THE QUEST FOR THE IDEAL PHOTODETECTOR FOR THE NEXT GENERATION OF NEUTRINO TELESCOPES

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J.Learned, L.Bezrukov, A.Roberts et al. 70-80s requirements for pmts for deep underwater neutrino experiments. DUMAND, BAIKAL **GRANDE**, NEVOD AMANDA, NESTOR, ANTARES, NEMO, ICECUBE, KM3net

Citius, Altius, fortius

Faster, More Sensitive, Smarter

- High sensitivity to Cherenkov light bialakali photocathode.
- Large sensitive area and 2π acceptance hemispherical photocathode
- High time resolution (as low jitter as possible) hemispherical cathode
- Good SER (as good as possible) to suppress background due to K40.
- Low dark current bialkali photocathode
- Fast response (~10 ns width or less)

History of deep under water neutrino telescopes spans more than 30 years.

So far the Baikal Neutrino Telescope is the only deep under water neutrino telescope in the world.

Water transparency

Light dispersion in deep Baikal water



Cherenkov light spectrun transformation in water

Baikal

Mediterranean



Water parameters play crucial role

Light dispersion in water smears photons arrival times

e.g. 100 m - $\Delta t \sim 5$ ns for Mediterranean PMT's jitter of ~3-4 ns (fwhm) is enough

sensitivity in wider range than conventional bialkali cathode (Ultra/Hyper Multialkali Cathode?)

Disadvantages of classical PMTs

- Poor collection and effective quantum efficiencies
- Poor time resolution???
- prepulses
- late pulses
- afterpulses
- sensitivity to terrestrial magnetic field
- larger PMT size larger dynode system (Dph/Dd1), practically impossible to provide 2π acceptance

Hybrid phototubes with luminescent screen

A.E.Chudakov 1959 - hybrid tube with luminescent screen

- Van Aller et al. 1981 prototypes of «smart tube»
- Van Aller et al. 1981-1986 XP2600
- L.Bezrukov, B.Lubsandorzhiev et al. 1985-1986 Quasar-300 and Quasar-350 tubes
- L.Bezrukov, B.Lubsandorzhiev et al. 1987 Tests of XP2600 and Quasar -300 tubes in Lake Baikal

L.Bezrukov, B.Lubsandorzhiev et al. 1990 - Quasar-370 tube.

Hybrid phototube with luminescent screen

Light amplifier + small conventional type PMT

XP2600

QUASAR-370



Quasar-370 phototube has excellent time and very good single electron resolutions

- no prepulses
- no late pulses in TTS
- low level of afterpulses
- ~100% effective collection effiency
- 1 ns TTS (FWHM)
- very good SER (competitive to HPD)
- immunity to terrestrial magnetic field
 >2π sensitivity
 ageing

Quasar-370LSO with LSO crystal 1996-97, ICRC1997



The PMT used in the Quasar-370LSO had ~17% $\eta(eff)$

$G = Y \times k \times \eta(eff)$

Y - scintillator light yield

k - colection efficiency of photons on small PMT's cathode

$\eta(eff)$ - effective quantum efficiency of small PMT

PMT with higher $\eta(eff)$ will provide even better parameters of Quasar-370!

Studies of Quasar-370 at low thresholds



M1 - single pe peak of small PMTM2 - single pe peak of Quasar-370

naige

Studies of R1463 at a low threshold





- Threshold ~ 0.005 p.e.!
- Green spectra measured with cathode camera switched off, i.e. cathode and 1 dynode are short circuited

Big hemispherical PMTs at low thresholds Hamamatsu R8055 (13'')





SER ~70% (fwhm) 0.005 pe threshold!

Jitter ~ 1.8 ns (fwhm)

Photonis XP1807 (12'')





point like illumination at the pole of the PMT's photocathode

Afterpulses - ~20%

New parameter to evaluate PMT's quality - its ability to work at low threshoulds?

XP2600



C.Wiebusch, RWTH-Aachen 1995





H.Miyamoto, Chiba University

Quasar-370















A.Braem et al. NIMA 570 (2007) 467 Photonis measurements

1987-1990 XP2600 30-50% effect Quasar-370 - 10-20%

XP2600 and Quasar-370 sensitivity vs photons incident angle





OM's response to plane wave vs incident angle of plane wave





Quasar-370 - ~2000 cm**2 R7081 - ~500 cm**2







1995-1997 ICRC1997, COMO2001



- high mechanical precision
- diode protection
- terr. magnetic field infl.

The quest for the ideal scintillator for Quasar-370 like tubes

Requirements:

- Hight light yield
- Fast emission kinetics
- vacuum compatibility
- compatibility with phocathode manufacturing procedure: high temperature, aggressive chemical atmosphere etc.

Scintillators must be:

Inorganic scintillators

Nonhygroscopic?

Time resolution of hybrid phototubes and scintillator parameters

- $$\begin{split} W(t) &\sim \exp(-(G/\tau)t) \\ G &- \text{ the first stage amplification factor} \\ G &= n_{p.e.} / N_{p.e.} \\ G &\sim Y(E_e) \\ Y &- \text{ scintillator light yield} \end{split}$$
- τ scintillator decay time

Figure of merits - F $F_1 = (Y/\tau) \times a$ $F_2 = (Y/\tau) \times a \times b$ Y - light yield, τ - decay time, a - detectibility by small PMT or SiPM b - compatibility with photocathode manufacturing YSO YAP SBO LSO LS Bril350 Bril380 F₁ 1 1.3 1.3 1.8 4* 4.6 6.4 1.3 1.8 4* 1.3 0? 0? \mathbf{F}_2 1 * - using a photodetector with A3B5 photocathode B.K.Lubsandorzhiev, Ringberg 31 Castle 23-28 September 2007



F1 = F2 = 250!

Challenge:

the material should be extremelly pure
problems with monocrystal growth but phosphor will be O'K for luminescent screen

Hamamatsu J9758 phosphor, τ~1ns, Y ~3 Y(YAP)



Hamamatsu news 2006, pp18-19

E.D.Bourret-Courchesne et al. NIMA

Conclusions

What is the ideal photodetector for the next generation neutrino telescopes? Spherical (up to 50 cm dia) with $>2\pi$ angular acceptance High sensitivity in a wider region than conventional bialkali cathode <u>High effective quantum efficiency - good SER</u> Time resolution - ~3-4 ns (fwhm)

Quasar-370 like tubes are very good competitor to HPDs with diodes

Cheap manufacturing cost, higher gain, simpler in operation etc.

good prospects for further developments cheap production?

CLOSE TO THE IDEAL VACUUM PHOTODETECTOR? (Eilat 2006)



Photonis and INR collaboration on smart tube R&D is currently underway



