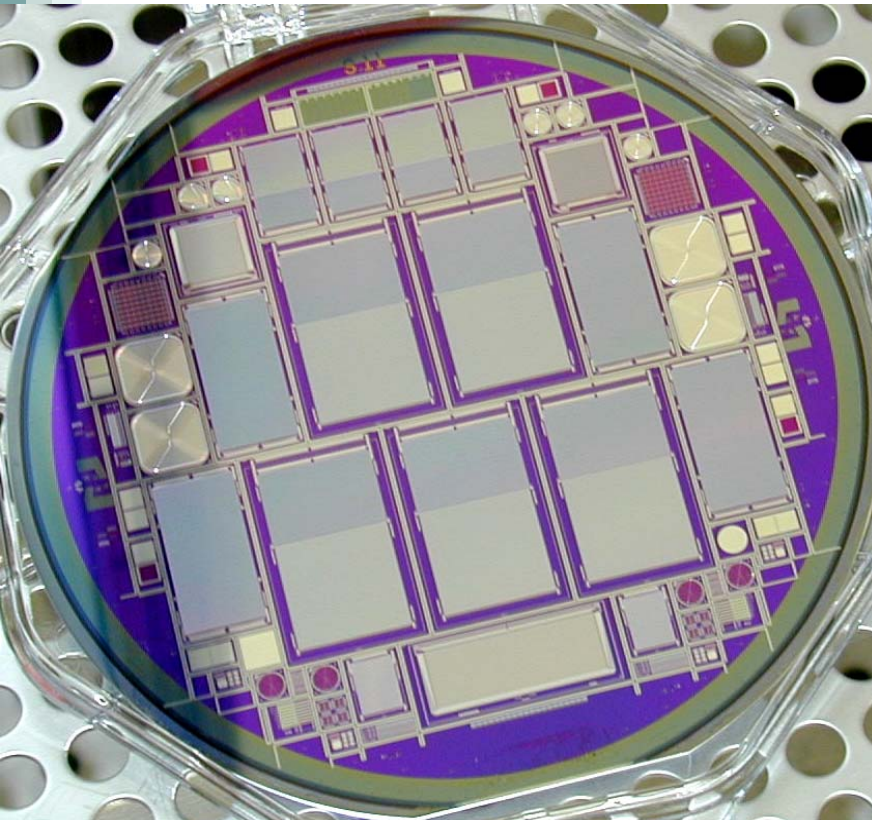
The thickness of Silicon !!



Monolithic Integration of optical blocking filters

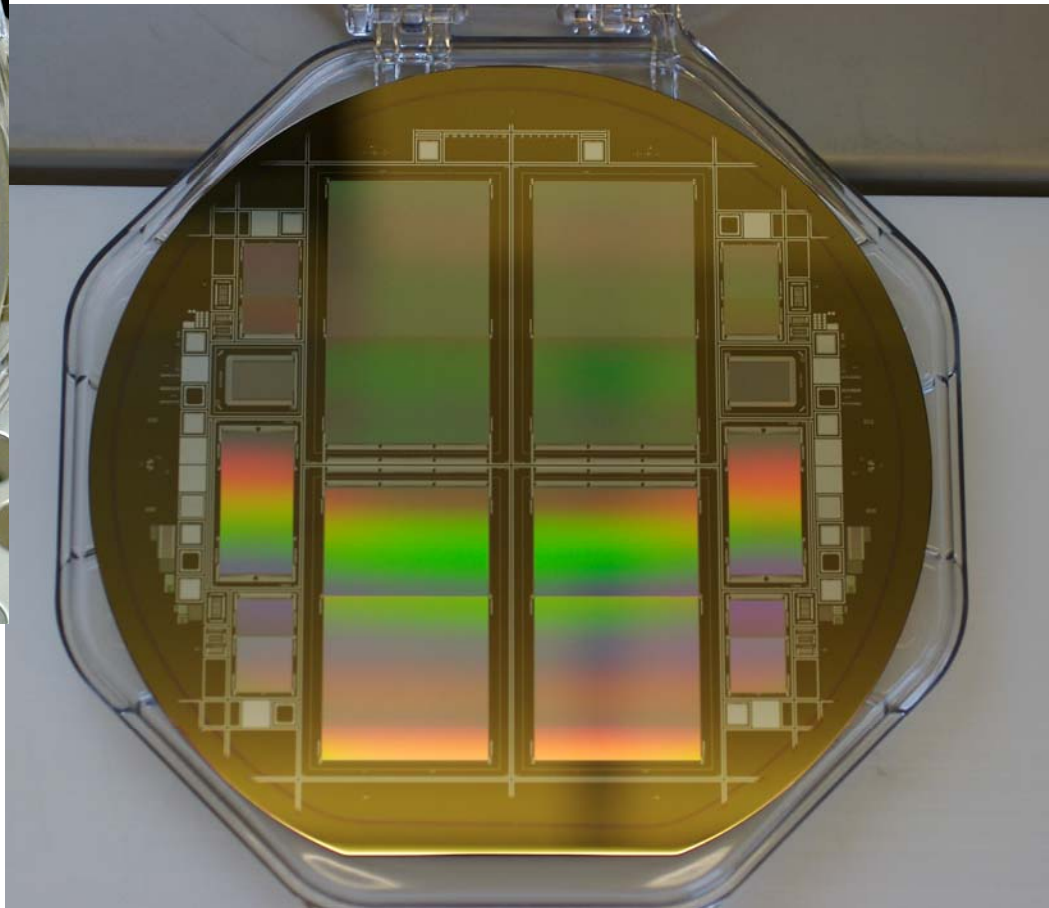


Recent pnCCDs fabrications



prototype eROSITA version
format: $256 \times 256 \times 2$ pixel
area: $3.7 \text{ cm}^2 + 2.5 \text{ cm}^2$
pixel size: $75 \times 75 \mu\text{m}^2$

flight type eROSITA version
format: $384 \times 384 \times 2$ pixel
area: $8.4 \text{ cm}^2 + 5.6 \text{ cm}^2$
pixel size: $75 \times 75 \mu\text{m}^2$

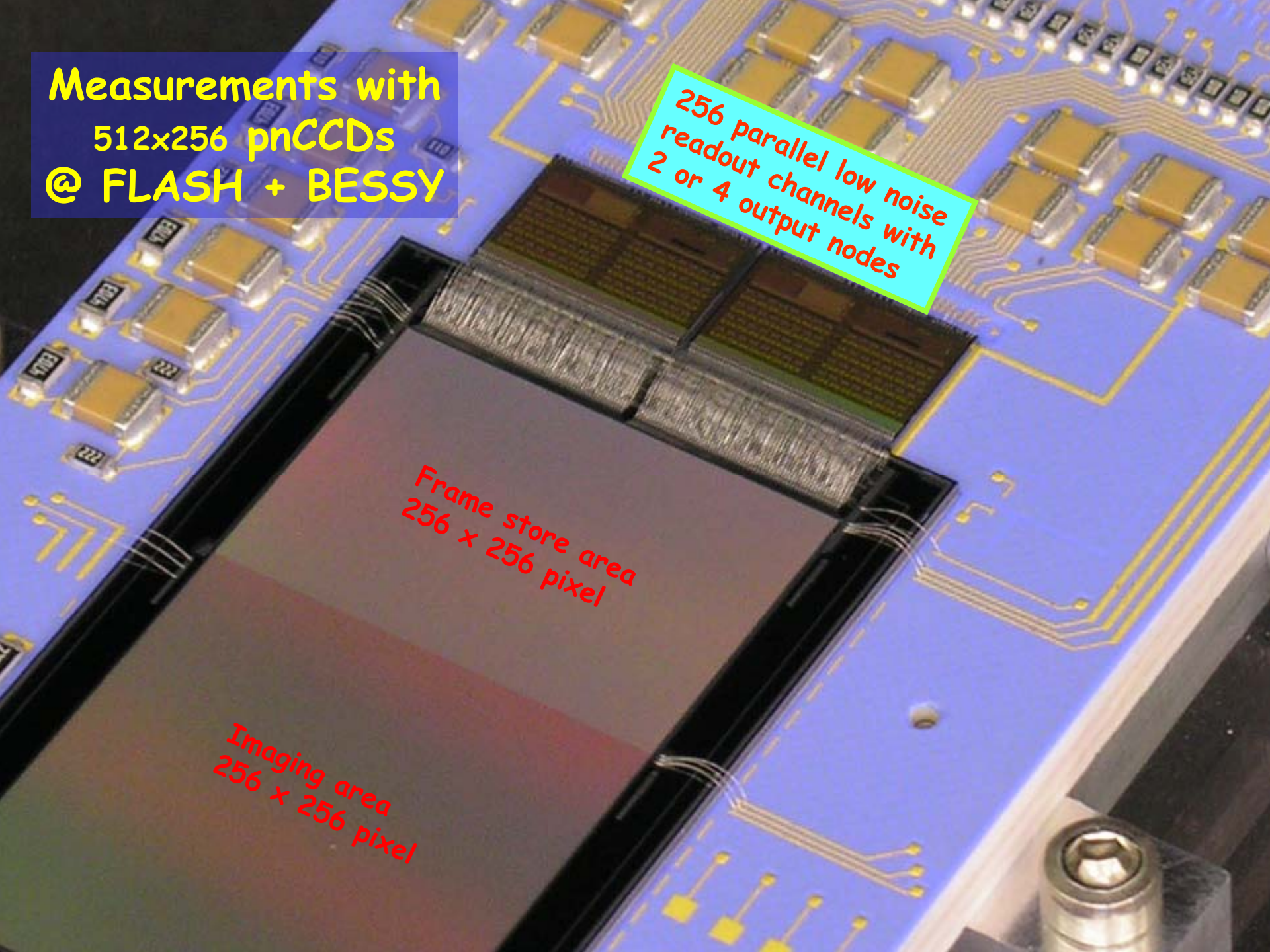


Measurements with
512x256 pnCCDs
@ FLASH + BESSY

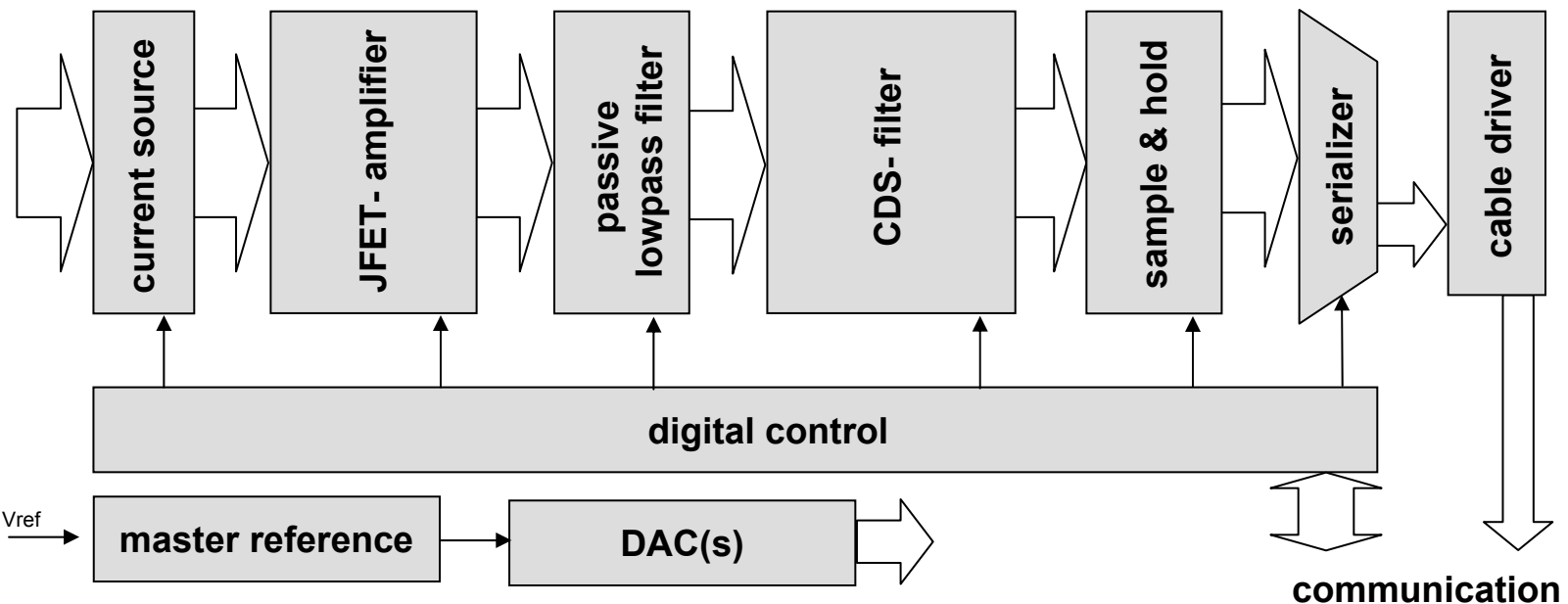
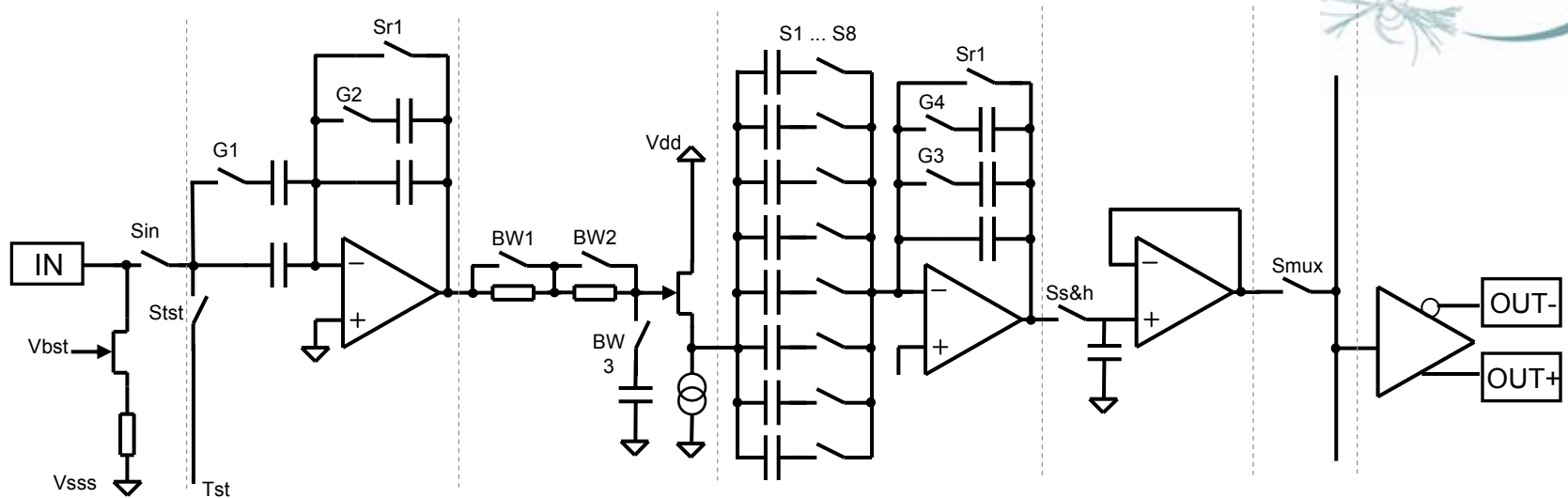
256 parallel low noise
readout channels with
2 or 4 output nodes

Frame store area
256 x 256 pixel

Imaging area
256 x 256 pixel

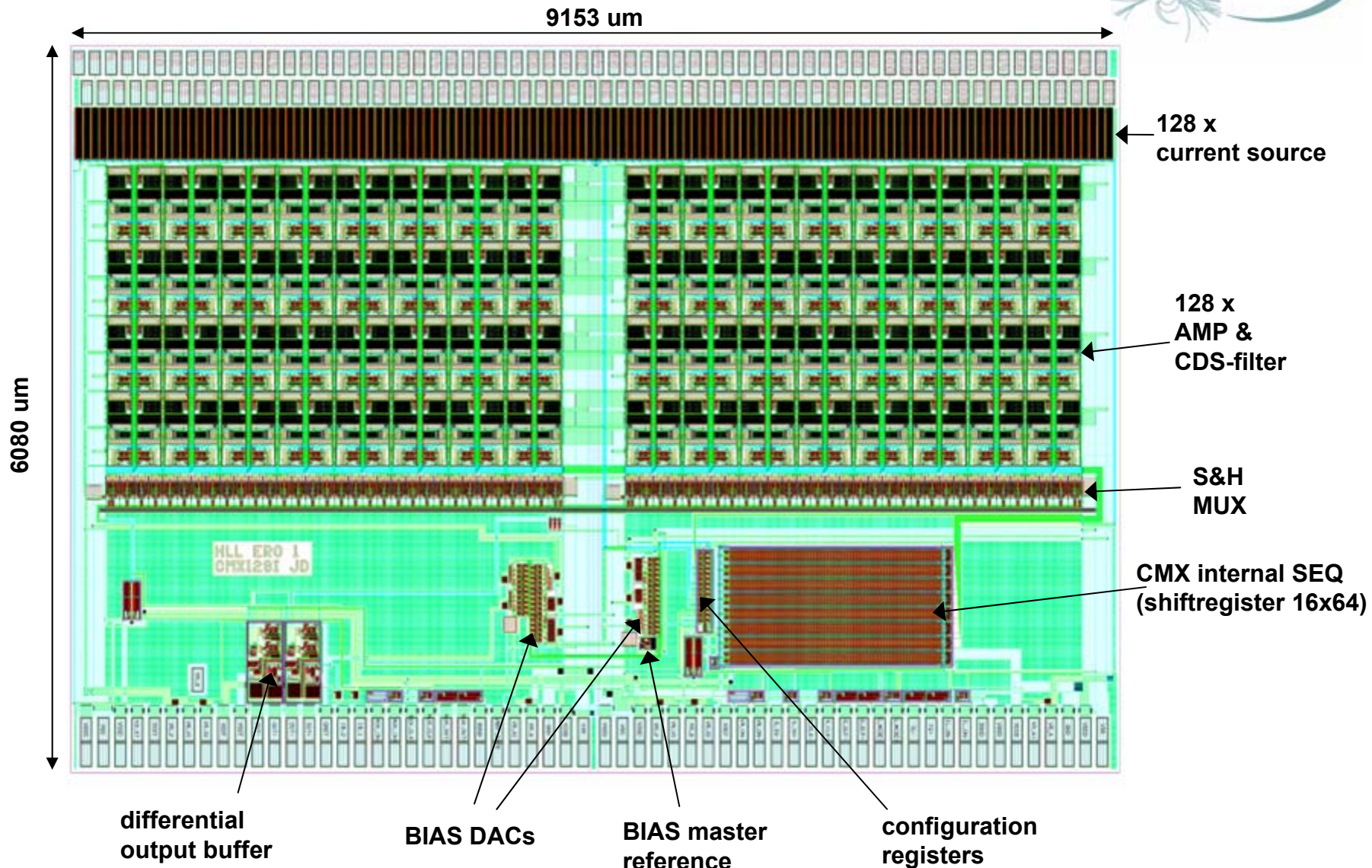


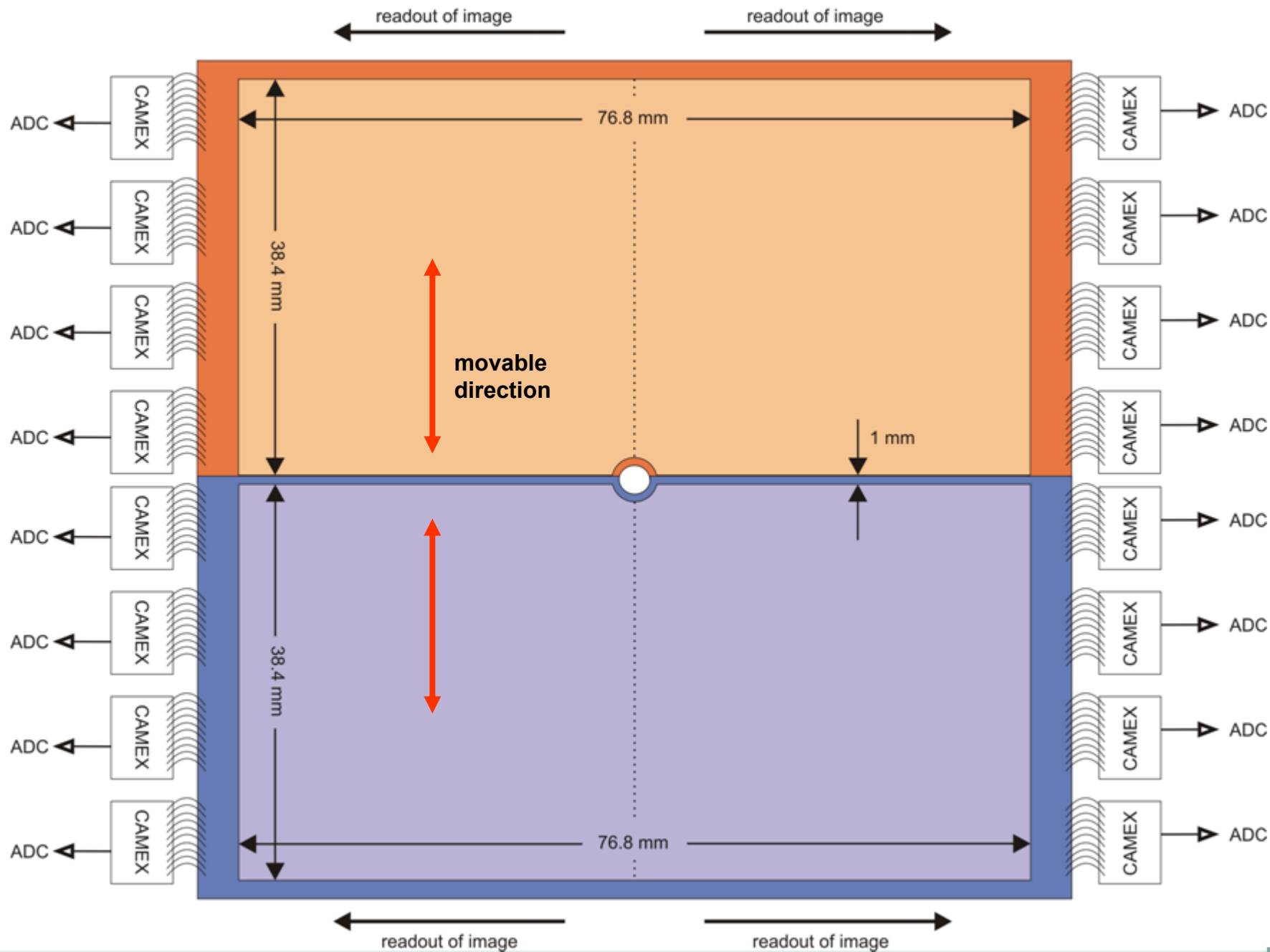
CAMEX block diagram



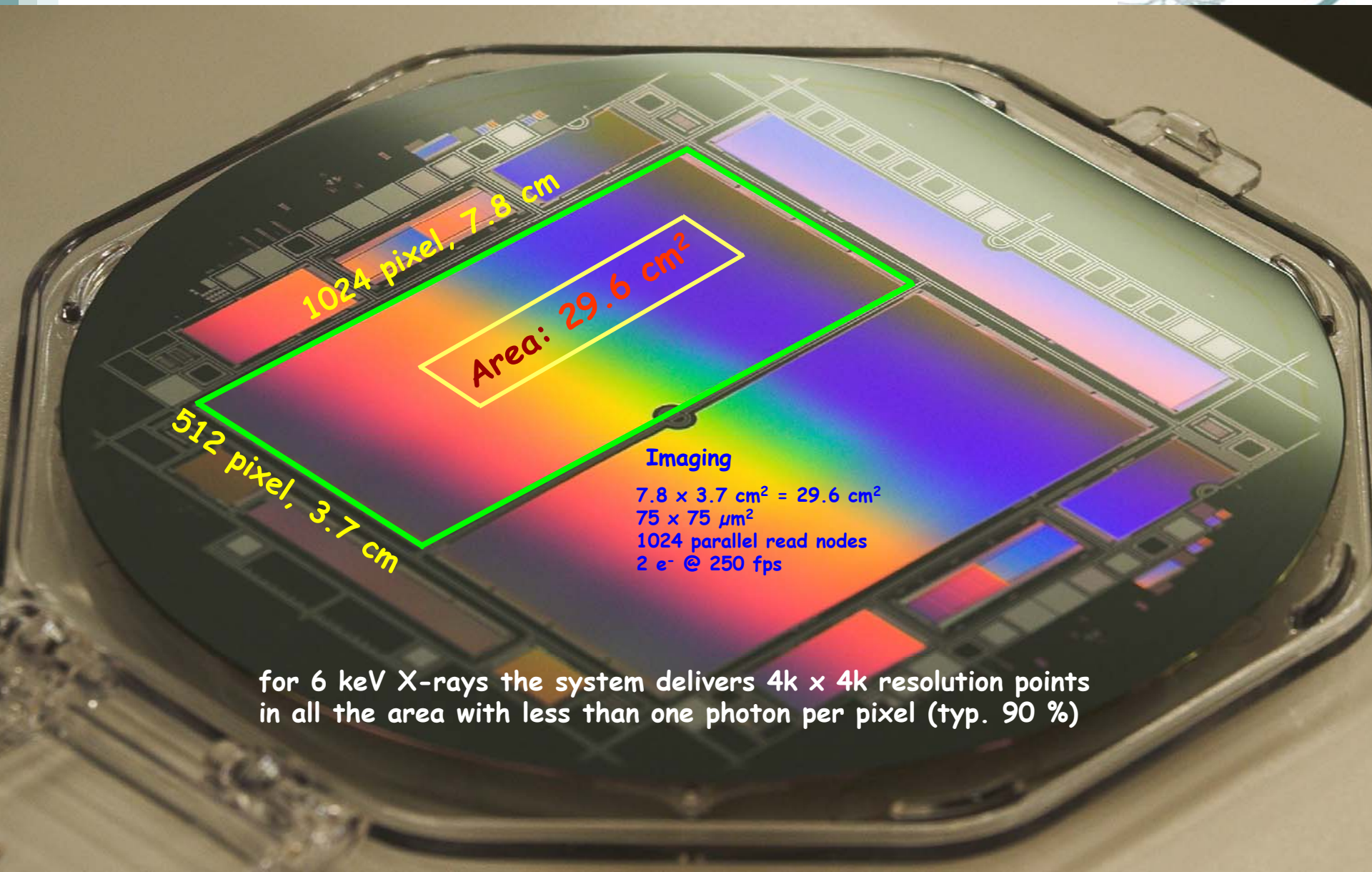
communication

CAMEX eROSITA 128I-JD





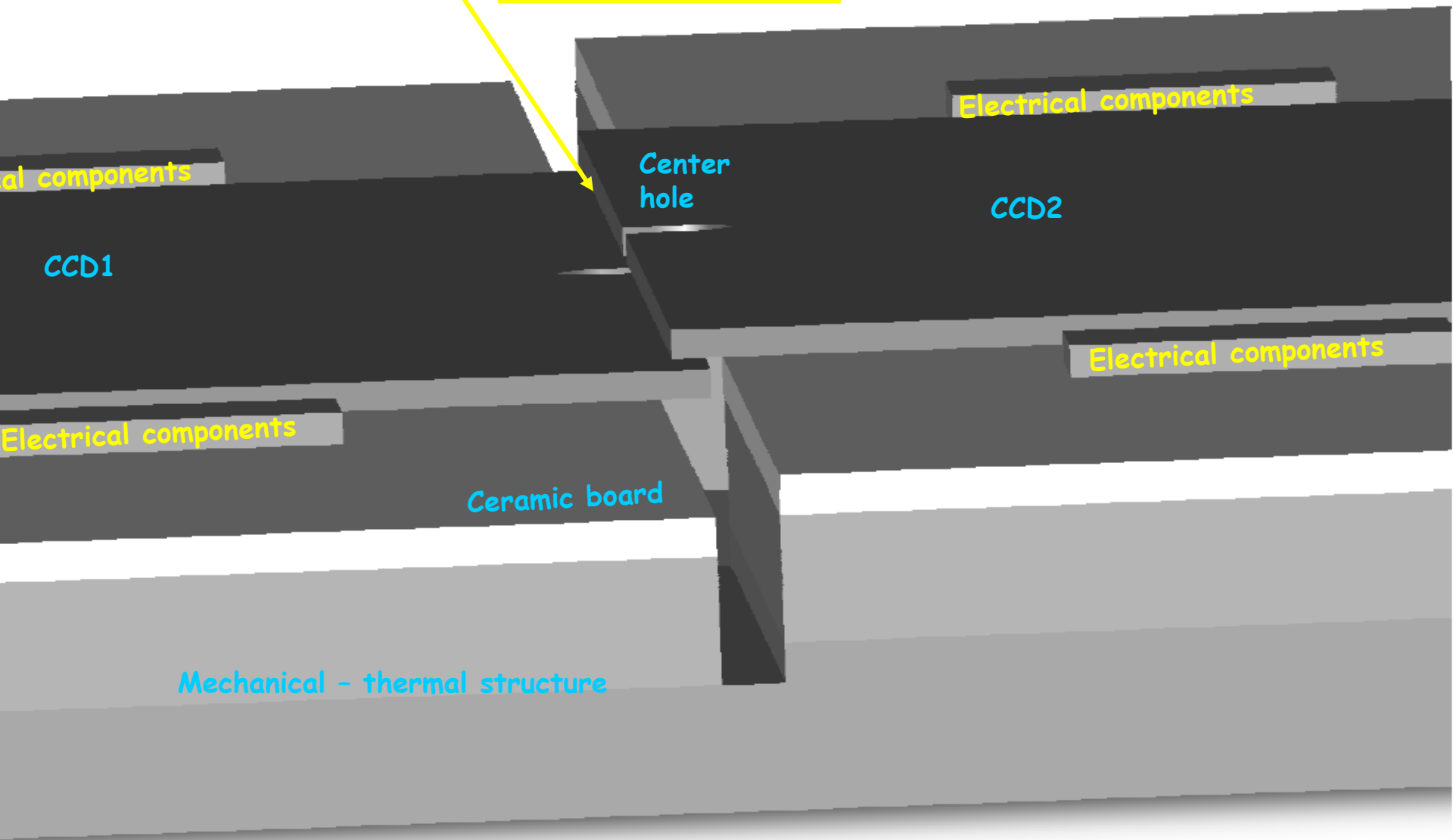
pnCCD: 1024 x 512, 30 cm²

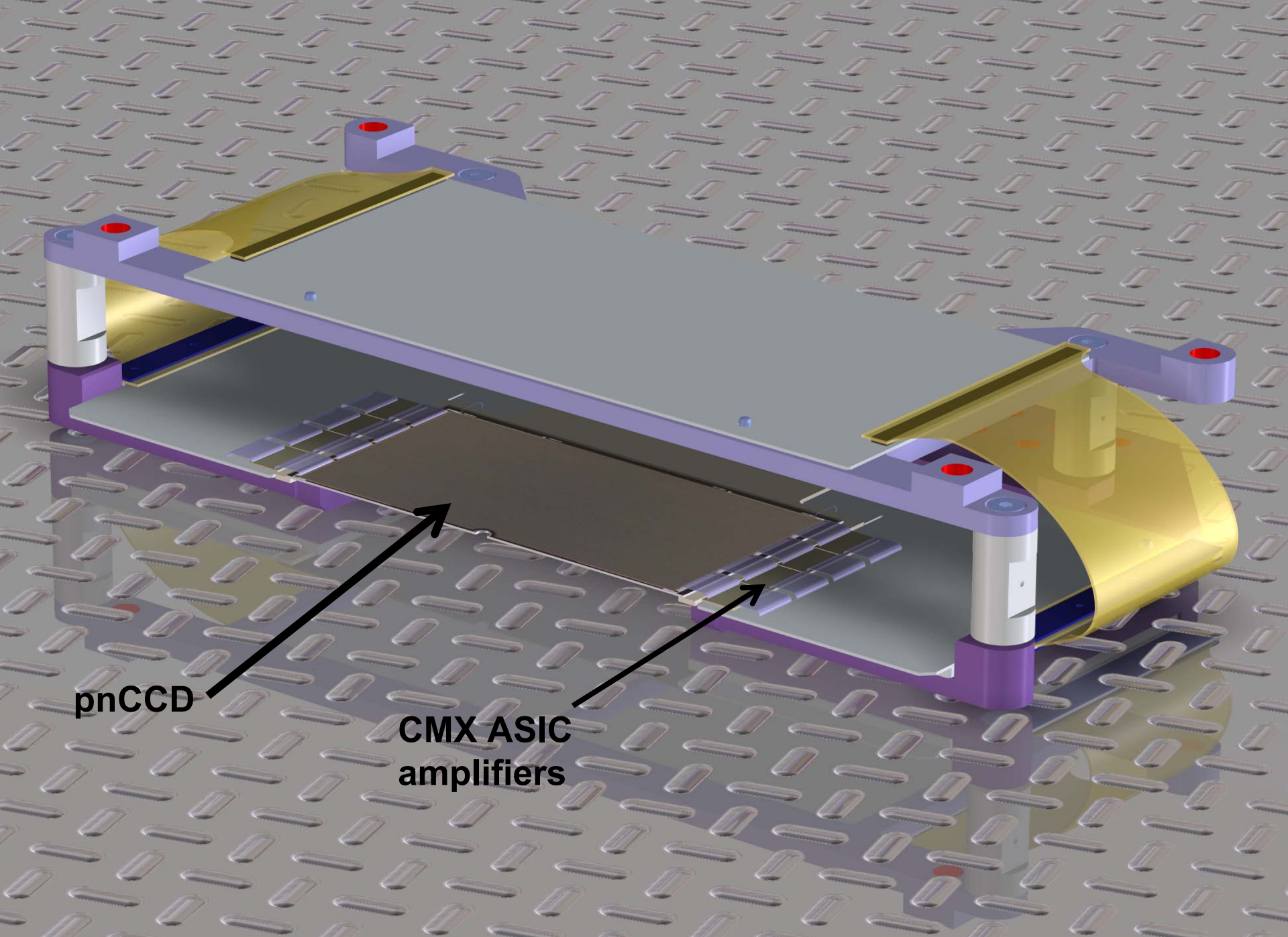


for 6 keV X-rays the system delivers 4k x 4k resolution points
in all the area with less than one photon per pixel (typ. 90 %)

pnCCDs overlapping in the center

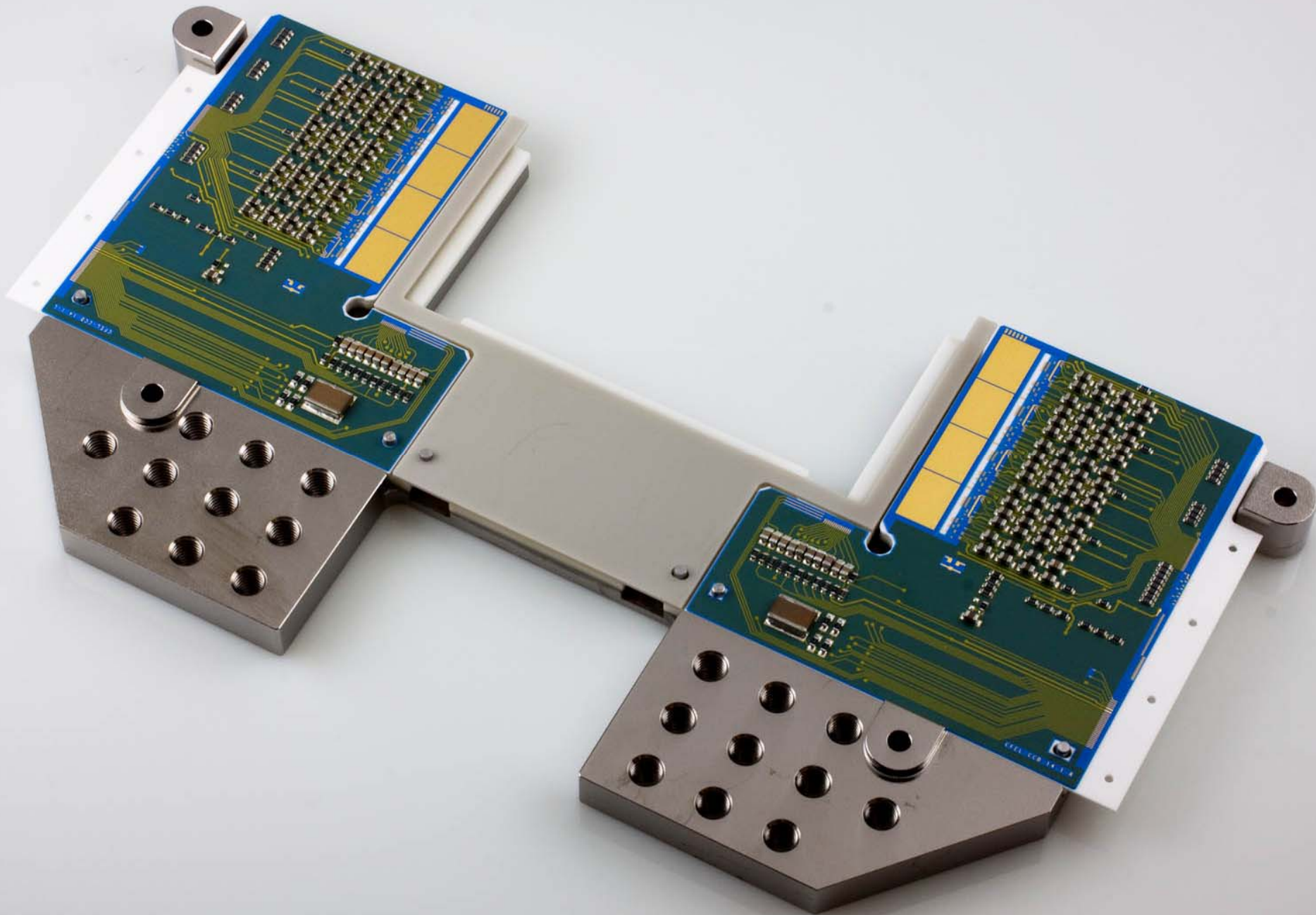
800 μm insensitive gap



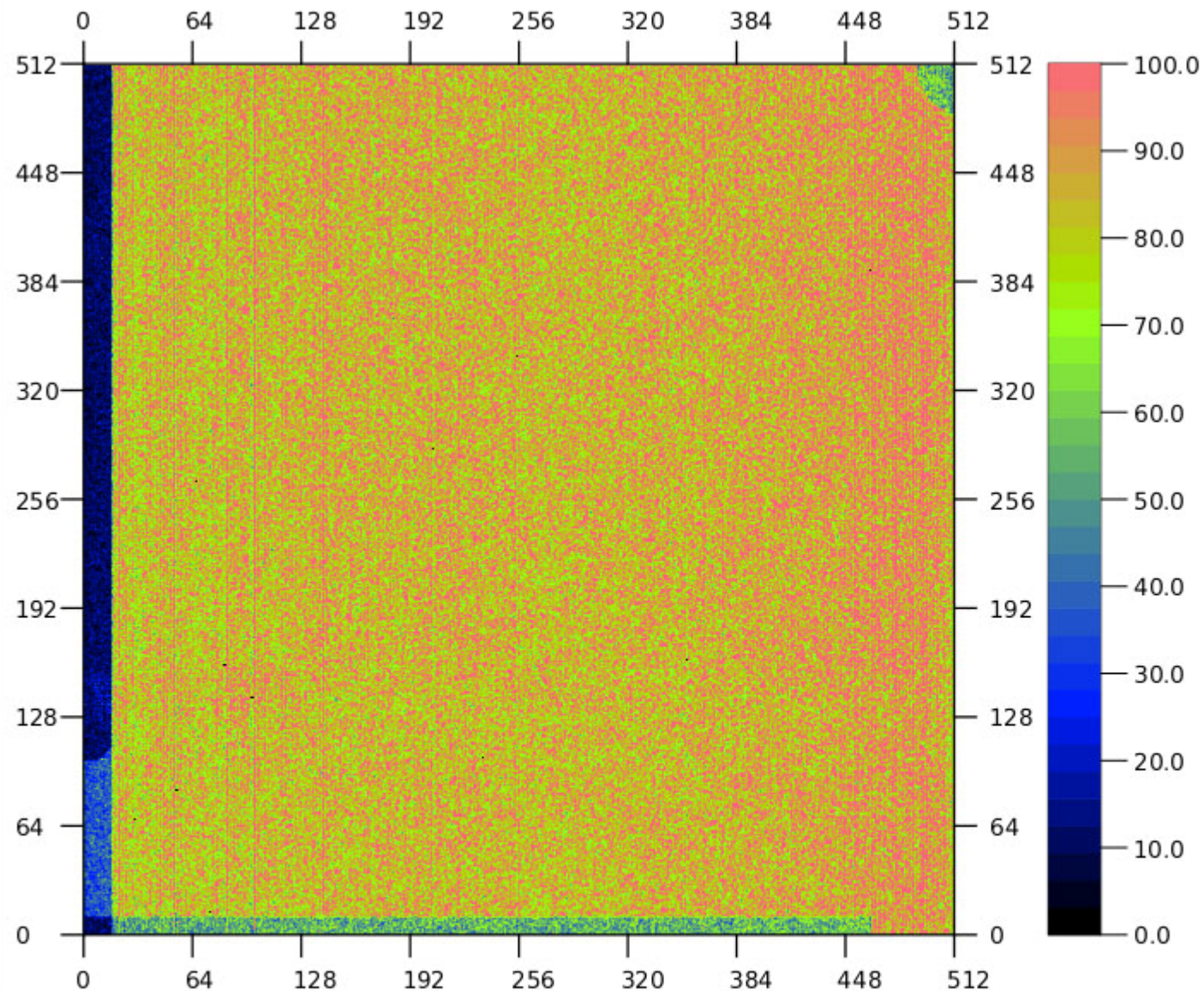


pnCCD

CMX ASIC
amplifiers







**4 "halves" (1024x1024)
were successfully tested**

**operation in the coherence
experiments starts on
August – 5, 7 am (CET)**

**noise floor at – 40 ° C
ENC = 6 el. (rms)**

QE @ 2 keV = 98 %

Frame rate: up to 150 Hz

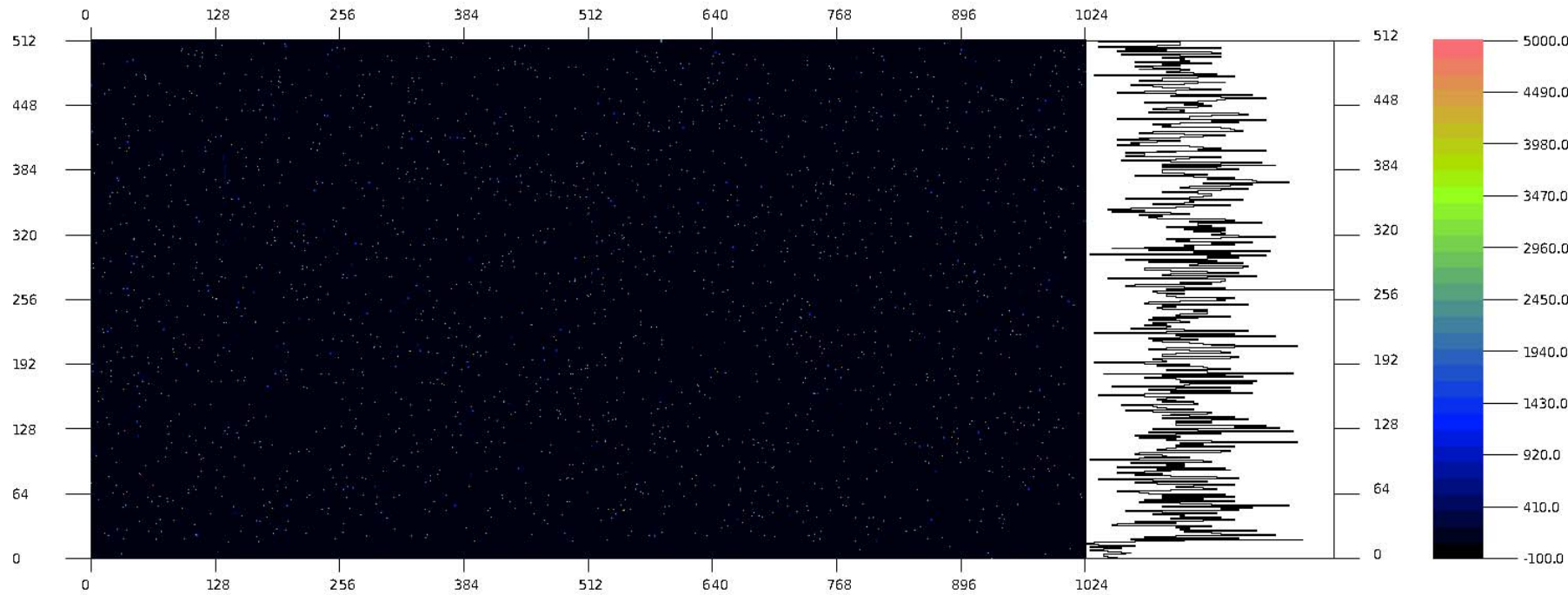
Zoom Out

☐ Auto Range Lower Cut: 0

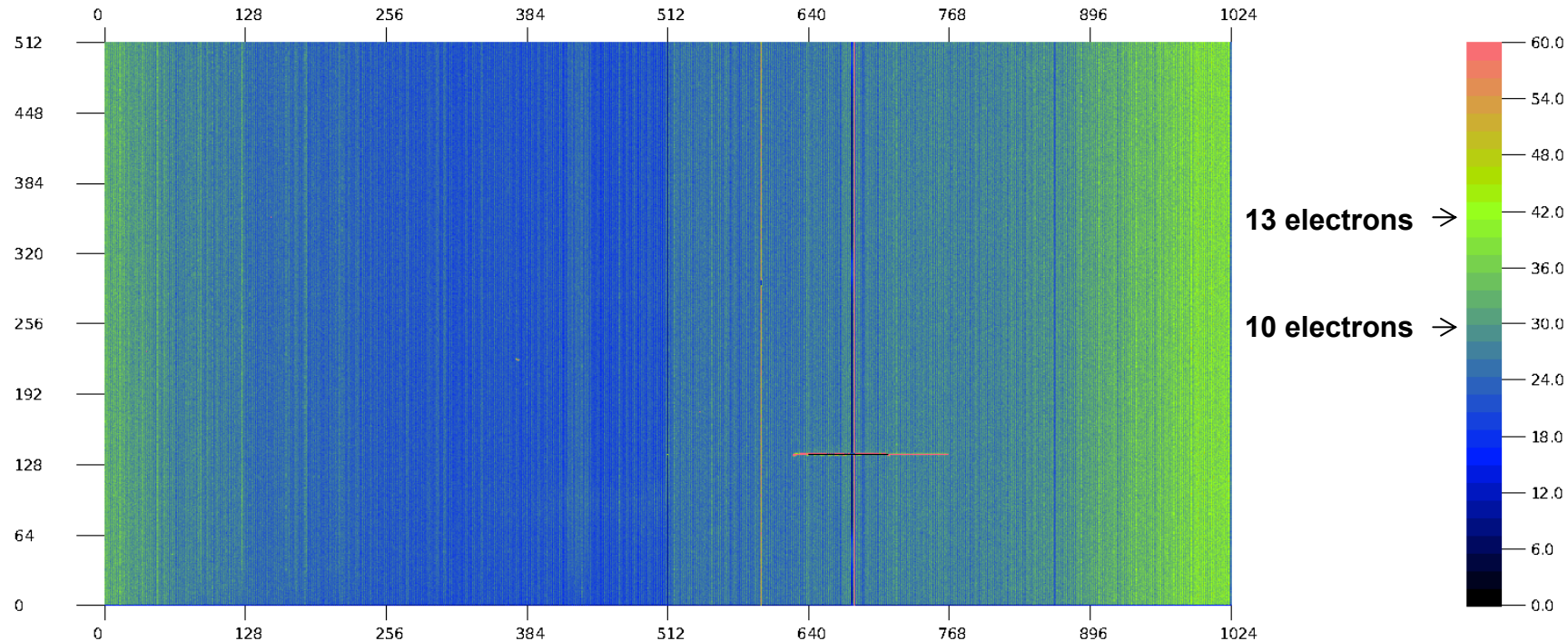
Upper Cut: 100

Clear

☒ Integrate Event Counts ☐ Show Pulse Height

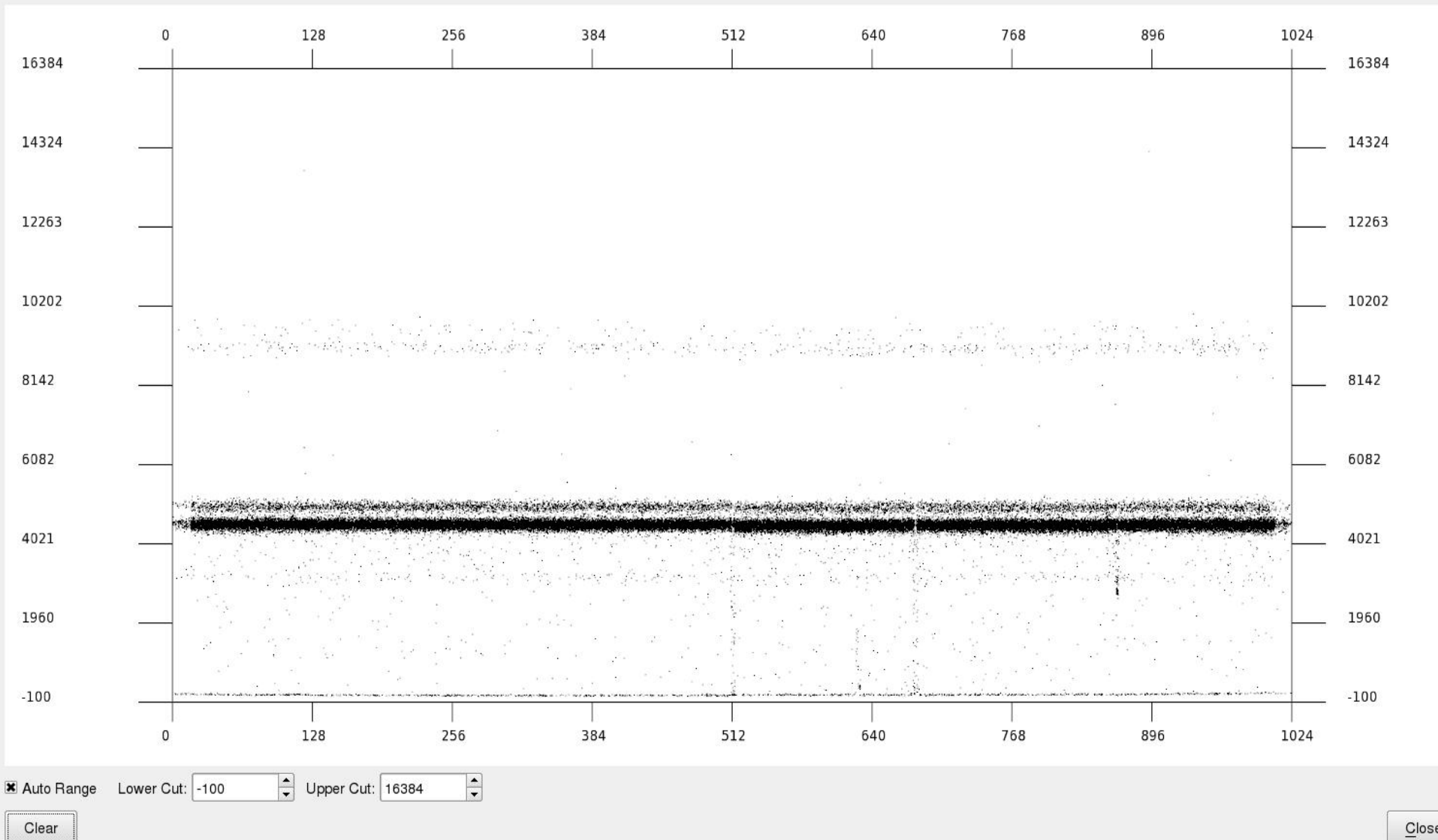


CAMP pnCCD module performance (typical module)

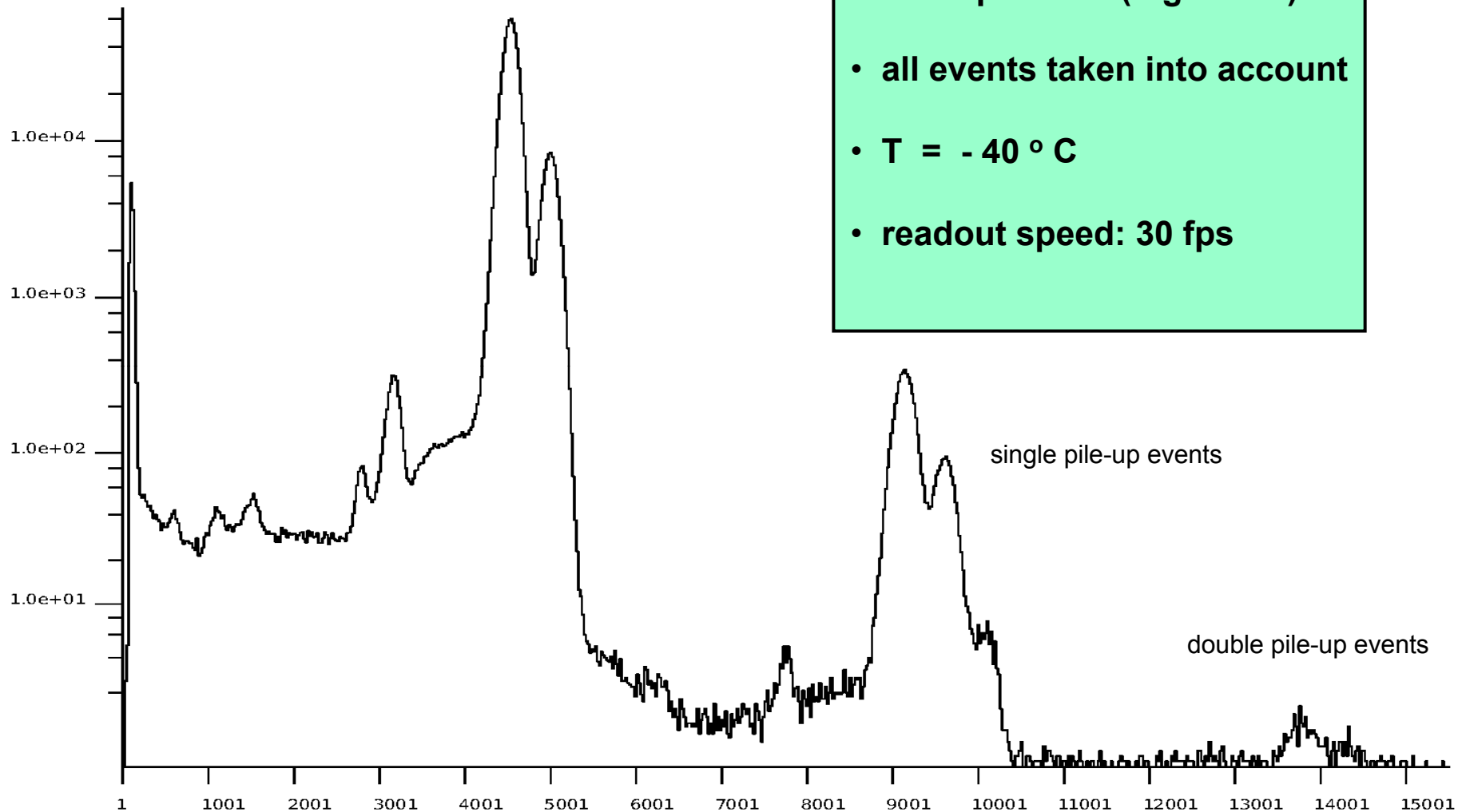


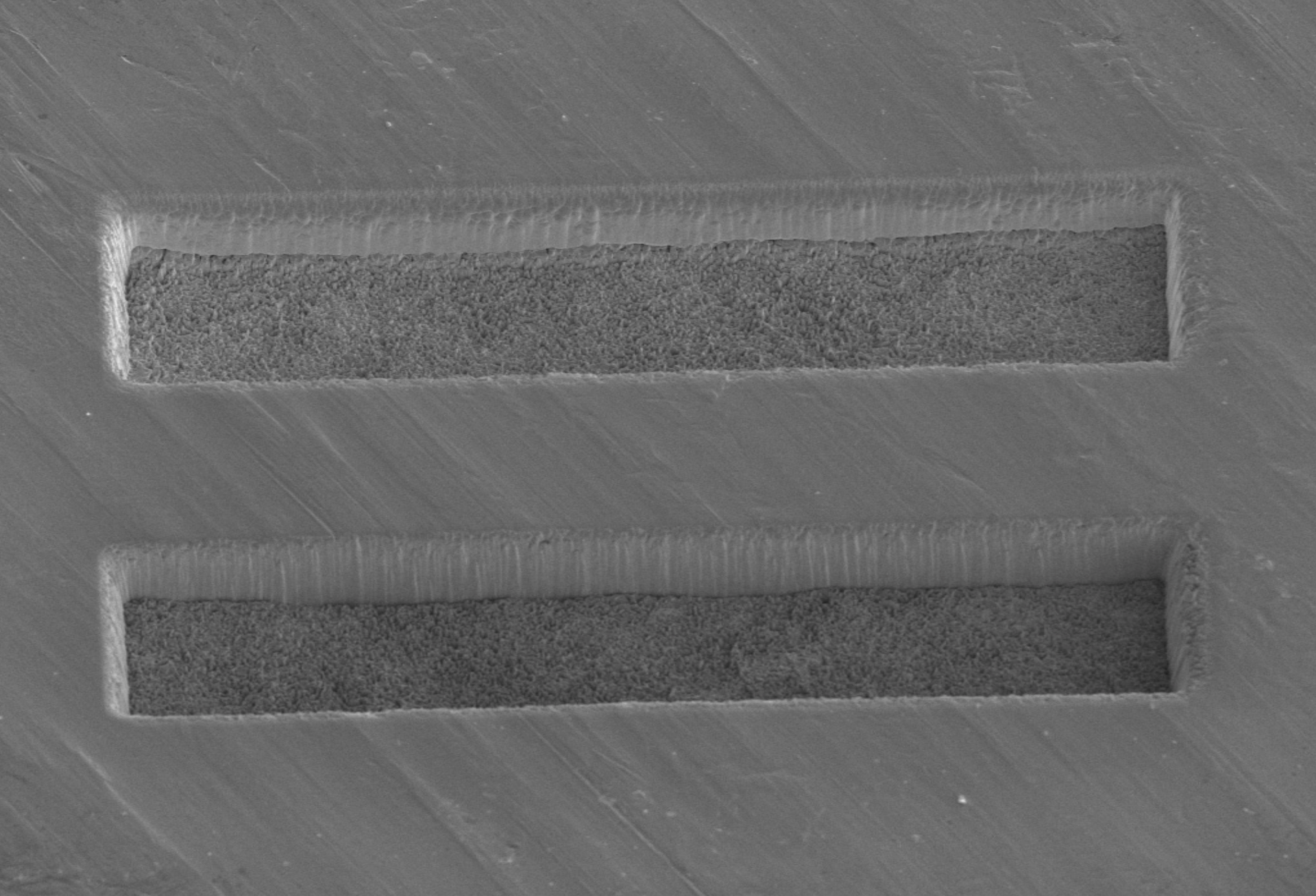
- 588 noisy or bad pixel out of 524.288 pixel
- average noise is 9 electrons (rms) at -40°C
- min. X-ray energy: 850 eV, i.e. 240 electrons. $S/N = 27/1$

CAMP pnCCD module performance (typical module)



CAMP pnCCD module performance (typical module)





	3/2/2009	HV	WD	tilt	mag	spot	 40 μ m
	1:23:19 PM	3.00 kV	10.7 mm	25 °	1 200 x	5.0	

GEANT4 simulations

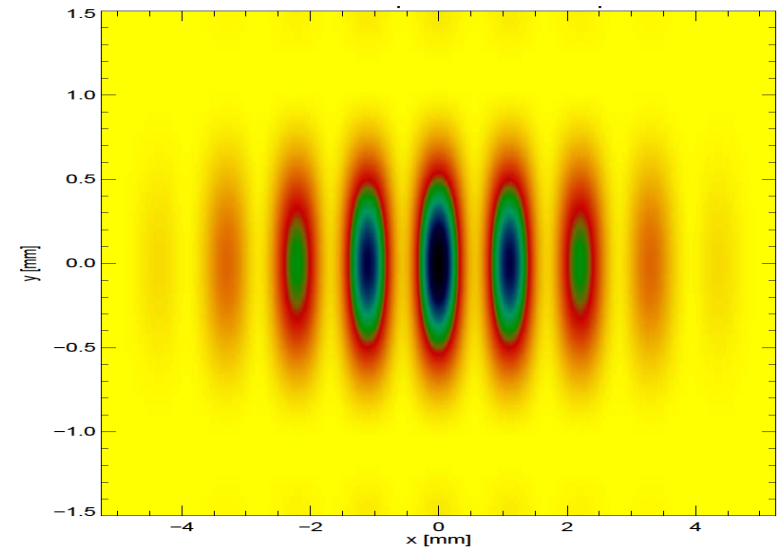
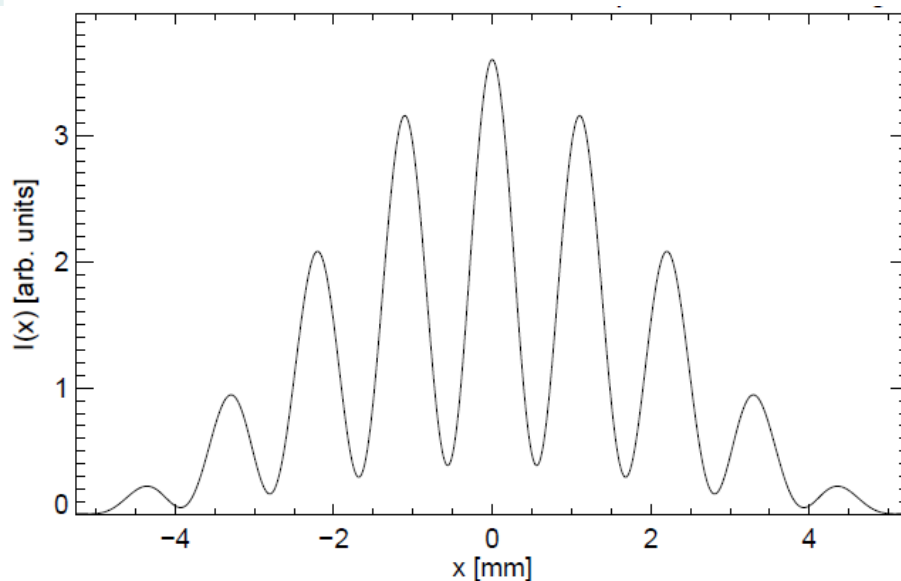
Upcoming FLASH Experiments: Double Slit Beam Geometries



Implemented according to
Singer et al., 2008:

$$I(P) = I_1(P) + I_2(P) + 2\sqrt{I_1(P)I_2(P)}|\gamma_{12}(\tau)| \times \cos[\omega\tau - \alpha_{12}(\tau)],$$

Slit size $10 \mu\text{m} \times 50 \mu\text{m}$, $d = 8 \text{ m}$, $\gamma = 0.8$, $\lambda = 7 \text{ nm}$

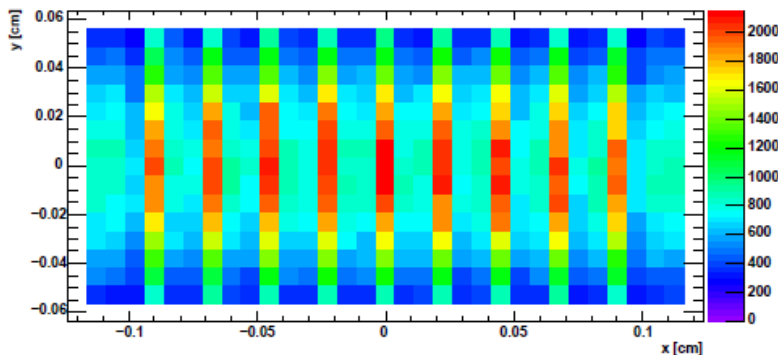


Slit distance $50 \mu\text{m}$

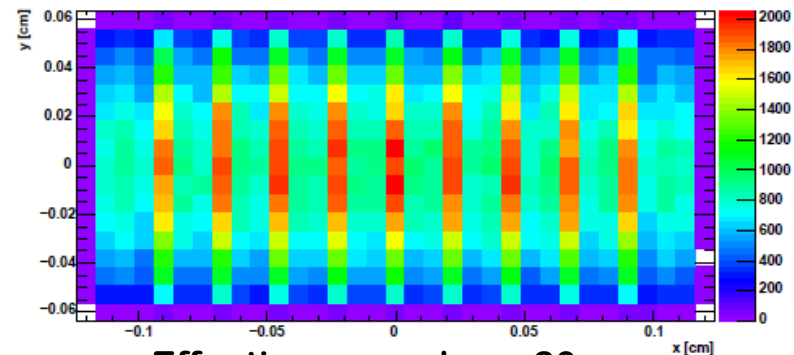
Upcoming FLASH Experiments: Pixel Distribution of Measured Photons

Spread of signal electrons depends on:

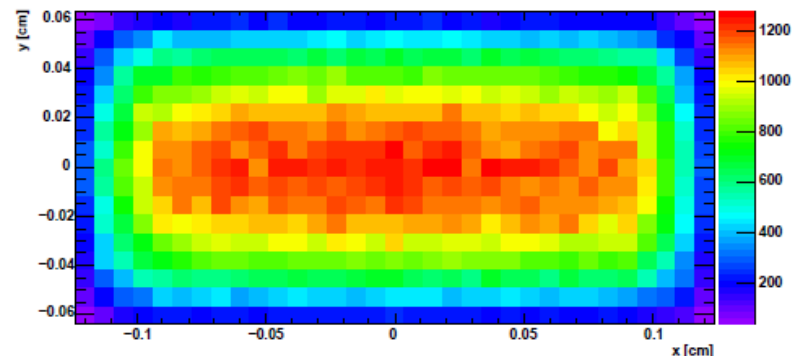
- initial electron number and density
- number of electrons collected per pixel
(overflow if charge handling capacity exceeded)



No spread, ideal detector



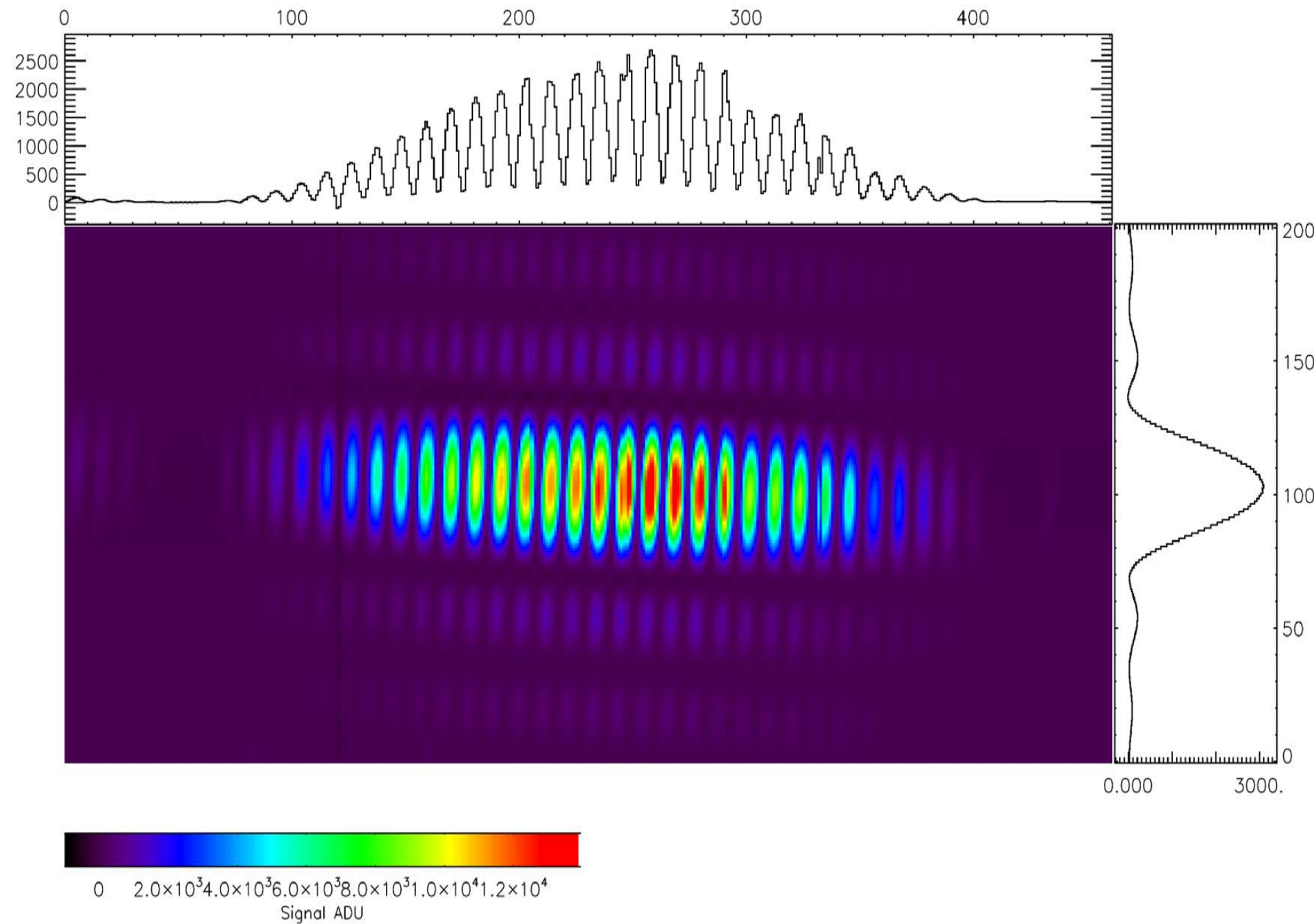
Effective spread: $\sigma = 20 \mu\text{m}$

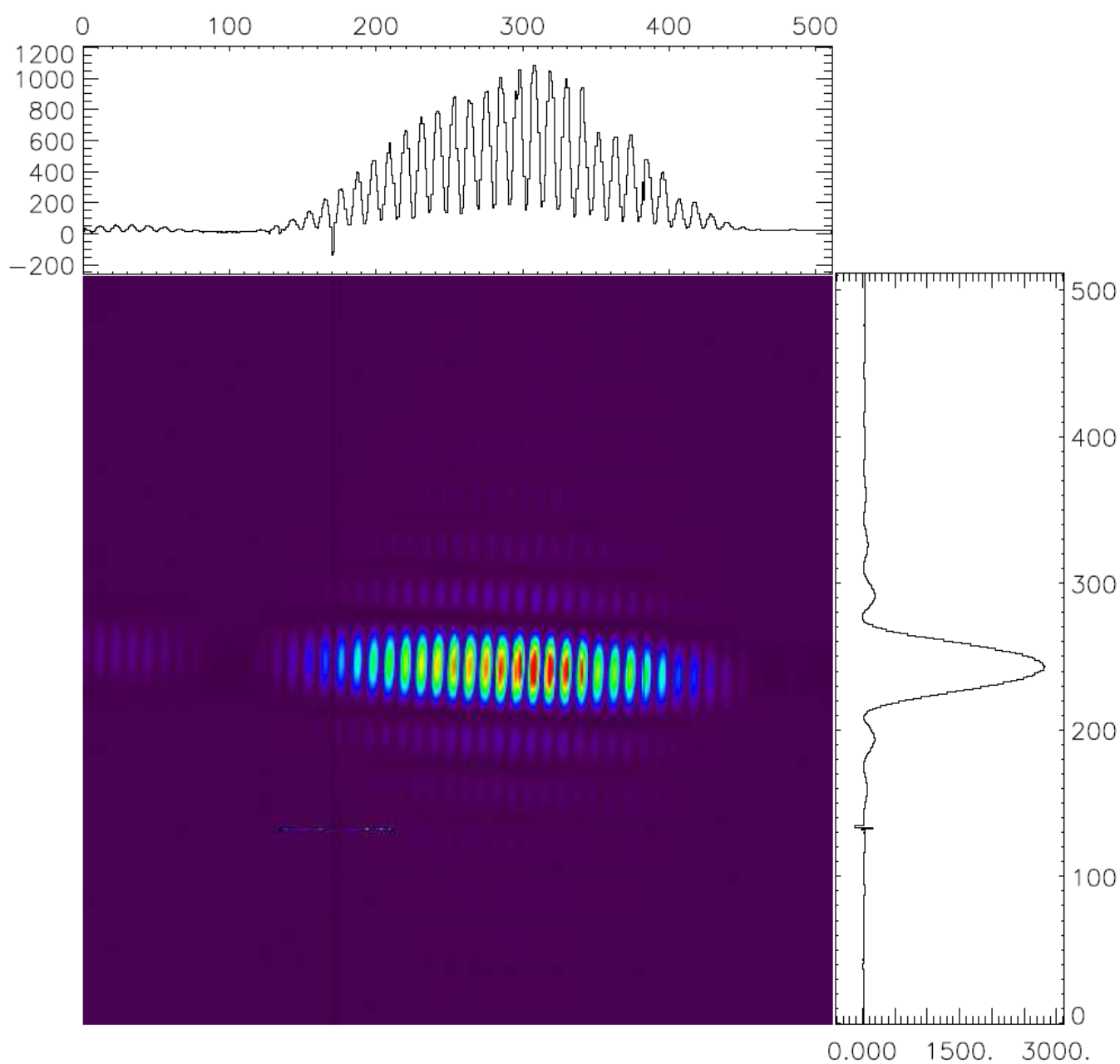


Effective spread: $\sigma = 100 \mu\text{m}$ ("overflow")

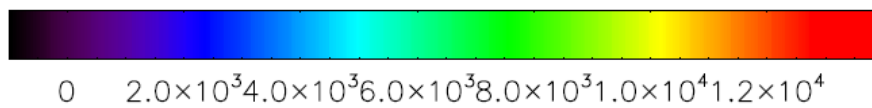
Slit size $10 \mu\text{m} \times 50 \mu\text{m}$,
 $d = 8 \text{ m}$, $\gamma = 0.8$, $\lambda = 7 \text{ nm}$

Pattern for 10^{12} photon pulse,
unfocused beam





beam energy: 90 eV
double slit: 10 x 50 μm^2

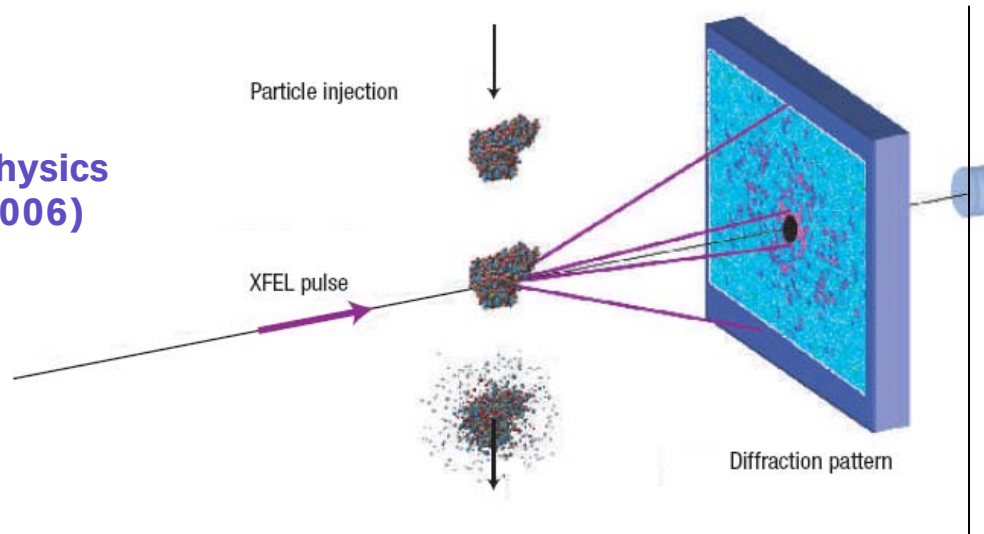


Time resolved structure determination with X-ray FELs

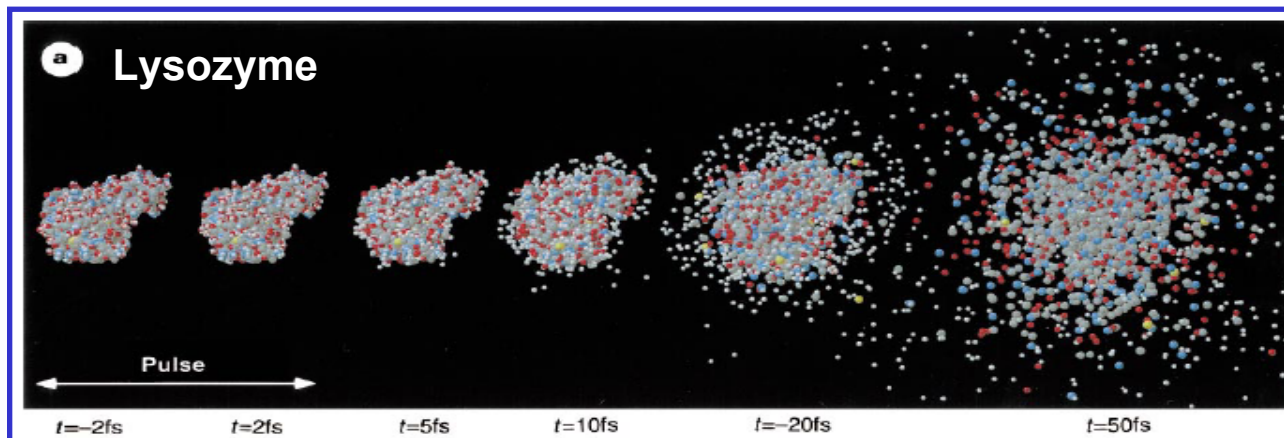
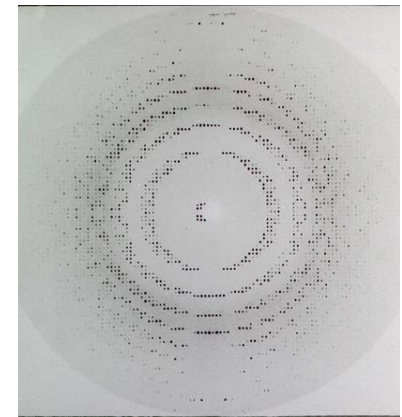


Molecules atomic resolution

J. Kirz,
Nature Physics
2, 799 (2006)



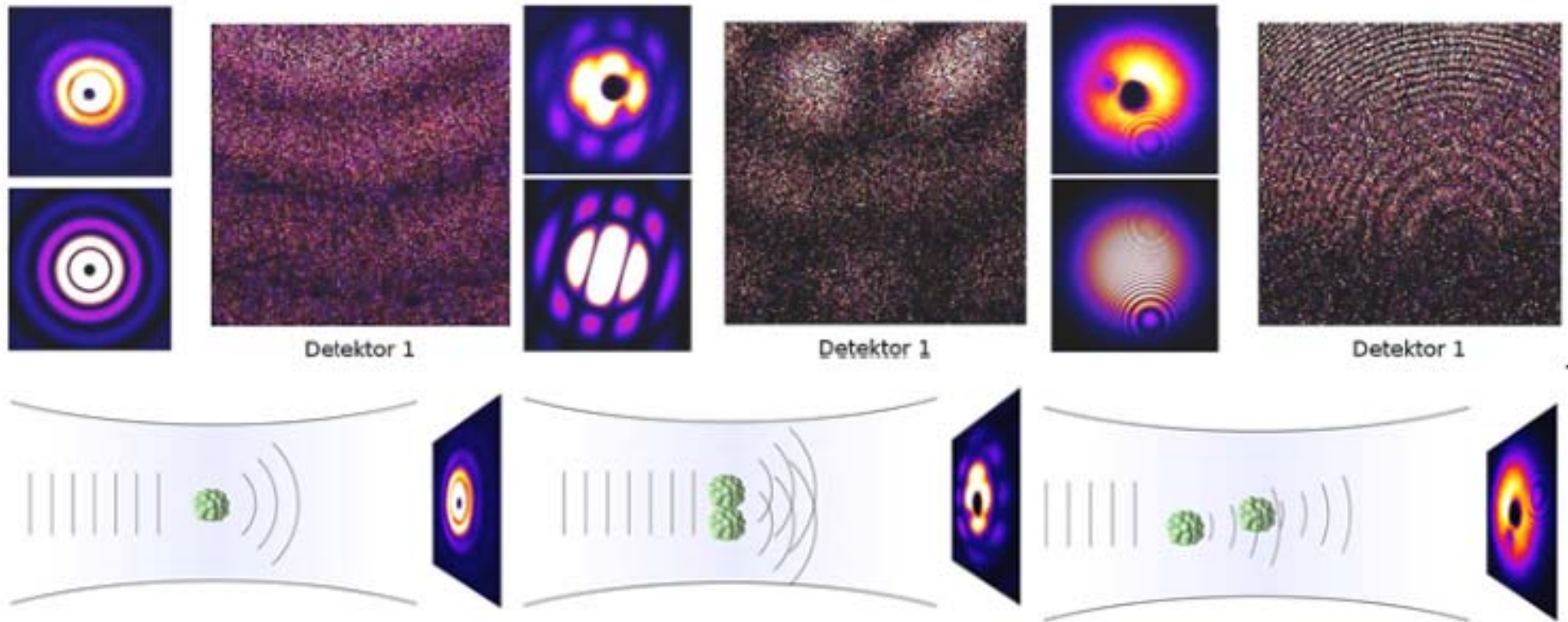
Crystal



R. Neutze, J.
Haidu et al.,
Nature 406, 752
(2000)

**Radiation
damage
and Coulomb
explosion**

X-ray scattering: 30 nm Xe - Cluster

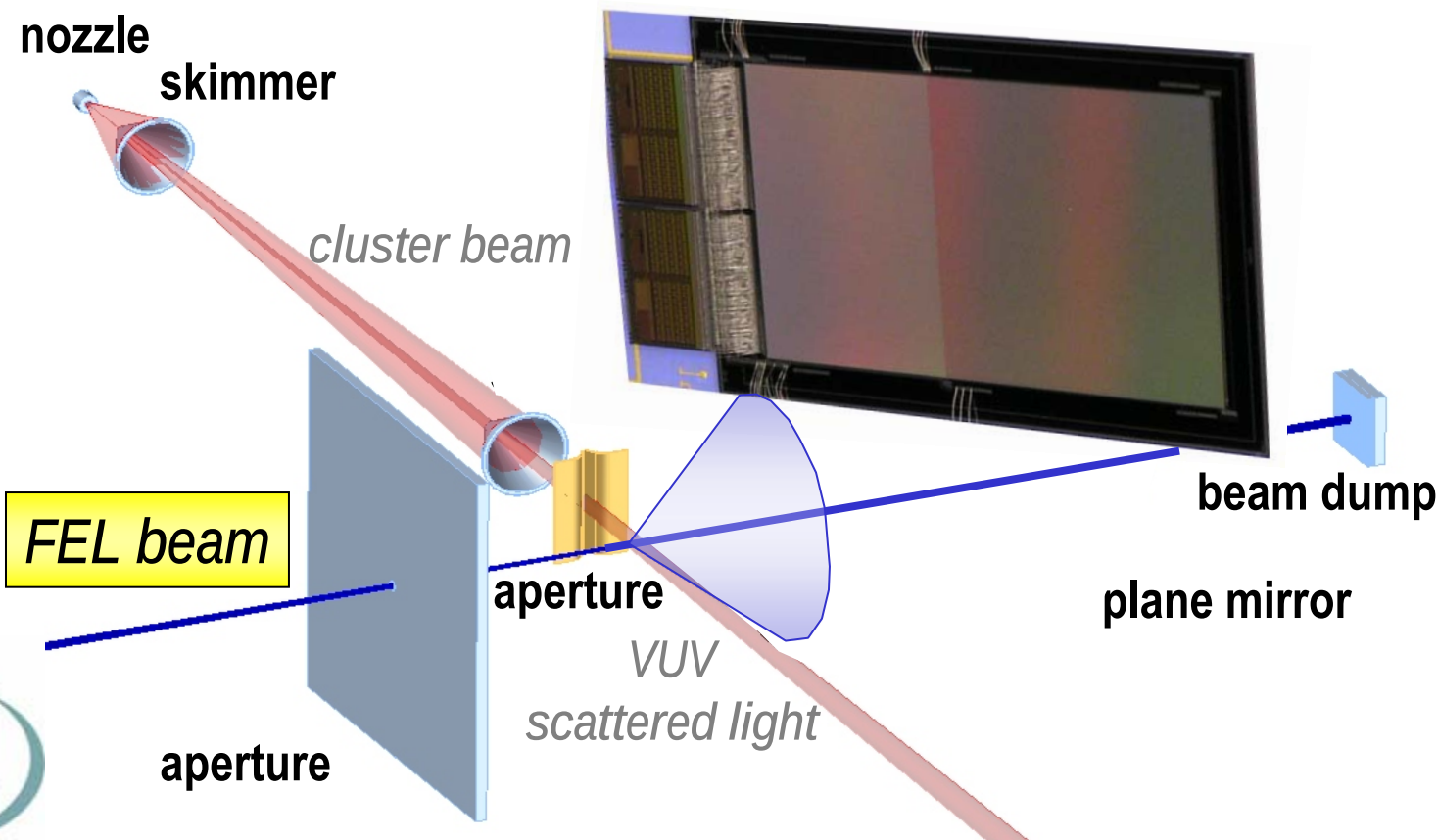


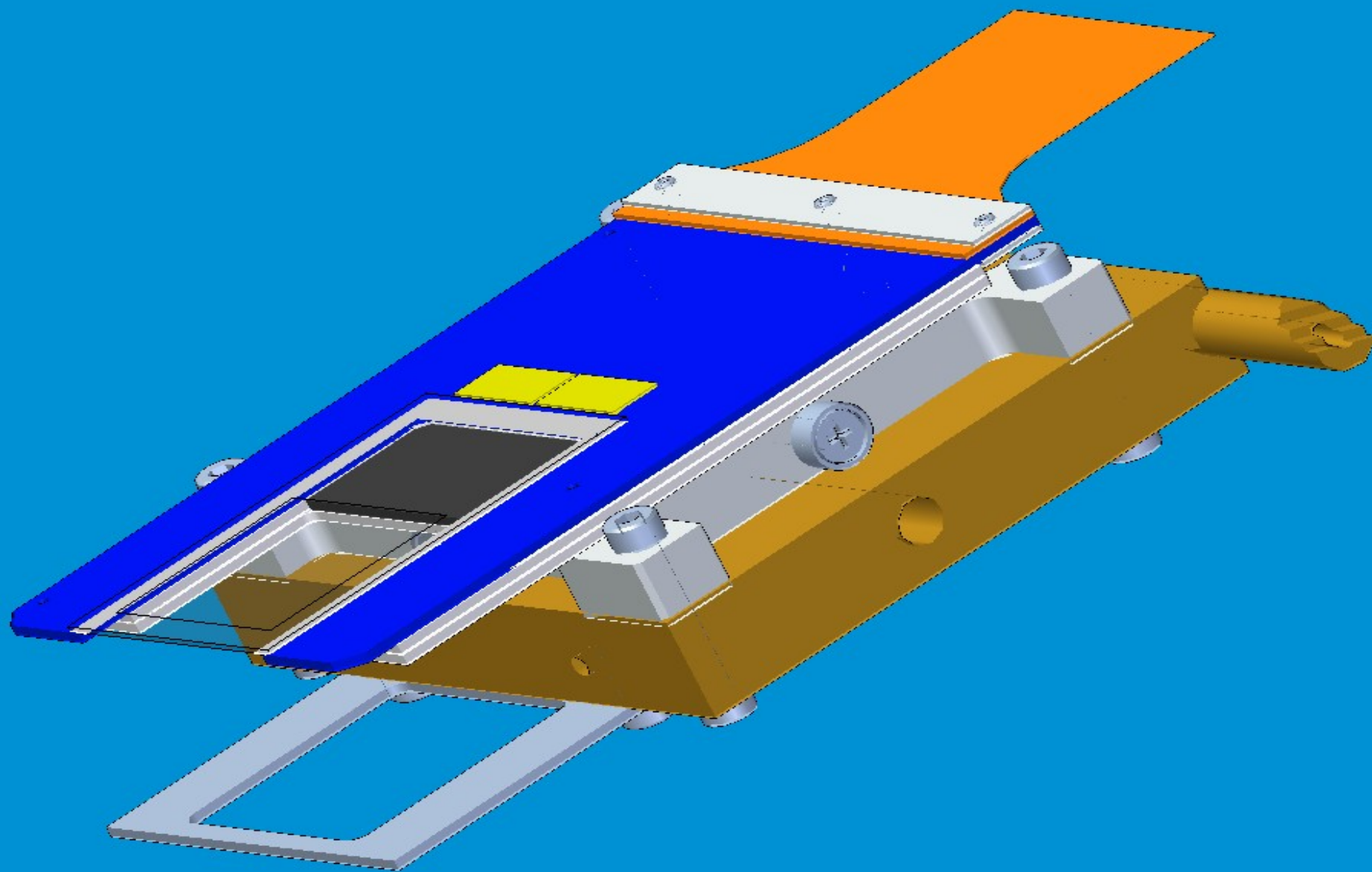


Clusters “in the Flight”

Xe clusters

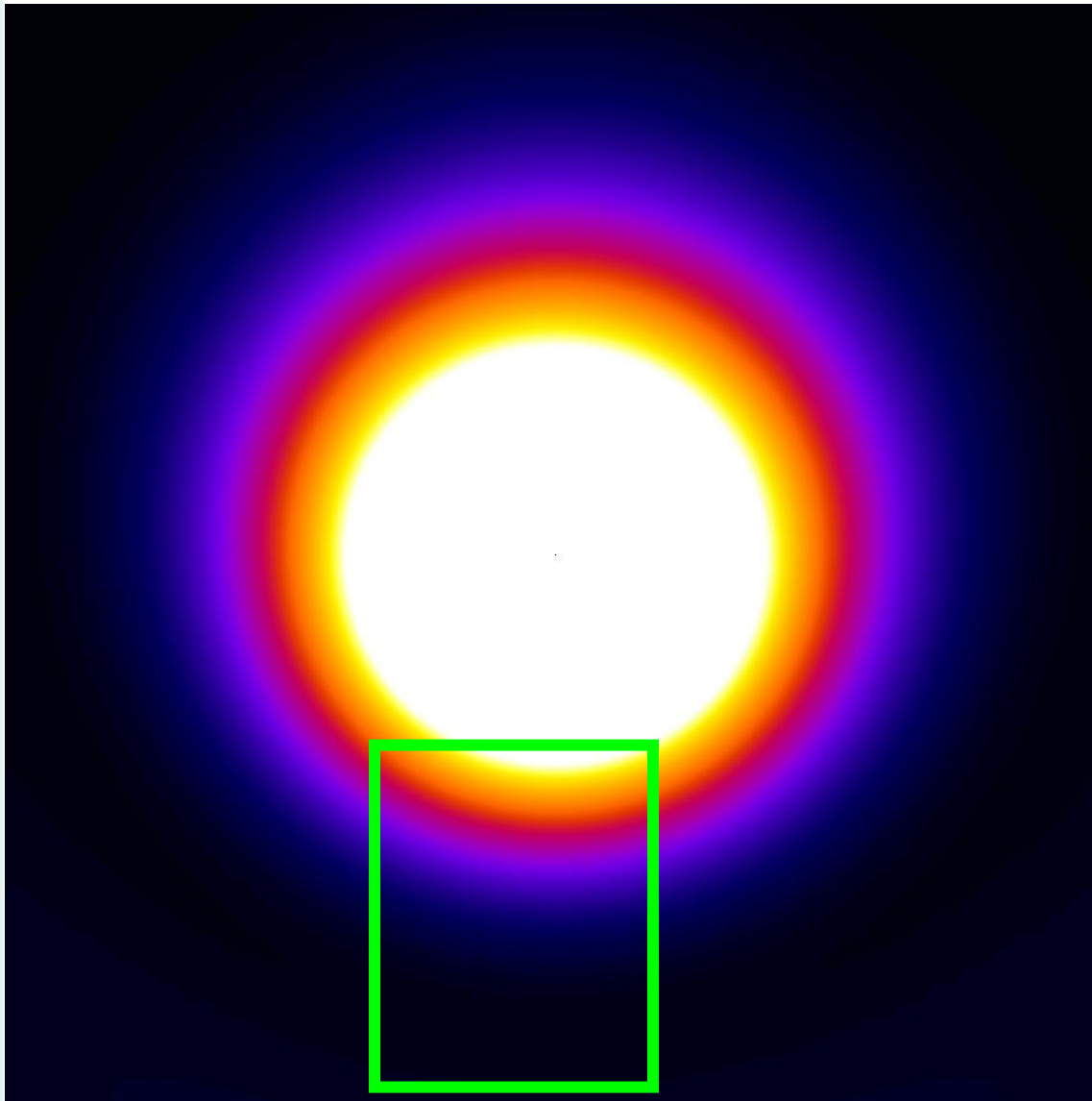
CFEL Detectors





First observation of 32 nm Xe - nanocluster in direct detection mode

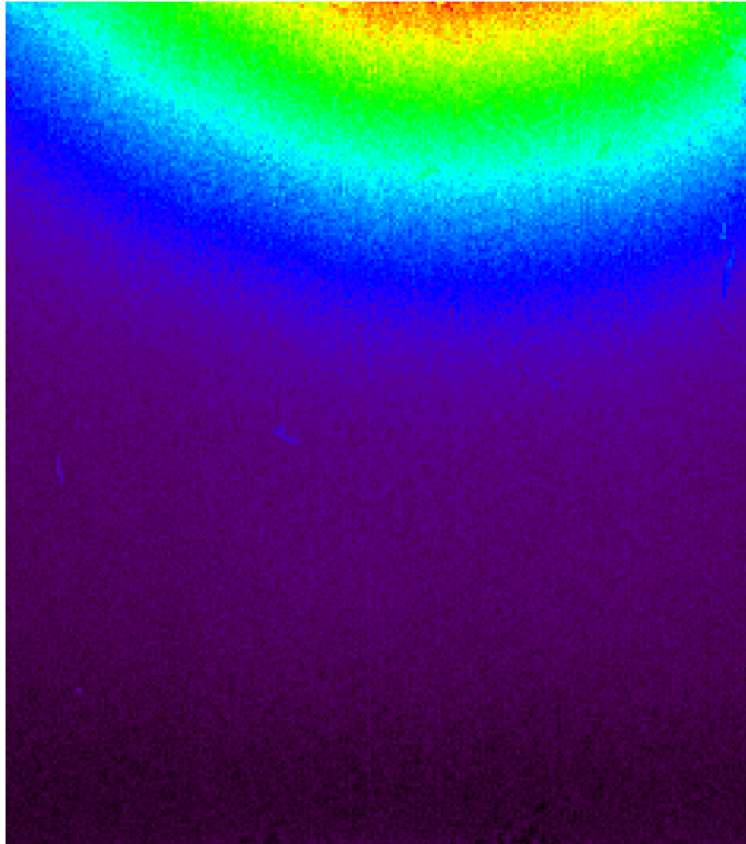
MPI
Halbleiterlabor



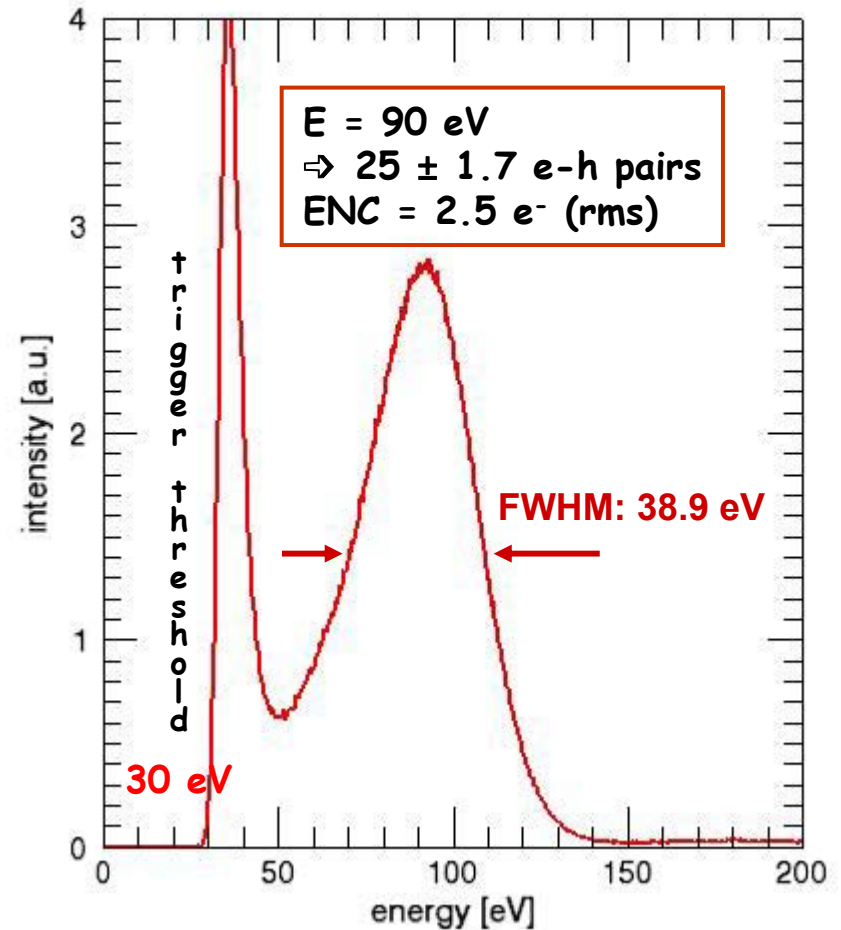
simulation of 90 eV X-rays
scattering at single Xe
nanoclusters with 32 nm size
(C. Bostedt et al.)

field of view of the pnCCD
during the experiment at
the FLASH X-ray free electron
laser facility at DESY, Hamburg

90 eV X-rays in single photon counting mode !!!

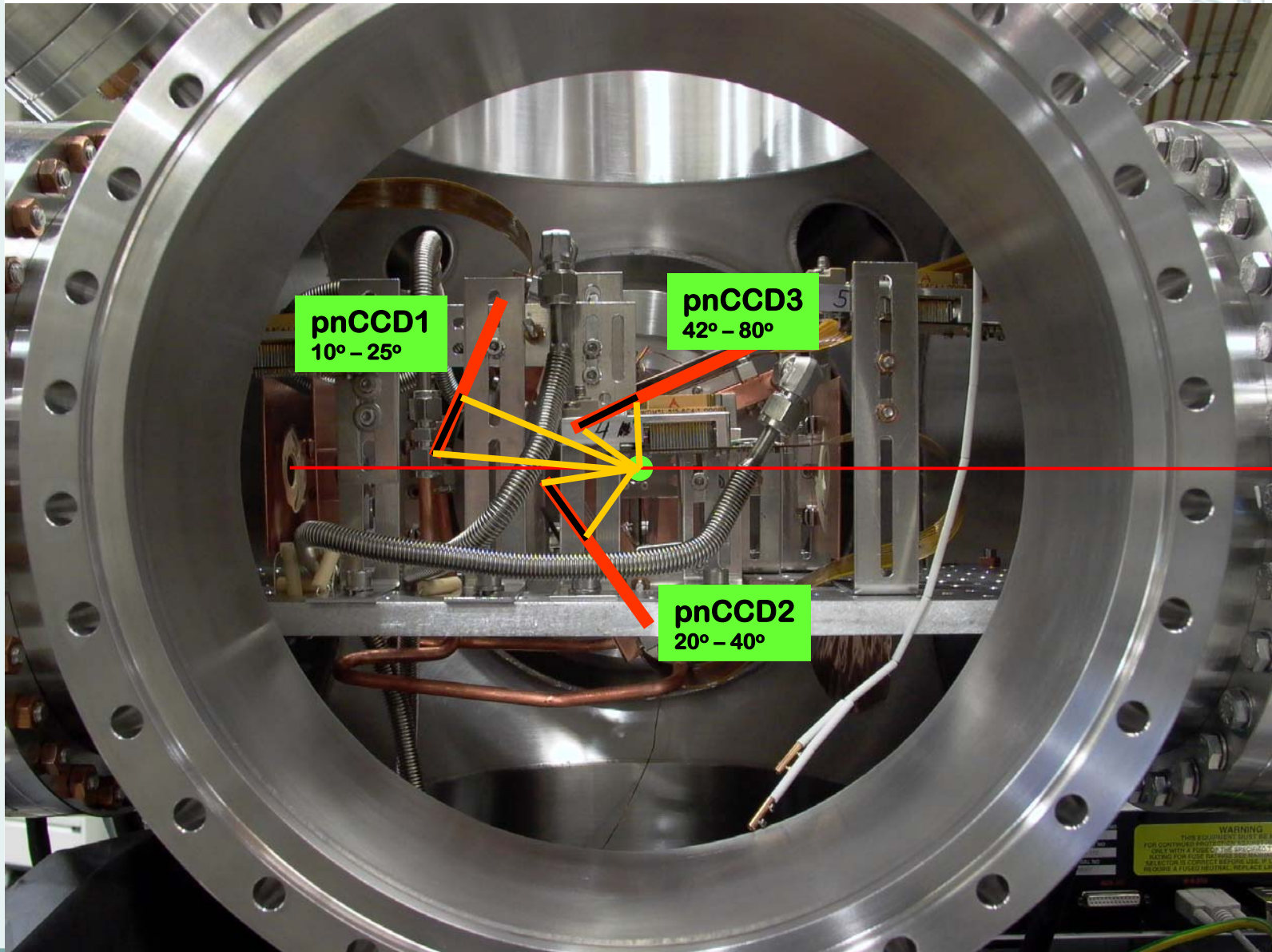


$T = -50^{\circ}\text{C}$



Spectrum from 4.000 frames
with 0.01 photons/pixel/frame

Set-up in T.M. Beamline @ FLASH



WARNING
THIS EQUIPMENT MUST BE
FOR CONTINUOUS PROTECTION
ONLY WITH A FUSE CIRCUIT
ACTING FROM AC POWER
REFLECTION TO CORRECT REPAIRS
NEEDING A TYPED INSTRUCTIONS



Clusters “in the Flight”



MPI-HLL Detectors

FLASH

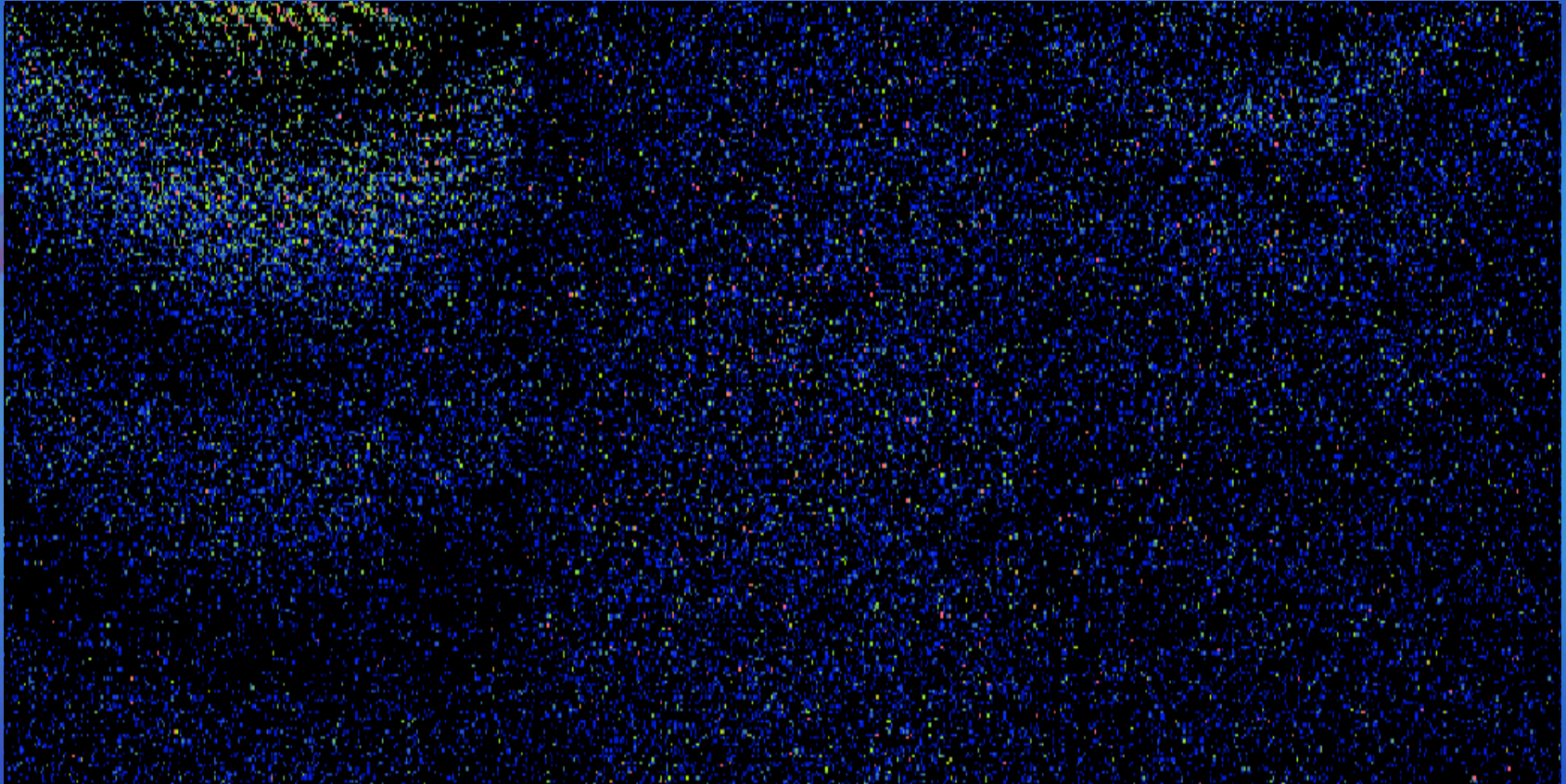


MPI-mF Samples

Collaboration: TU Berlin, MPI-HLL, MPI-K:
Bostedt, Rupp, Adolph, Möller, Hartmann,
Strüder, Rudenko, et al.



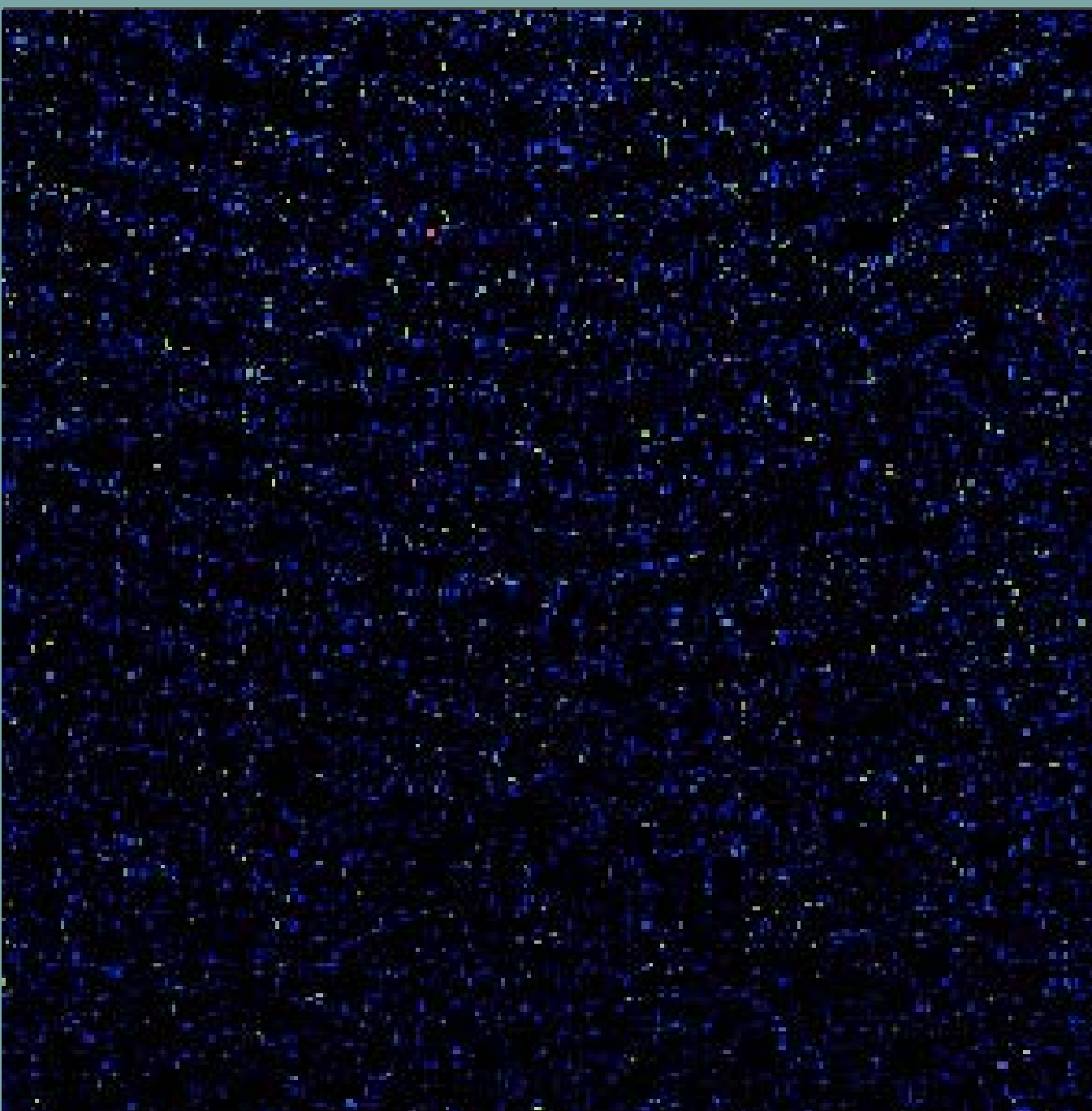
Free Electron Lasers



typ. 5.000 X-rays
per shot per detector,
i.e. ≈ 0.1 photon per pix.

Screen-Shots: To be evaluated!





Measurement @ FLASH
 $\lambda \approx 40 \text{ nm}$
on Xe clusters

E = 30 eV

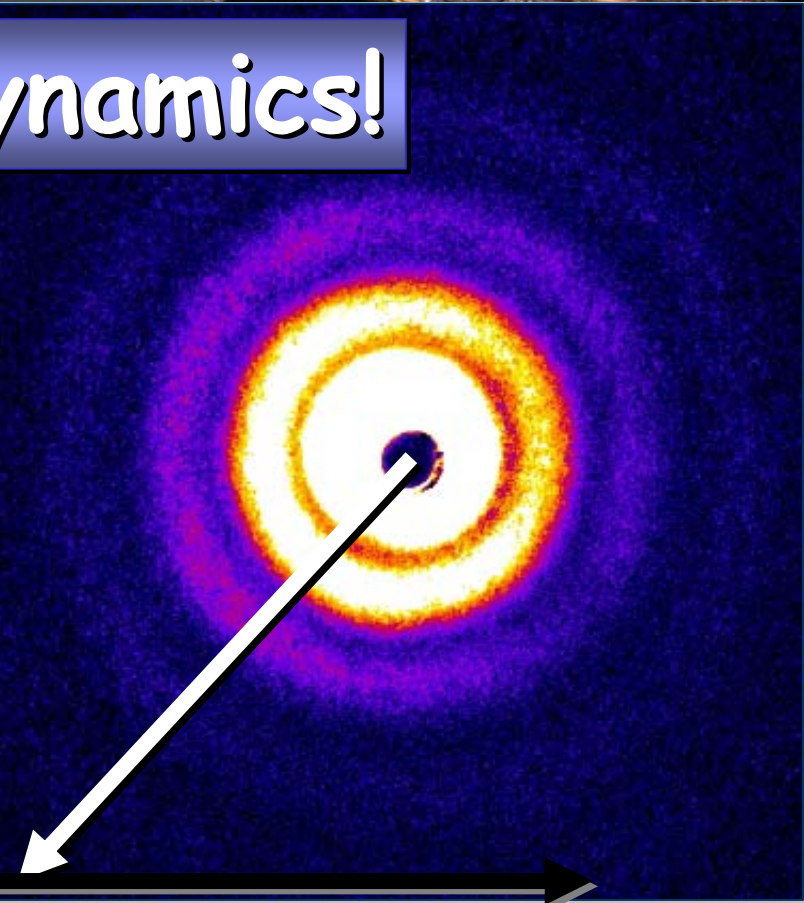
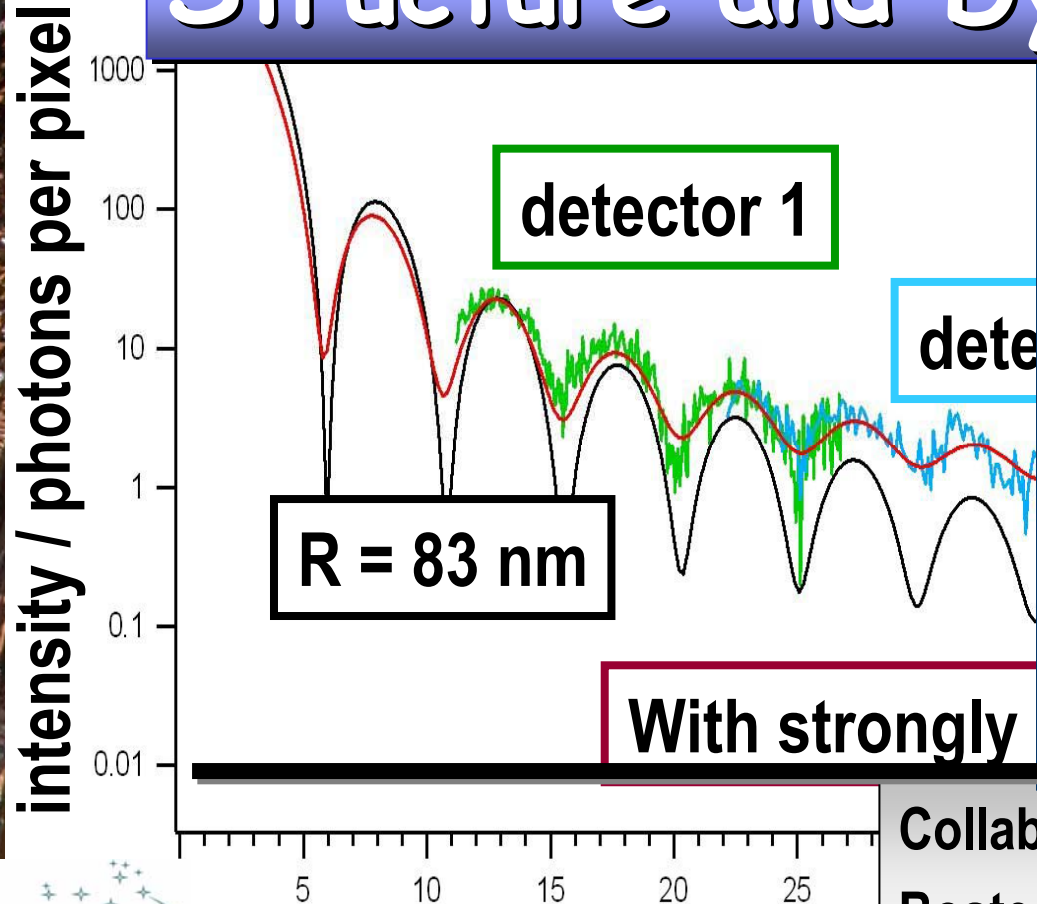
i.e. 8.1 e/h pairs

ENC = 2.7 el. (rms)



Clusters “in the Flight”

Structure and Dynamics!



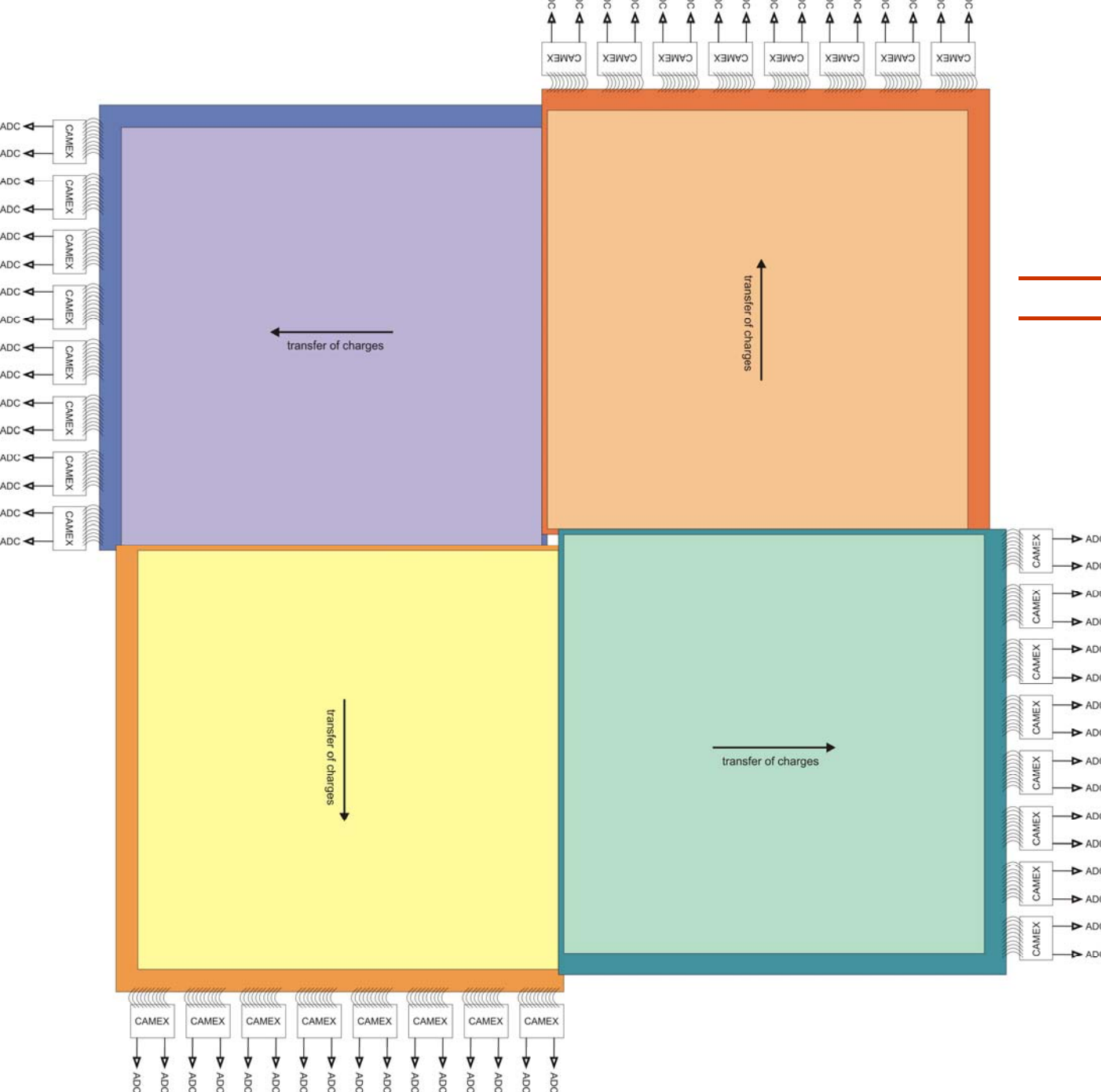
Collaboration: TU Berlin, MPI-HLL, MPI-K:
Bostedt, Rupp, Adolph, Möller, Hartmann,
Strüder, Rudenko, et al.

Achievements of the FLASH, LCLS and XFEL

Photon Counting and Integrating X-ray Imaging Detectors



	FLASH, LCLS + XFEL	pnCCD system
single photon resolution	yes	yes
energy range	$0.05 < E < 24$ (keV)	$0.05 < E < 25$ [keV]
pixel size (μm)	100	75
sig.rate/pixel/bunch	10^3 (10^5)	$10^3 - 10^4$
quantum efficiency	> 0.8	> 0.8 from 0.6 to 12 keV
number of pixels	512×512 (min.)	1024×1024 and 2048×2048
frame rate/repetition rate	10 Hz - 120 Hz	up to 250 Hz
Readout noise	$< 50 e^-$ (rms)	$< 5 e^-$ (rms) ($2 e^-$ possible)
cooling	possible	around -40°C room temperature possible
vacuum compatibility	yes	yes
preprocessing	no (yes) ?	possible upon request



Expected system properties

pixel size: $50 \times 50 \mu\text{m}^2$
 or $75 \times 75 \mu\text{m}^2$
 resolution: $\sigma_{x,y} \leq 10 \mu\text{m}$
 frame rate: 150 Hz
 noise: 5 el. (rms)
 CHC: 500.000 el.
 PSF: typ. 3×3
 EDpP: 20 MeV
 thickness: 450 μm

Parameter	Expected DSSC performance
Energy range	0.5 ... 25 keV (optimized for 0.5 ... 4 keV)
Number of pixels	1024 x 1024
Sensor Pixel Shape	Hexagonal
Sensor Pixel pitch	~ 204 x 236 μm^2
Dynamic range / pixel / pulse	> 10.000 photons @1 keV
Resolution (S/N >5:1)	Single photon @ 1 keV (5 MHz) Single photon @ 0.5 keV (≤ 2.5 MHz)
Electronics noise	< 40 (25) electrons r.m.s.
Frame rate	1-5 MHz
Stored frames per Macro bunch	≥ 512
Operating temperature	-30°C optimum, RT possible

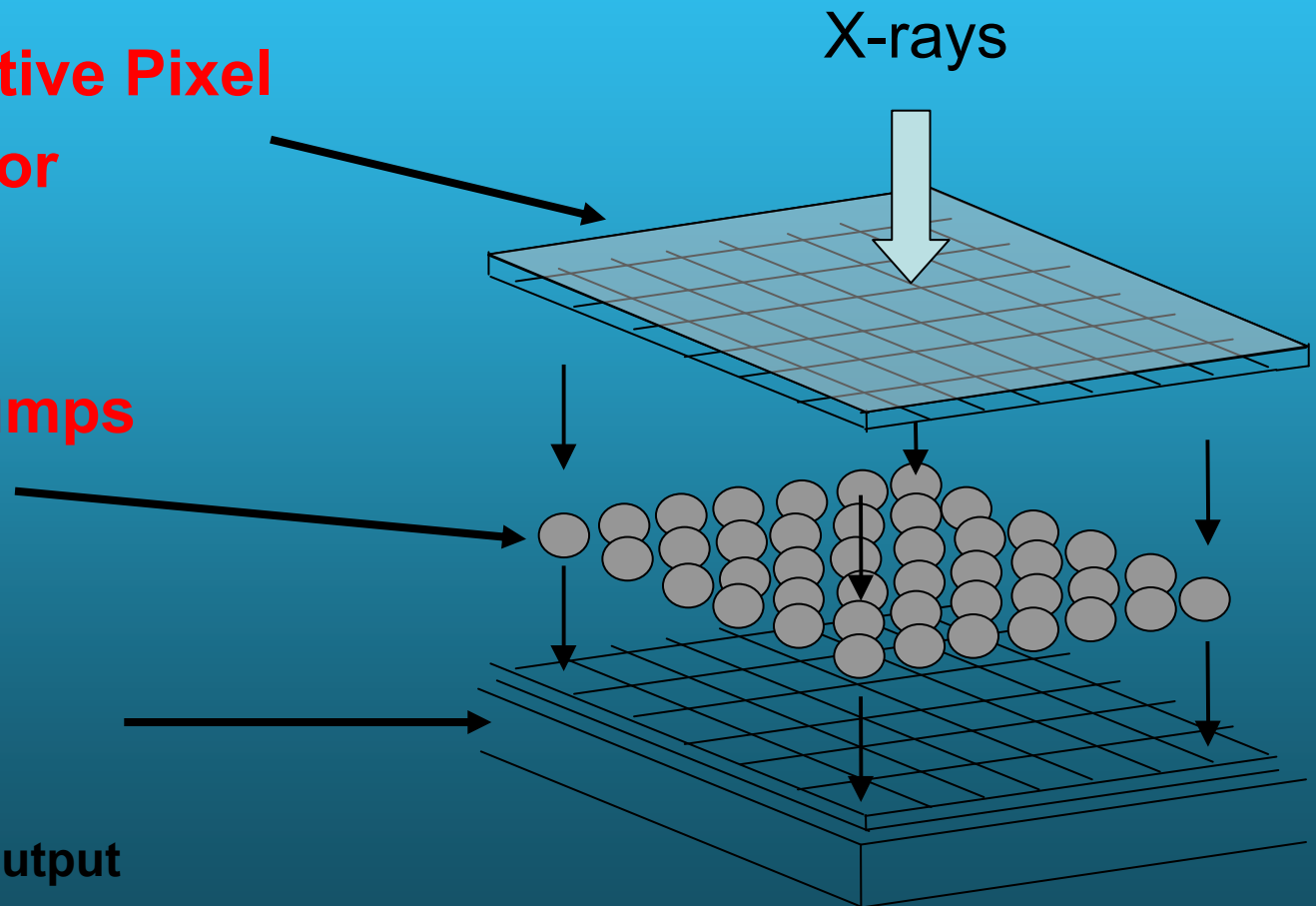
DePMOS Active Pixel Sensor

Connecting Bumps

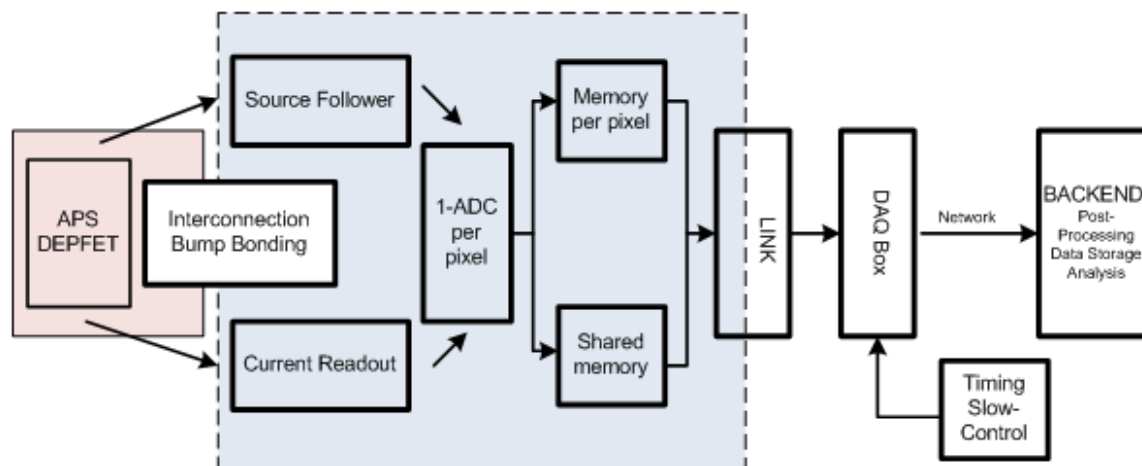
- 1 per pixel

CMOS Layer

- Signal processing
- Signal storage & output



- DEPFET Active Pixel Sensor
- Every DEPFET pixel provides detection and amplification with:
 - **Low noise**
 - **Signal compression at the sensor level**
 - **High speed** (fully parallel readout at 1-5 MHz)



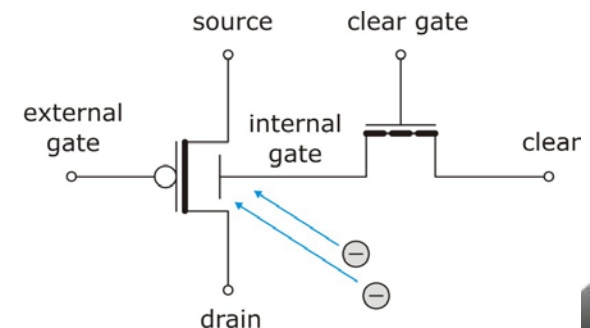
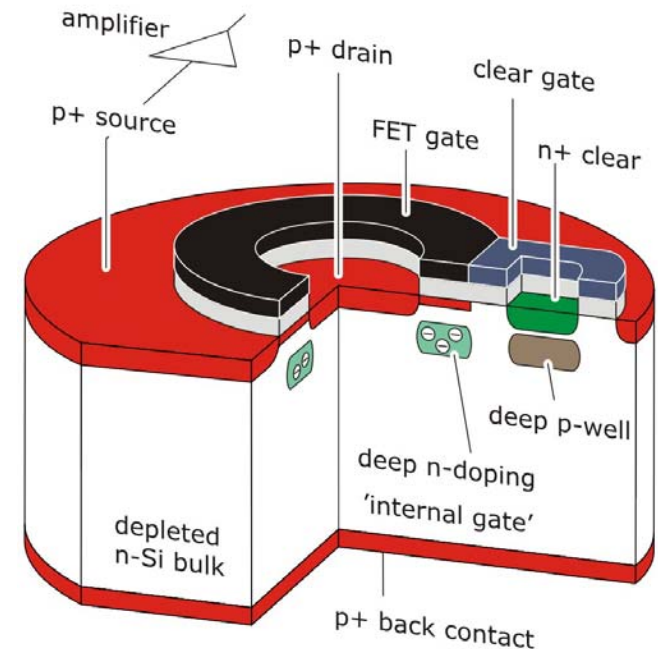
- Readout Concept
 - Immediate 8 Bit digitization
 - In-Pixel digital memory
 - Readout during macro bunch gaps

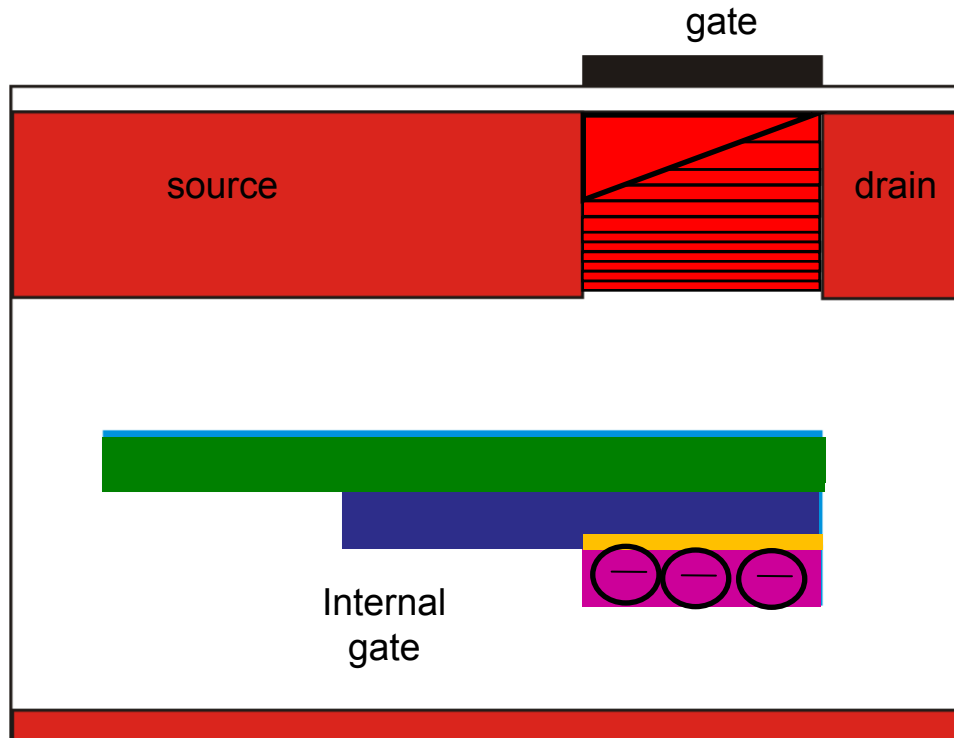
p-FET on depleted n-bulk

- signal charge collected in potential minimum below FET channel
- steers transistor current (1 el. $\sim$ 300 pA)

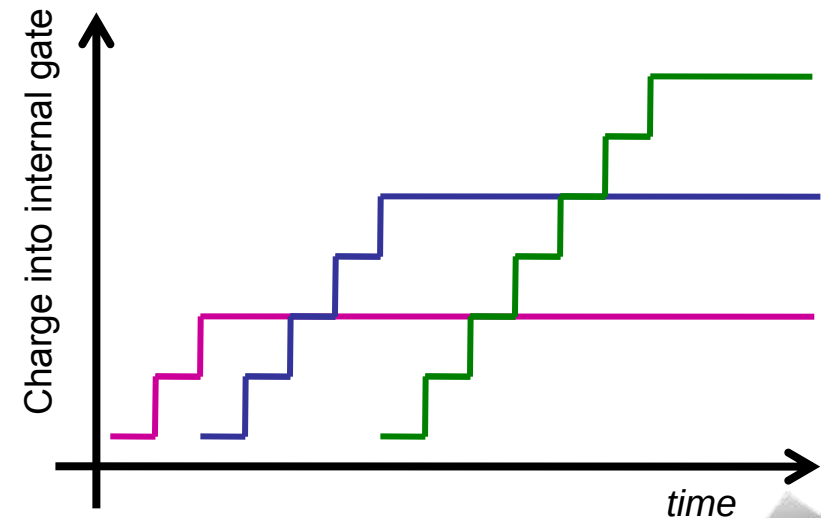
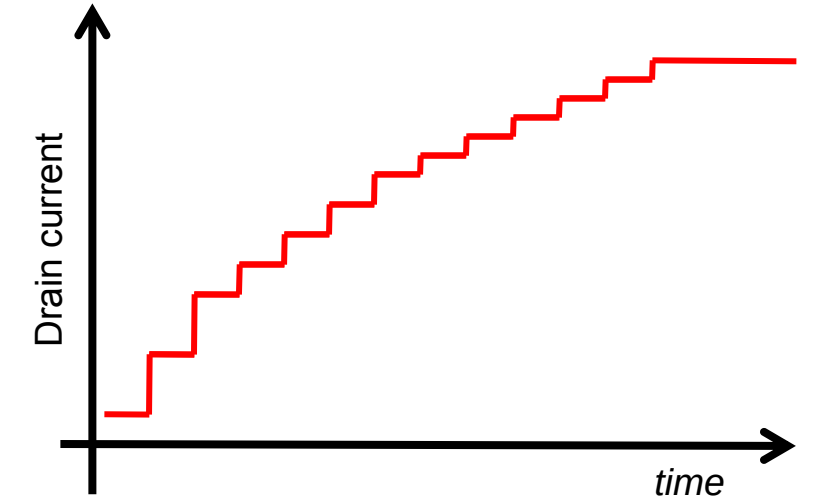
combined sensor & amplifier

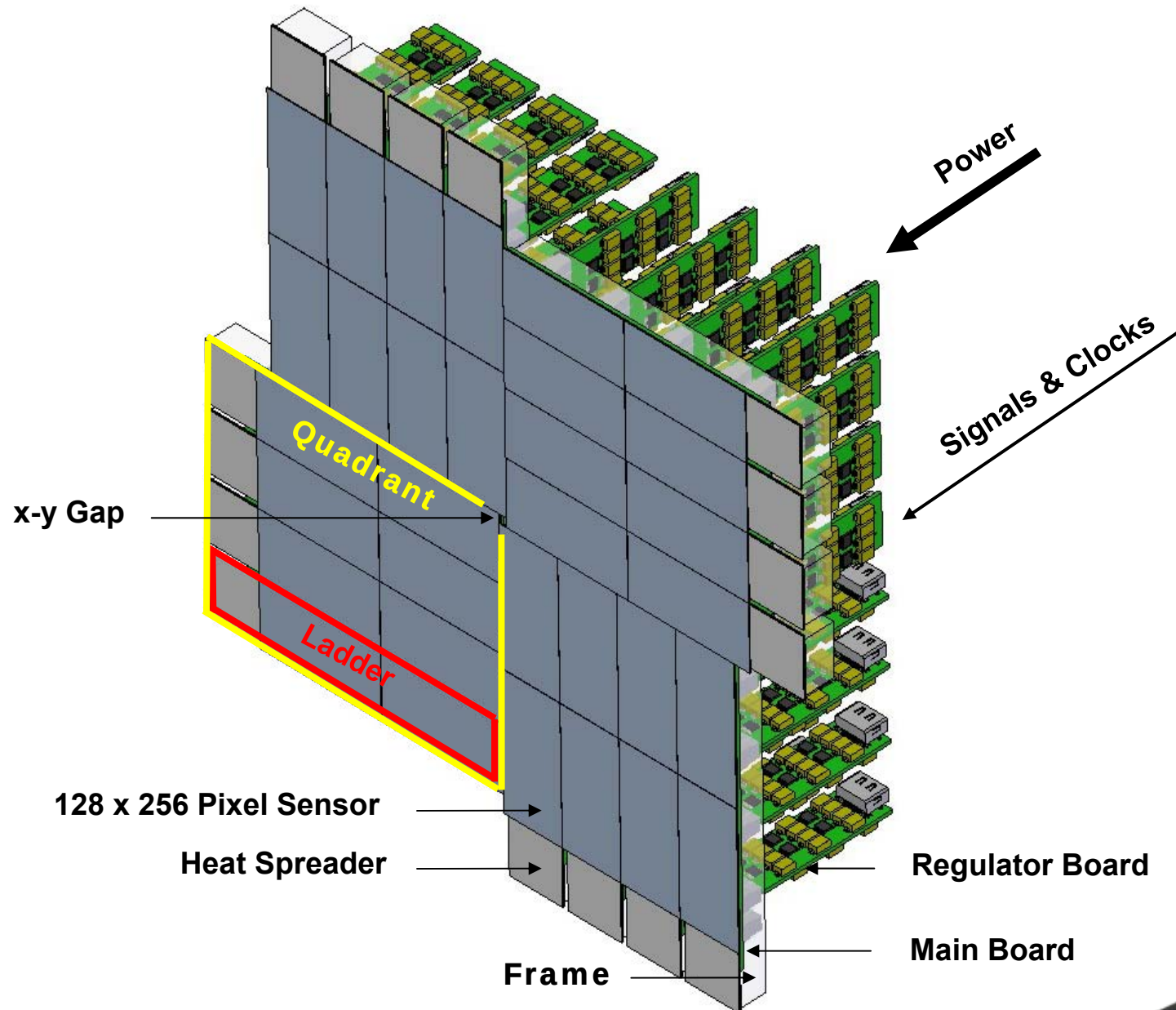
- low capacitance and noise
 - » **excellent spectroscopic performance**
- complete clearing of signal charge
 - » **no reset noise**
- non-destructive readout
 - » **potential of repetitive readout**
- charge storage capability
 - » **readout on demand**
- full depletion
 - » **backside illumination**
 - » **thin entrance window**

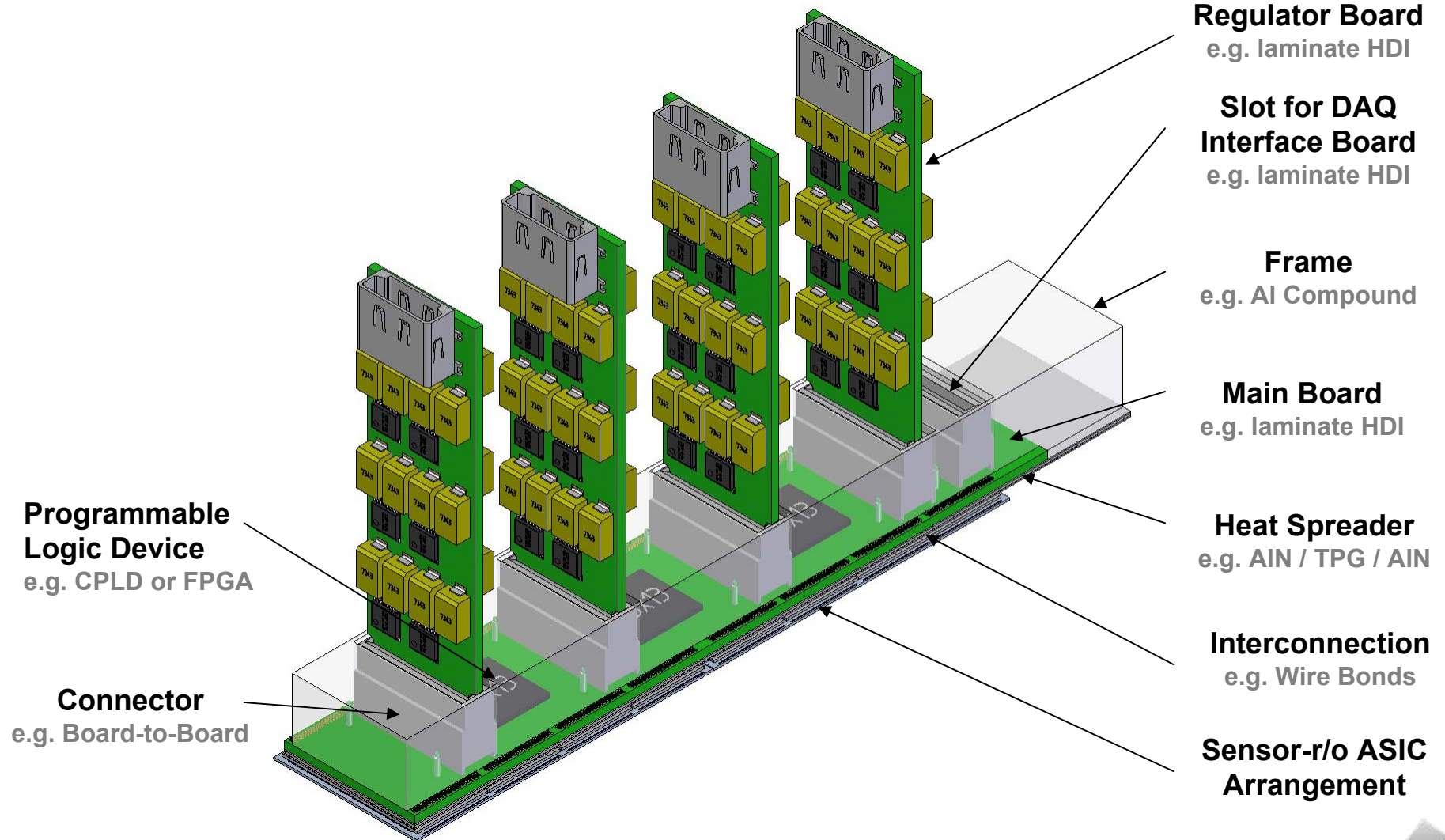




- The internal gate extends into the region below the source
- Small signals assemble below the channel, being fully effective in steering the transistor current
- Large signals spill over into the region below the source. They are less effective in steering the transistor current.

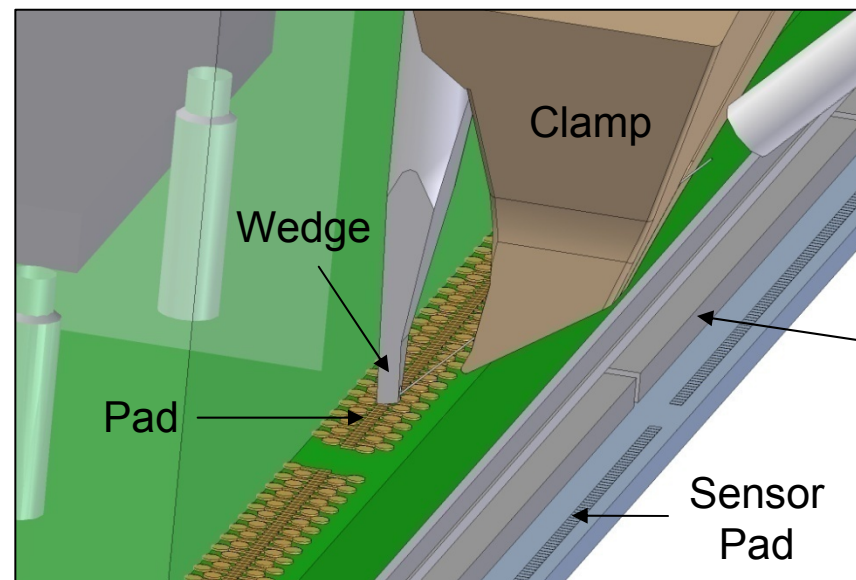
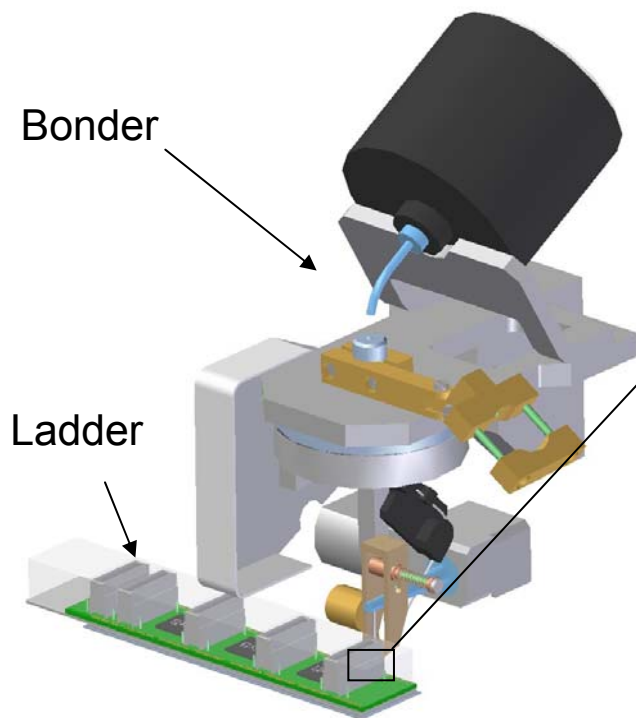






Feasibility Check of

- geometry
- process Steps



Wire Bonding

- Pitch 150 μm
- Step Height 3.3 mm
- Distance 2 mm
- Total number ~24.000 for 1MPix

Pads on Hybrid

- Area 100 x 200 μm^2
- min. Gap 50 μm
- out. Via Diam. 300 μm
- Cut. Edge Dist. 300 μm

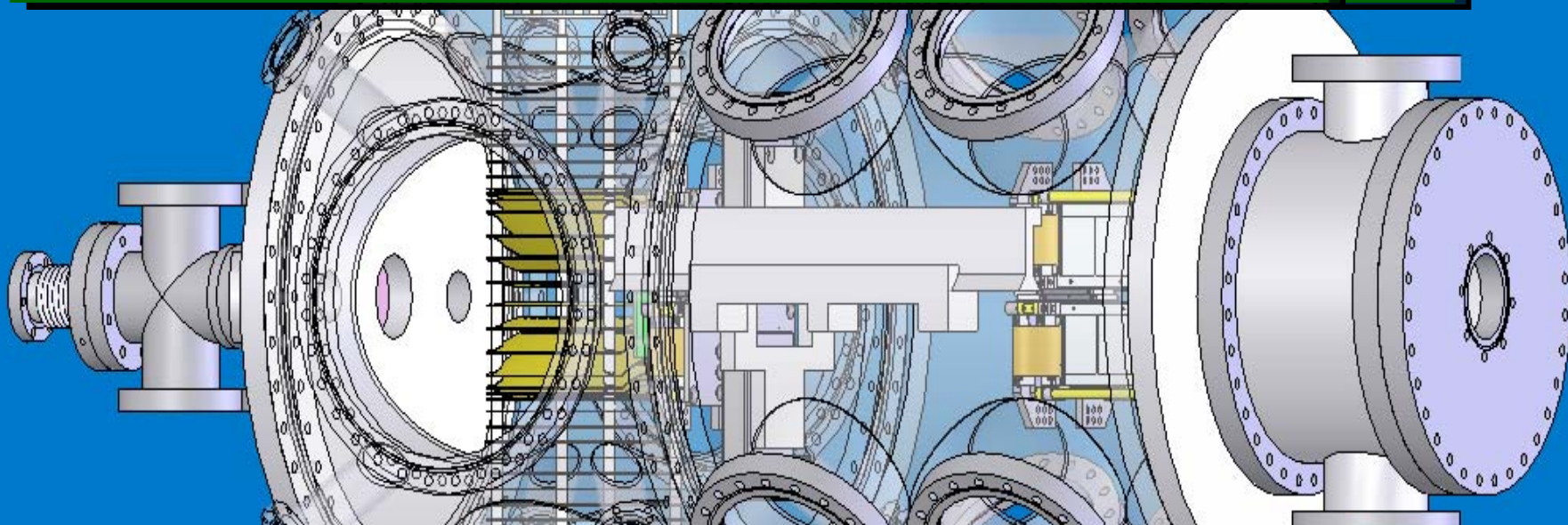
Readout
ASIC



CFEL-ASG Chamber



Plus 20 proposals for 2010! S!



- Fluorescence from atoms: Rolles,..+ ASG
- Clusters: Bostedt, Möller,.. +ASG
- Aligned molecules: Küpper, Chapman,.. + ASG
- Bio-samples: Chapman, .. + ASG



Future Focus:

- ❑ High speed, low noise X-ray pnCCD imagers
- ❑ X-ray pnCCDs with large formats
- ❑ New high speed analog signal processors
- ❑ DePFETs for X-ray imaging and vis./NIR imaging
- ❑ DePFETs with
 - analog signal compression
 - repetitive non-destructive readout
 - "gatable" internal gate
- ❑ SDDs in various shapes and electronic configurations

This provides a scientific horizon of 30 years for innovative detector developments at the APS, LBL



End