



# Status and Prospects for the CRESST Experiment

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Max-Planck-Institut für Physik



MPP Project Review  
Dec 16<sup>th</sup>, 2018

SFB 1258

Neutrinos  
Dark Matter  
Messengers



# CRESST @ MPP

- **Staff:**

- F. Petricca (PI)
- D. Hauff
- A. Bento (visiting)

- **PostDoc:**

- L. Canonica
- M. Mancuso
- R. Strauss (up to 10.2018)

- **Ph.D student:**

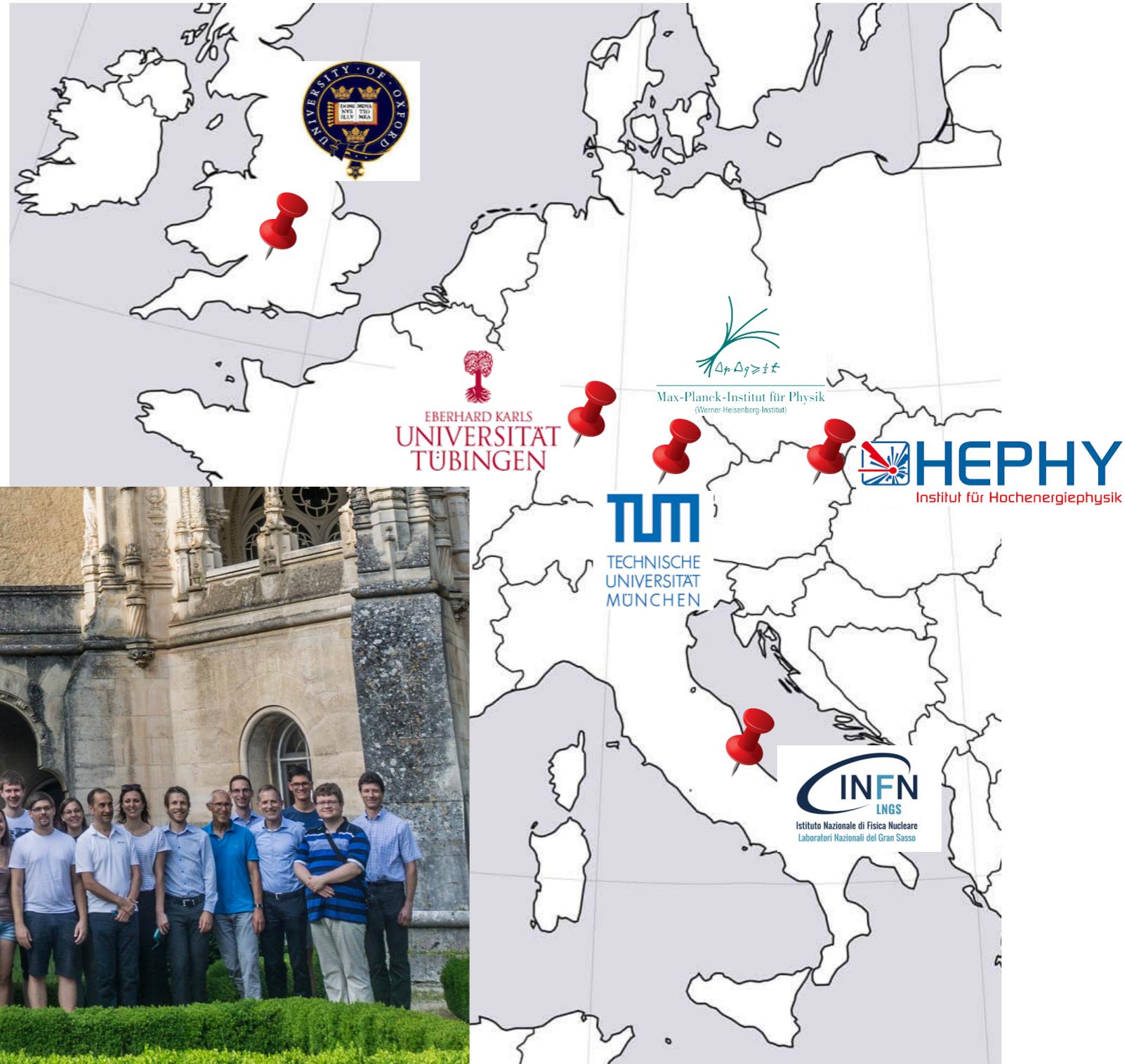
- A. Abdelhameed
- P. Bauer
- E. Bertoldo
- N. Ferreira\*
- J. Rothe
- M. Wüstrich\*

\*thesis completed in 2018

# The CRESST Collaboration

## ~45 Collaborators:

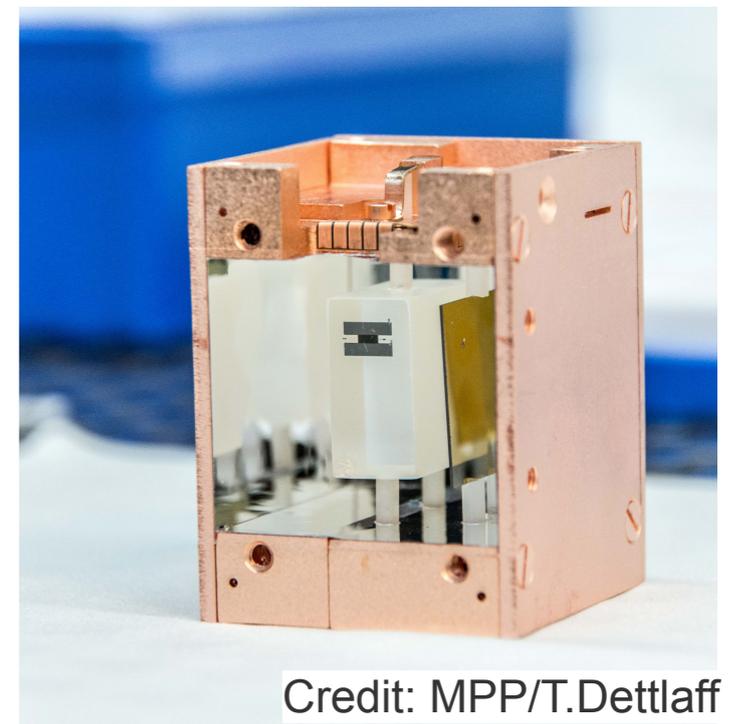
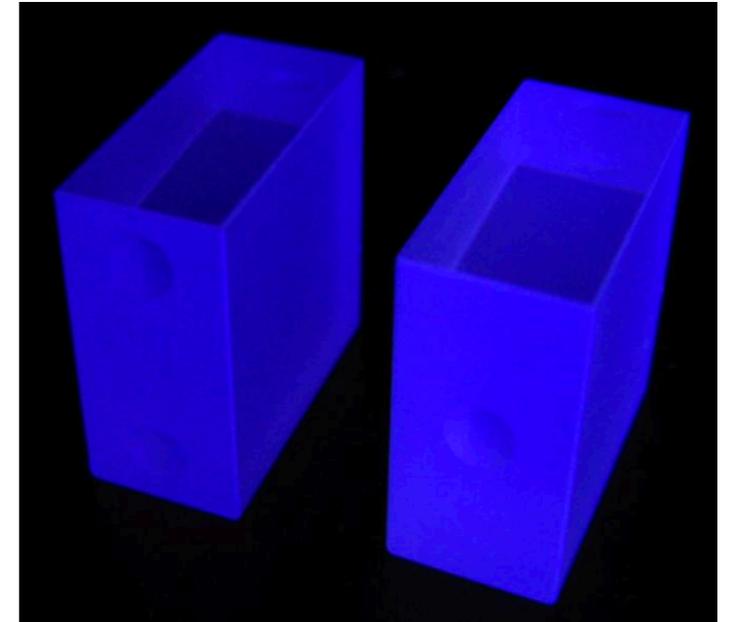
- 14 MPP, DE
- 10 TUM, DE
- 4 Tübingen, DE
- 8 HEPHY, AT
- 8 LNGS, IT
- 1 Oxford, UK



# The CRESST experiment

## Cryogenic Rare Event Search with Superconducting Thermometers

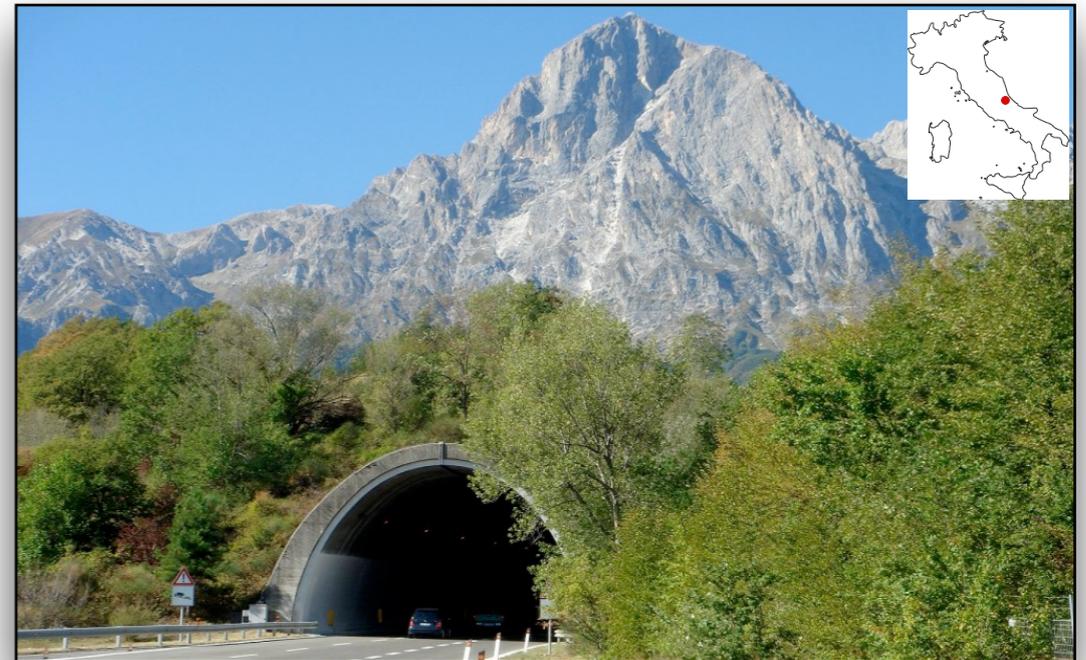
- Direct detection of Dark Matter particles via their scattering off target nuclei
  - Signal: low energy nuclear recoil
- Target: Scintillating  $\text{CaWO}_4$  crystals
  - Composite detector  $\text{CaWO}_4$  + Light Detector
  - Efficient particle ID at low energy



Credit: MPP/T.Dettlaff

# The CRESST experiment @ LNGS

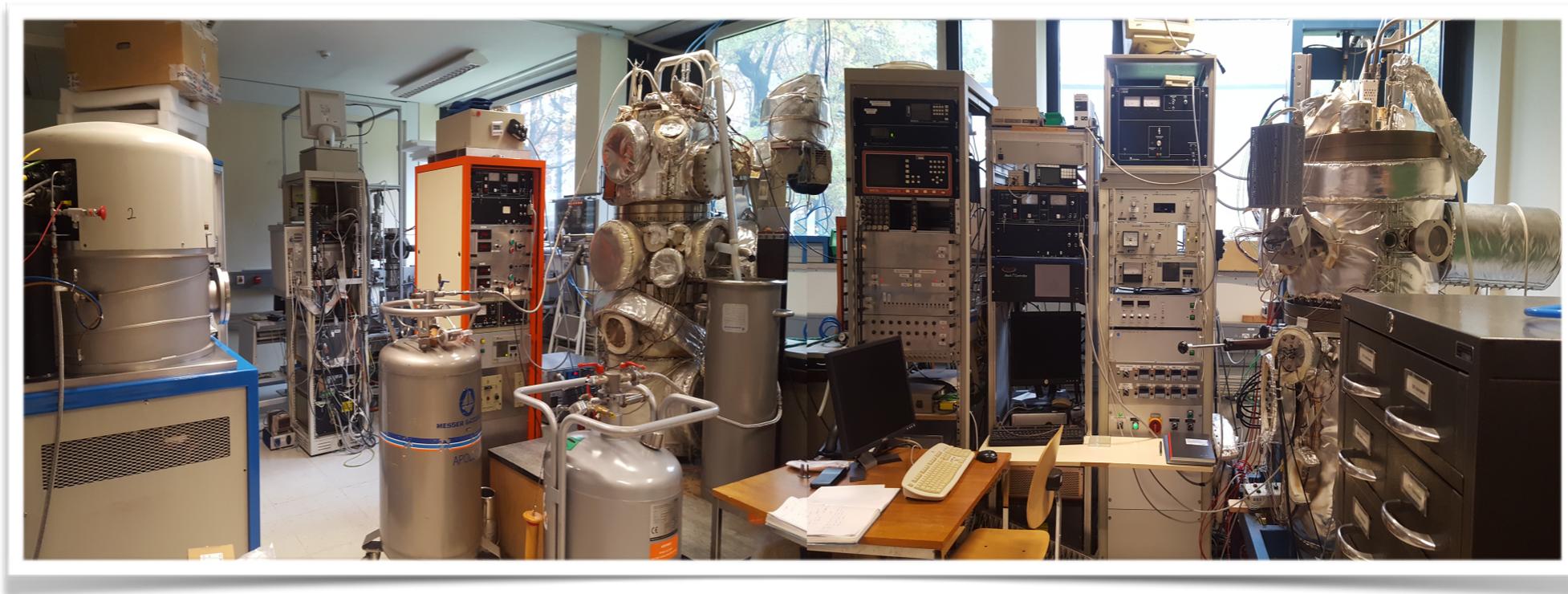
Cryogenic Rare Event Search with Superconducting Thermometers



**Laboratori Nazionali  
del Gran Sasso (Italy)**

Experimental location:  
Average depth  $\sim 3600$  m w.e.  
Muon flux  $\sim 2.6 \times 10^{-8}$   $\mu/s/cm^2$   
Neutrons ( $<10$  MeV):  $<10^{-6}$  n/s/cm $^2$

# The CRESST experiment @ MPP



**Sensor  
production**

# The CRESST experiment @ MPP



**Sensor  
production**



**Detector  
assembly**



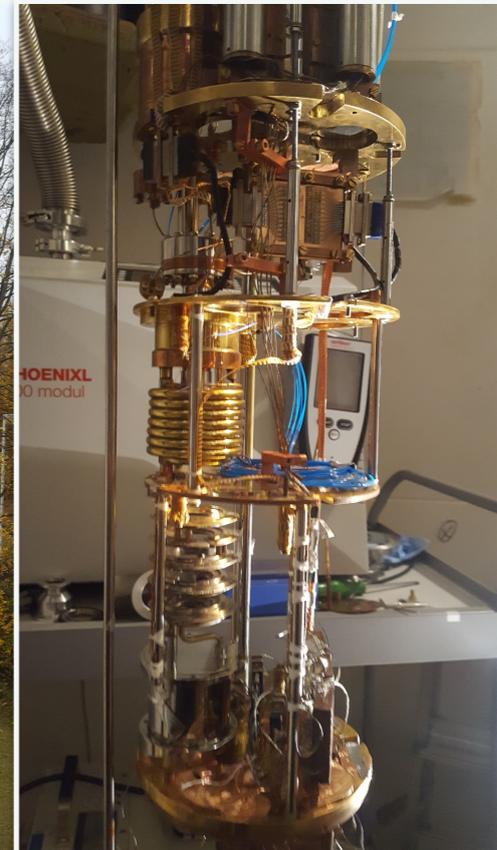
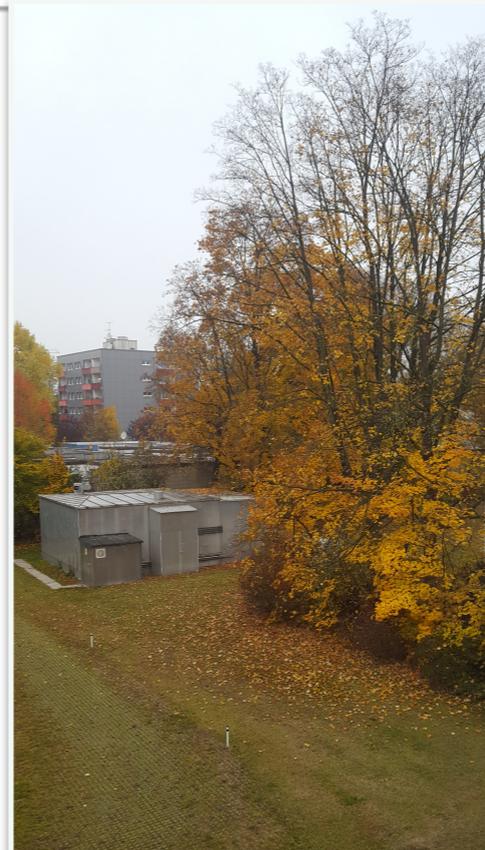
# The CRESST experiment @ MPP



**Sensor  
production**

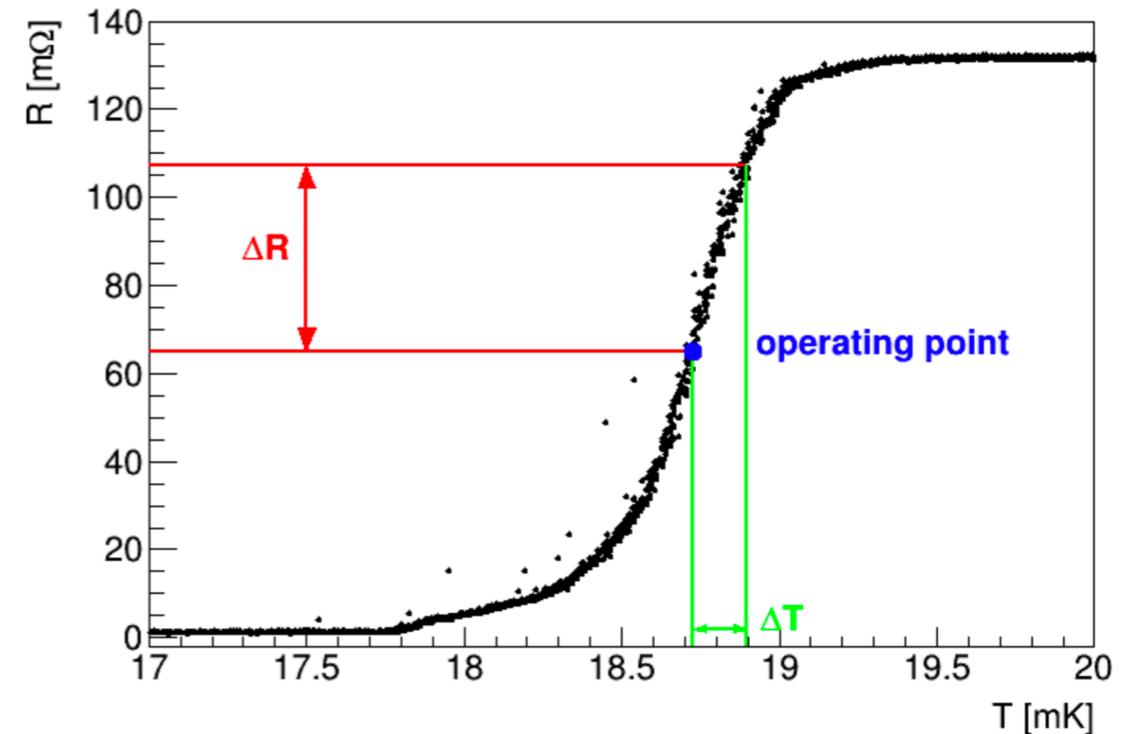
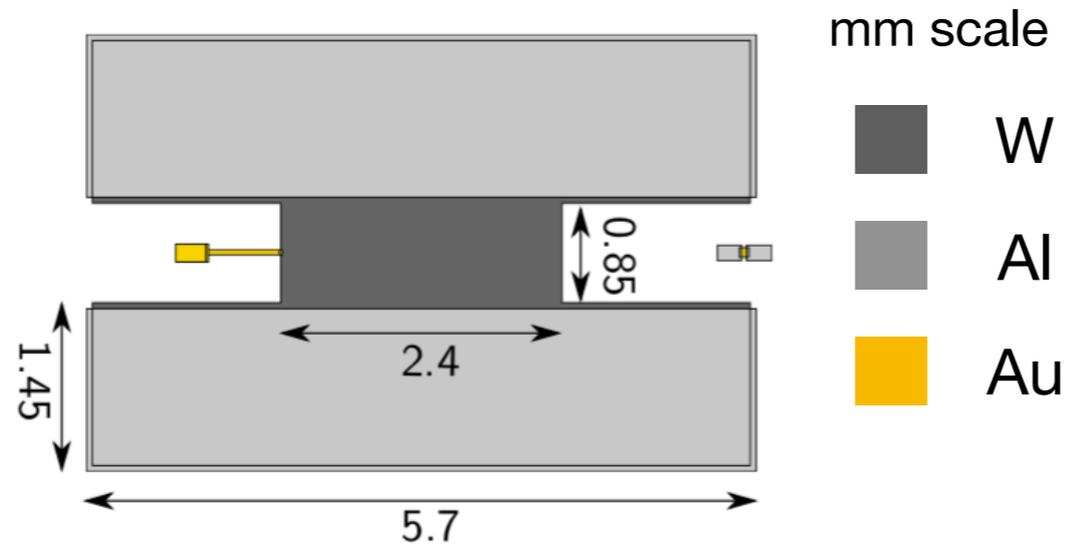


**Detector  
assembly**

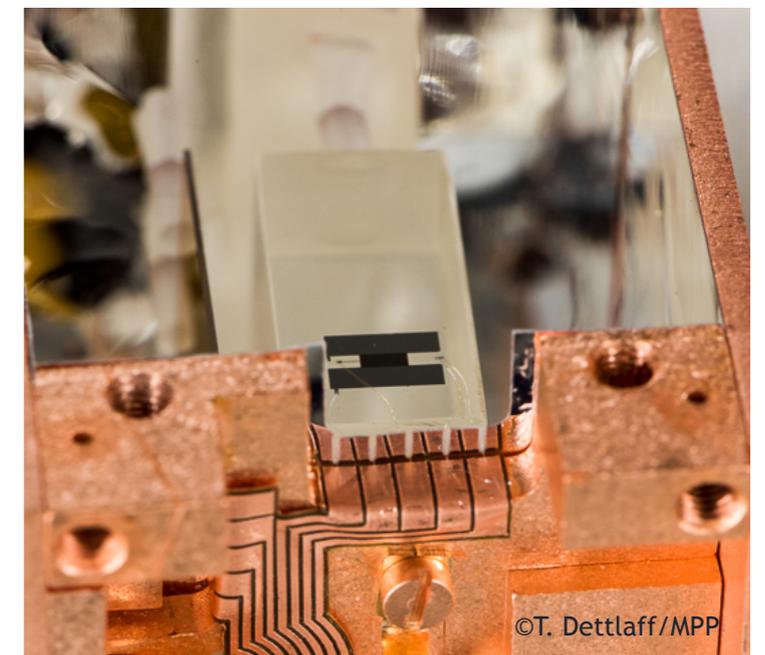


**Detector  
Testing**

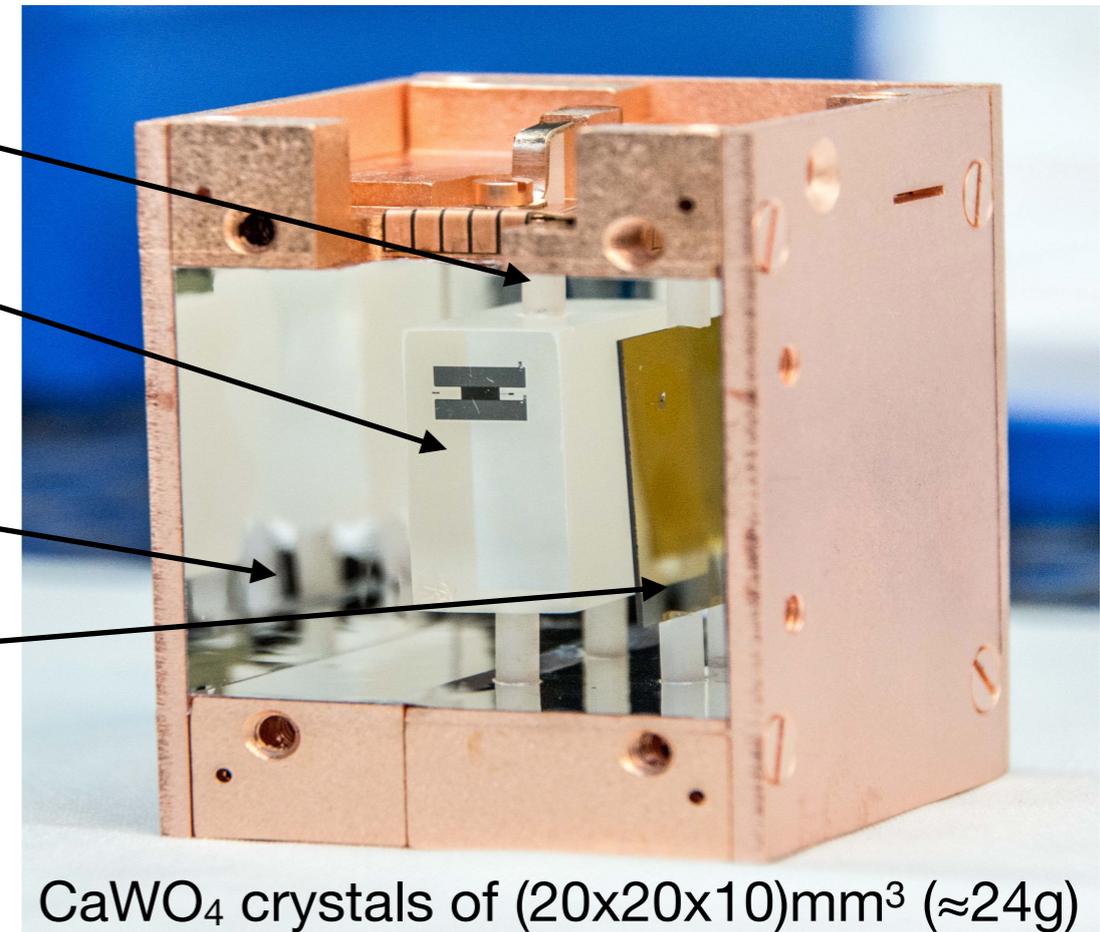
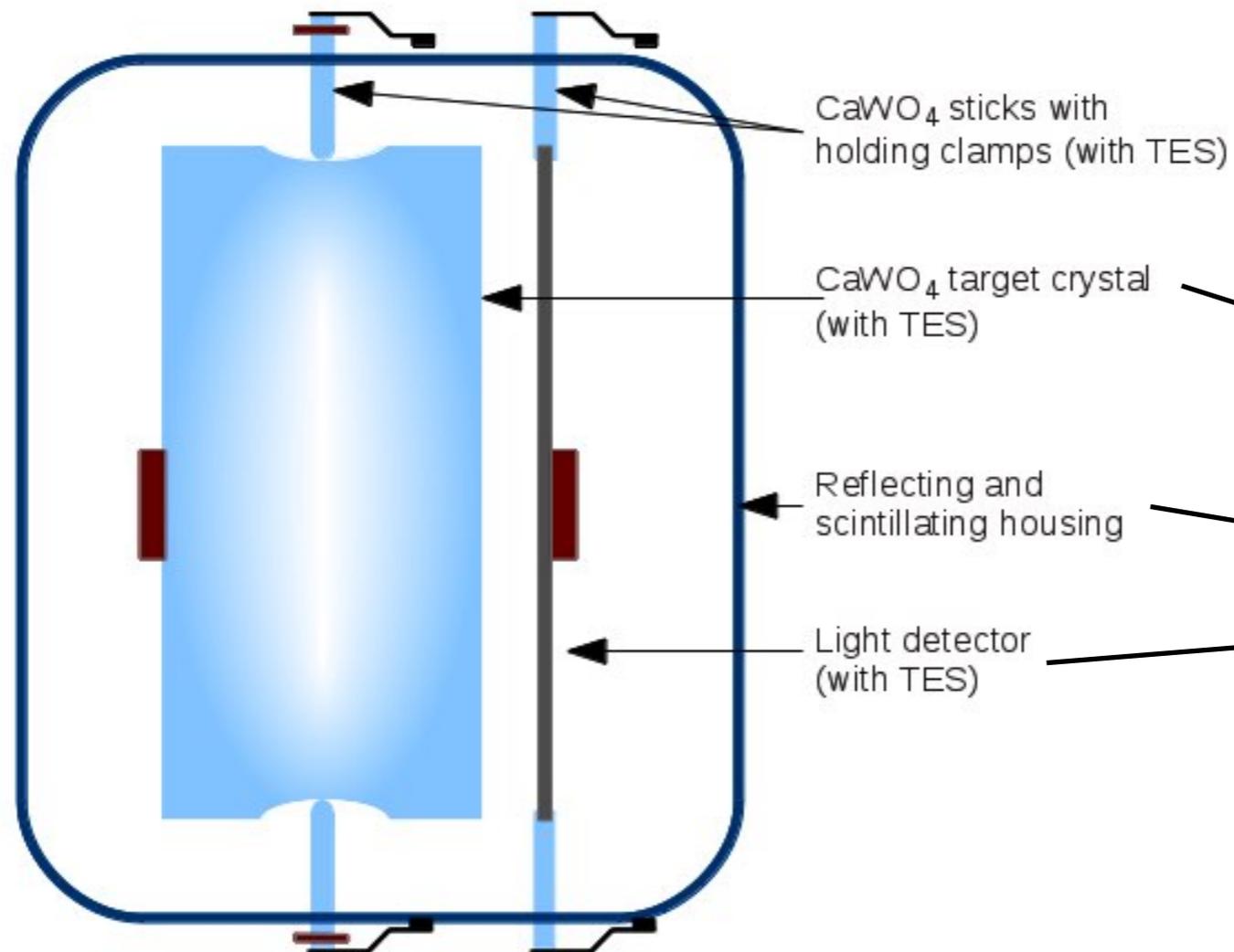
# Transition Edge Sensors



- Thin ( 200 nm) Tungsten film evaporated on the crystal.
- Transition temperature  $\sim$  [10 - 20] mK
- **Key technology for detecting very small recoiling energies.**
- **TES R&D and fabrication at MPP.**



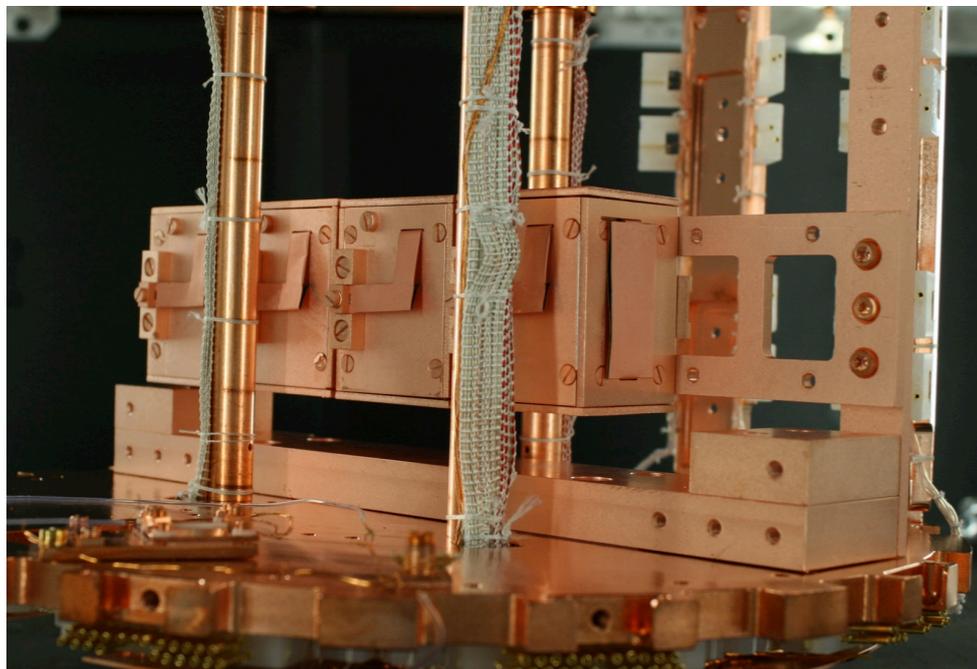
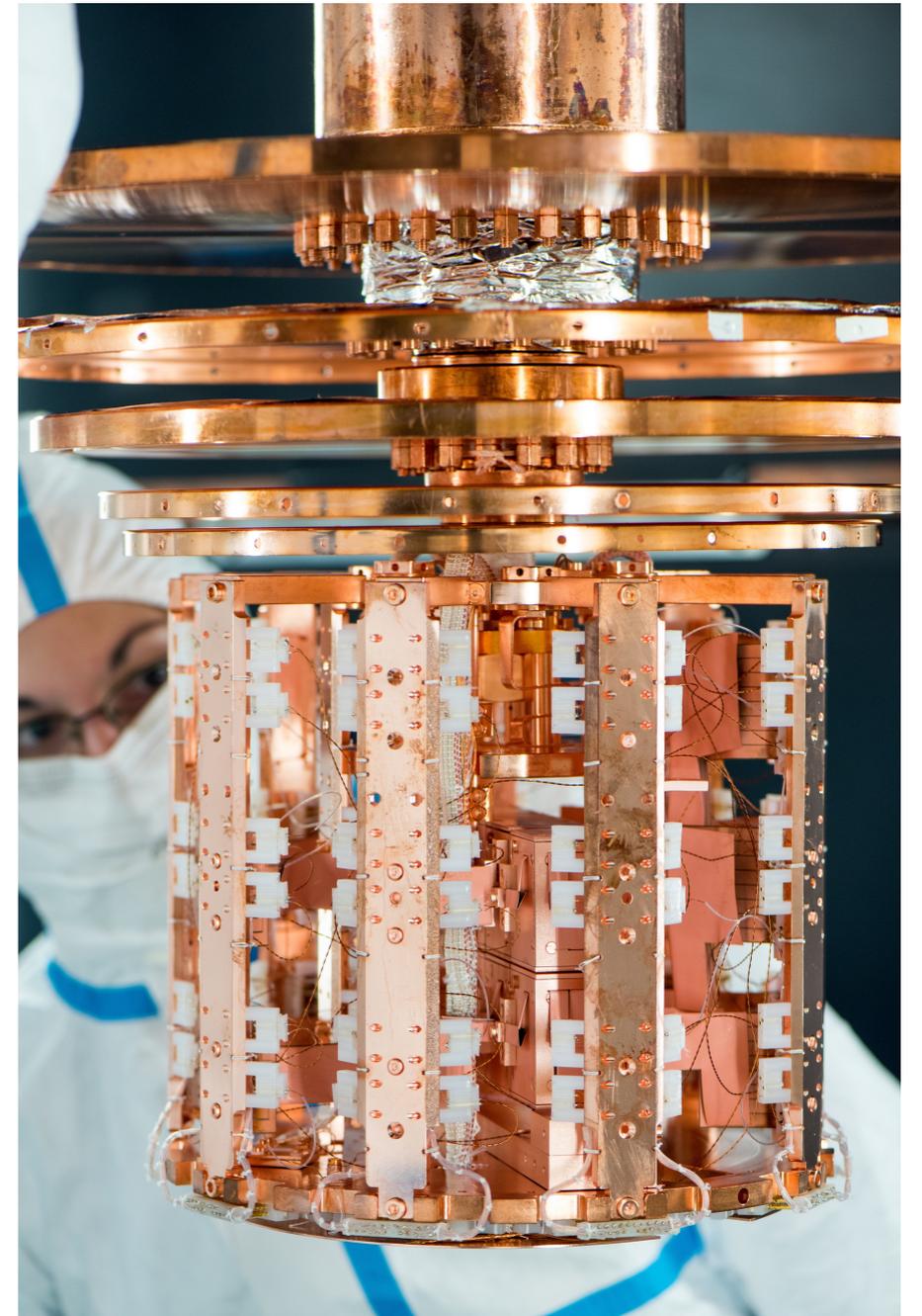
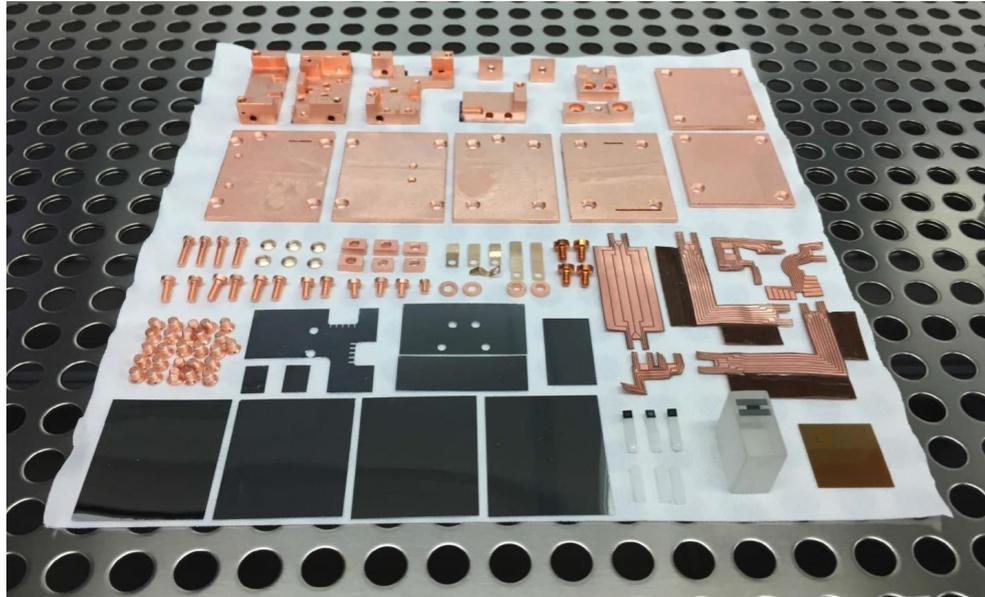
# CRESST III detector module



- Detector layout optimized for low-mass dark matter
- Fully scintillating housing
- Instrumented holders

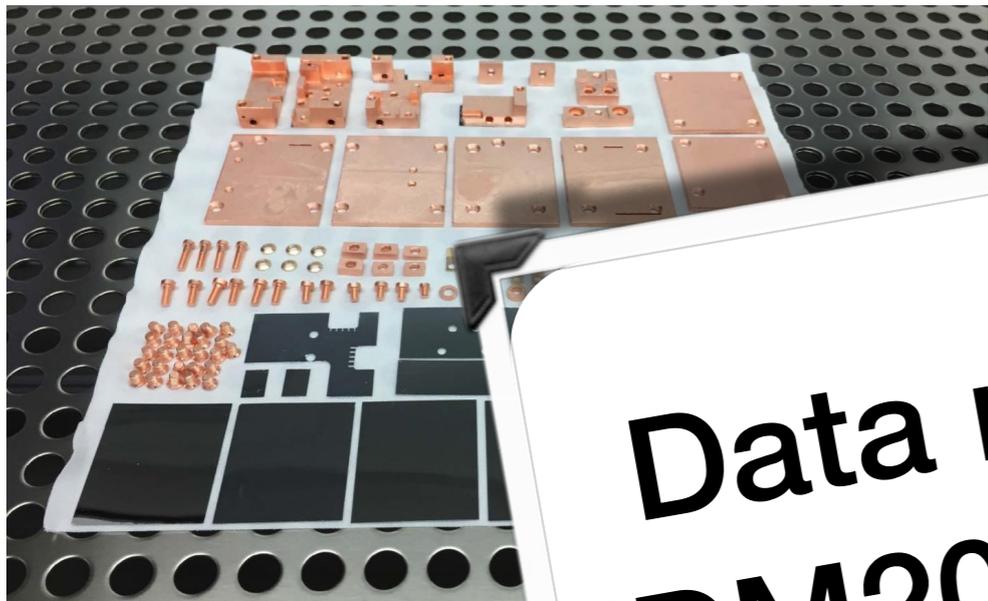
**Threshold design goal:  
100 eV**

# CRESST-III detectors



**10 detectors operating in Gran Sasso from July 2016 to February 2018**

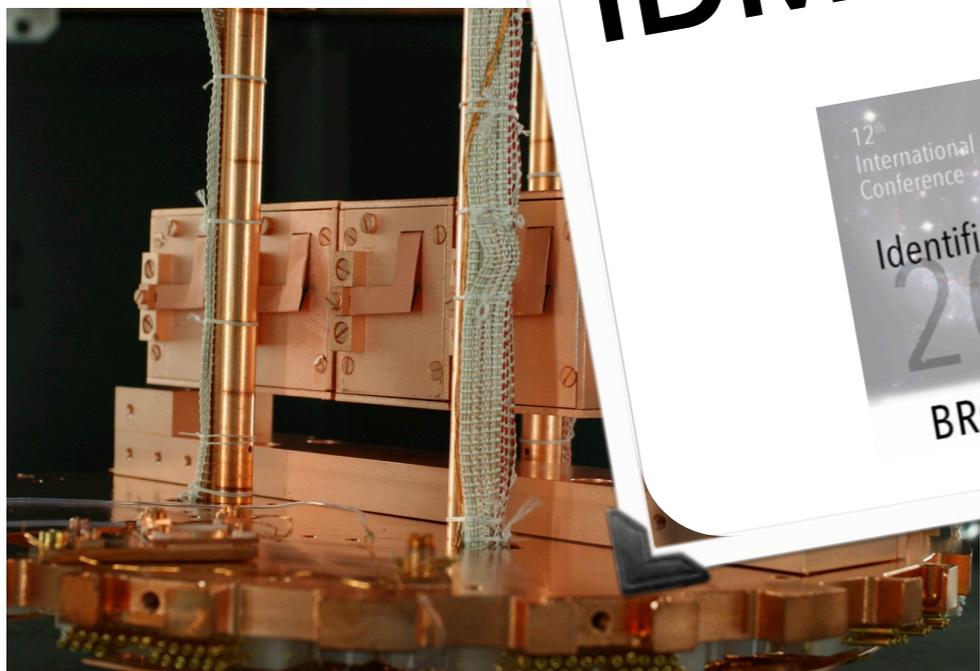
# CRESST-III detectors



Data release at the  
IDM2018 conference

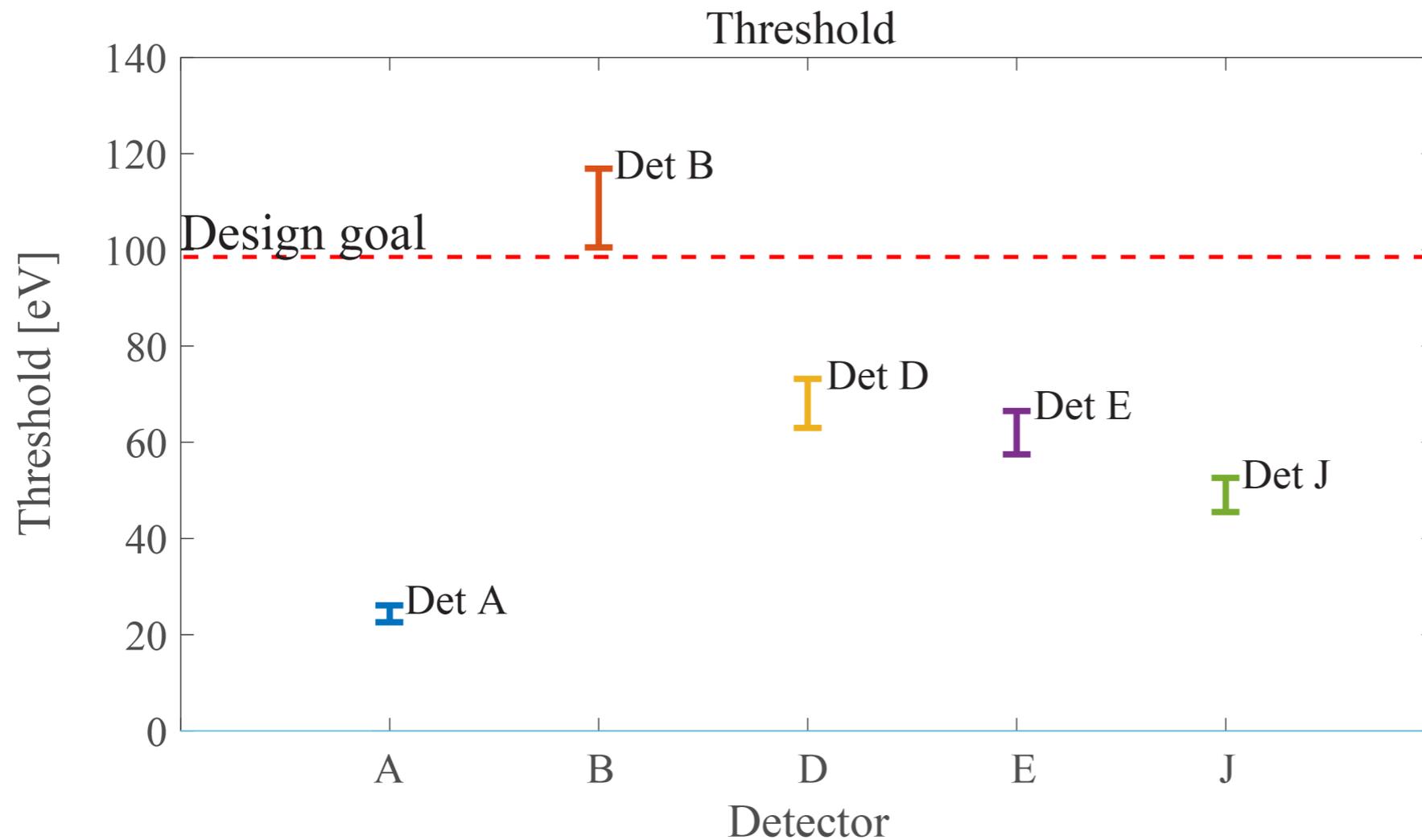


©T. Dettlaff/MPP



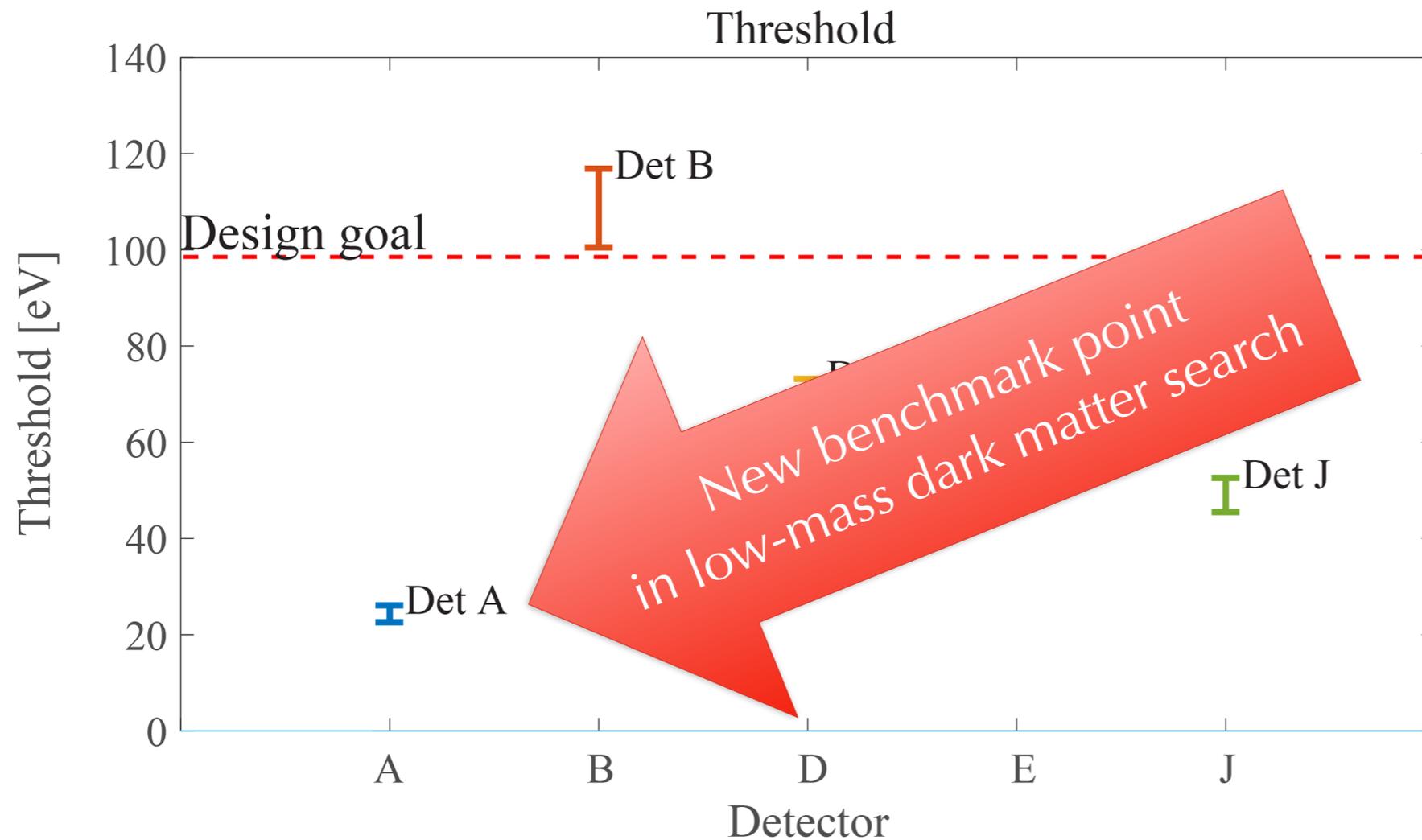
**10 detectors operating in Gran Sasso from July 2016 to February 2018**

# Optimum thresholds



5 detectors reach/  
exceed the  
CRESST-III design  
goal

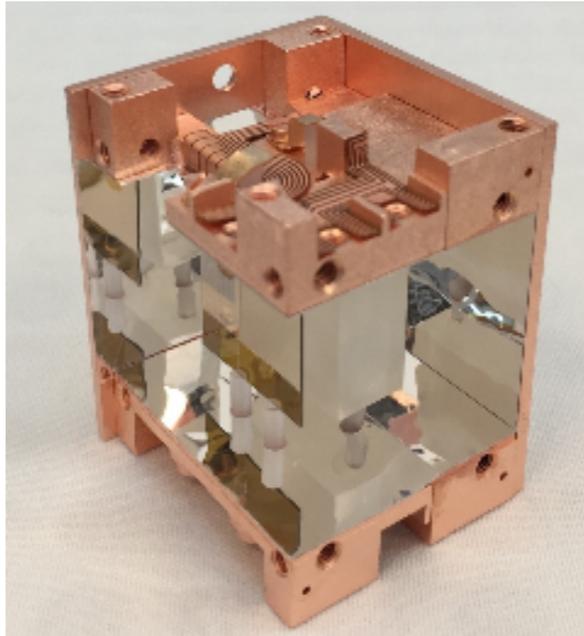
# Optimum thresholds



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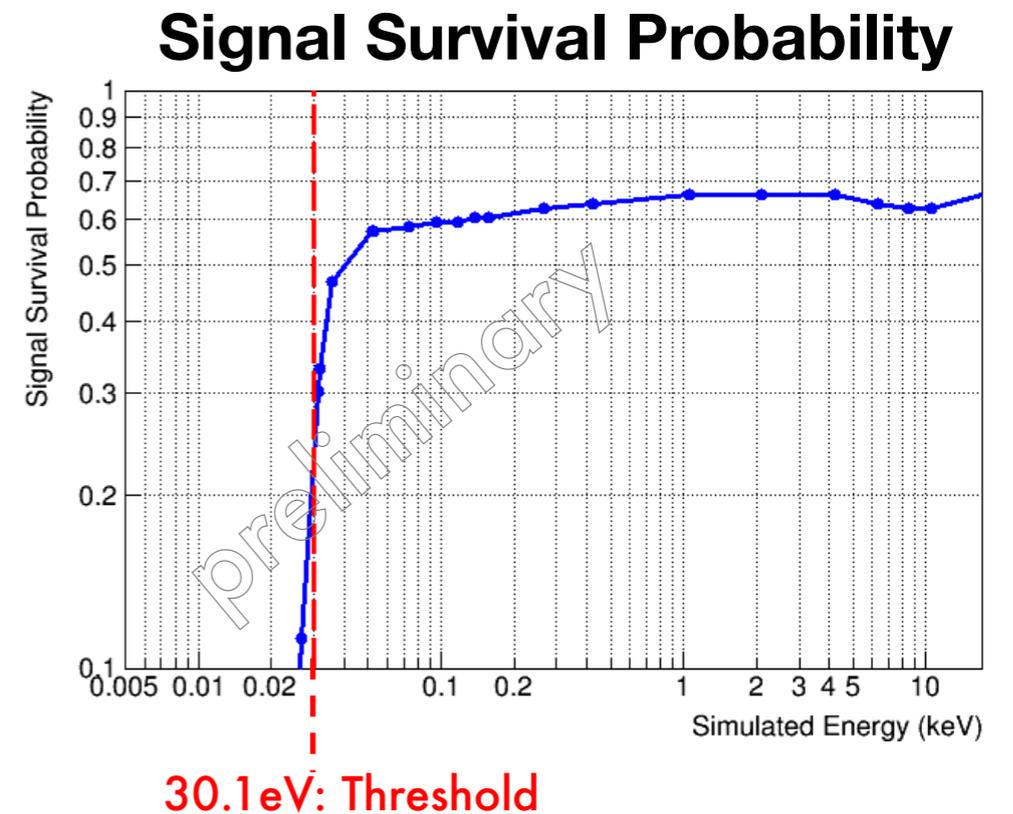
**NEW FRONTIER IN DIRECT DM DETECTION**

# Detector A

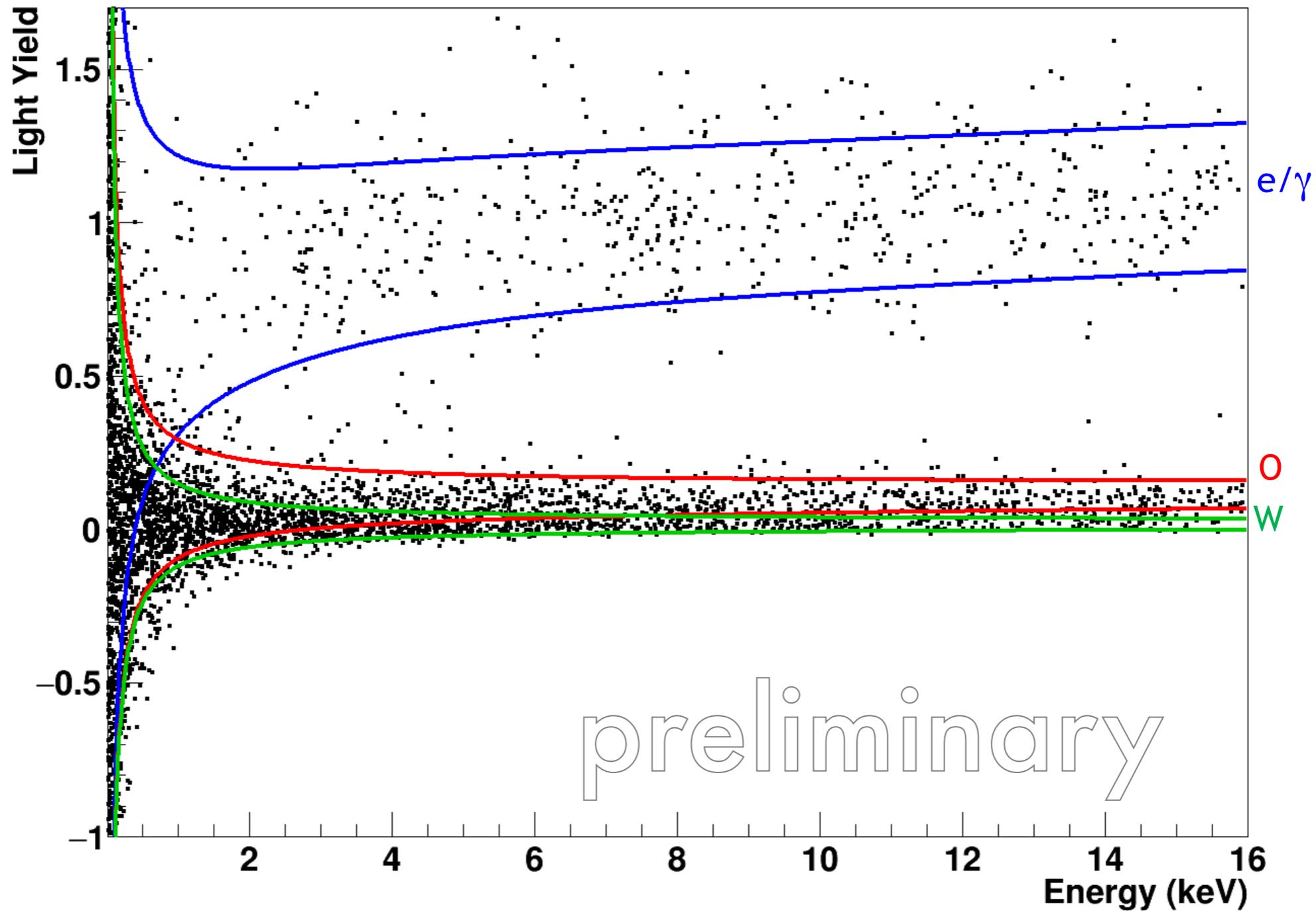


Data taking period: 10/2016 – 01/2018  
Target crystal mass: 23.6 g  
Gross exposure (before cuts): 5.7 kg days  
**Energy threshold: 30.1 eV**

- Analysis chain includes selections on:
  - *Rate*: to select stable noise conditions
  - *Stability*: to select detector(s) in operating point
  - *Data quality*: Non-standard pulse shapes are discarded
  - *Coincidences*: rejected events in coincidence with iSticks, with other detectors and with muon veto

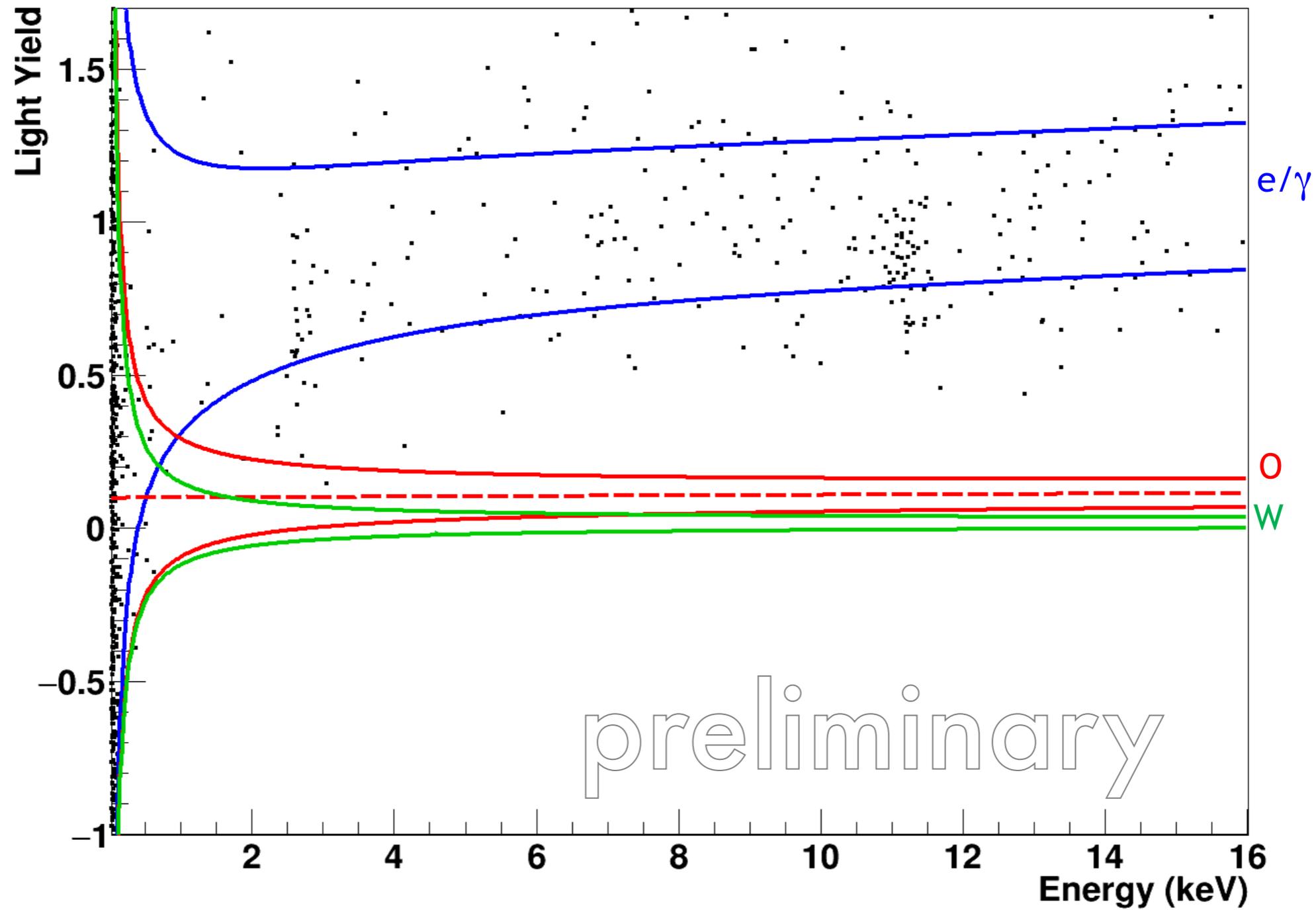


# Neutron calibration data

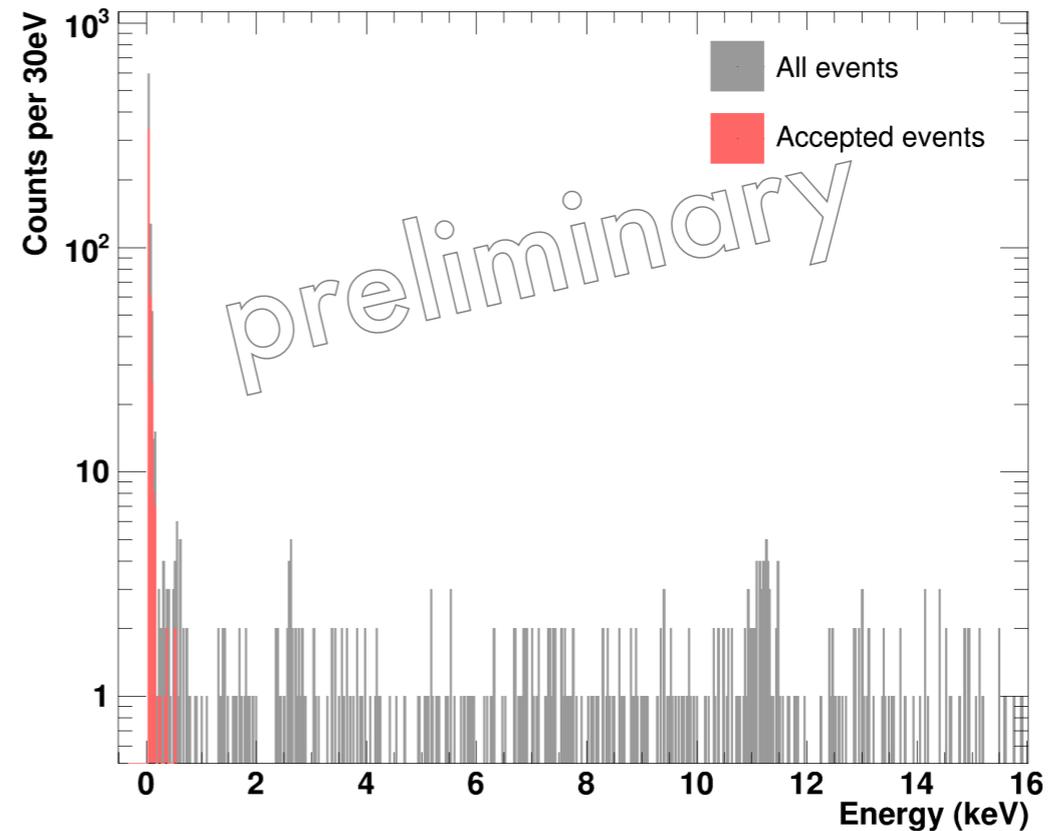
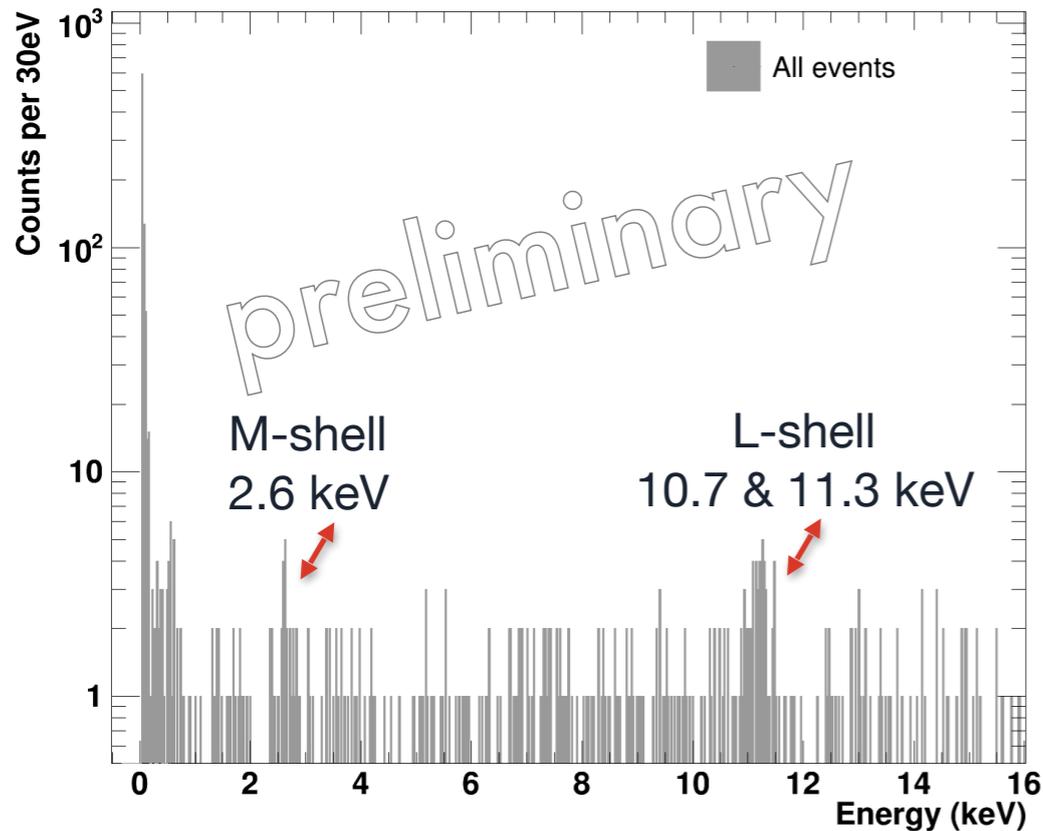


Quenching factors measured with neutron beam  
Unbinned maximum likelihood fit

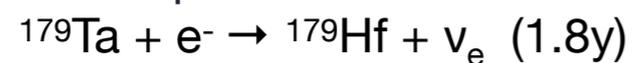
# Dark matter data



# Energy spectra

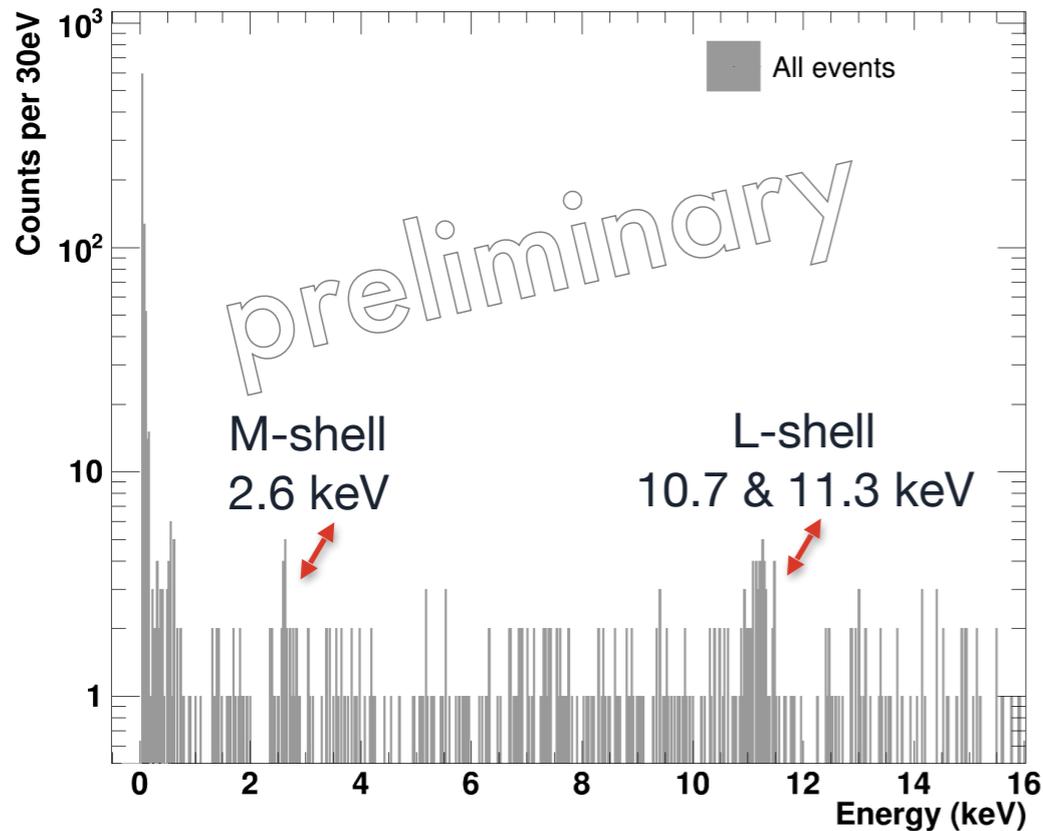


Cosmogenic activation



- 445 events in the acceptance region

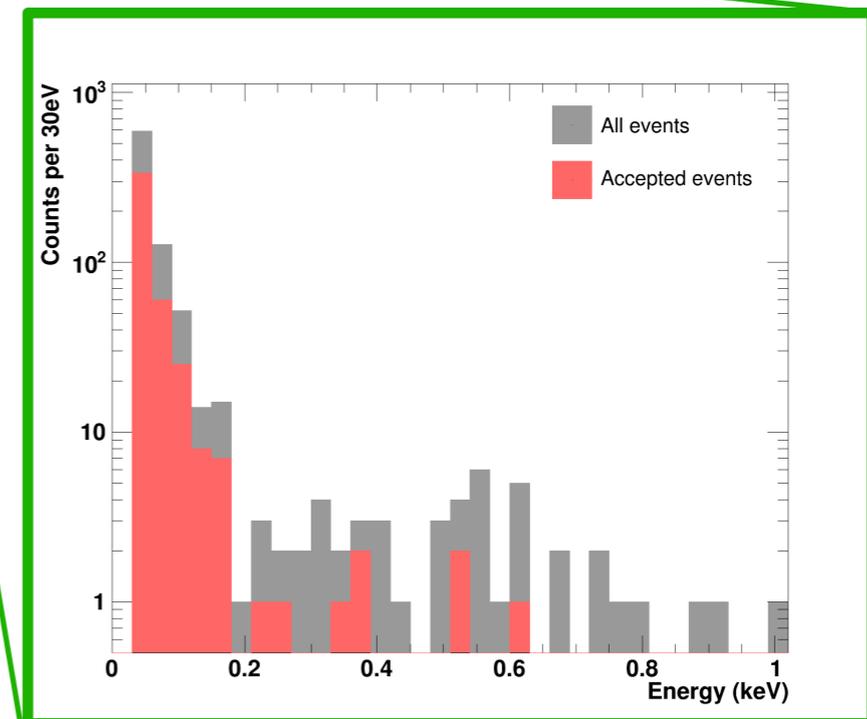
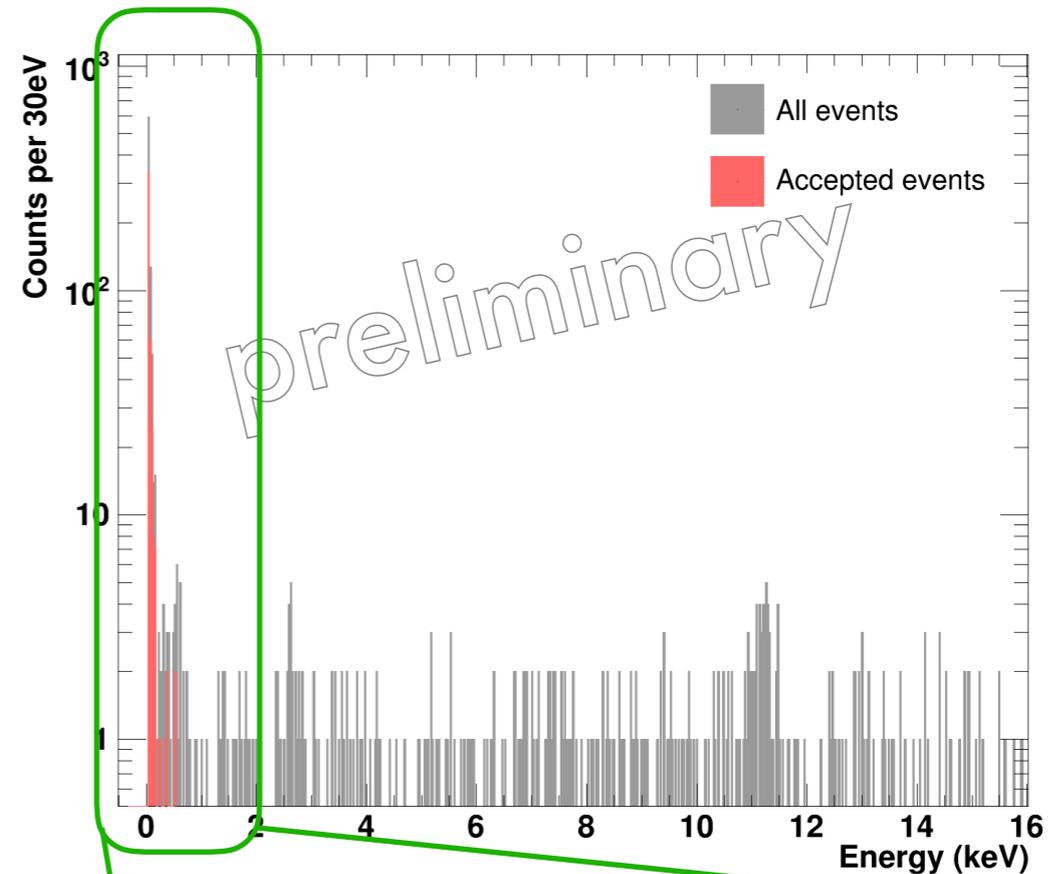
# Energy spectra



Cosmogenic activation



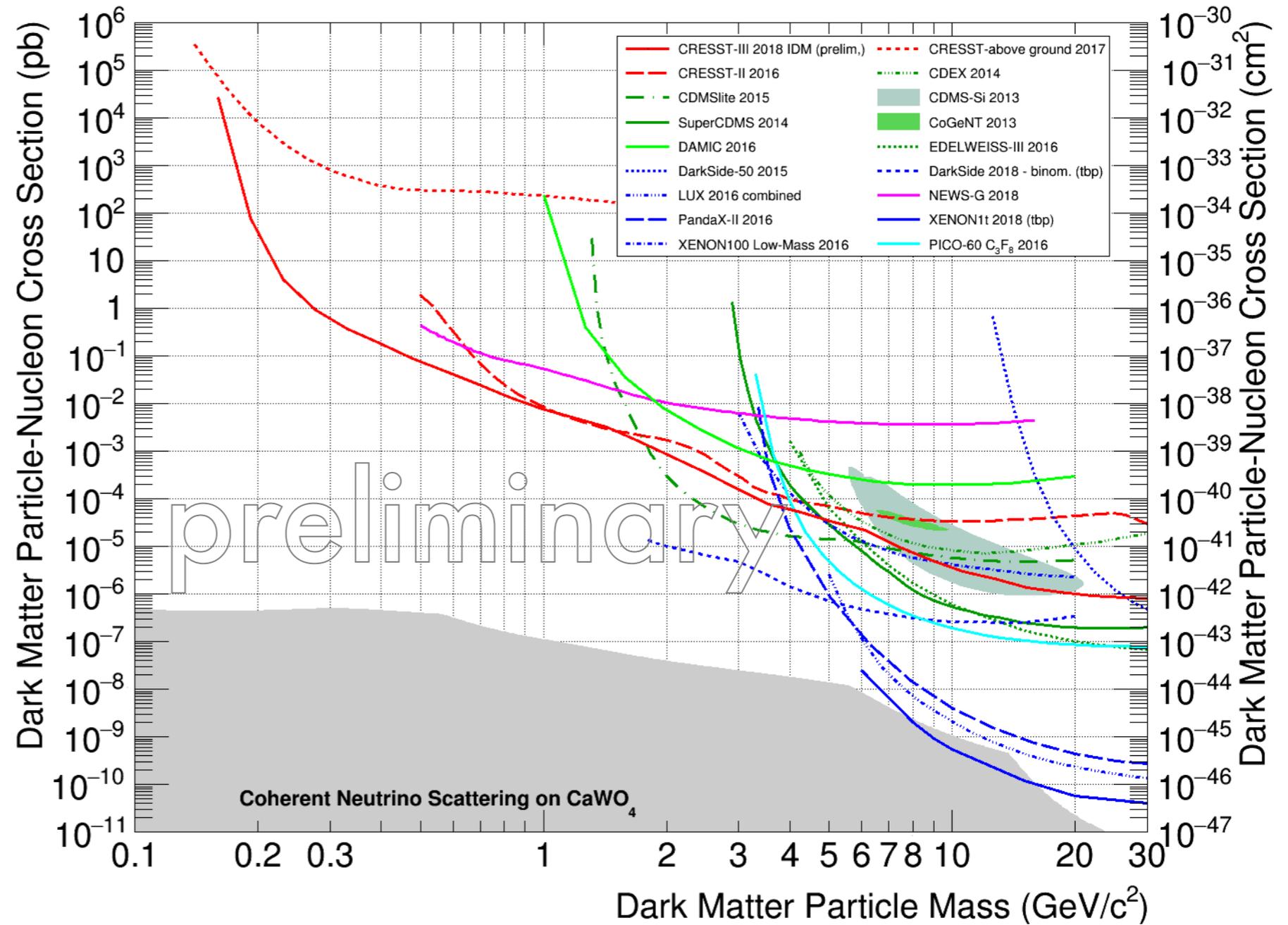
- 445 events in the acceptance region
- Unexpected rise of event rate <200 eV



# Result

1D Yellin optimum interval method to compute the exclusion limit:

Energy spectrum of accepted events  
+  
Expected DM energy spectrum

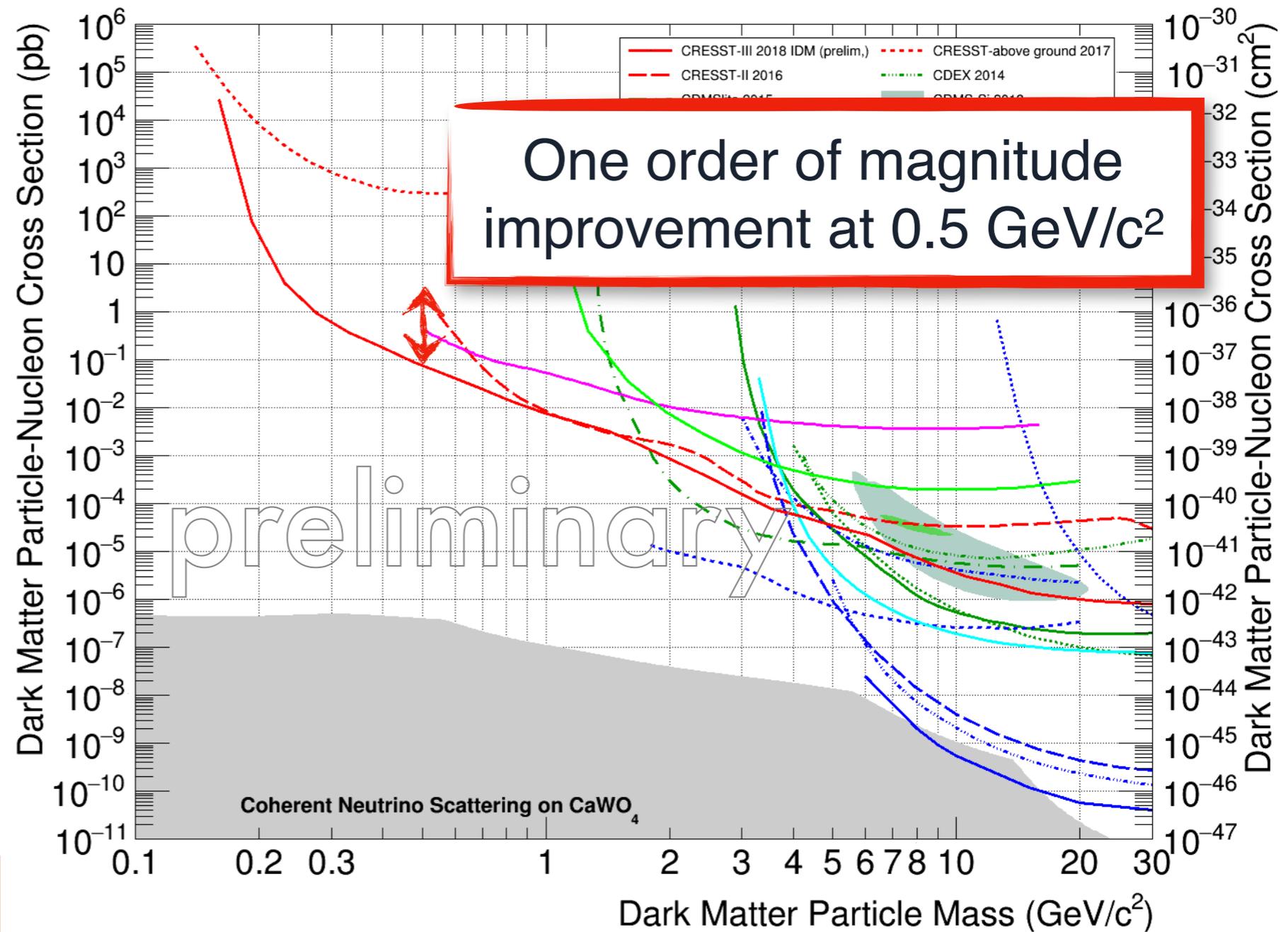


# Result

1D Yellin optimum interval method to compute the exclusion limit:

Energy spectrum of accepted events  
+  
Expected DM energy spectrum

World leading limit at low-mass  $< 1.7 \text{ GeV}/c^2$



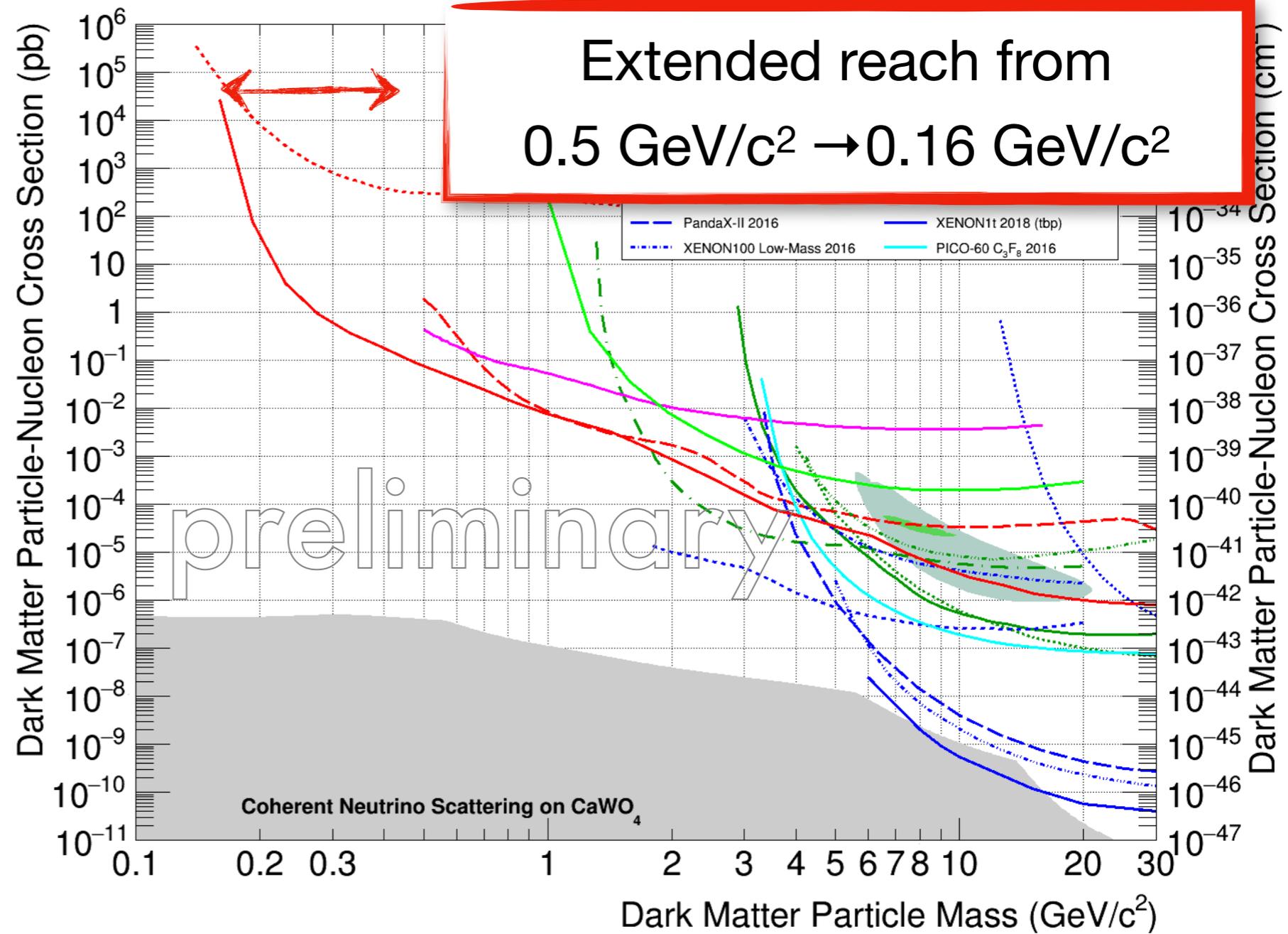
↳ Background limited

# Result

1D Yellin optimum interval method to compute the exclusion limit:

Energy spectrum of accepted events  
+  
Expected DM energy spectrum

Lowest limit  
>0.16 GeV/c<sup>2</sup>



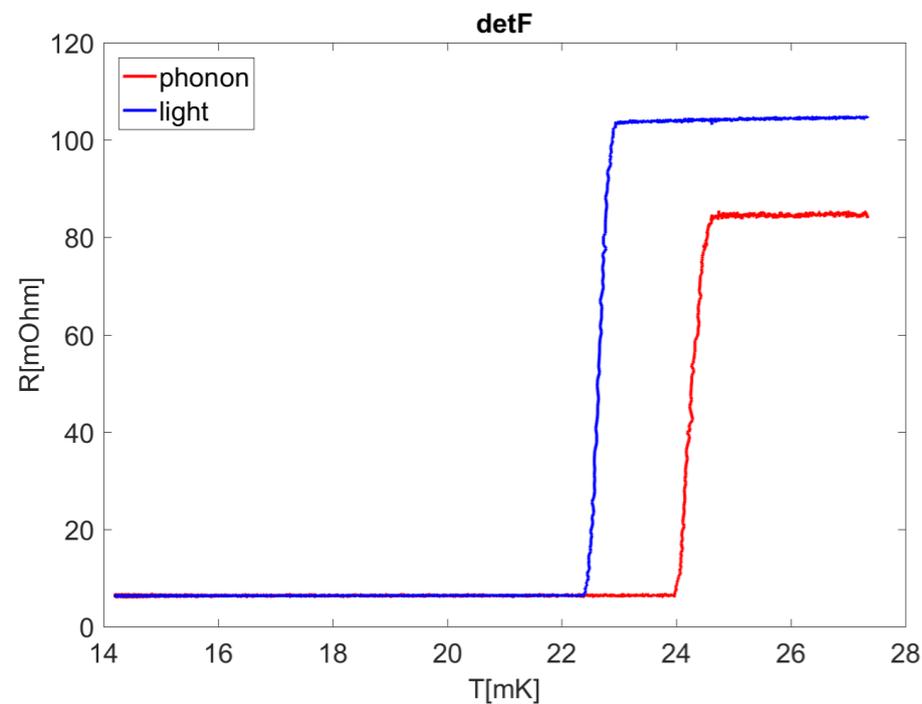
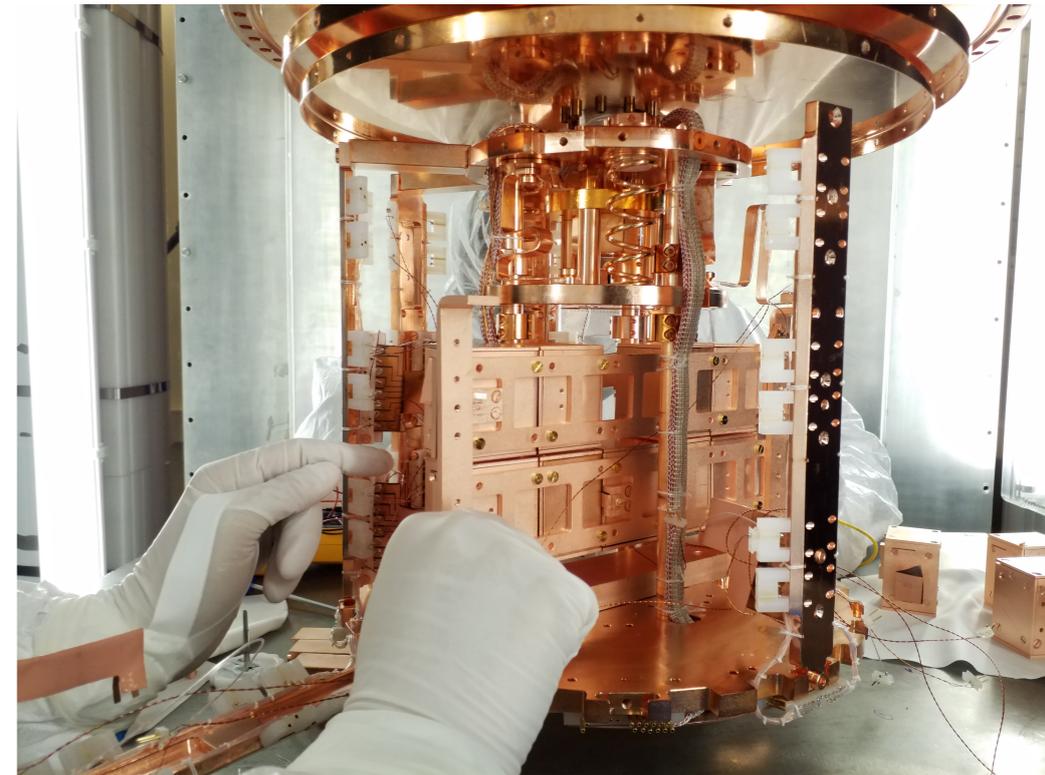
↳ Performance "limited"

# Current status

**April-May 2018:**

“Upgraded” detector modules have been prepared at MPP (crucial contribution from MPP Mechanical Workshop) and installed in Gran Sasso.

Dedicated hardware changes to understand source of excess events (different crystal absorbers, different detector holders, etc..)

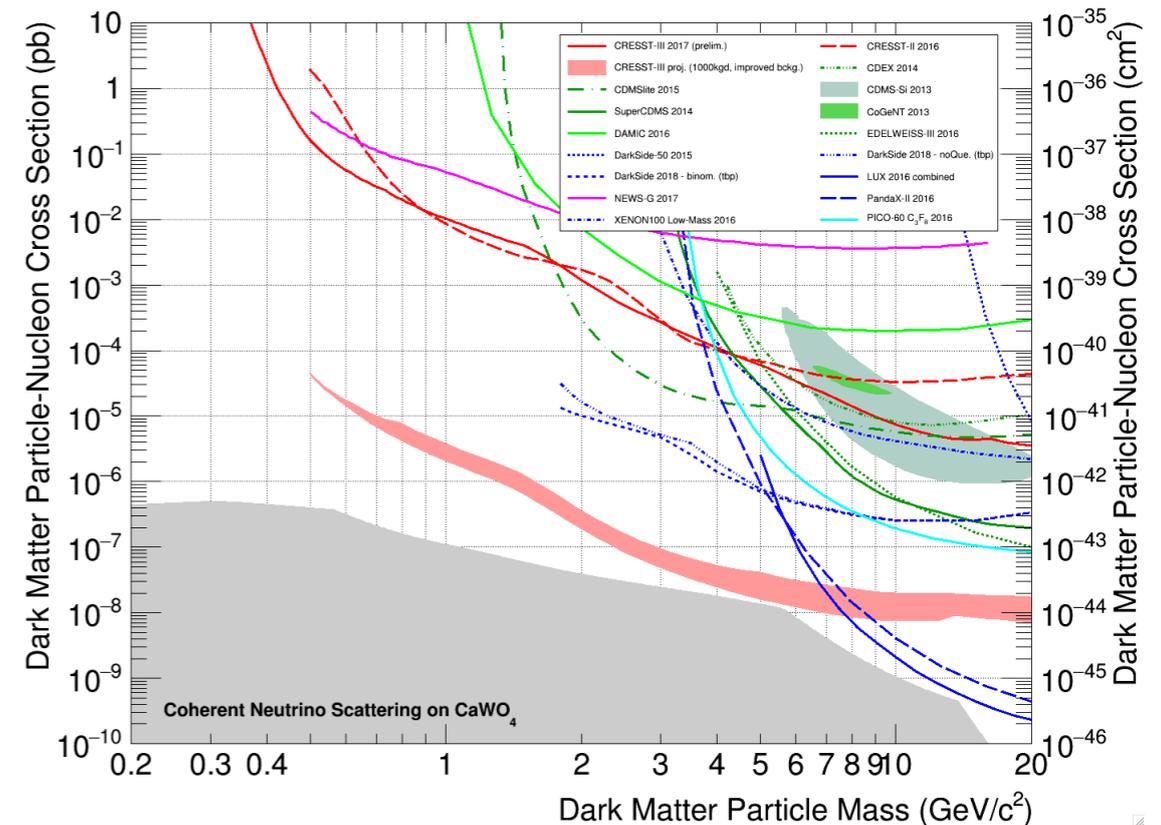


**First data from Gran Sasso  
in these days.**

# What's next? (2)

**Long term:** Major upgrade of the experiment is foreseen to start next year.

Goals: increase the number of channels in the LNGS setup to 100 and further improve threshold and background.



**Final planning, prototyping and testing of wiring and read out electronics**

**Detectors R&D:**

- Further **threshold reduction**
- Improved **reproducibility**
- Increased **production rate**
- Reduced **background**
- Different **materials**

**2019**

**Upgrade of the setup at LNGS**  
**Production of detectors**

**2020**

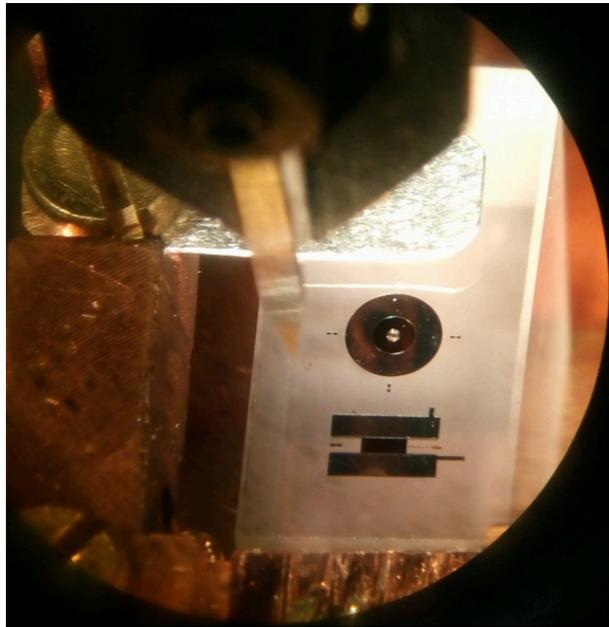
**Restart data taking**

**2021**

# MPP activities (non exhaustive list...)

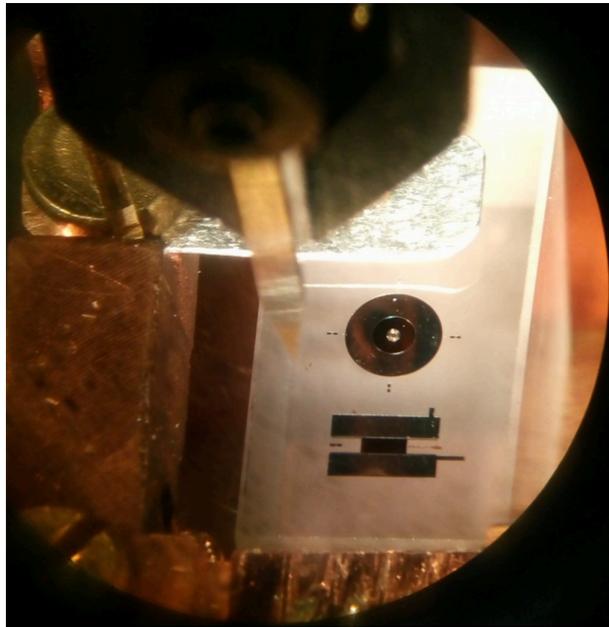
# MPP activities (non exhaustive list...)

## Sensor optimisation

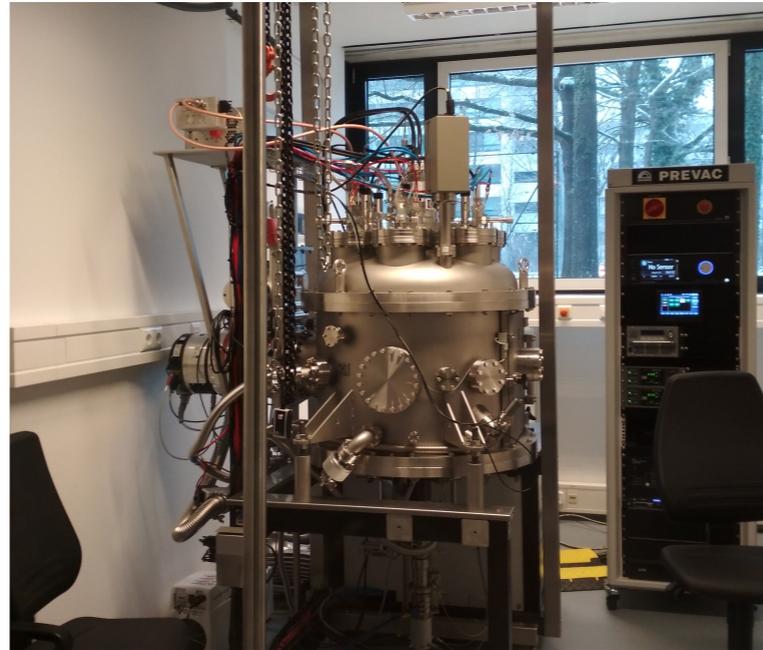


# MPP activities (non exhaustive list...)

**Sensor optimisation**

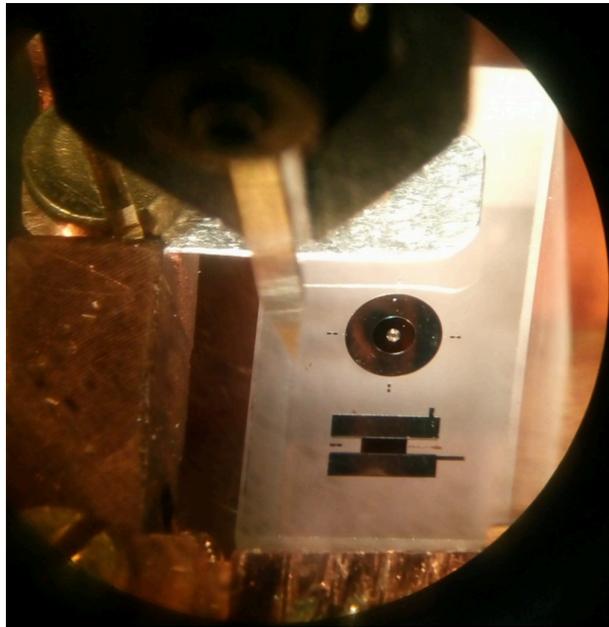


**Sensor reproducibility**

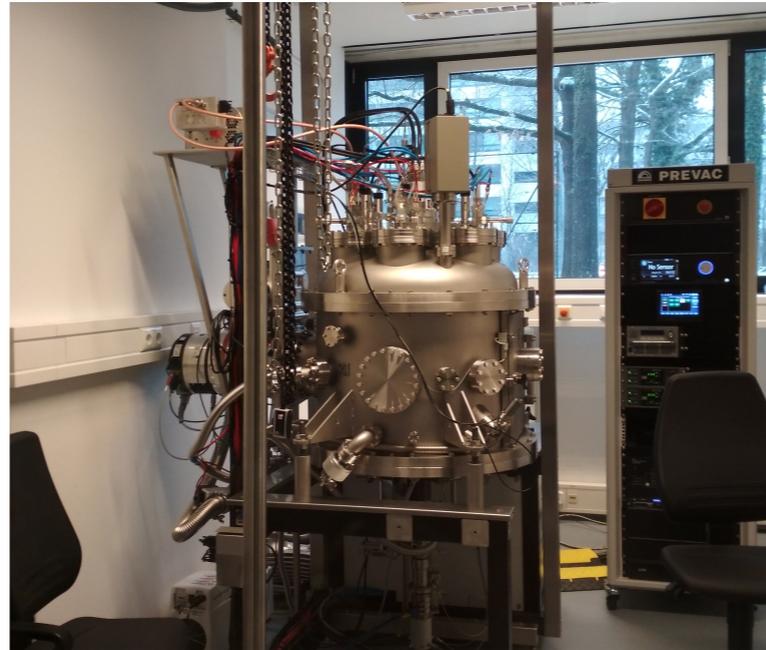


# MPP activities (non exhaustive list...)

**Sensor optimisation**



**Sensor reproducibility**

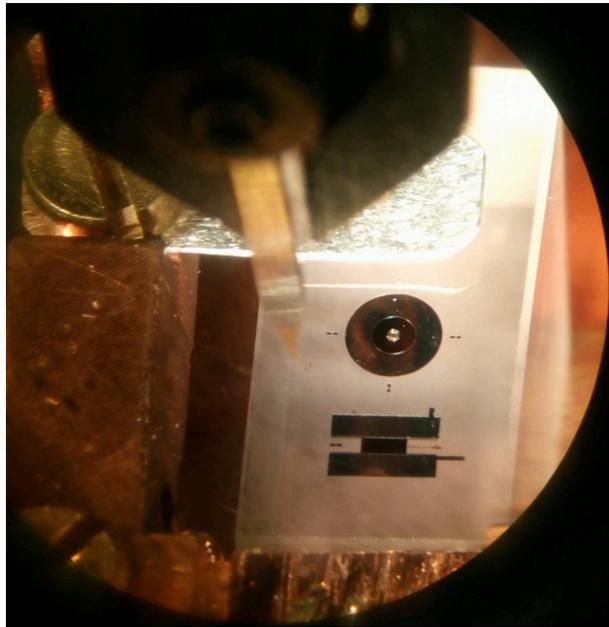


**New materials**

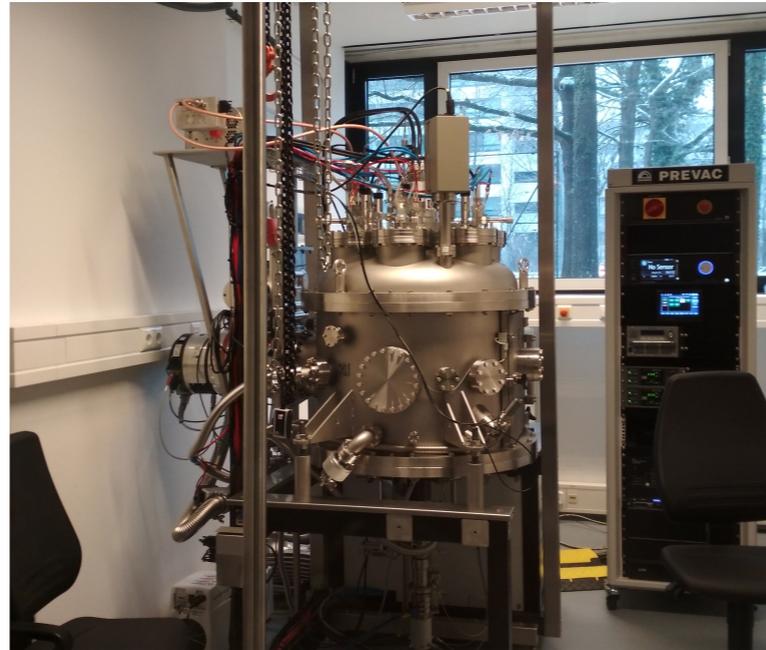


# MPP activities (non exhaustive list...)

**Sensor optimisation**



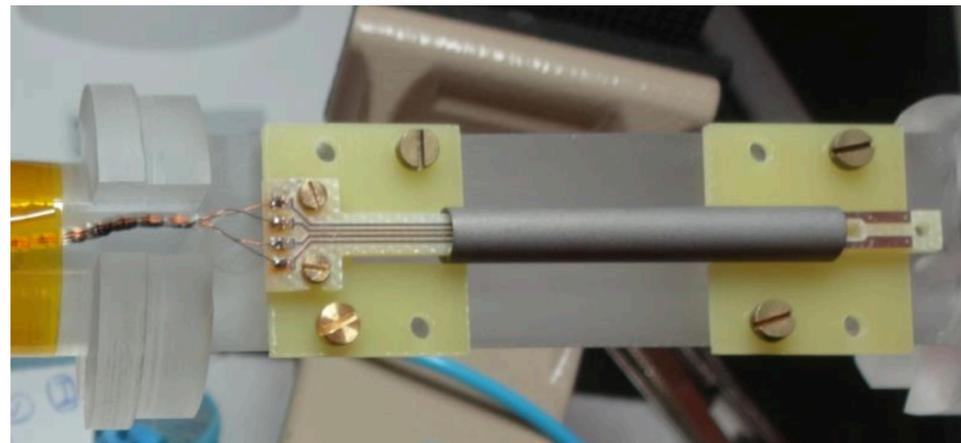
**Sensor reproducibility**



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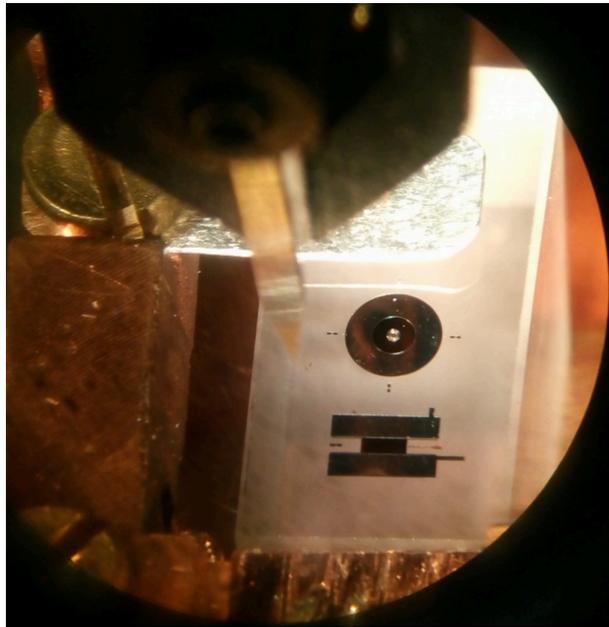


## **SQUID**

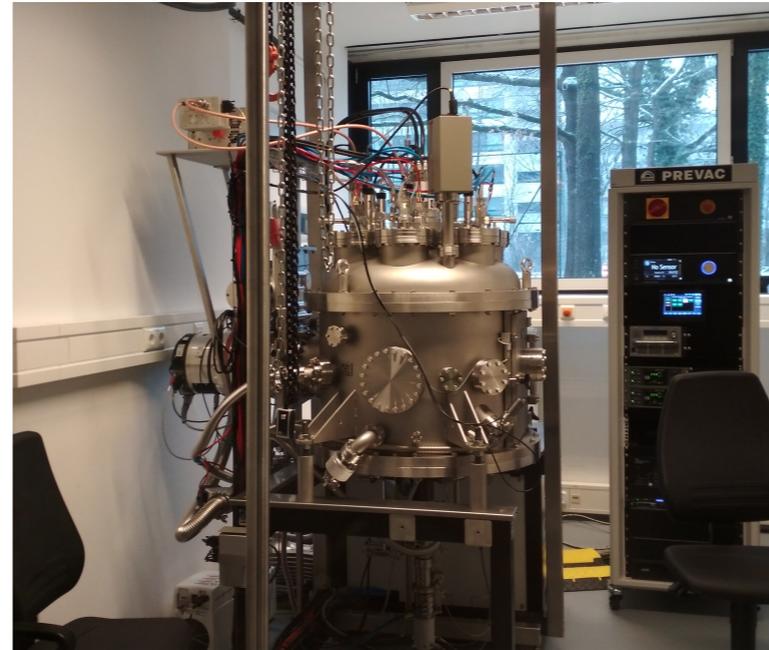


# MPP activities (non exhaustive list...)

**Sensor optimisation**



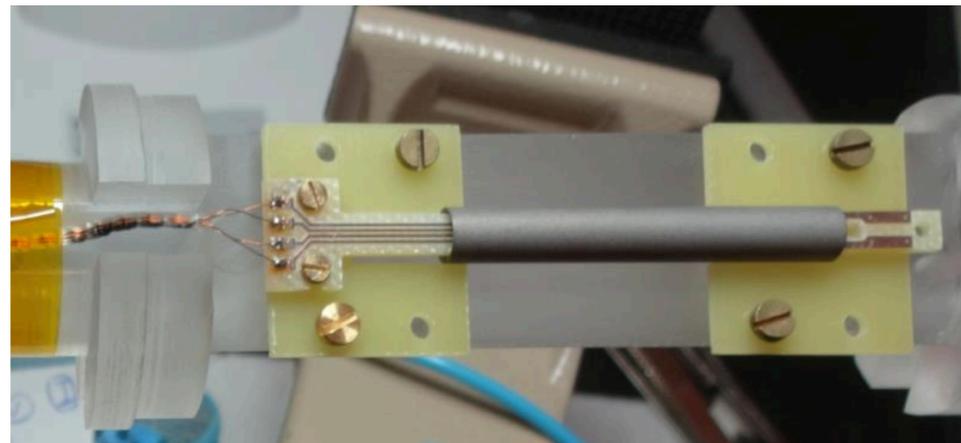
**Sensor reproducibility**



**New materials**



**SQUID**

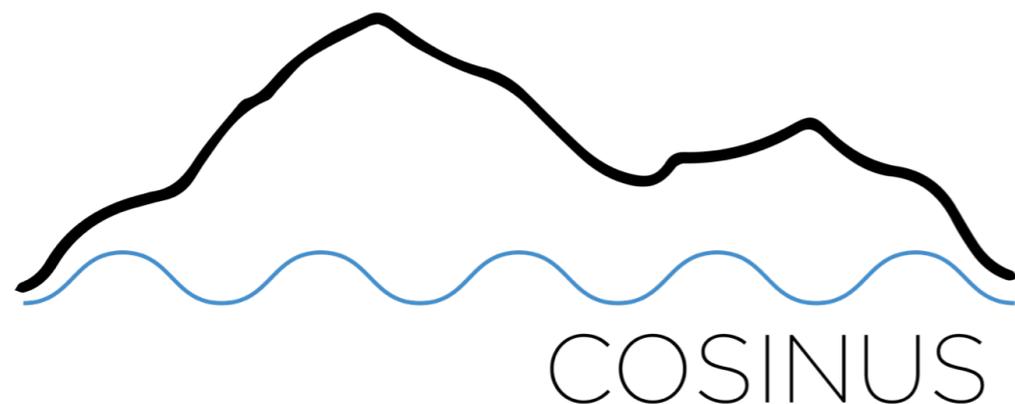


**Read-Out**



# CRESST “Spin-offs”

Spin-offs of the CRESST-MPP technology successfully funded in 2018  
→ from R&D to experiments



MPRG @ MPP, K. Shaeffner  
CRESST-like cryogenic detectors  
made of NaI to study the DAMA  
signal.



ERC Starting Grant 2018, 804228, R. Strauss  
Miniaturized CRESST-like detectors for  
studying CNNS at reactors.

# Conclusions

- Cryogenic calorimeters represent a well established technology for the investigation of dark matter and other rare event searches.
- **CRESST** has reached an unprecedented low nuclear recoil threshold of 30 eV and is leading sensitivity for light DM searches at 160 MeV/c<sup>2</sup>.
- **CRESST - MPP** is the leading institution in the experiment, developing highly sensitive cryogenic particle detectors.
- A new explorative run is ongoing at LNGS to investigate the source of excess events.
- Preparation for the major upgrade of the LNGS experiment.