

1. What is the "distance ladder"?

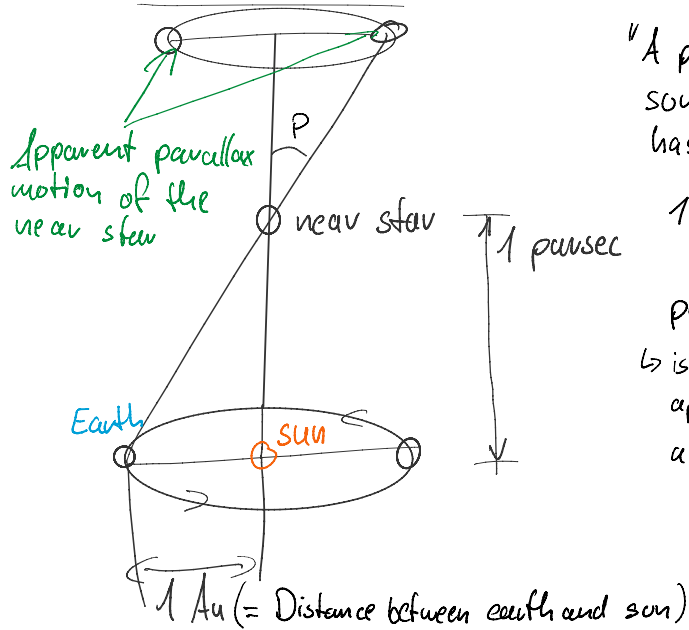
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- Also called "extragalactic distance scale"

↳ It's the succession of methods to determine distances in space
e.g. from here to a celestial object or between objects

- A direct measurement is only possible for close objects: $d < 1000 \text{ pc}$

Extra: The unit parsec: $1 \text{ pc} = 1 \text{ "parsec"} \sim 31 \text{ peta meter} \sim 3,26156 \text{ ly}$



"A parsec is the distance between the sun and an astronomical object that has a parallax angle of one arcsecond"

$$1 \text{ arcsecond} = \frac{1}{3600} \text{ degree}$$

parallax \approx alternation

↳ is a displacement or difference in the apparent position of an object viewed along two different lines of sight.

(src: wikipedia)

- Several of these methods rely on "Standard Candles"

↳ known properties like luminosity

↳ calculate the distance by the absolute magnitude in brightness

$$D = 10^{(m - M - 10)/5}$$

m : apparent magnitude
 M : absolute magnitude
 D : distance in [kpc]

↳ Measured distance is called **luminosity distance**

- Another class is the "Standard Ruler"

↳ astrophysical object where the actual size is known (e.g. by trigonometry)

$$\vartheta \approx \frac{r}{D}$$

ϑ : angular diameter [rad]
 r : actual diameter
 D : distance

↳ The distance measured by the standard ruler is called **angular diameter distance**

- Next: Standard Siren

- ↳ Gravitational waves originating from compact binary systems which are in their inspiral phase
- ↳ A shrinking orbit is directly observable as an increase in frequency
- ↳ Just as with the standard candles, given the emitted and received amplitudes, the inverse-square law determines the distance

So... in general:

- close enough: direct measurement (< 1000 pc)
- not close enough: search for something close to it which is recognizable and homogeneous enough
(depends upon physical assumptions)
- ↳ There are many ways, we just named the standard ones!