

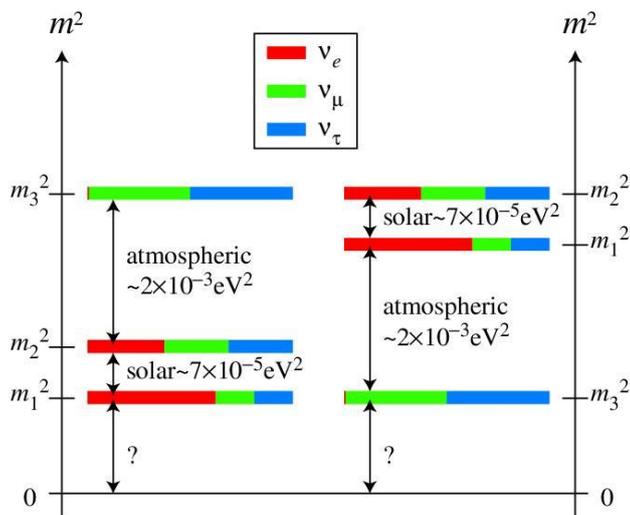
Massive Neutrinos:

What we know about neutrinos:

- Neutrino flavors mix with each other
  - At least two neutrino flavors must have a small mass
  - Contain CP violating phase
- They have no electric charge
- Weak interaction couples only to left handed (Chirality=-1) neutrinos and right handed (Chirality=+1) anti-neutrinos (V-A) structure of the interaction

What we would like to know about neutrinos:

- What is the mass scale of neutrinos?
  - So far: limits from
    - cosmology on sum of neutrino masses (structure formation)
    - end point measurements of electron spectrum (KATRIN,...)
- Why are neutrino masses so small?
- Which mass hierarchy do neutrinos obey, i.e. which flavor is the lightest?



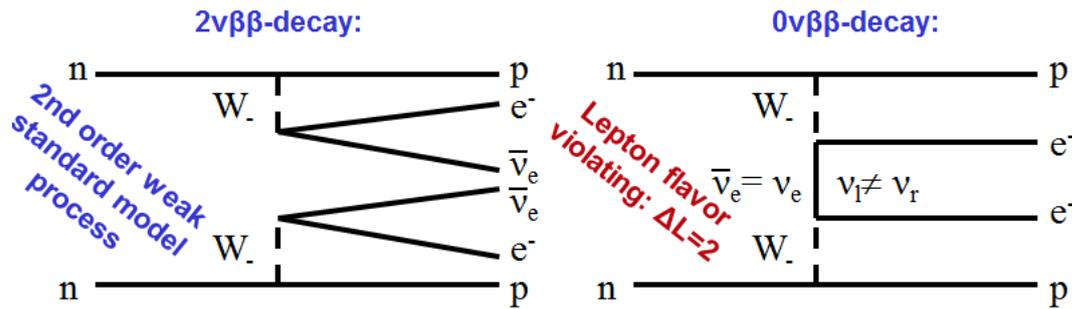
- As neutrinos have no charge: what does CPT-conjugation do to neutrinos?
  - Are neutrinos and anti-neutrinos identical? I.e. is the neutrino a Majorana particle?
  - This would imply Lepton number violation!

Also remember: Majorana neutrino could help in BAU (decay of heavy neutrinos CP violating)!

Neutrinoless Double Beta Decay:

To tackle these questions → Neutrinoless double beta decay with  $|\Delta L| = 2$ :

$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$



$2\nu\beta\beta$  – decay can be observed if single  $\beta$  – decay is energetically forbidden, while coherent decay of two neutrons into two protons is allowed as  $2^{\text{nd}}$  order weak process  
 $\rightarrow 2\nu\beta\beta$  Half-lives of isotopes are of the order  $10^{18} - 10^{21} \text{yr}$ .

$0\nu\beta\beta$  –decay can only occur if:

- Neutrino has Majorana nature  $\rightarrow$  Lepton number violation!
- Helicity flip occurs in the vertex Standard assumption: if decay exists: dominated by exchange of Majorana neutrino. In general: all Lepton flavor violating processes leading to helicity flip can induce  $0\nu\beta\beta$  –decay.

Observation is possible for even even nuclei due to pair binding term in Weizsäcker formula:  
 Two parabolas of binding energy vs. Z:

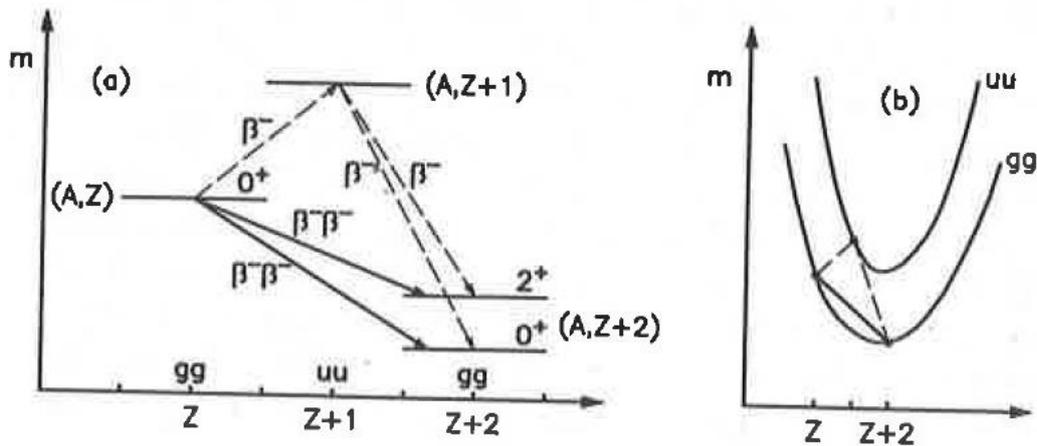
Half-life limits:  $\sim 10^{26} \text{yr}$ !

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Left: Scheme for decay of a gg nucleus- single beta decay is forbidden

Candidate isotopes (among others):  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{130}\text{Te}$ ,  $^{100}\text{Mo}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$

For exchange of light Majorana neutrino:

$$\left(T_{1/2}^{0\nu\beta\beta}\right)^{-1} = G(Q^5) |\mathcal{M}_{Nucl}|^2 \langle m_{ee} \rangle^2$$

with  $G(Q^5)$  the phase space factor of the decay,  $\mathcal{M}_{Nucl}$  the nuclear matrix element and

$$\langle m_{ee} \rangle = \left| \sum_i U_{ei}^2 m_i \right|$$

Nuclear matrix element calculations uncertain to factor of  $\sim 2!$  Phase space factor uncertainty due to axial vector coupling.....

The effective Majorana electron neutrino mass (note that phases appear in  $U_{ei}$ ):

$$\begin{aligned} |\langle m \rangle| &= \left| m_1 U_{e1}^2 + m_2 U_{e2}^2 + m_3 U_{e3}^2 \right| \\ &= \left| \left( m_1 c_{12}^2 + m_2 s_{12}^2 e^{i\alpha_{21}} \right) c_{13}^2 + m_3 s_{13}^2 e^{i(\alpha_{31} - 2\delta)} \right| \end{aligned}$$

→ Can express this in terms of neutrino mixin parameters

$$\Delta m_{21}^2, \sin^2 \theta_{12}, \Delta m_{31}^2 \text{ and } \sin^2 \theta_{13}$$

→  $|\langle m \rangle|$  related to effective mass of lightest neutrino (assessed with beta-end point measurements):

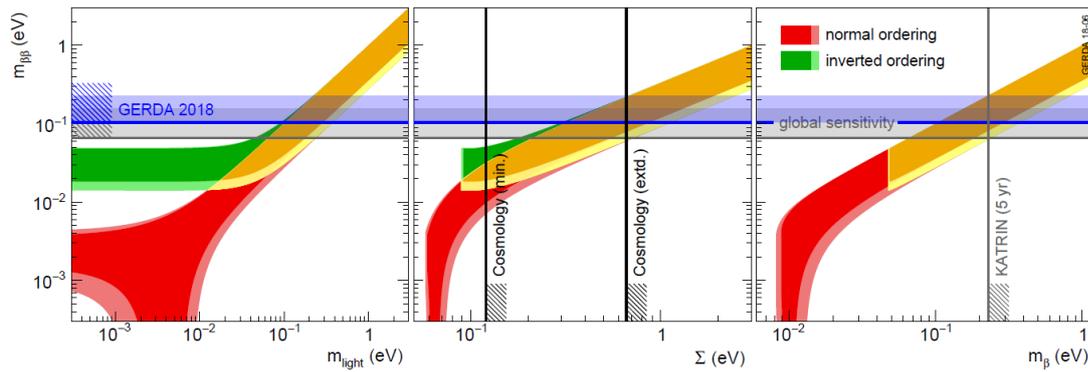
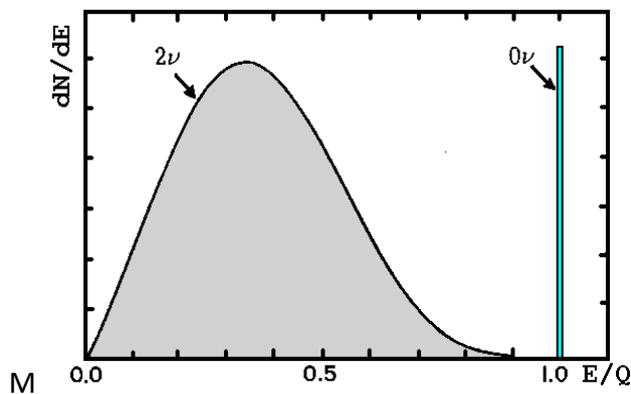


Figure 3: Constraints of the parameter space for  $m_{\beta\beta}$  in the scenario of three light Majorana neutrinos as function of the lightest neutrino mass  $m_{light}$ , the sum of neutrino masses  $\Sigma$ , and the effective neutrino mass  $m_{\beta}$ . Contours follow from a scan of the Majorana phases (dark colours) and the  $3\sigma$  intervals of the neutrino oscillation data (light colours) within NuFIT (4). The blue horizontal band shows the upper limits on  $m_{\beta\beta}$  obtained by GERDA, the grey band those from combining sensitivities of all leading experiments in the field (see Table 1). Vertical lines denote  $\Sigma = 0.12$  eV and  $\Sigma = 0.66$  eV, a stringent limit from cosmology (18) and an extended model bound (7), as well as  $m_{\beta} = 0.23$  eV, the 5 yr sensitivity of the KATRIN experiment (17).

Effective Majorana neutrino mass vs. mass of lightest neutrino allowed regions for different hierarchies

Experimental search: "Signature":



*Schematic sketch of energy spectrum of  $2\nu\beta\beta$ - and  $0\nu\beta\beta$  decays. While for the neutrino accompanied mode the spectrum is continuous, for the neutrinoless mode a sharp peak at the Q-value of the decay is expected. Peak height not to scale!*

Expected rates are extremely low: for  $\langle m_{ee} \rangle = 10$  meV expect  $\lesssim 1$  Events/(ton yr)!

Half life [years]	~Signal [cts/(t·yr)]
$1 \cdot 10^{26}$	50
$5 \cdot 10^{27}$	1
$1 \cdot 10^{29}$	0.05

Figure of merit for an experiment: Limit setting sensitivity for an experiment with, abundance  $a$  of the double beta emitting isotope signal detection efficiency  $\varepsilon$ , total detector mass  $m$ , measuring time  $t$ , background index  $B$  and energy resolution  $\Delta$  in the energy region of interest :

$$T_{1/2}^{0\nu\beta\beta} \propto a\varepsilon \sqrt{\frac{mt}{\Delta B}}$$

$a$ :Enrichment fraction,  $\varepsilon$ :detection efficiency,  $m$ :mass of the detector,  $\Delta$ :RoI used (related to energy resolution),  $B$ :Background index.

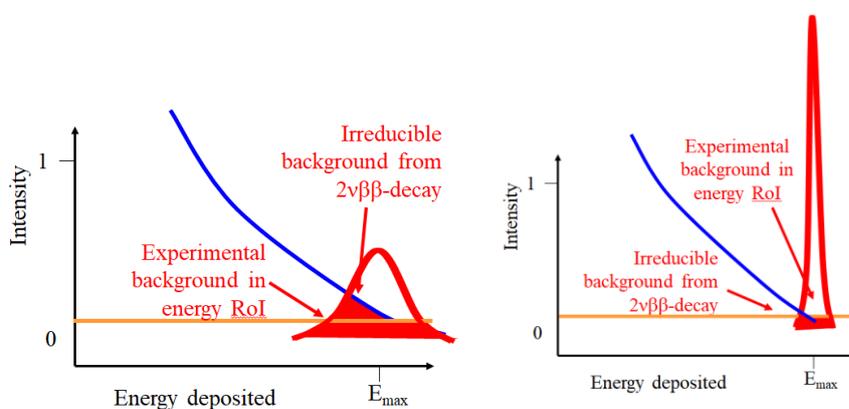
Strategy: look for peak at the Q-value of the decay.

- avoid and identify backgrounds:
- excellent energy resolution (background)
- Detector made from double beta emitting isotope (detection efficiency)
- Maximize detector mass  $\rightarrow$  isotopically enrich material

#### Example: The GERDA experiments:

Use high purity germanium radiation detectors: reverse biased diode: energy deposition create electron-hole pairs that can drift in E-field to contacts. Read- out using charge sensitive preamplifiers (Field Emission transistors – FETs):

- Excellent energy resolution (0.15% @ Q-value),

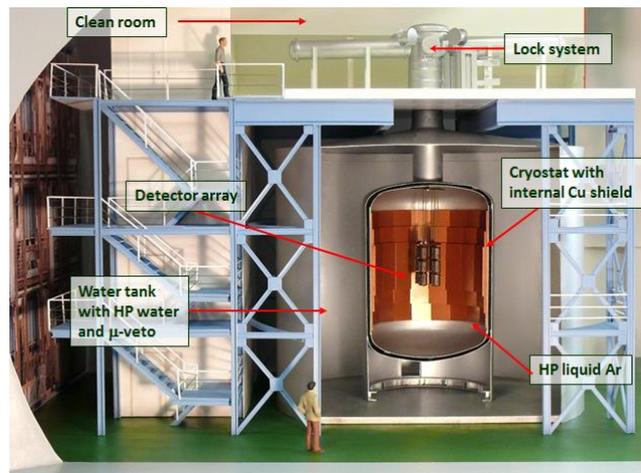


In order to deplete large volumes: very clean material needed (no electrically active impurities) → Only around  $10^{10}$  impurities per  $\text{cm}^3$  allowed (one in  $5 \cdot 10^{12}$ )

- Intrinsically low in long lived radioactive isotopes

Germanium material can be enriched in isotope  $^{76}\text{Ge}$  to  $\sim 87\%$ . Centrifugation of germanium tetrafluorid (same process as for uranium enrichment)

→ Detector from source material → high detection efficiency



*Schematic of GERDA experiment at LNGS.*

The GERD experiment is located at LNGS underground laboratory in Italy.

→ Shield cosmic rays by 1.4km of rock overburden:  $\sim 3400$  mwe. Hard component (neutrons, protons, pions,...) shielded completely,  $\mu$ - component reduced by factor  $\sim 10^6$ .

- Avoid background:

All materials used selected for very natural low background:  $^{238}\text{U}$ ,  $^{232}\text{Th}$  chains +  $^{222}\text{Rn}$  from  $^{226}\text{Ra}$  sub-chain: Use HPGe screening, ICPMS characterization, Rn emanation measurements.

Germanium and copper stored underground, careful logistics: Avoid cosmogenically produced long lived isotopes:  $^{60}\text{Co}$ ,  $^{68}\text{Ge}$ , ...

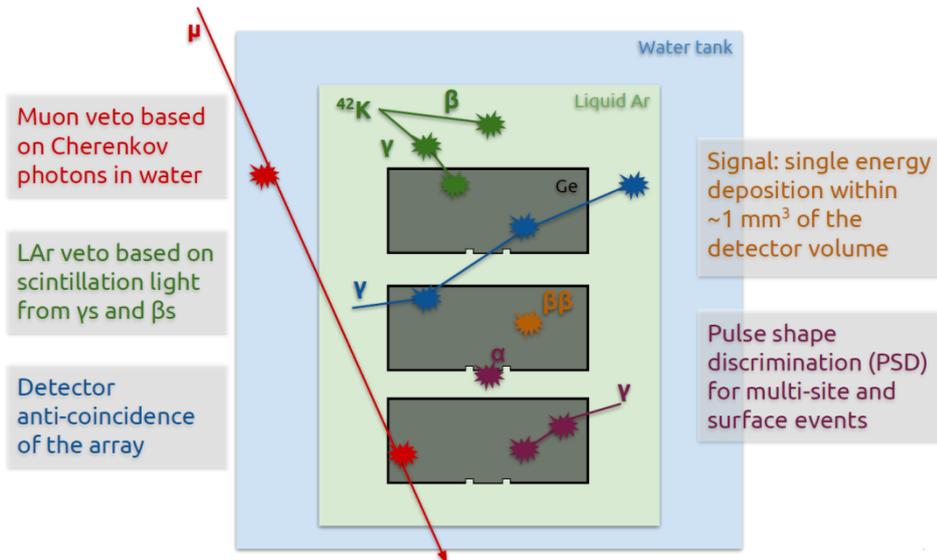
Water tank to shield external natural radioactivity + thermalize and capture neutrons.

10m tank of ultra clean water: equipped with PMTs to detect Čerenkov light from through-going muons.

LAr Cryostat with 4m diameter placed inside. LAr for shielding of natural radioactivity from steel walls (specially selected)

Deploy array of 40 detectors (10 coaxial – 20kg, and 30 BEGe detectors – 20kg) directly in liquid argon (LAr): Use liquid argon as cooling medium for detectors and as ultra clean shield against background radiation.

Identification of background:



- *Myon veto:*

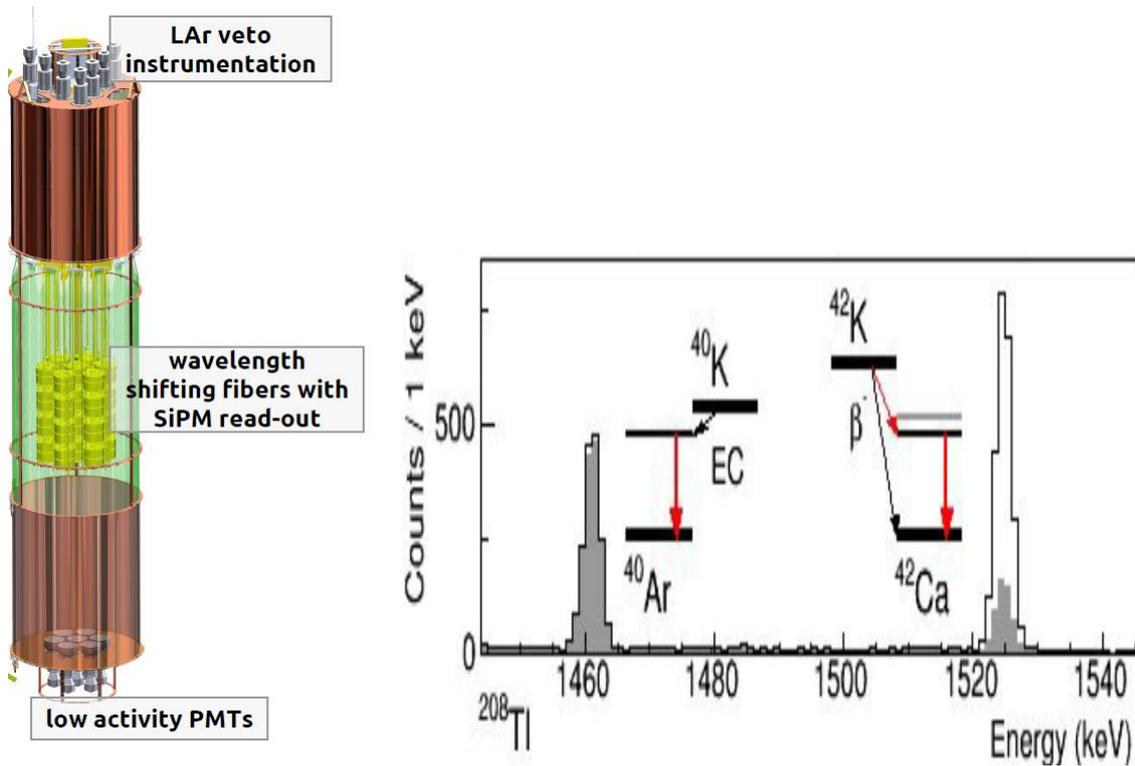
Cerenkov light detection in water around LAr tank

→ Water tank is equipped with PMTs → >99% myon detection probability

- *LAr veto:*

LAr scintillates at  $\sim 128\text{nm}$  → Detect scintillation light for Bkg tagging!

→ Deploy optical fibre curtain equipped with SiPMs around detector array + view array from top and bottom using PMTs

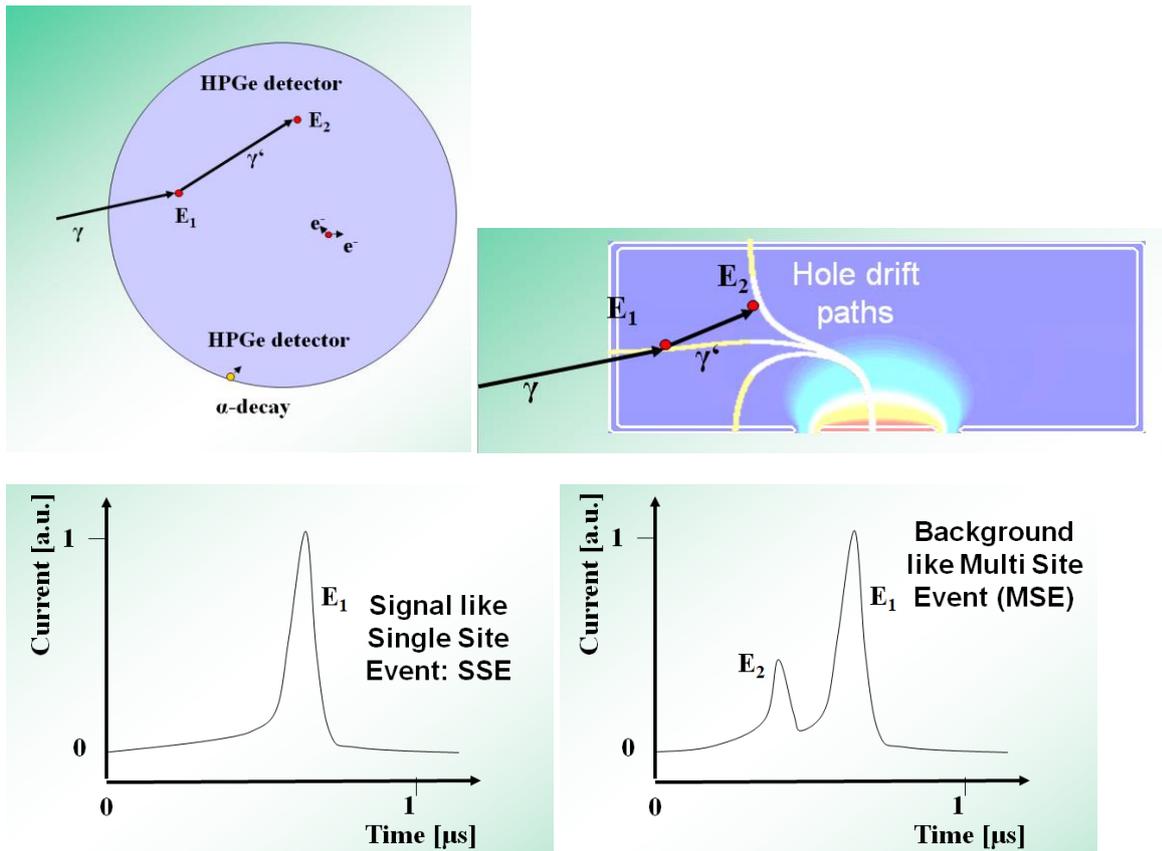


Left: Conceptual drawing of GERDA detector string inside LAr veto system. Right: Spectrum around the two peaks from decay of  $^{40}\text{K}$  and  $^{42}\text{K}$  measured in the GERDA setup. While the  $^{42}\text{K}$  is reduced considerably by the LAr veto cut, the  $^{40}\text{K}$  is not, as expected.

- *Pulse shapes to distinguish background from signal:*

Distinction between

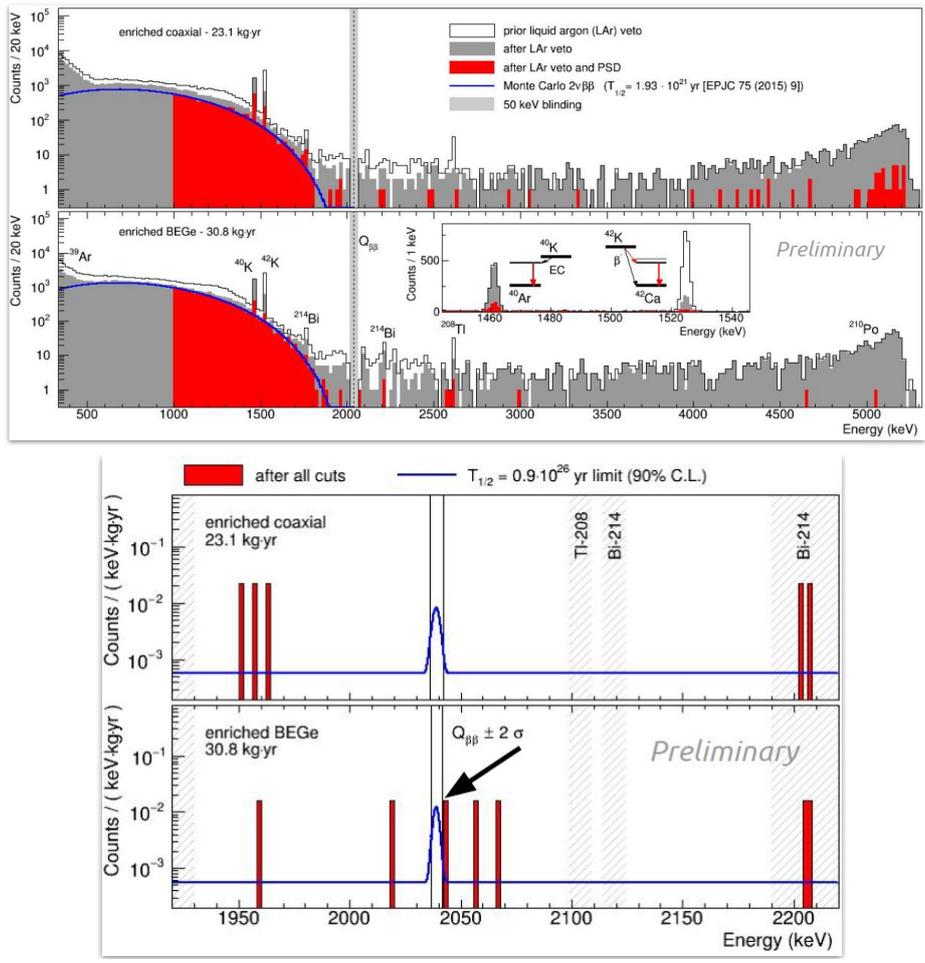
- signal like Single Site Events: two betas deposit energy within  $\sim 1\text{mm}^3$
- background from gamma rays: 2 MeV gammas mean free path after Compton scattering  $\sim \text{cm}$ )  $\rightarrow$  multiple energy depositions within detector
- p+ surface events: very fast,
- n+ surface: slow due to low field region and partly diffusion into active volume



Differences in event topologies can lead to differences in pulse shapes of events:  $\gamma$  -rays at 2 MeV are most likely to interact via Compton scattering with mean free path of few cm in germanium  $\rightarrow$  events with more than one energy deposit inside detector can be identified as background events (mean free path of electrons with 1 MeV around mm).

GERDA Results:

Took data from Nov 2011 to June 2019 (without LAr veto): BI in RoI: 0.01 Cts/(kg yr keV) with  $\Delta@RoI$ : 4.8 keV for coaxial detectors, 3.2 keV for BEGe detectors.



Top: Overall spectrum of Phase II measurement Result from Phase II showing the low background and the efficiency of the background identification. Bottom zoom into the region around the Q-value of <sup>76</sup>Ge at 2040 keV of the GERDA experiment. No peak like structure could be found.

World record low background for 0νββ decay experiments:

~0.5 Counts/(ton keV yr)

No peak observed

→  $T_{1/2}^{0\nu\beta\beta} > 9 \cdot 10^{25} \text{ yr (90\% C.L.)}$ ,  $\langle m_{ee} \rangle < 0.1 - 0.2 \text{ eV}$  (depending on matrix element).

Phase II data taking resumes in December 2019.

→ follow up experiment: LEGEND 200 with ~200kg germanium detectors

Sensitivity estimate: ~10<sup>27</sup> yr

Final goal: LEGEND 1000 with 1ton of enriched germanium detectors

Sensitivity reach: ~10<sup>28</sup> yr → Discovery if IH!

Other  $0\nu\beta\beta$  decay experiments:

**Majorana** ( $^{76}\text{Ge}$ ), use germanium detectors (~40kg) in “conventional” copper cryostat with copper shield, use ultra-pure copper for cryostat and shield. Started measurements end of 2015

Expected sensitivity:  $\sim 10^{26}$  yr until 2019

→ Afterwards merging with GERDA (and others) → LEGEND

**CUORE:**  $^{130}\text{TeO}_2$  crystals in low temperature bolometry. Energy resolution expected  $\sim 6\text{keV}$  at RoI, BI expected: total detector mass: 741kg →  $\sim 200\text{kg}$  of  $^{130}\text{Te}$ .

Started data taking with full mass at 20mK!

Expected sensitivity:  $\sim 10^{26}$  yr until 2020

**EXO, nEXO, NEXT, Kamland ZEN** all use  $^{136}\text{Xe}$  (noble gas)

EXO, EXO200 and nEXO: LAr TPC. Energy resolution  $\sim 2\text{-}3\%$  at RoI

NEXT: pressurized Xe TPC: improved energy resolution

KamlandZEN:  $^{136}\text{Xe}$  doped liquid scintillator

→ large mass, moderate energy resolution

Latest results: limit:  $T_{1/2}^{0\nu\beta\beta} > 1.1 \cdot 10^{26}\text{yr}$  with sensitivity  $3.7 - 4.9 \cdot 10^{25}\text{yr}$

→ background underfluctuation seems to help (?)

Had problems with background from man-made anthropogenic isotope  $^{110\text{m}}\text{Ag}$

Now resolved. Taking data. Expecting results at TAUP conference in September with sensitivity  $\sim 3 \cdot 10^{26}\text{yr}$ !

.....

**Helicity** is defined as the expectation value of spin in the direction of momentum:

$$\hat{h} = \frac{\vec{\sigma} \cdot \vec{p}}{|\vec{p}|}$$

For massive particles this operator is NOT invariant under Lorentz transformation! This means that you can always find a Lorentz boost into an inertial system that is changing the helicity of a particle. This means that helicity is not a good quantum number!

**Chirality** describes the “handedness” of a particle, i.e. the asymmetry of an object (wave function) under reflection.

Every wave function can be separated in left- and right handed components:

$$|\psi\rangle = |\psi_R\rangle + |\psi_L\rangle = \widehat{P}_R |\psi\rangle + \widehat{P}_L |\psi\rangle$$

with the projection operators  $\widehat{P}_L = \frac{1}{2}(1 + \gamma_5)$  and  $\widehat{P}_R = \frac{1}{2}(1 - \gamma_5)$ .

$\widehat{P}_L$  and  $\widehat{P}_R$  have the following properties:

$$\widehat{P}_L \cdot \widehat{P}_R = \widehat{P}_R \cdot \widehat{P}_L = 0, \quad \widehat{P}_L \cdot \widehat{P}_L = \widehat{P}_L, \quad \widehat{P}_R \cdot \widehat{P}_R = \widehat{P}_R$$

Chirality is invariant under Lorentz transformation, i.e. there is no Lorentz boost that can change the chirality of a system.

Charge conjugation transformation does not alter chirality of a particle, i.e.  $\hat{C}|\nu_L\rangle = |\bar{\nu}_L\rangle$ . So far no observation of right-handed neutrinos or left handed anti-neutrinos could be made. This means that the weak interaction couples only to left handed particles and right handed anti-particles  $\rightarrow$  Charge conjugation is maximally broken in weak interaction!