Particle Detectors I

Physics Cases and Detector Challenges

- Particle Interactions
- Tracking Detectors
- Calorimeters
- Reconstruction and Detector Concepts

can only give an incomplete overview of existing detectors and detector technology with slight focus on LHC detectors

Day 1

Day 2

Physics at Hadron Colliders – Particle Detectors

How (different) Physicists See the World



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Physics Cases and Detector Challenges

The High Energy Frontier
 LHC and ILC Physics Cases
 Detector Challenges

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The High Energy Frontier

Accelerator projects on the market (in very different states of development)

-> LHC

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- nominal LHC with 2 x 7 TeV and L = 10^{34} cm⁻²s⁻¹
- Super-LHC (SLHC) with 2 x 7 TeV and L = 10^{35} cm⁻²s⁻¹
 - needs major detector upgrades and R&D on radiation hardness
- **Double-LHC (DLHC)** with 2 x 14 TeV and $L = 10^{34} \text{ cm}^{-2}\text{s}^{-1}(?)$
 - needs new superconducting magnets with B \sim 15 T, no R&D yet
- → e+e- Linear Collider
 - **ILC** with $\sqrt{s} = 500$ GeV and L = 2 5 x 10³⁴ cm⁻²s⁻¹, upgradable to 1 TeV
 - Global Design Effort (GDE) under way to prepare Technical Design Report until end of 2008
 - CLIC with $\sqrt{s} = 3 5$ TeV and L = 10³⁵ cm⁻²s⁻¹
 - feasibility study under way, hope for positive results until 2010
- ⊸ µ+µ⁻ Collider
 - multi-TeV with L = 10^{31} cm⁻²s⁻¹(?)
- LHeC (recent new proposal)
 - LHC + 70 GeV e⁻ ring, $\sqrt{s} = 1.4$ TeV (4.5x HERA), L = 10^{33} cm⁻²s⁻¹ (20x HERA)

the only approved big project in high energy physics at present, under construction

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The Physics Case (LHC)

Possible discoveries with O(30) fb⁻¹ (~2009/10)



more easy...

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...more challenging

The Physics Case (LHC)

"Later" discoveries with O(100) fb⁻¹

in general more difficult: non-resonant hadronic or very rare leptonic final states will need more luminosity and better detector understanding



Strong EW Symmetry Breaking

deviations from SM due to new interaction in W_LW_L in absence of Higgs:

very challenging sign. 2 lept + forw. jets + Emiss possibly non-resonant

possibly only ${\sim}3\sigma$ with 100 fb^{-1}



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The Physics Case (ILC)

Guaranteed and needed: Top mass

- top-quark could play a key role in the understanding of flavour physics
- m_{top} fundamental parameter
- Δm_{top} will limit many predictions



requires precise determination of its properties

Energy scan of top-quark threshold:

 $\Delta M_{top} \approx 100 \text{ MeV}$ (dominated by theory)

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The Physics Case (ILC)

Precision Higgs physics





- decay-mode-independent observation
- mass (50 MeV)
- absolute couplings (Z,W,t,b,c,τ) (1-5%)
- total width (model-independent)
- spin, CP
- top Yukawa coupling (~5%)
- self coupling (~20%, 120-140 GeV)
- $\Gamma_{\gamma\gamma}$ at photon collider (2%)

fully establish Higgs mechanism!

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pp Cross Section

• Total pp cross section at $\sqrt{s} = 14$ TeV is about ~100 mb



physics interest ~10...100 pb (and some < 1 pb)

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pp Physics Cross Sections

Many orders of magnitude between signal and background



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LHC Parameters

Energy = 7 x Tevatron, luminosity = ~ 50-100 x Tevatron

	LHC (2007)	Tevatron (1987)	SppS (1981)
max. Energy (TeV)	(7)		0.450
circumference (km)	26.7	6.3	6.9
luminosity (10 ³⁰ cm ⁻² s ⁻²)	10000	210	6
time between collisions (µs)	0.0025	0.396	3.8
crossing angle (µrad)	300	0	0
p/bunch (10 ¹⁰)	11	27/7.5	15/8
number of bunches	2808	36	6
beam size (µm)	16	34/29	36/27
filling time (min)	7.5	30	0.5
acceleration (s)	1200	86	10

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Detector Challenges at LHC

Physics case (the driving force)

- defines what type of accelerator/facility is needed
 - e.g. hadron/lepton collider, energy, luminosity
- Physics + accelerator parameters
 - define what type of detector is needed
- Detector requirements
 - High energy collisions
 - sufficiently high momentum resolution up to TeV scale
 - High luminosity
 - high rate capabilities and fast detectors because of high interaction rate
 - Large particle density
 - high granularity, sufficiently small detection units to resolve particles
 - At hadron colliders

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- radiation-hard detectors and electronics
 - radiation mainly due to many protons/neutrons emerging from the interactions not due to LHC machine backgrounds

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LHC Detectors



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What we have to expect

One bunch crossing every 25 ns with ~20 interactions

- 1000 tracks per bunch crossing = 4 * 10¹⁰ tracks per second...
- ... and very often you're interested in a few tracks only...



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The Perfect Detector...

Inshould reconstruct any interaction of any type with 100% efficiency and unlimited resolution (get "4-momenta" of basic physics interaction)

- → efficiency:
 - not all particles are detected, some leave the detector without any trace (neutrinos), some escape through not sensitive detector areas (holes, cracks for e.g. water cooling and gas pipes, cables, electronics, mechanics)



Magnet Concepts at LHC experiments



- large homogenous field inside coil
- needs iron return yoke (magnetic shortcut)
- limited size (cost)
- coil thickness (radiation lengths)

CMS, ALICE, LEP detectors



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(air-core) toroid



- + can cover large volume
- + air core, no iron, less material- needs extra small solenoid for general tracking
- non-uniform field
- complex structure



ATLAS and CMS Coils





CMS Solenoïde

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CMS solenoid (5 segments)

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Particle Interactions

- Photon Interactions
- Charged Particle Interactions
 - Scattering, Ionization
 - Photon Emission: Čerenkov + Transition Radiation

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Particle Interactions

 Particles cannot be seen/detected directly, we only can observe the result of their interactions with the detector (material)

- Interactions are mainly electromagnetic
 - exceptions: strong interactions in hadronic showers (hadron calorimeters) weak interactions at neutrino detection (not discussed here)



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Particle Interactions – Photons

Photo effect

- used at various photo detectors to create electrons on photo cathodes in vacuum and gas or at semi conductors (surface)
 - Photo Multiplier Tubes (PMT)
 - photo diodes

Compton scattering (e⁻ γ scattering)

- not used for particle detection

but was/is used for polarization measurement of beams at e^+e^- machines and could be used to create high energy photons in a $\gamma\gamma$ - collider

 $\gamma + e^- \rightarrow e^+e^- + e^- + \gamma$

• Pair production ($\gamma \rightarrow e^+e^-$)

 initiates electromagnetic shower in calorimeters, unwanted in tracking detectors





 γ + atom \rightarrow atom⁺ + e⁻



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Photon Interactions - Overview



- Photo effect dominating at low γ energies (< some 100 keV)
- Compton scattering regime ~some 100 keV up to ~10 MeV
 - exact energy domain depends on Z
 - low Z: wide energy range of Compton scat.
 - large Z: small energy range of Compton scat.
- Pair production dominating at high energies (> ~10 MeV)
- $\sigma_{\text{p.e.}} = \text{Atomic photoelectric effect (electron ejection, photon absorption)}$ $\sigma_{\text{Rayleigh}} = \text{Rayleigh (coherent) scattering-atom neither ionized nor excited}$ $\sigma_{\text{Compton}} = \text{Incoherent scattering (Compton scattering off an electron)}$ $\kappa_{\text{nuc}} = \text{Pair production, nuclear field}$
 - $\kappa_e =$ Pair production, electron field
 - $\sigma_{g.d.r.}$ = Photonuclear interactions, most notably the Giant Dipole Resonance [4]. In these interactions, the target nucleus is broken up.

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Radiation Length

- Main energy loss of high energy electrons/photons in matter

 Bremsstrahlung (e[±]) and pair production (γ)
- Can characterize any material by its radiation length X₀
 - 2 definitions (for electrons and for photons)
 - X_0 = length after an electron looses all but 1/e of its energy by Bremsstrahlung
 - $X_0 = 7/9$ of mean free path length for pair prodution by the photon

Very convienient quantity

- Rather than using thickness, density, material type etc. detector
 - often expressed as % of X₀
- tracking detectors should have X_0 as low as possible (<< 1 X_0)
 - ATLAS and CMS trackers: 30% 130% X₀
 - not really "transparent", high probability to initiate electromagnetic showers in tracker far before electrons/photons reach calorimeters
 - "pre-shower" detectors in front of calorimter should detect and correct measured ECAL energy for such early showers
- --- calorimeters should have X_0 as high as possible (typically 20...30 X_0)

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Electromagnetic Cascades

Starting from the first electron/photon an electromagnetic shower (cascade) develops in thick materials

- shower maximum (peak of energy deposition) slightly energy dependent



Nuclear Interaction Length

- Similar as radiation length but for hadrons
- Strong interaction of hadron with nucleus
 - Development of hadronic cascade (shower)

 $\pi, \mathbf{p}, \mathbf{K}$

multiplicity $\propto \ln(E)$

Hadronic showers have two main components

- hadronic
 - charged hadrons, breaking up of nuclei (binding energy) nuclear fragments, neutrons
- electromagnetic
 - decay of neutral pions: $\pi^0 \rightarrow 2 \gamma$ (100% branching ratio)

"invisible" energy = large energy fluctuations

C. Grupen

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Radiation and Nucl. Interaction Lengths

Typical radiation length

- gases, e.g. Argon ~100 m
- light materials, e.g. Aluminum, Silicon ~10 cm
- heavier metals, e.g. Iron, Copper, Lead ~ 0.5 1.5 cm

For Z > 6: $\lambda_{I} > X_{0}$



Charged Particle Interactions

- (Multiple) elastic scattering with atoms of detector material
 - mostly unwanted, changes initial direction, affects momentum resolution
- Ionization
 - the basic mechanism in tracking detectors
- Photon radiation
 - -> Bremsstrahlung
 - initiates electromagnetic shower in calorimeters, unwanted in tracking detectors
 - Čerenkov radiation
 - hadronic particle identification (ALICE), also in some homogeneous electromagnetic calorimeters (lead glass)
 - Transition radiation
 - electron identification in combination with tracking detector

Excitation

- creation of scintillation light in calorimeters (plastic scintillators, fibers)

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Elastic Scattering

Most basic interaction of a charged particle in matter

elastic scattering with a nucleus
 = Rutherford (Coulomb) scattering

Ζ

 $\frac{d\sigma}{d\Omega} = 4zZr_e^2 \left(\frac{m_e c}{\beta p}\right)^2 \frac{1}{\sin^4 \Theta/2}$



Hans Geiger



- non-relativistic
- no spins

Scattering angle and energy transfer to nucleus usually small

- No (significant) energy loss of the incoming particle
- Just change of particle direction

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Ernest Rutherford

Science Museu

Multiple Scattering

If thick material layer: Multiple scattering

- after passing layer of thickness L particle leaves with some displacement r_{plane} and some deflection angle Θ_{plane}



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lonization

Primary number of ionizations per unit length is Poissondistributed

- typically ~30 primary interactions (ionization clusters) / cm in gas at 1 bar

However, primary electrons sometimes get large energies

- can make ionizations as well (secondary ionization)
- can even create visible secondary track ("delta-electron")
- large fluctuations of energy loss by ionization



- Typically: total ionization = 3 x primary ionization
 - on average ~ 90 electrons/cm

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Ionization Cluster Size



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Energy Loss Distribution

1 cm sampling length

- Cluster size fluctuations cause large variations of energy loss from track to track
 - Landau distribution
 - large broad peak (single or few el. clusters)
 - soft collisions, interaction with whole gas molecule
 - small energy transfer
 - **looong tail** (multi el. clusters, δ -electrons)
 - hard collisions, semi-free shell electrons
 - large energy transfer



tracks in CERN 2m bubble chamber



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Energy Loss by Photon Emission

Ionization is one way of energy loss

- emission of photons is another...



Čerenkov Radiation

• Čerenkov radiation is emitted when a charged particle passes though a dielectric medium with velocity $\beta \ge \beta_{thres} = \frac{1}{n}$ speed of light - classical picture: wave front or cone under Čerenkov angle



- number of emitted photons per unit length and unit wave length interval



Transition Radiation

Predicted by Ginzburg and Franck in 1946

- emission of photons when a charged particle traverses through the boundary of two media with different refractive index
- (very) simple picture
 - charged particle is polarizing medium
 - polarized medium is left behind when particle leaves media and enters unpolarized vacuum
 - formation of an electrical dipol with (transition) radiation

Radiated energy per boundary $W \propto \gamma$

- only very high energetic particles can radiate significant energy

need about $\gamma > 1000$

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- in our present energy range reachable with accelerators only electrons can radiate
 - but probability to emit photons still small

$$N_{photons} \propto \alpha_{EM} \approx \frac{1}{137}$$

need many boundaries (foils, foam) to get a few photons

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