## MPI Young Scientists Workshop October 27<sup>th</sup>-29<sup>th</sup>, 2004

# Deformed Spaces and Symmetries

Frank Meyer

Max-Planck-Institut für Physik München Universität München, Fakultät für Physik

Joined work with Marija Dimitrijević and Julius Wess

## Why Noncommutative Geometry?

- Divergences in QFT:
   Discrete space-time may lead to a finite or at least renormalisable theory (natural cutoff)
- Poor understanding of physics at short distances/high energies:
   Rich mathematical structures of deformed symmetries

Rich mathematical structures of deformed symmetries (Hopf algebras) give rise to new features

Localization with extreme precision cause gravitational collapse ⇒ space-time below Planck scale has no operational meaning

 $(\rightarrow K. Fredenhagen, S. Doplicher, J. E. Roberts)$ 

• String Theory:

Open strings in a magnetic background field

→ endpoints of open strings move on noncommutative D-branes

 $(\rightarrow N. Seiberg, E. Witten)$ 

Loop Quantum Gravity:
 Discretization of space-time in spin-foam models

## Noncommutative Spaces

Underlying idea: Noncommutative Coordinates
 At very short distances: coordinates do not commute (Heisenberg 1930)

$$[\hat{x}^i, \hat{x}^j] = C^{ij}(\hat{x}) \neq 0 \tag{1}$$

- This could be valid at very short distances
- At large distance we do not experience Noncommutative coordinates
- $\bullet \implies C^{ij}(\hat{x})$  must be complicated functions that vanishes at large distances

- ullet We do not know such a function  $C^{ij}(\hat{x})$  which leads to a mathematical meaningful commutation relations
- ullet Consider power series expansion of the unknown functions  $C^{ij}(\hat{x})$ :

$$C^{ij}(\hat{x}) = \theta^{ij} + iC^{ij}{}_k \hat{x}^k + R^{ij}{}_{kl} \hat{x}^k \hat{x}^l + \dots$$

- Look at processes that take place at very short distances where the first terms of such an expansion lead to a good approximation for  $C^{ij}(\hat{x})$
- ullet Understand in detail commutation relations which are constant or linear/quadratic in  $\hat{x}^i$

## The Algebra of Functions

■ give up differentiable space-time manifold
 → consider algebra of noncommutative coordinates:

$$\widehat{\mathcal{A}}_{\hat{x}} = \mathbb{C}\langle\langle \hat{x}^1, \dots, \hat{x}^n \rangle\rangle/([\hat{x}^i, \hat{x}^j] - C^{ij}(\hat{x}))$$

- Roughly speaking, these are "functions" that respect the commutation relations (1)
- ullet THE COMMUTATIVE ANALOGUE: Take all formal power series generated by the coordinates  $x^i$ :

$$\mathbb{C}\langle\langle x^1,\ldots,x^n\rangle\rangle$$
.

 In the commutative setting, the coordinates satisfy the commutation relations

$$[x^i, x^j] = 0.$$

Devide by the Ideal generated by the above commutation relations

$$A_x = \mathbb{C}\langle\langle x^1, \dots, x^n \rangle\rangle/([x^i, x^j] - 0) = \mathbb{C}[[x^1, \dots, x^n]].$$

- "Good" examples (Poincare Birkhoff Witt property):
   Freely generated algebras with the following commutation relations
- 1. canonical structure:  $[\hat{x}^i, \hat{x}^j] = i\theta^{ij}$
- 2. Lie algebra structure:  $[\hat{x}^i, \hat{x}^j] = iC^{ij}{}_k\hat{x}^k$  (e.g.  $\kappa$ -Deformed Minkowski Spacetime)
- 3. Quantum Space structure:  $\hat{x}^i\hat{x}^j=q\hat{R}^{ij}{}_{kl}\hat{x}^k\hat{x}^l$

## **Symmetries**

★ over to noncommutative algebras

 $\pmb{x}$   $i,\hat{x}^j]=i\theta^{ij}$  breaks Lorentz symmetry.

 $\Downarrow$ 

X

Commutative
Space
+
Commutative
Symmetry

Deformation

Deformed Space + Deformed Symmetry

**✗** on a Lie-group within the category of *Hopf algebras*.

metries, i.e. we have a deformed symmetry acting on the deformed space

## Canonically Deformed Spaces

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$$[\hat{x}^i, \hat{x}^j] = i\theta^{ij}, \, \theta^{ij} = \text{const}$$

breaks Lorentz-symmetry.

X

deformed space was known.

X

heta-deformed Poincaré bialgebra

$$[\hat{\partial}_{\mu}, \hat{\partial}_{\nu}] = 0, \quad [\hat{\delta}_{\omega}, \hat{\partial}_{\rho}] = \omega_{\rho}^{\mu} \hat{\partial}_{\mu},$$

$$[\hat{\delta}_{\omega}, \hat{\delta}'_{\omega}] = \hat{\delta}_{\omega \times \omega'}, \quad (\omega \times \omega)'_{\mu}{}^{\nu} = -(\omega_{\mu}^{\sigma} \omega'_{\sigma}{}^{\nu} - \omega'_{\mu}{}^{\sigma} \omega_{\sigma}{}^{\nu}),$$

$$\Delta \hat{\partial}_{\mu} = \hat{\partial}_{\mu} \otimes \mathbf{1} + \mathbf{1} \otimes \hat{\partial}_{\mu},$$

$$\Delta \hat{\delta}_{\omega} = \hat{\delta}_{\omega} \otimes \mathbf{1} + \mathbf{1} \otimes \hat{\delta}_{\omega} + \frac{i}{2} (\theta^{\mu\nu} \omega_{\nu}{}^{\rho} - \theta^{\rho\nu} \omega_{\nu}{}^{\mu}) \hat{\partial}_{\rho} \otimes \hat{\partial}_{\mu}.$$

#### **Derivatives**

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$$\hat{\partial}^{\widehat{\mathcal{A}}}: \widehat{\mathcal{A}}_{\hat{x}} \to \widehat{\mathcal{A}}_{\hat{x}}$$
.

Thus, they have to be consistent with commutation relations of the coordinates.

X

$$[\hat{\partial}_{\mu}^{\widehat{\mathcal{A}}}, \hat{x}^{\nu}] = \delta_{\mu}^{\nu} + \sum_{j} A_{\mu}^{\nu\rho_{1}\dots\rho_{j}} \hat{\partial}_{\rho_{1}} \dots \hat{\partial}_{\rho_{j}}.$$

X

duced by requiring that the derivatives should be a module with respect to the deformed background symmetry

X

$$\hat{\partial}_{\mu}(\hat{f}\hat{g}) = \hat{\partial}_{\mu}(\hat{f})\hat{g} + \hat{O}_{\mu}{}^{\nu}(\hat{f})(\hat{\partial}_{\nu}\hat{g}).$$

 $\hat{\partial}_i$  act as in the commutative setting and satisfy the usual Leibniz rule

$$[\hat{\partial}_{\mu}, \hat{x}^{\nu}] = \delta^{\nu}_{\mu}$$

$$\hat{\partial}_{\mu}(\hat{f}\hat{g}) = \hat{\partial}_{\mu}(\hat{f})\hat{g} + \hat{f}\hat{\partial}_{\mu}(\hat{g}).$$

### Towards a Physical Theory

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abstract algebra to complex numbers.

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- 1. Study Representations of the NC algebra (cp. QM)
- 2. Star product approach and Seiberg-Witten map:

Represent the noncommutative algebra on the algebra of commutative functions by a *star product* (next slide)

and express noncommutative fields in terms of commutative ones by Seiberg-Witten map

#### Star Products

Vector space of formal power series in commutative coordinates is isomorphic to the vector space of formal power series in noncommutative coordinates

$$\rho: \mathbb{R}[[x^0, \dots, x^n]] \to \widehat{\mathcal{A}}$$
$$f(x^\mu) \mapsto \widehat{f}(\widehat{x}^\mu)$$

To transmit the noncommutativity to the algebra of commutative functions we define a new product, called star product by pulling back the product of the noncommutative algebra:

$$f(x^{\mu}) \star g(x^{\mu}) := \rho^{-1} (\hat{f}(\hat{x}^{\mu}) \cdot \hat{g}(\hat{x}^{\mu}))$$

- ullet The star product is *not* unique since the isomorphism ho (called *ordering prescription*) is not unique
- Example: Canonical Structure  $[\hat{x}^i, \hat{x}^j] = i\theta^{ij}$ MOYAL-WEYL PRODUCT

$$f \star g = \mu \circ e^{i\theta^{ij}\partial_i \otimes \partial_j} (f \otimes g) = fg + \frac{i}{2} \theta^{ij} (\partial_i f) (\partial_j g) + \dots,$$

where  $\mu(f\otimes g):=fg$  is just the multiplication map.

## Seiberg-Witten Map

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in terms of commutative ones

$$A_i = A_i[a_i], \ \psi = \psi[\phi, a_i], \ \Lambda = \Lambda_{\alpha}[a_i].$$

X

X

explicitly read off in orders of h where in the limit  $h \to 0$  the commutative theory is obtained

1. Requirement: commutative gauge transformations induce noncommutative ones, i.e.

$$A_i[a_i] + \delta A_i[a_i] = A_i[a_i + \delta a_i]$$
  
$$\psi[\phi, a_i] + \delta \psi[\phi, a_i] = \psi[\phi + \delta \phi, a_i + \delta a_i]$$

2. Requirement: consistency condition

$$(\delta_{\alpha}\delta_{\beta} - \delta_{\beta}\delta_{\alpha})\psi = \delta_{-i[\alpha,\beta]}\psi$$
  

$$\Leftrightarrow i\delta_{\alpha}\Lambda_{\beta} - i\delta_{\beta}\Lambda_{\alpha} + [\Lambda_{\alpha} * \Lambda_{\beta}] = i\Lambda_{-i[\alpha,\beta]}$$

## Summary

- Concept of deformed spaces: spacetime is discrete at very short distances resp. at high energies
- ◆ differentiable manifold ⇒ noncommutative algebra commutative product ⇒ star product Leibniz rule ⇒ deformed Leibniz rule (important for gauge theory via Seiberg-Witten map)
- We found a  $\theta-$ Deformed Poincaré Algebra as symmtery for the canonical case where  $[\hat{x}^i,\hat{x}^j]=i\theta^{ij}$

#### Outlook

- Study implications of the new symmetry for canonically deformed spaces for Noncommutative Gauge Field Theories
- Formulate a theory of Gravity on Noncommutative Spaces
- Study phenomenological consequences of noncommutativity

Noncommutative Spaces ⇒ Interesting concepts for a better understanding of physics at short distances