

Low-Background Experiments

The GERDA Experiment in a nutshell

Felix Fischer

July 26th, 2019

IMPRS Young Scientist Workshop, Ringberg Castle

Max-Planck-Institut
für Physik

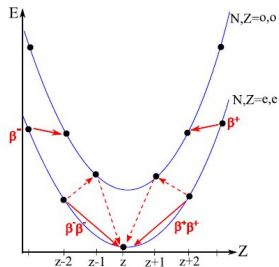


LEGEND

Large Enriched
Germanium Experiment
for Neutrinoless $\beta\beta$ Decay



Neutrinoless double beta decay



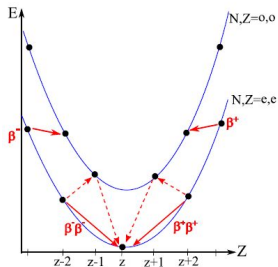
Normal beta decay is strongly suppressed for some isotopes

→ Double beta decay, $2\nu\beta\beta$ -Decay

¹ Fig. 1, http://www.cobra-experiment.org/double_beta_decay/

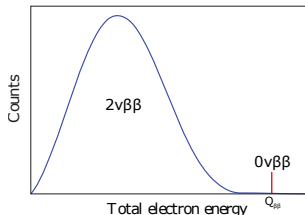
² Fig. 2, arXiv:1601.07512

Neutrinoless double beta decay



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(Particle = Anti-Particle)

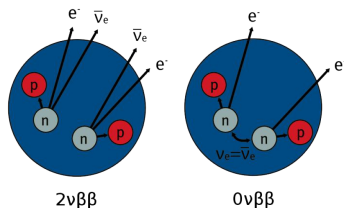
If neutrinos are majorana particles

→ Neutrinoless double beta decay, $0\nu\beta\beta$ -Decay

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Neutrinoless double beta decay



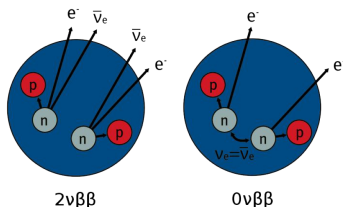
Discovery of $0\nu\beta\beta$ -decay would

- demonstrate lepton number violation
- give information about the nature of neutrinos
→ Majorana or Dirac particle
- give information about absolute neutrino mass

³ Fig. 3, AJ Zsigmond, INPA, Dec 2017

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Neutrinoless double beta decay



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The experimental limit on the $0\nu\beta\beta$ -decay half-life is very long:
> $0.9 \cdot 10^{26}$ years (for ^{76}Ge)⁴

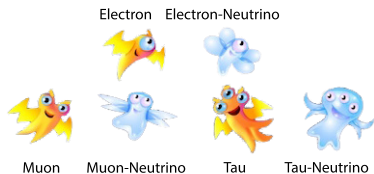
Known double beta decay isotopes:

^{48}Ca , ^{76}Ge , ^{78}Kr , ^{82}Se , ^{86}Kr , ^{96}Zr , ^{100}Mo , ^{136}Xe , ^{130}Te , ...

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But why $0\nu\beta\beta$ -decay?



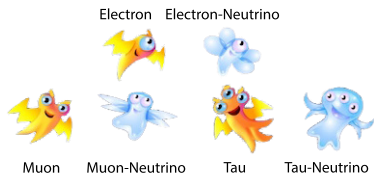
What we **know** about neutrinos:

- 2nd most abundant known particles in the observable universe
- They love to oscillate
- Some neutrinos must have mass

⁵ Fig. 4, <http://faculty.wcas.northwestern.edu/> (11/2018)

⁶ Fig. 5, minutephysics, YouTube (11/2018)

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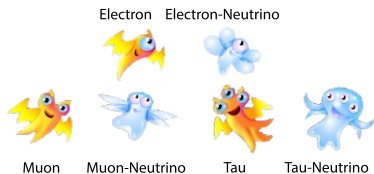
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What we **don't** know about neutrinos:

- Does the neutrino relate to matter-antimatter asymmetry?
- Absolute neutrino mass scale?
- Neutrino mass hierarchy?
- Why is their mass tiny?
- ...

Neutrinos are SUPER weird



But whats the problem with $0\nu\beta\beta$ -decay?



The insane half-life of $> 0.9 \cdot 10^{26}$ years..

⁷ Fig. 6, www.joiyon.co.uk, (11/2018)

Half-Life sensitivity

How can we measure a decay with a half-life much greater than the age of the universe?

Sensitivity on half-life:

$$T_{1/2}^{0\nu} \propto \sqrt{\frac{m \cdot t}{BI \cdot \Delta E}} \text{ with } \begin{cases} m \cdot t & , \text{ exposure} \\ BI & , \text{ background index} \\ \Delta E & , \text{ energy resolution} \end{cases}$$

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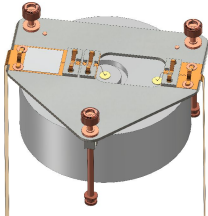
$m \cdot t$ → is limited by funding

BI → can be improved!

Germanium Detector Array (GERDA)

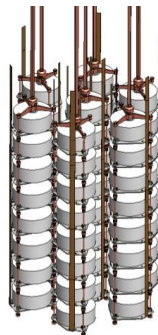
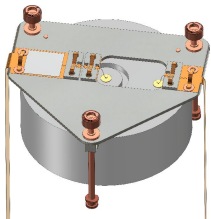


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- Energy resolution $\Delta E_{Q_{\beta\beta}} < 0.1\%$
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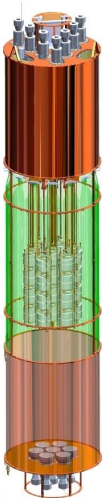


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- 40 detectors (35.6 kg enriched Germanium)

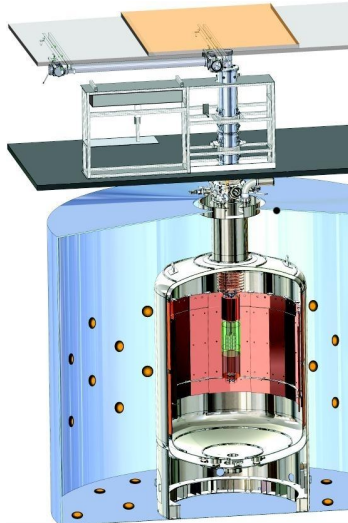
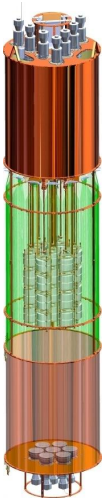
The GERDA-Experiment

Then reduce/reject the background:



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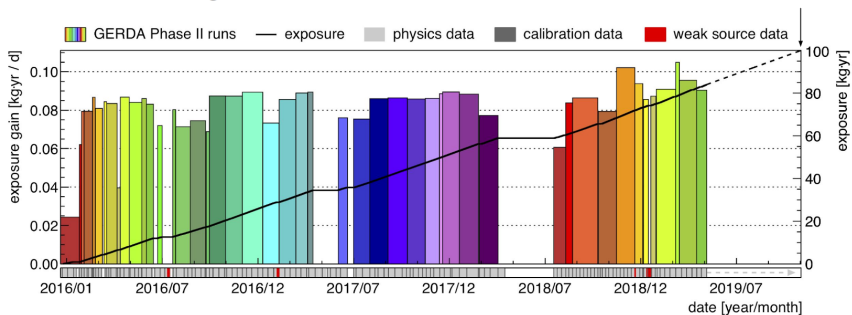


⁹ Fig. 12, Paolo Zavarise, 1st June 2012, VULCANO Workshop

Phase II data taking

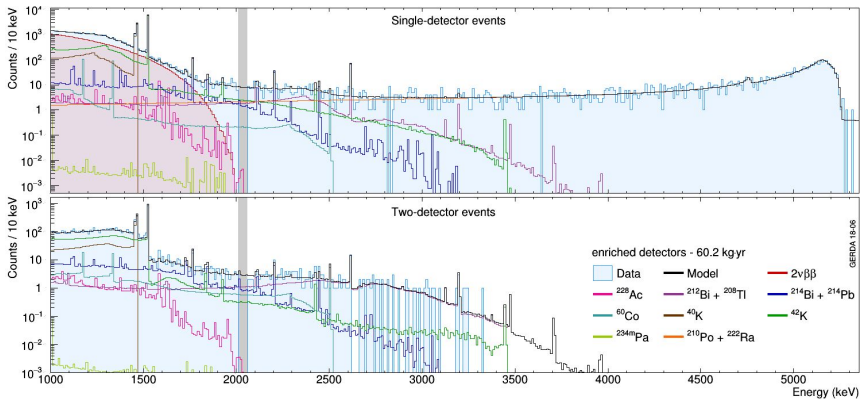
data taking

02-Nov-2019
11:31:32



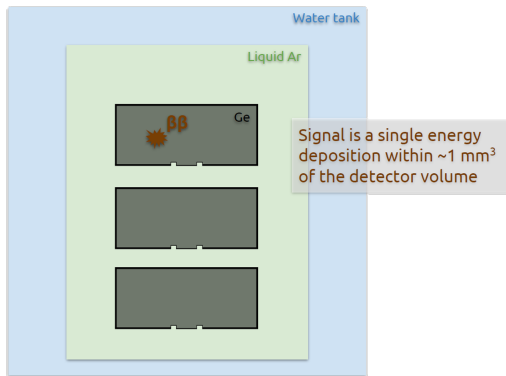
10 Fig. 13, C. Wiesinger – 24th June 2019 - GERDA Meeting

Background before analysis cuts



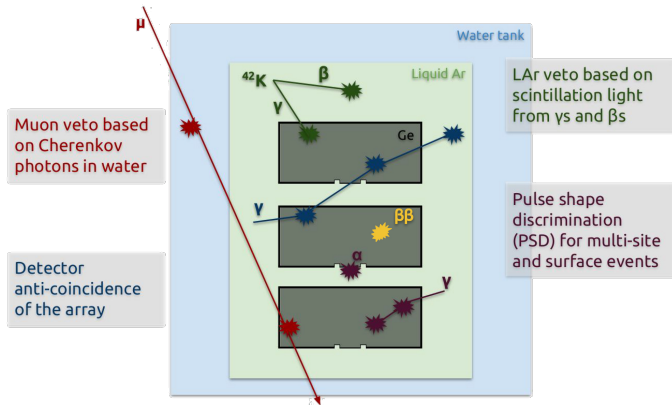
Background at $Q_{\beta\beta}$: α from ^{210}Po , γ from $^{208}\text{Ti}/^{214}\text{Bi}$ and β from ^{42}K

The low background challenge



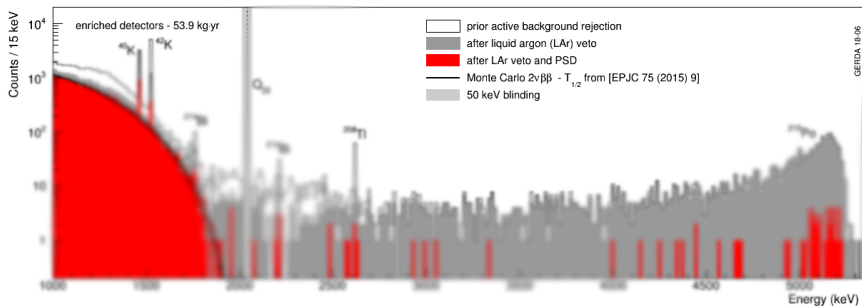
See following talks for details!

The low background challenge



See following talks for details!

Active background suppression - PSD LAr veto



Background in the ROI:

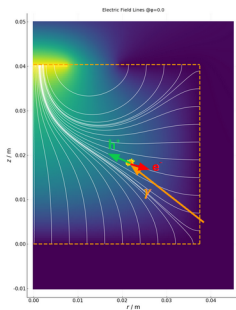
$$BI \approx 5.6 \cdot 10^{-4} \frac{\text{cts}}{\text{keV} \cdot \text{kg} \cdot \text{yr}}$$

Spoiler alarm!



Spoiler alarm!

Electric Potential & Charge Drift



Potential: Solve Gauss' Law

$$\nabla(\epsilon_r(\vec{r})\nabla\varphi(\vec{r})) = \frac{\rho(\vec{r})}{\epsilon_0}, \quad \vec{r} = (r, \phi, z)$$

Electric Field:

$$\vec{\mathcal{E}}(\vec{r}) = -\nabla\varphi(\vec{r})$$

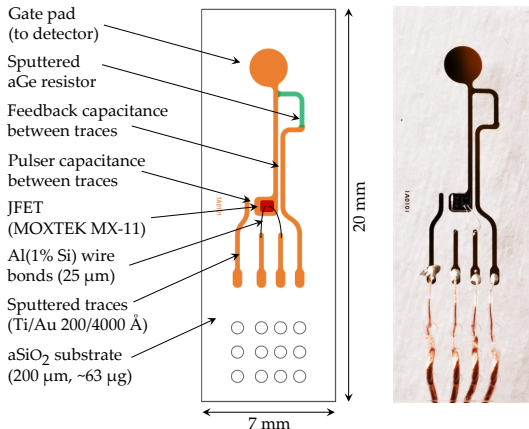
Charge Drift Model:

- Effects of Crystal Axes
- Different Mobilities for e^- and h^+
- ...

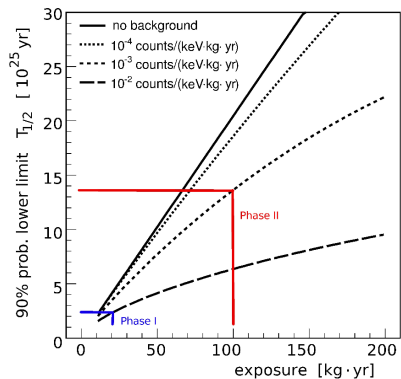
<https://github.com/JuliaHEP/SolidStateDetectors.jl>

15

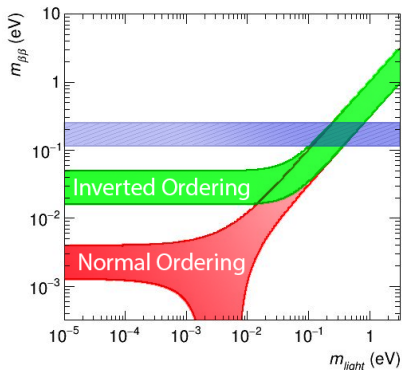
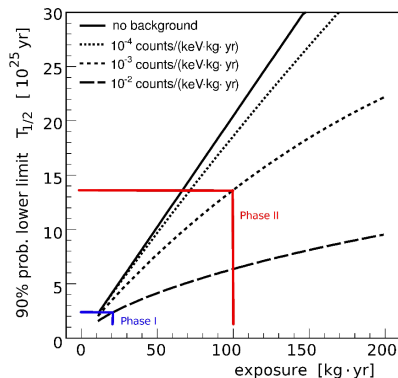
Spoiler alarm!



Implication for neutrino physics



Implication for neutrino physics



- $T_{1/2}^{0\nu} > 0.9 \cdot 10^{26} \text{ yr}$ (90% C.L.)
- $m_{\beta\beta} < (0.11 - 0.25) \text{ eV}$

- Next target: inverted ordering band

¹⁴ Fig. 17, M. Agostini, ICHEP 2018 – July 2018

Outlook: LEGEND

MAJORANA and GERDA merge to form the LEGEND collaboration

- Increase exposure $m \cdot t$ in steps:
 1. LEGEND-200 with 200 kg of enriched germanium
 - Add detectors to existing GERDA environment
 - In preparation to reach $T_{1/2}^{0\nu}$ above 10^{27} yr (starting in 2021)



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 2. LEGEND-1000 with 1 ton of enriched germanium
- Decrease the background index BI^{12}
 - Bigger detectors in order to reduce the number of cables and supports
 - Deep learning for event selection
 - Transparent, scintillating holding structures made of PEN



¹²AIP Conference Proceedings 1894, 020027 (2017); <https://doi.org/10.1063/1.5007652>

Backup

Why Germanium?

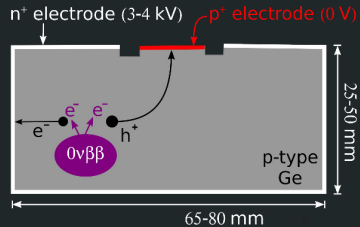
- ✓ Source can be used as a detector
 - High signal detection efficiency
- ✓ Detector material is very pure
 - Very low intrinsic internal background
- ✓ Very good energy resolution
 - Background due to $0\nu\beta\beta$ -decay is negligible
- ✓ Pulse shape discrimination is possible
 - Powerful tool to identify background
- ✓ Considerable experience
 - Industrial production, improvements possible
- X Natural abundance is just 7.8
 - Enrichment necessary
- X Individual detector mass 1 kg
 - Ton scale needs around 1000 detectors

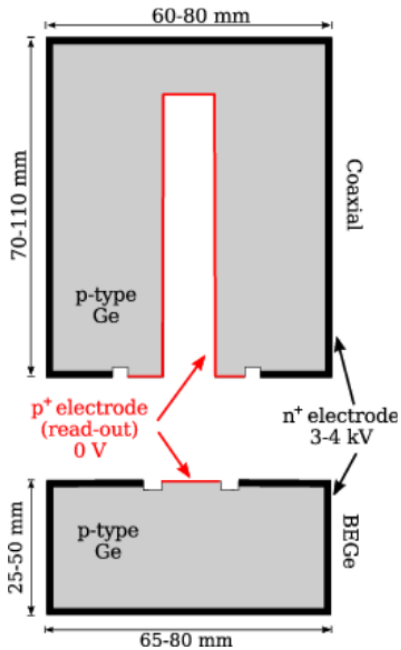
The Ge Detectors

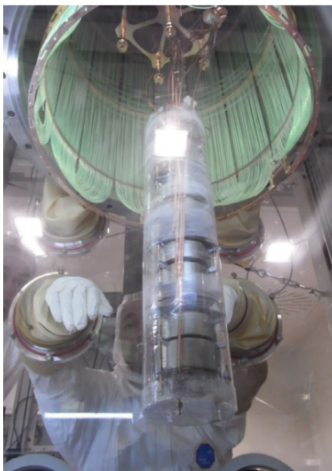
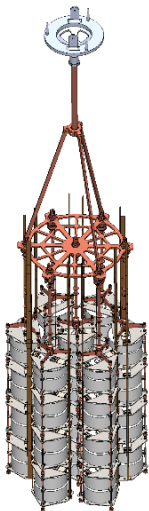
HPGe detector signals:

- signal induced by drift of electron-hole clusters
- time-projection chamber
- identification of events with multiple energy depositions
- identification of events on the surface

Signal/Background Discrimination!

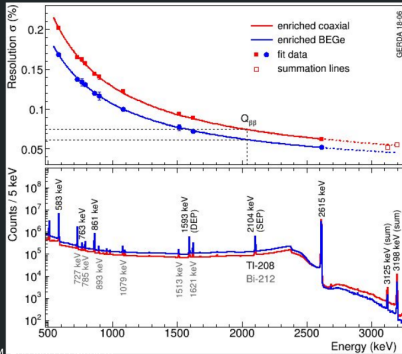




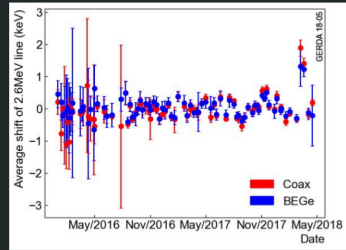


**with TPB COATED
nylon mini-shroud**

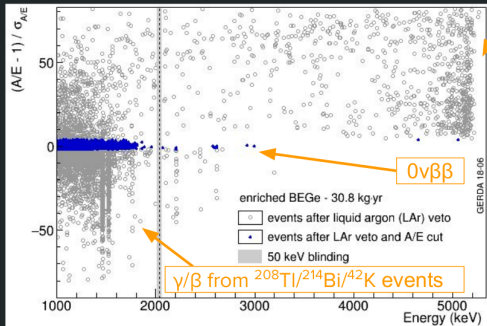
Energy Scale



- Weekly calibration with Th-228 sources
- Fluctuations between calibrations < 1 keV
- Resolution at Qbb better than 0.1% (3-4 keV FWHM)



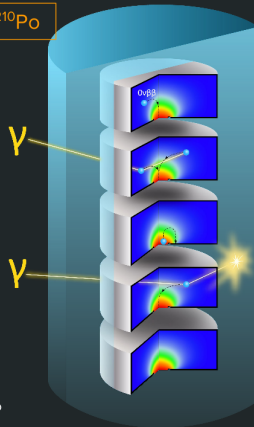
Pulse shape discrimination



α from ^{210}Po

$0\nu\beta\beta$

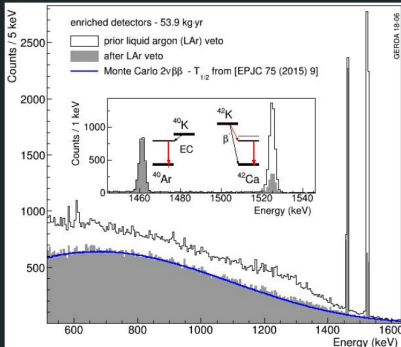
γ/β from $^{208}\text{Tl}/^{214}\text{Bi}/^{42}\text{K}$ events



$0\nu\beta\beta$ efficiency $\rightarrow \epsilon_{\text{BEGe}} = (87.6 \pm 2.5)\%$ $\epsilon_{\text{Coax}} = (71.2 \pm 4.3)\%$

M. Agostini (TU Munich)

LAr scintillation anti-coincidence



M. Agostini (TU Munich)

Dead time
~2%

