



VARIABILITY STUDY OF THE CRAB PULSAR ABOVE FEW TENS OF GeV WITH MAGIC

Giovanni Ceribella
Max-Planck-Institut für Physik

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Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)

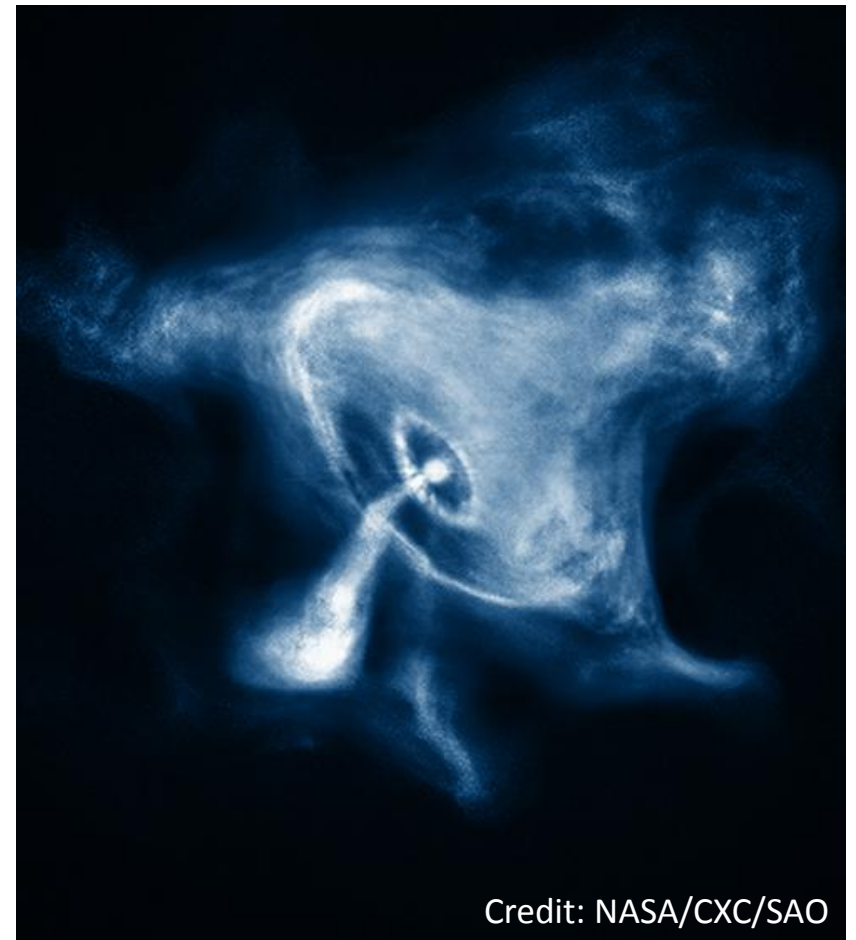


- 1. The Crab nebula and pulsar:**
Physical case of the nebula flares.
- 2. Analog Sum-Trigger II for MAGIC:**
Review of the lower energy trigger;
- 3. Preliminary results and future plans.**

THE CRAB NEBULA



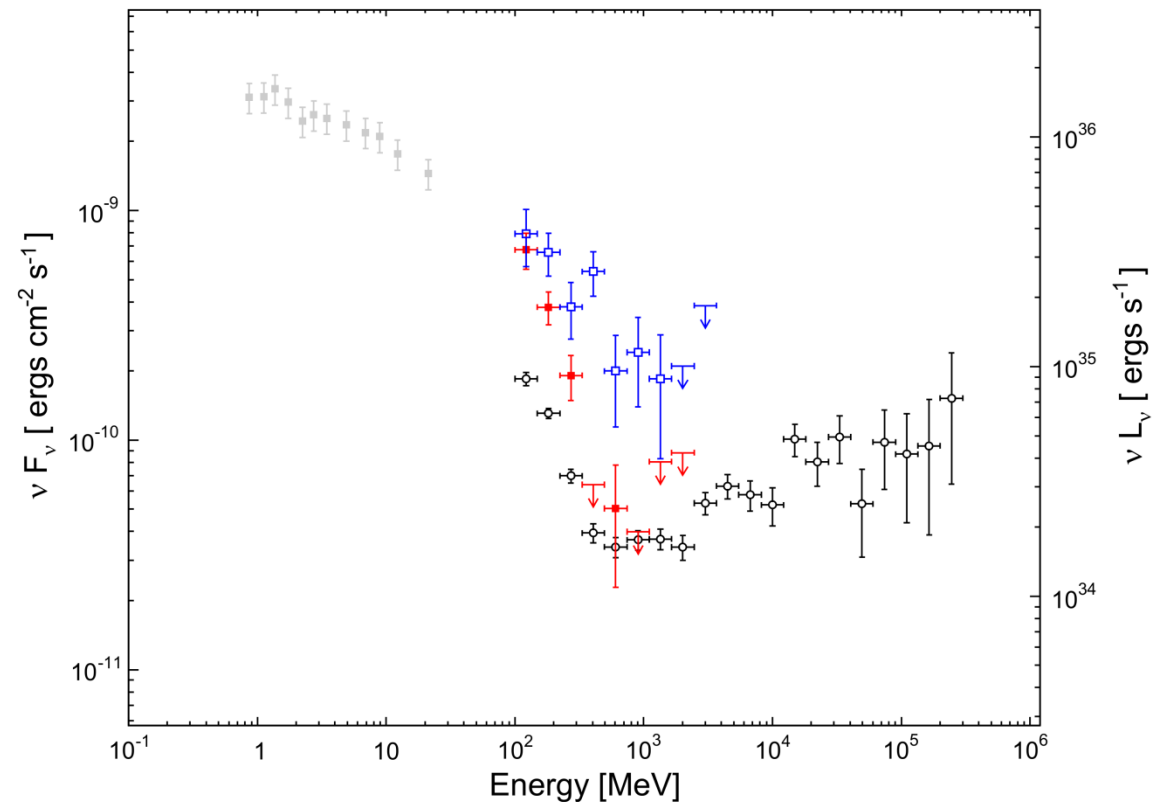
- ❖ Strongest stable VHE source of the sky and standard candle for gamma ray astronomy;
- ❖ Pulsar wind nebula powered by the young Crab pulsar;
- ❖ Synchrotron emission up to 1 GeV, Inverse compton up to tens of TeV.



CRAB FLARES

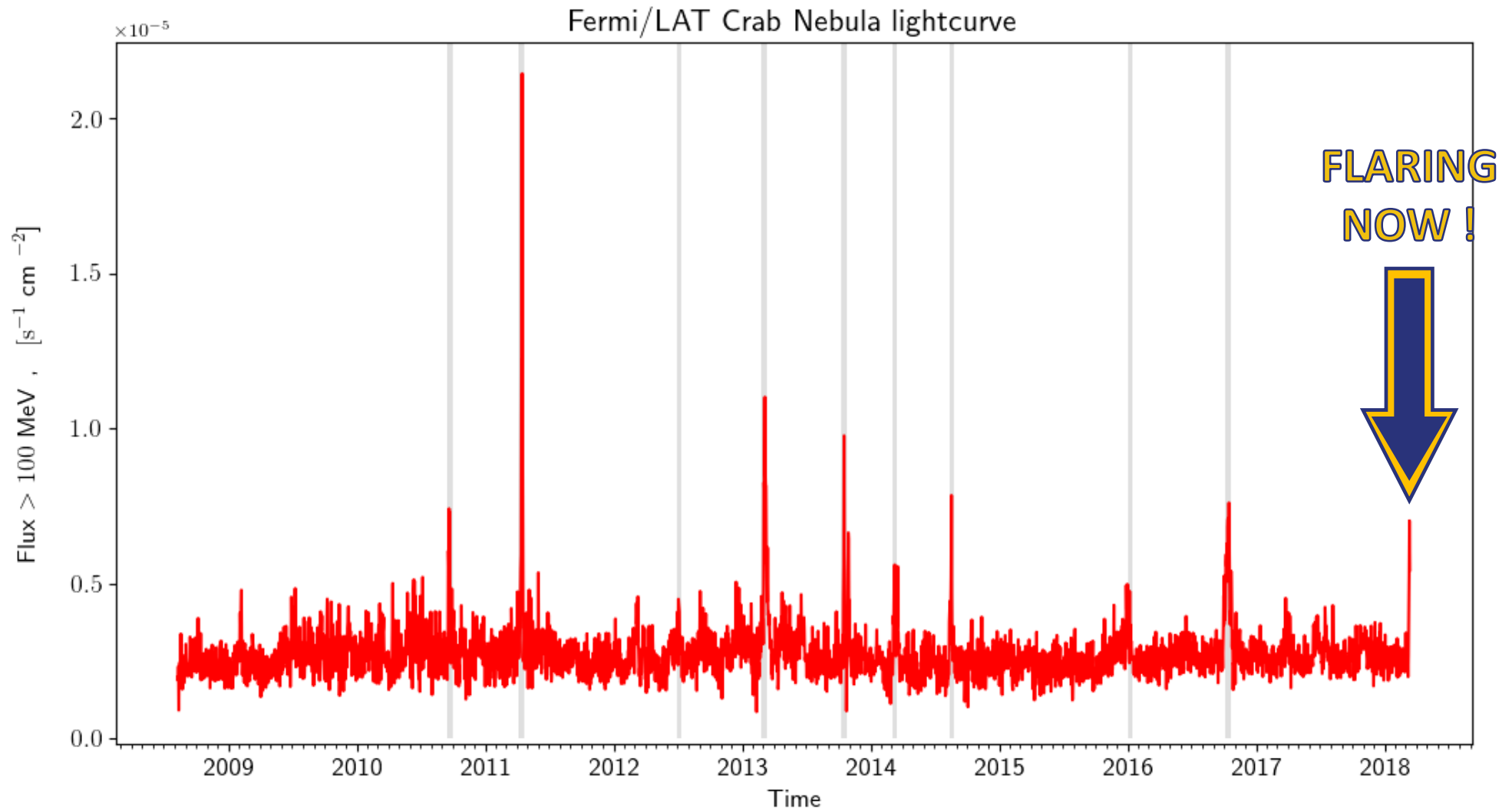


- ❖ First detected in 2010;
- ❖ Large increase of the flux in the 100 MeV to 1 GeV energy range;
- ❖ Short timescales of days to weeks:
 - Constrains the originating phenomenon to be acting in a small region.
- ❖ Physics still debated...



Gamma spectrum of the Crab Nebula from Fermi/LAT.
Black: long term average; **Blue/Red** (flare state). From [1].

DAILY FERMI/LAT LIGHTCURVE

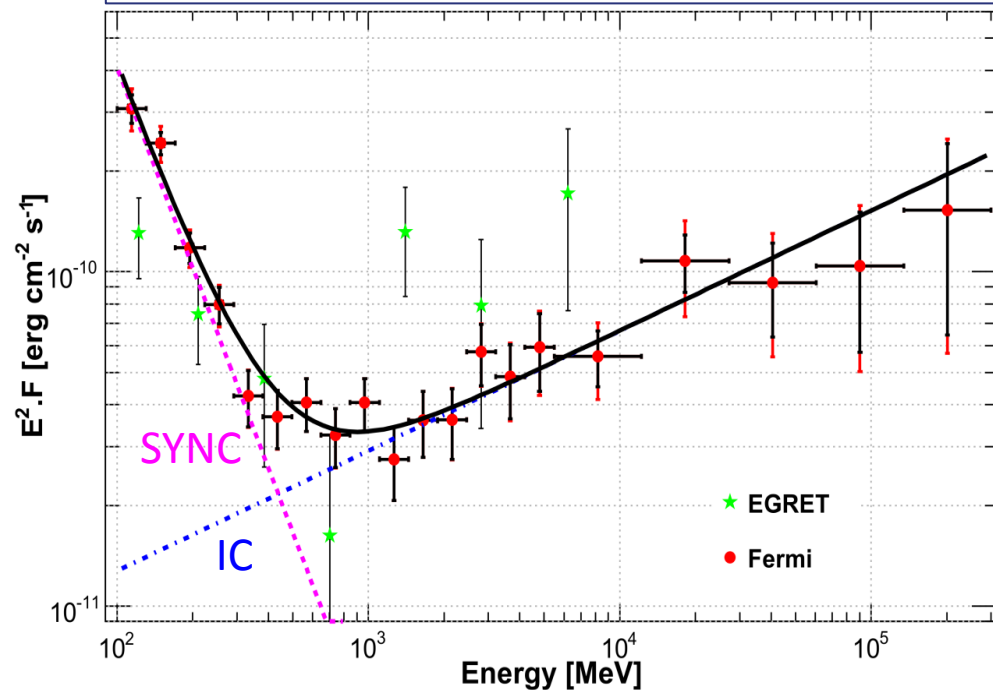


IS THE PULSAR INVOLVED?

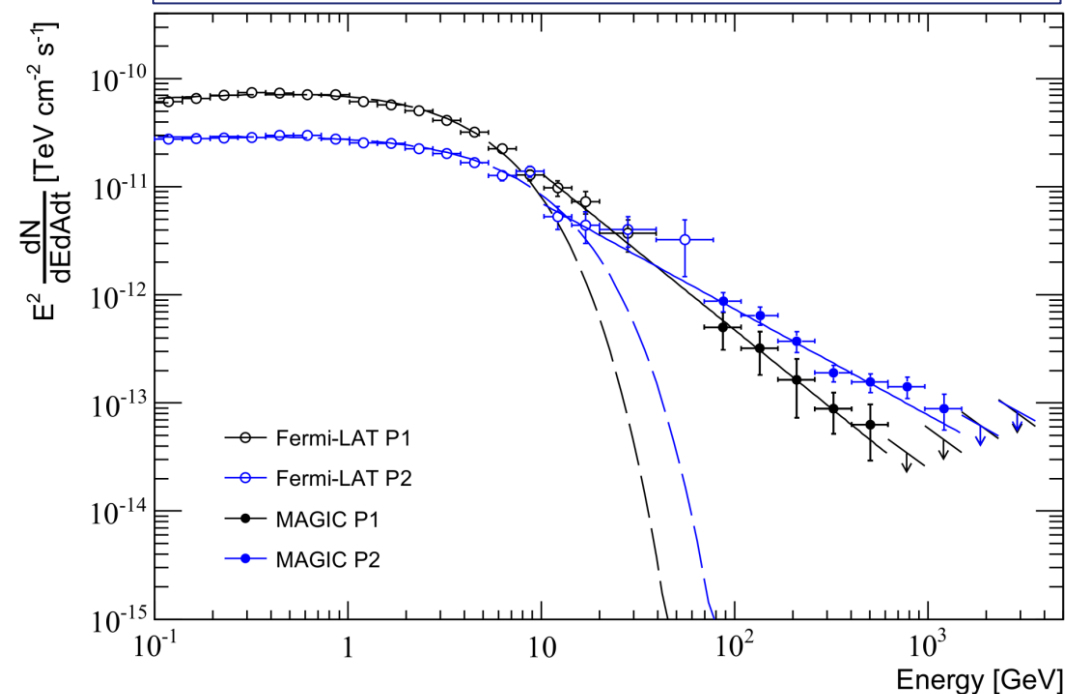


- ❖ Pulsar: central engine of the Nebula → related to flares?
- ❖ No significant increase of pulsar flux during flares (Fermi/LAT, [1]).
- ❖ What about the highest energy electrons?

Fermi/LAT Crab Nebula spectrum, [2].



MAGIC Crab Pulsar spectrum, [3].

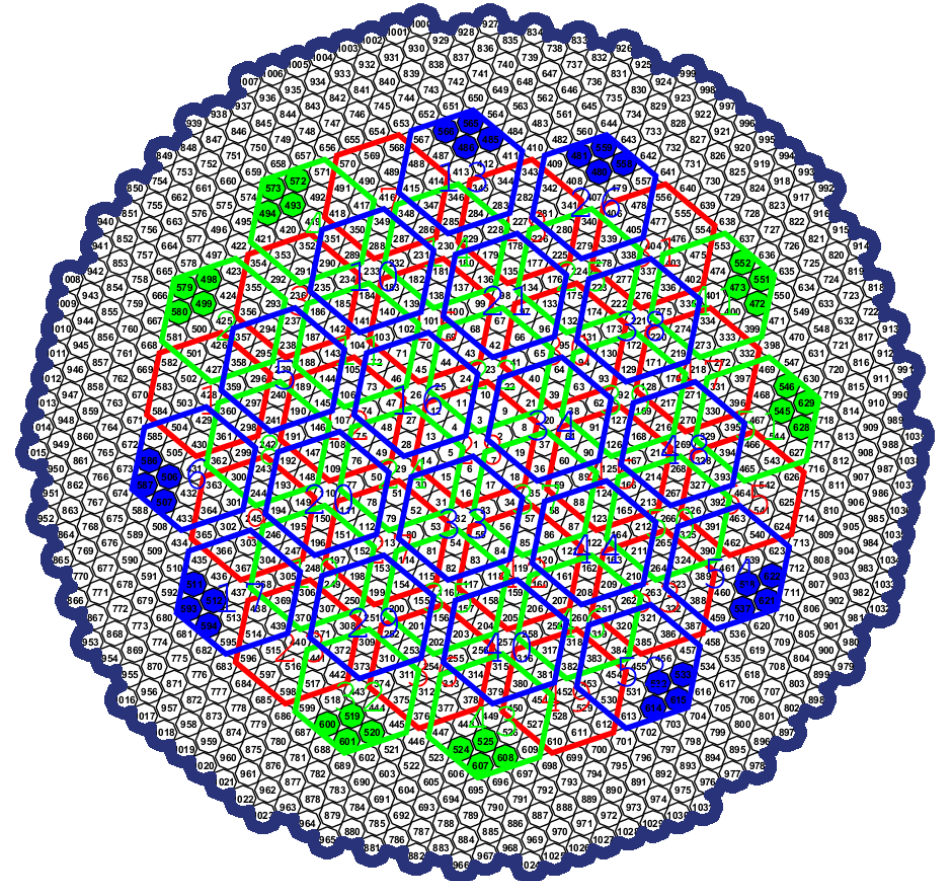


IS THE PULSAR INVOLVED?



- ❖ What about the highest energy electrons?
 - Tail of the **nebular** synchrotron emission:
 - the spectral region *where flares show up*.
 - Tail of the **pulsar** synchrotron and IC emission;
- ❖ Light distance between pulsar and the termination shock of the wind: *~5 months*;
- ❖ Fermi collection area is too small to probe the pulsar emission above few GeV at short timescales;
- ❖ **MAGIC** with the new **Sum-Trigger-II** can do that!

- ❖ An analog trigger for the lower energy air showers;
- ❖ Performing the **sum of signals** from neighbouring photomultiplier tubes;
- ❖ A **lower discriminator threshold** can be applied to the summed signal;
- ❖ A lower energy threshold is achieved.



Layout of the *macrocells* (sum regions) of the MAGIC Sum-Trigger-II. Adapted from [4].

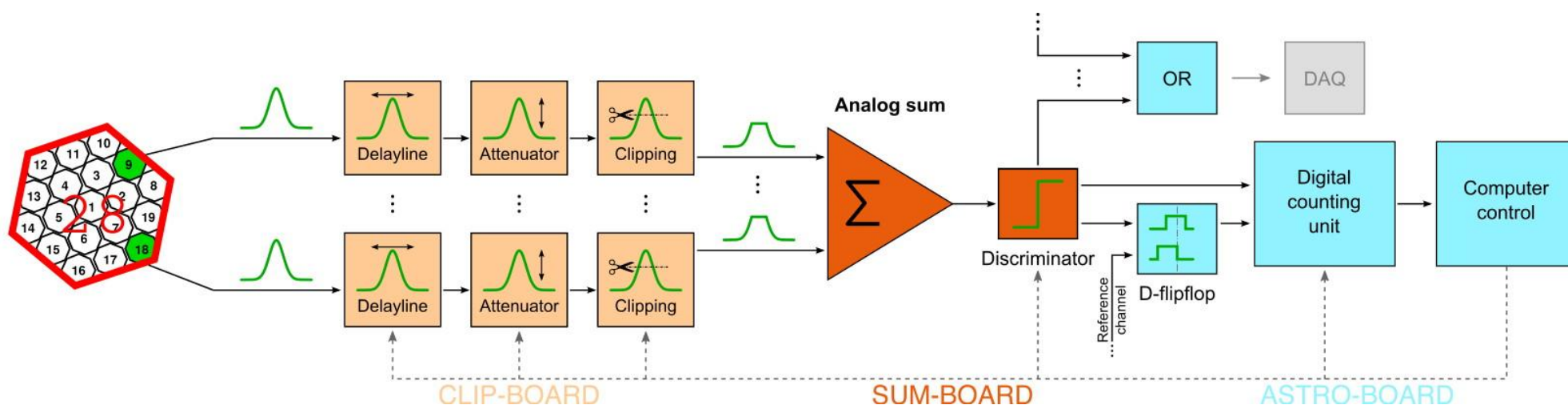
HOW DOES IT WORK?



Timing adjustment
and flat-fielding

Coping with
afterpulsing

Signals sum and
trigger formation



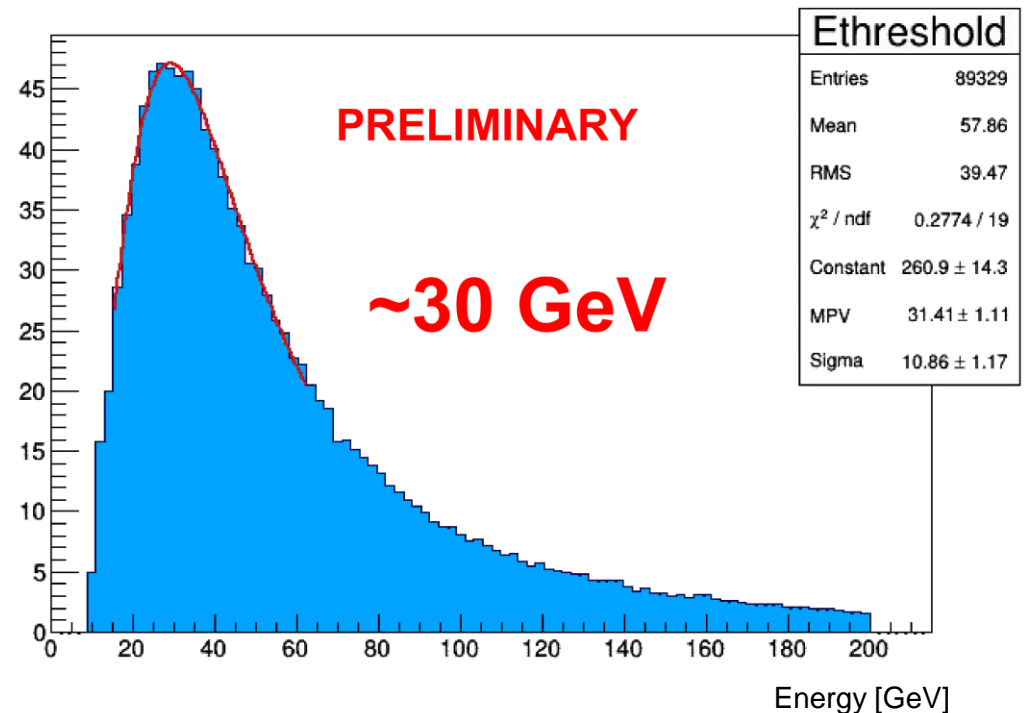
Workflow of the Sum-Trigger-II. Adapted from [4].

Threshold and
timing controller

PRELIMINARY RESULTS

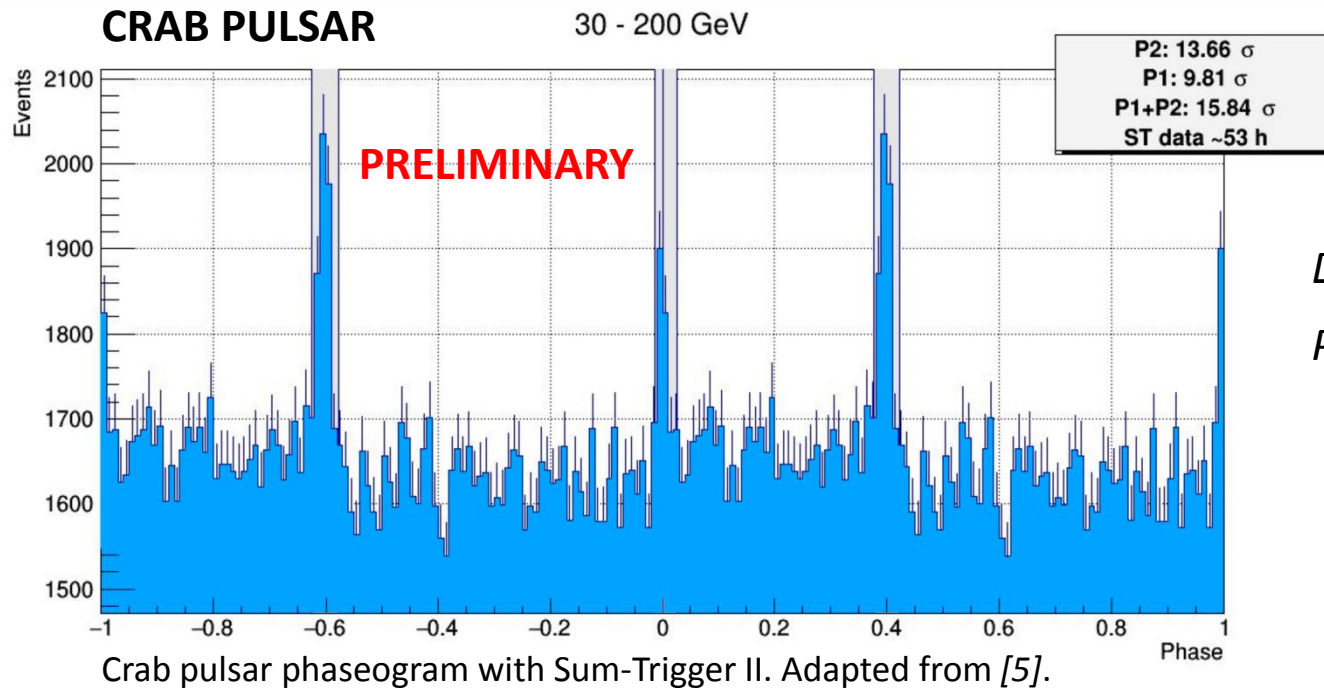


- ❖ Energy threshold at **30 GeV** computed from Monte Carlo simulations;
- ❖ The efficiency reduces to $\frac{1}{2}$ at **15 GeV**;
- ❖ The detection capability of low energy events is largely enhanced:
 - Previous result with the standard digital trigger: 50 GeV.



Global analysis efficiency on Monte Carlo simulated gamma events, energy threshold fit. Adapted from [5].

PRELIMINARY RESULTS



Detection efficiency **2.3 σ/\sqrt{h}**

Previous value 1.4 σ/\sqrt{h}

Same results in 2,7 times less observation time!

- ❖ Improved threshold \rightarrow probe the largely unexplored 10-50 GeV band;
- ❖ Enhanced detection efficiency \rightarrow fast determination of the phaseogram:
 - 5 σ detection in 4 hours of observation time.

- ❖ With the improved performance in the 20 – 50 GeV range, MAGIC became a unique instrument for gamma pulsar physics;
- ❖ We have proposed to monitor the flux of the Crab pulsar in the VHE regime:
 - Is it stable?
 - Can it be correlated with flares in the nebula?
- ❖ MAGIC with the Sum-Trigger II can do much more: new pulsars, distant AGNs, GRB.
- ❖ Stay tuned!

- 1) The Fermi/LAT collaboration, ***Gamma-Ray Flares from the Crab Nebula***, Science 2011.02, pp. 739-742;
- 2) The Fermi/LAT collaboration, ***Fermi Large Area Telescope Observations of the Crab Pulsar And Nebula***, ApJ 2010.01, pp. 1254-1267;
- 3) The MAGIC collaboration, ***Teraelectronvolt pulsed emission from the Crab Pulsar detected by MAGIC***, A&A 2016.01, A133;
- 4) Francesco Dazzi, ***A new stereoscopic “Sum-Trigger-II” for the MAGIC Telescopes***, PhD thesis 2012;
- 5) Jezabel Rodriguez Garcia, ***Pulsars with MAGIC***, TeVPa 2017.