

# Nested Sampling

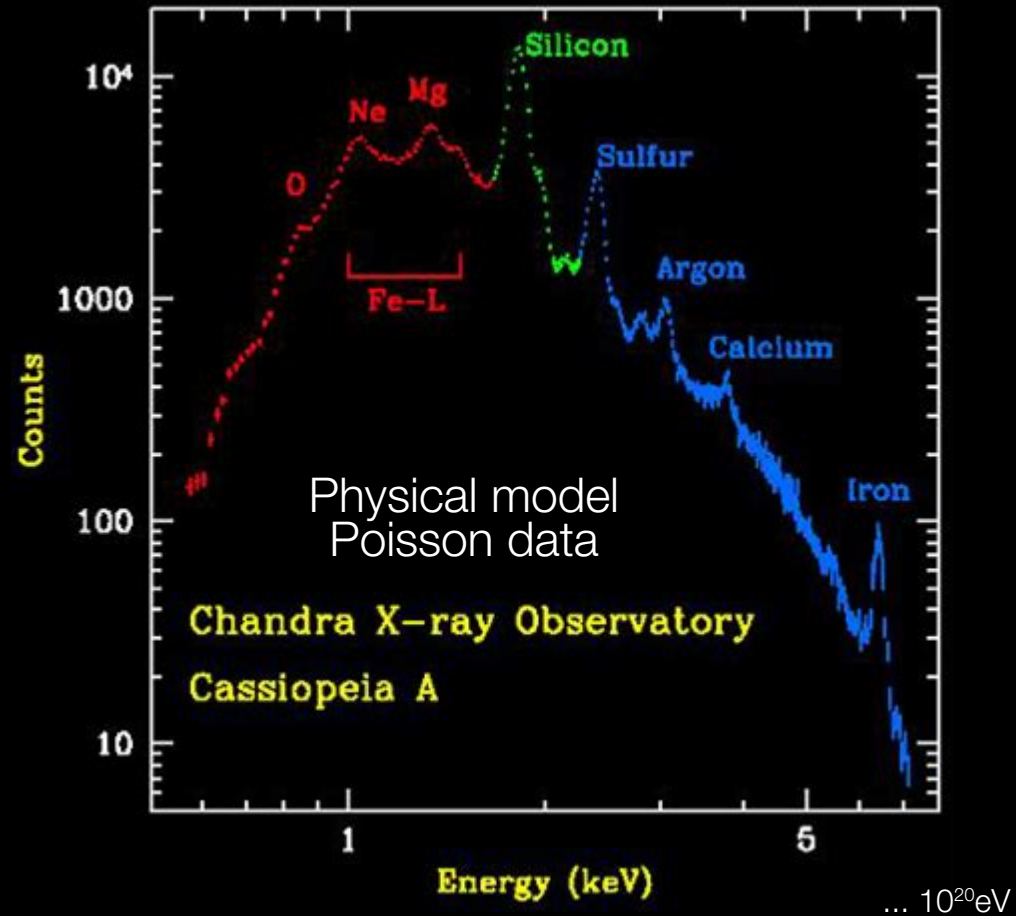


**Johannes Buchner**

Max Planck Institute for extraterrestrial Physics  
<http://astrost.at/istics/>

# Astronomy & Astrophysics

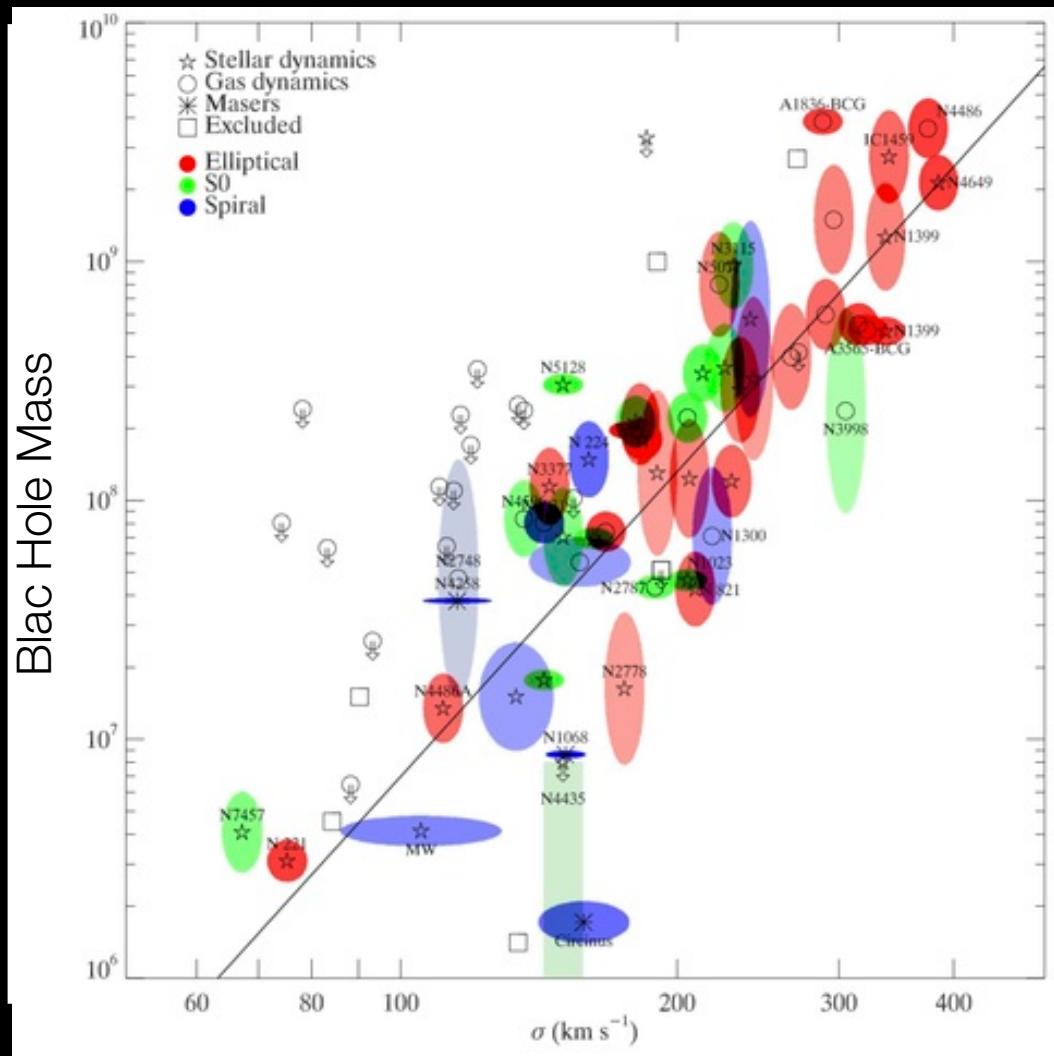
Understanding individual objects



In wavelength or energy →  
In time →  
In space →  
In multiple →  
Light & beyond

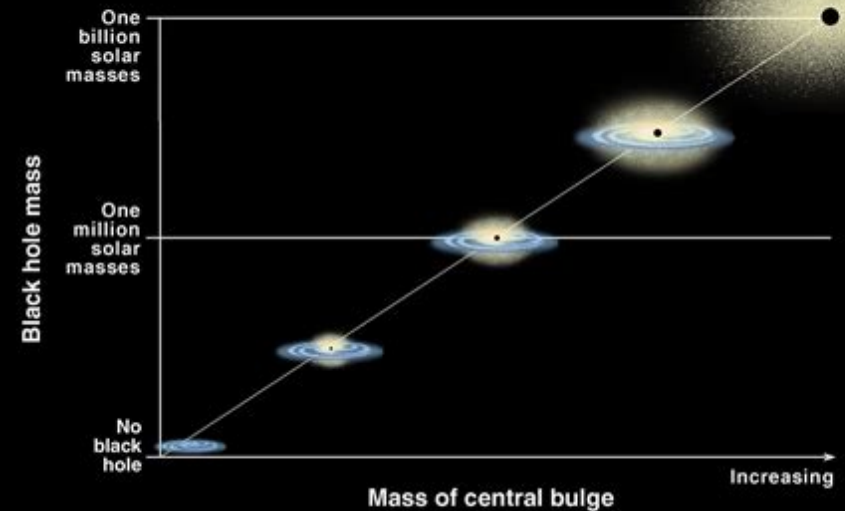
# Astronomy & Astrophysics

## Understanding samples



Velocity dispersion of stars

## Correlation Between Black Hole Mass and Bulge Mass



Empirical models, correlations

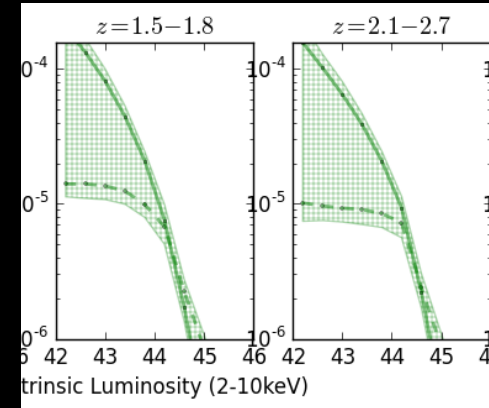
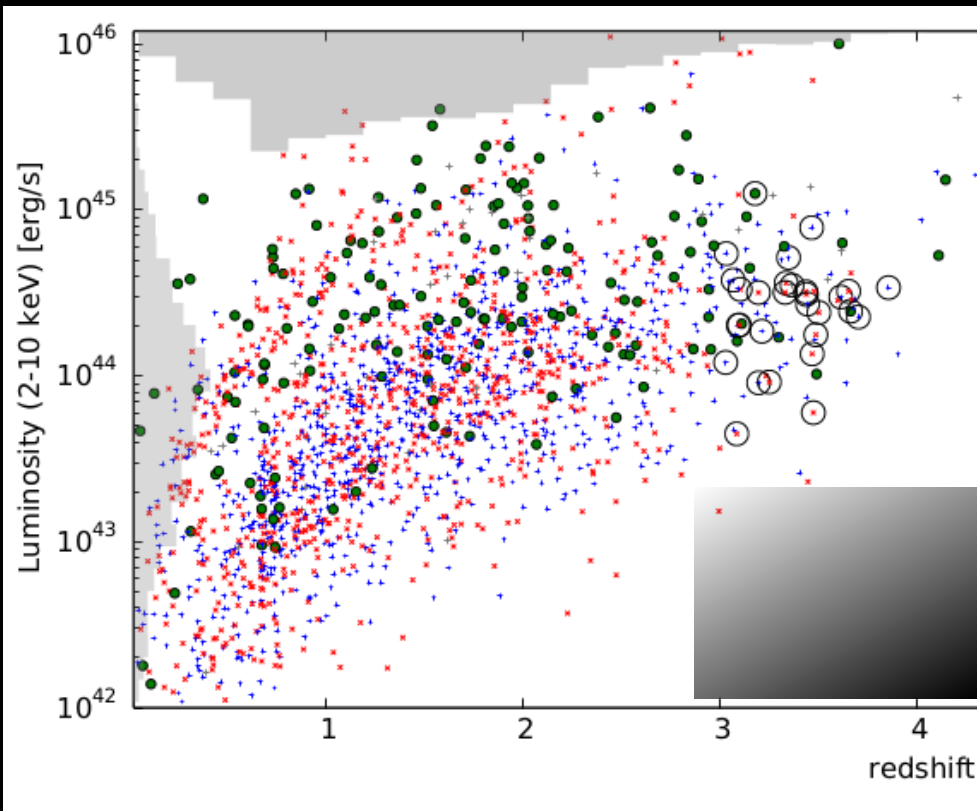
Physical interpretation

reproduction with galaxy  
simulations



# Astronomy & Astrophysics

Understanding underlying populations

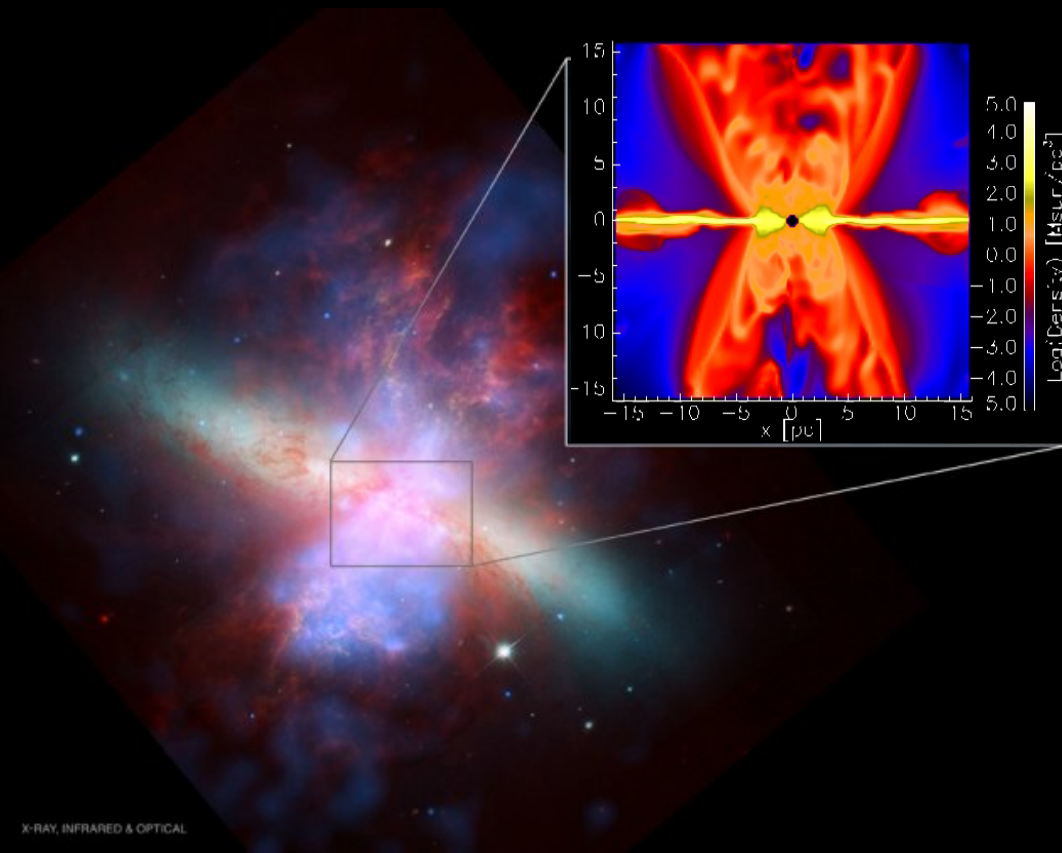


Censorship (but understood)

heavily heterogeneous  
uncertainties (but understood)

Hierarchical Bayesian Model  
with censorship (1983)





Galaxy clusters for Dark Energy  
3 million Active Galactic Nuclei



@eROSITA\_SRG

Johannes Buchner

<http://astrost.at/istics/>  
Max Planck Institute for Extraterrestrial Physics  
Garching, Munich



# Inference requirements

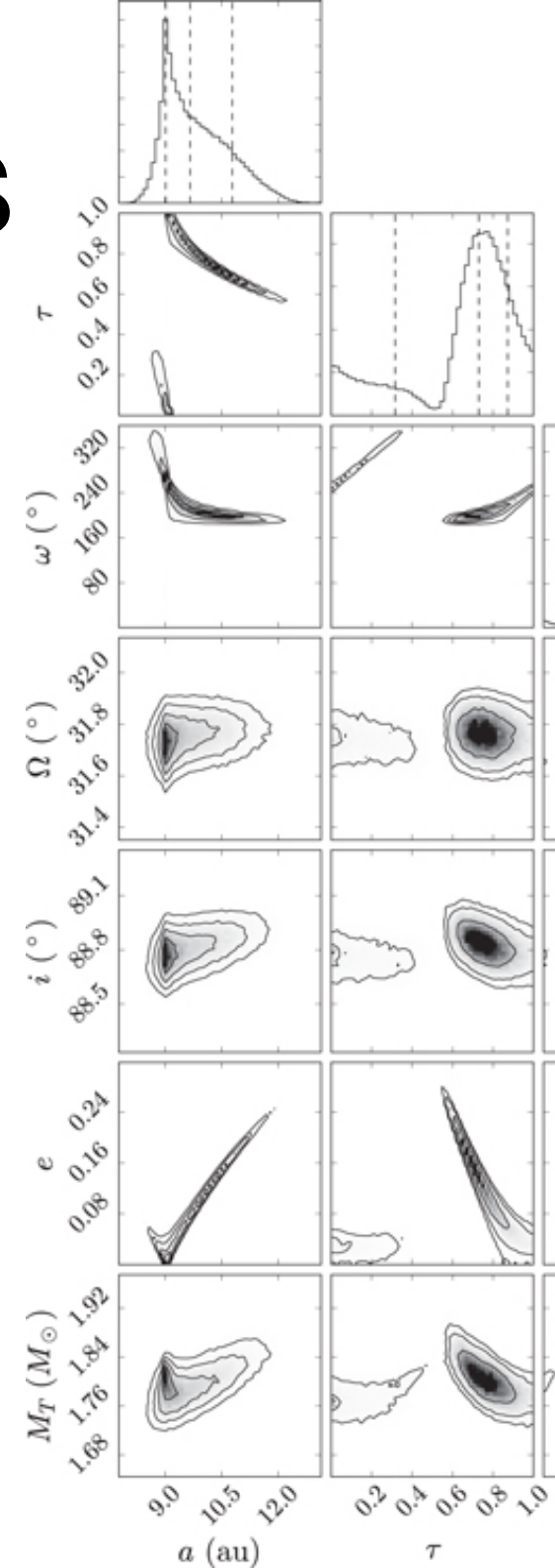
- Forward modeling with Bayesian inference is popular
- Physical models: can be complex & expensive Monte Carlo simulations
- Competing physical effects
- “degenerate” parameter constraints
- multiple solutions
- Sometimes little data, sometimes lots
- Often 2-20d (or 1e6d)

P

M

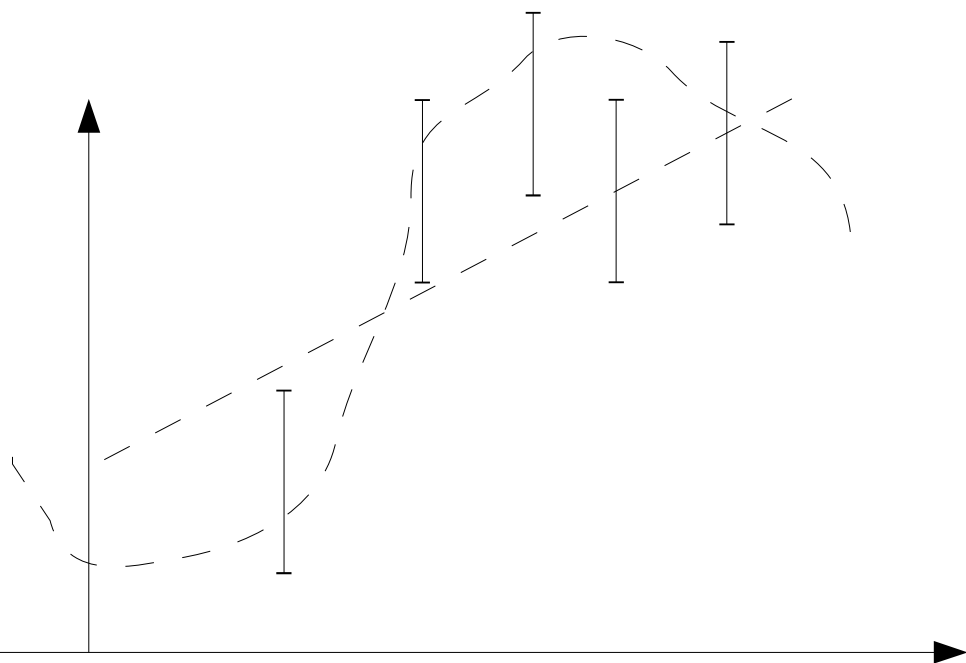
I

D



# Bayesian inference is integral

- What parameters are probable under a given model?
- What models are probable?



$$M_1 : y_{\text{pred}} = \theta_1 + x \cdot \theta_2$$

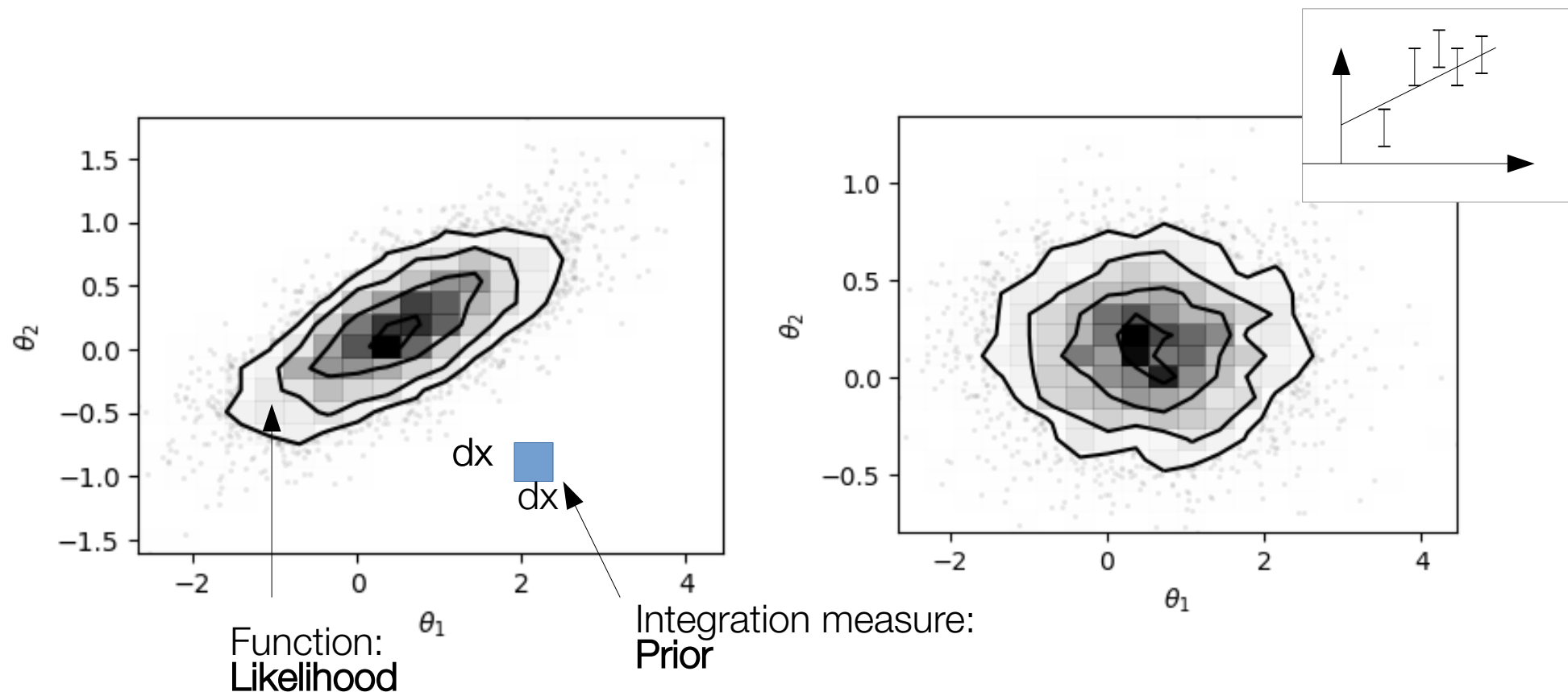
“physical model prediction”  
non-linear, complex, slow

$$L = \prod_i \frac{(y_i - y_{\text{pred}}(x_i))^2}{2 \cdot \sigma_i^2}$$

statistical measurement process  
data vs. model



# Bayesian inference is integral



- Parameter estimation: identify parameter space subset
  - Model comparison: identify model subset
- containing 99% of probability mass

# Bayesian tools

To handle multiple maxima, low state of information,  
peculiar posterior shapes, numerical likelihoods

(& have a life beyond convergence criteria)

## Parameter estimation

## Model comparison

low-d:

Nested sampling

Nested sampling

(~20d)

high-d:

HMCMC with multiple  
chains

open research problem

→ posterior samples

→ Bayes factors

# Nested sampling theory

- Idea Skilling 2004, 2006
- Convergence Evans 2007  
Chopin & Robert 2010  
Skilling 2009
- Point Process-based Monte Carlo Walter (2014)
- Sequential Monte Carlo Salomone (2018)  
Superset of SMC & NS algorithms  
with strong theoretical foundations  
  
Birge (2012)  
Polson & Scott (2014)

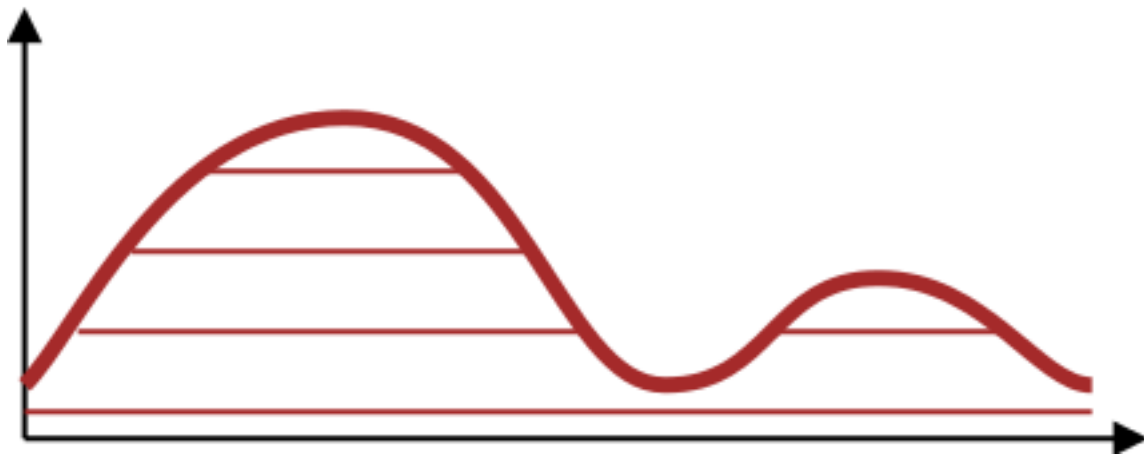


# Nested sampling in practice

- Inputs
  - Dimensionality of the problem
  - Prior density function
  - Likelihood function
- Outputs
  - Posterior samples (like in MCMC)
  - $\ln(Z)$  with uncertainties

MultiNest, PyMultiNest, nestle, dynesty, polychord, UltraNest

# Nested sampling idea (1)

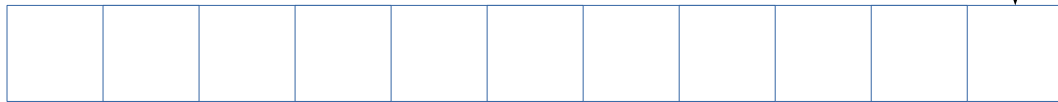


$$Z \approx \sum_i \Delta V_i \cdot L_i$$

Lebegue integral:  
Height \* dV

# Nested sampling idea (2)

Sorted live points

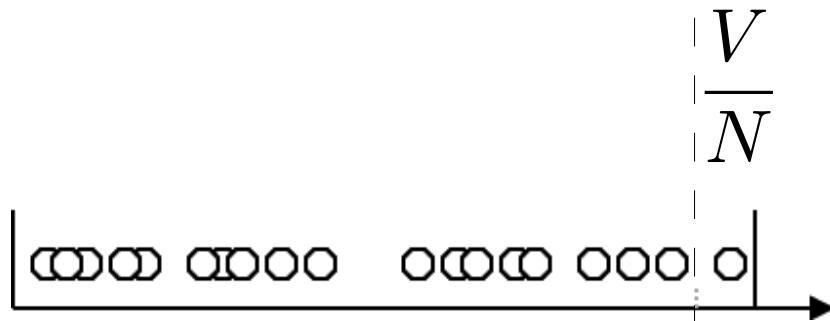


To be  
replaced



Keeping track of  
volume

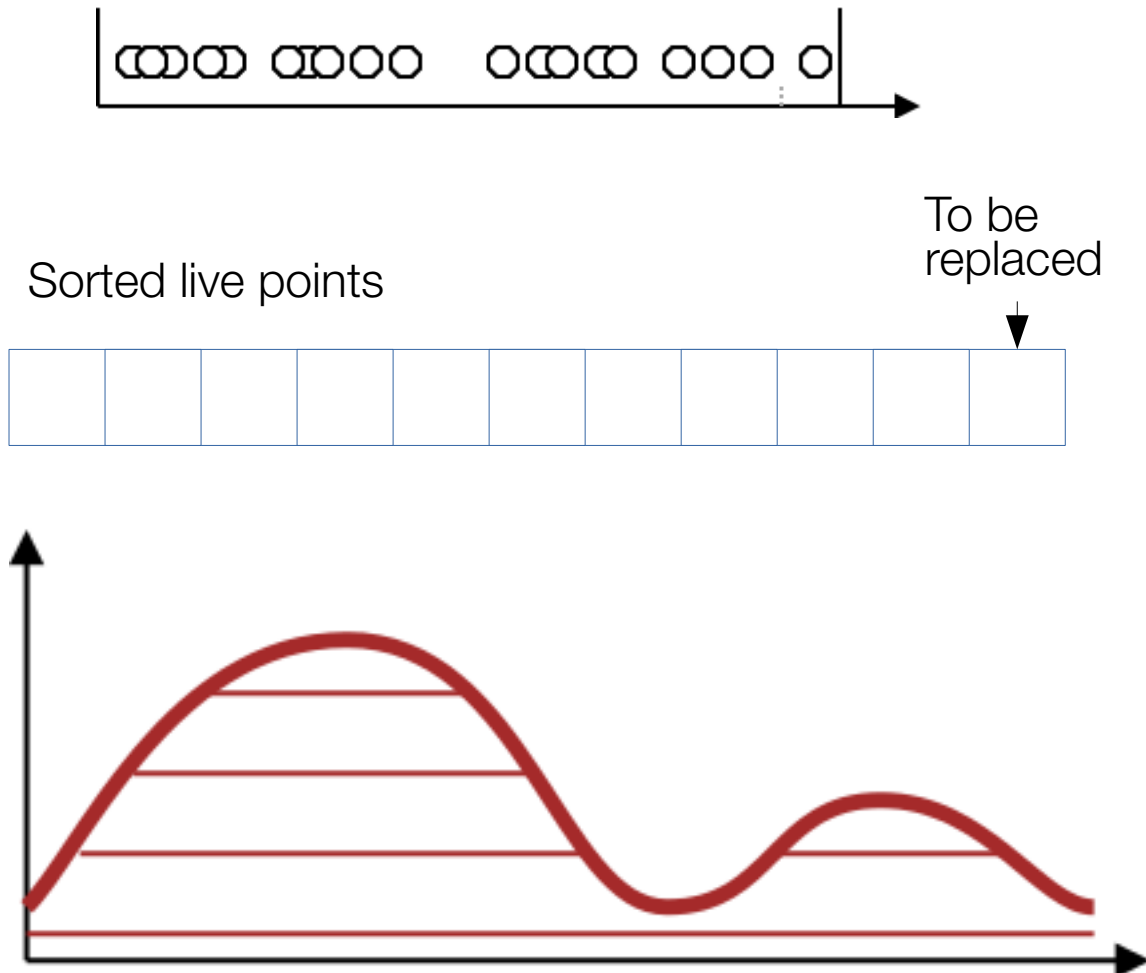
- Exponential shrinkage  
( $\sim 1/N$ )



$$V' = V \times \left(1 - \frac{1}{N}\right)$$



# Nested sampling idea (3)



- Keeping track of volume
  - Exponential shrinkage ( $\sim 1/N$ )
- Keeping track of height

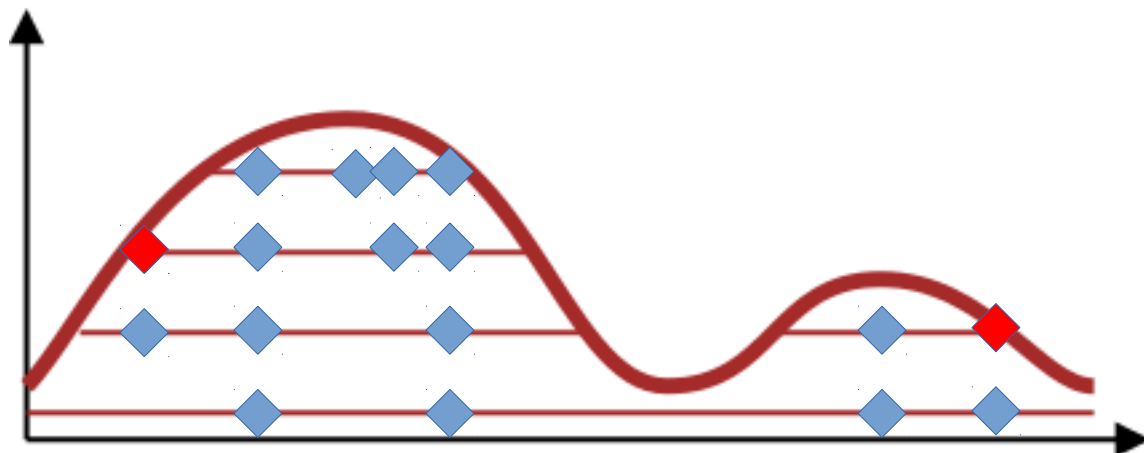
$$Z \approx \sum_i \Delta V_i \cdot L_i$$

Skilling '04,06,09  
Evans '07  
Chopin&Robert '07,10  
Walter '14

# Nested sampling idea (4)



Integration  
from low to high  
likelihoods



$$Z \approx \sum_i \Delta V_i \cdot L_i$$

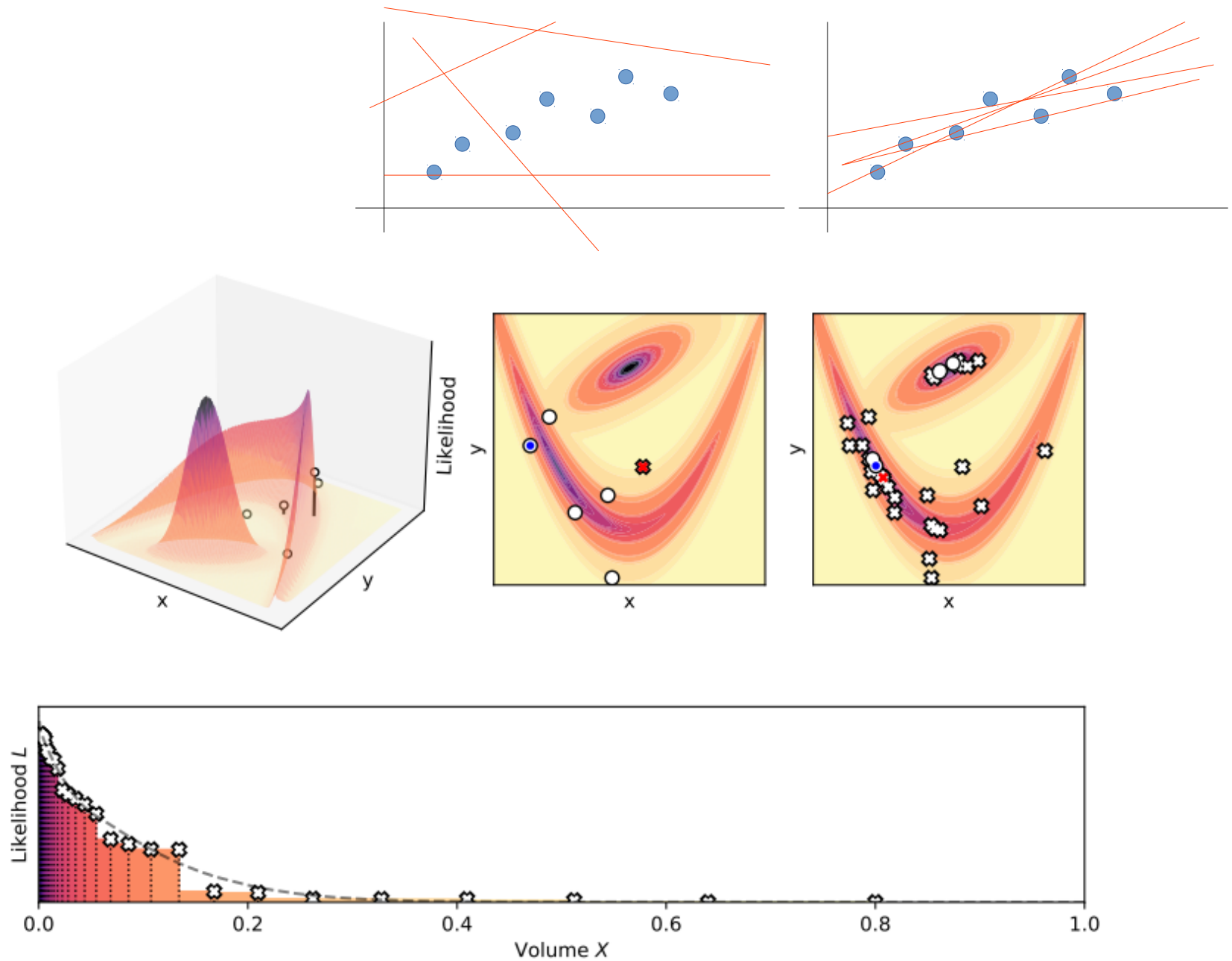
Skilling '04,06,09  
Evans '07  
Chopin&Robert '07,10  
Walter '14

# NS algorithm

- Generate initial live points from prior, evaluate  $L$
- Set  $V_1=1$
- Loop
  - Remove lowest  $L$  point =  $L_{\min}$
  - Dead point posterior weight:  $w_i = V_i \times 1/N \times L_i$
  - $V_{i+1} = V_i \times (1 - 1/N)$
  - sample new point, subject to  $L \geq L_{\min}$
- $Z = \sum_i w_i$



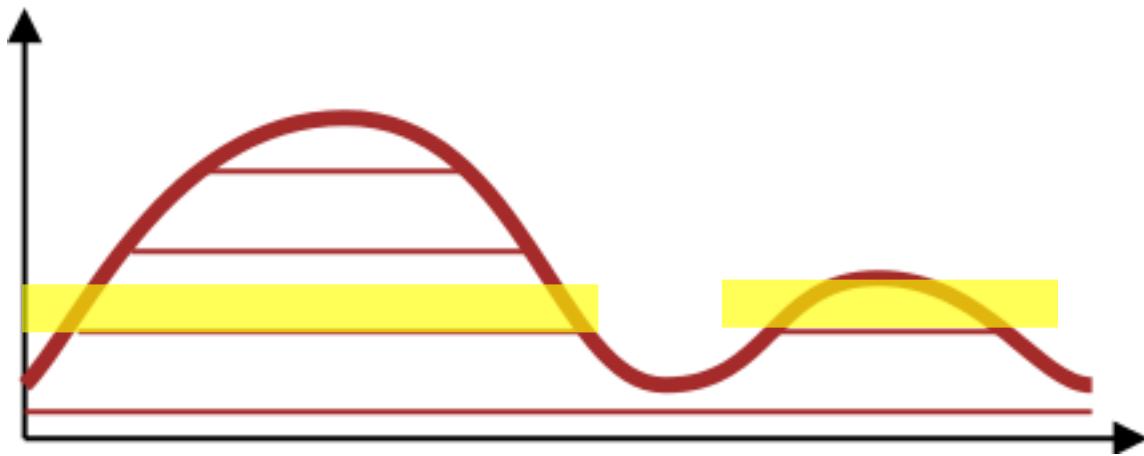
# A cooling scheme



# Nested sampling theory

- Idea Skilling 2004, 2006
- Convergence Evans 2007  
Chopin & Robert 2010  
Skilling 2009
- Point Process-based Monte Carlo Walter (2014)
- Sequential Monte Carlo Salomone+ 2018  
Superset of SMC & NS algorithms  
with strong theoretical foundations  
  
Birge (2012)  
Polson & Scott (2014)

# Constrained sampling

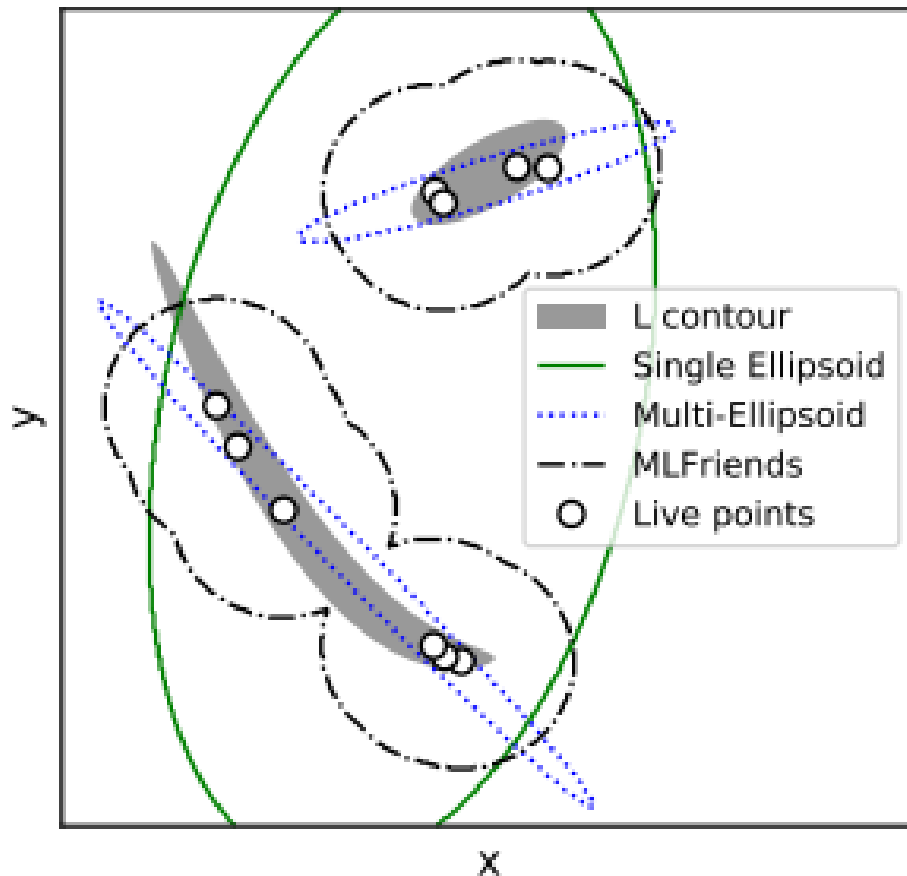


$$Z \approx \sum_i \Delta V_i \cdot L_i$$

Skilling '04,06,09  
Evans '07  
Chopin&Robert '07,10  
Walter '14

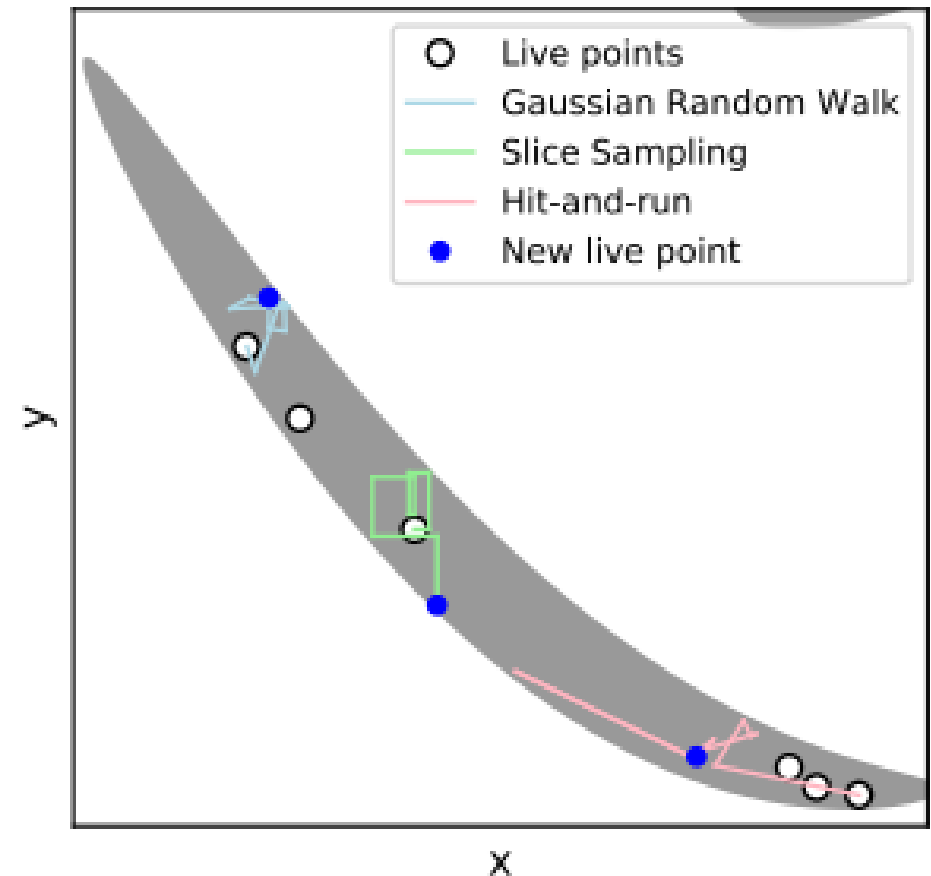
# L-restricted prior sampling

Region sampling LRPS



Non-volume preserving flows  
Moss+2019

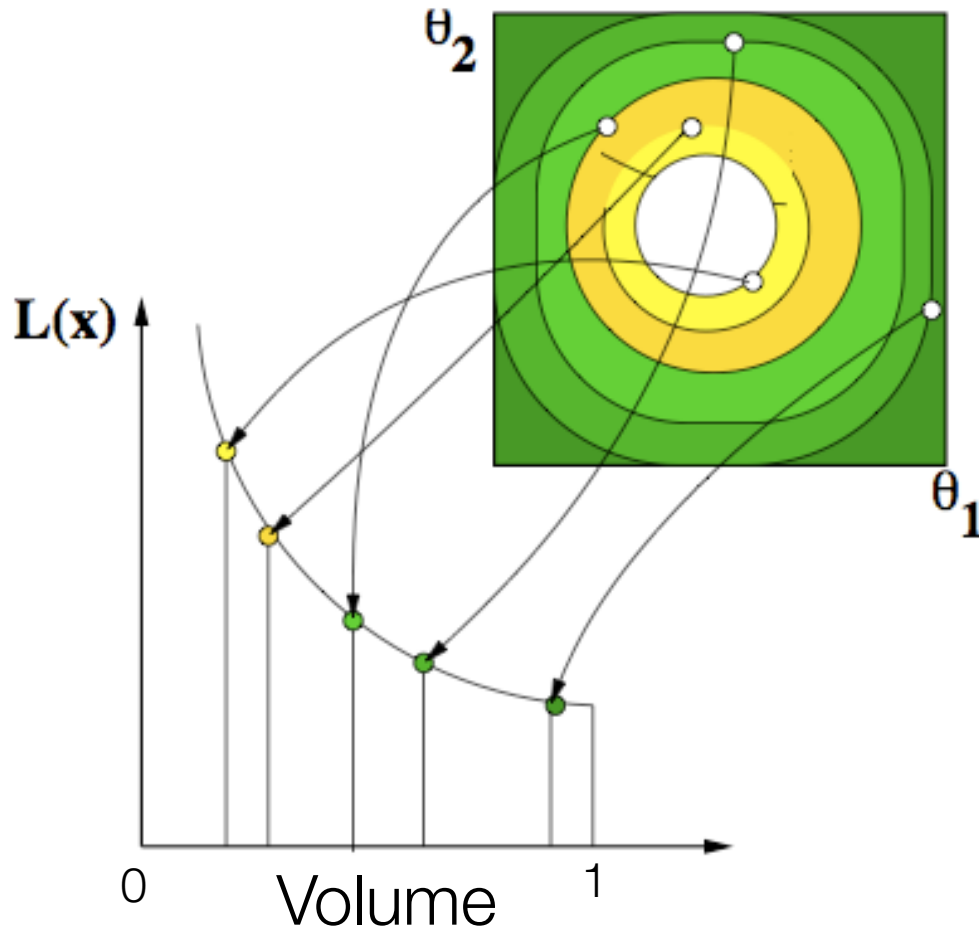
Local step LRPS



Slice sampling: Handley+2015

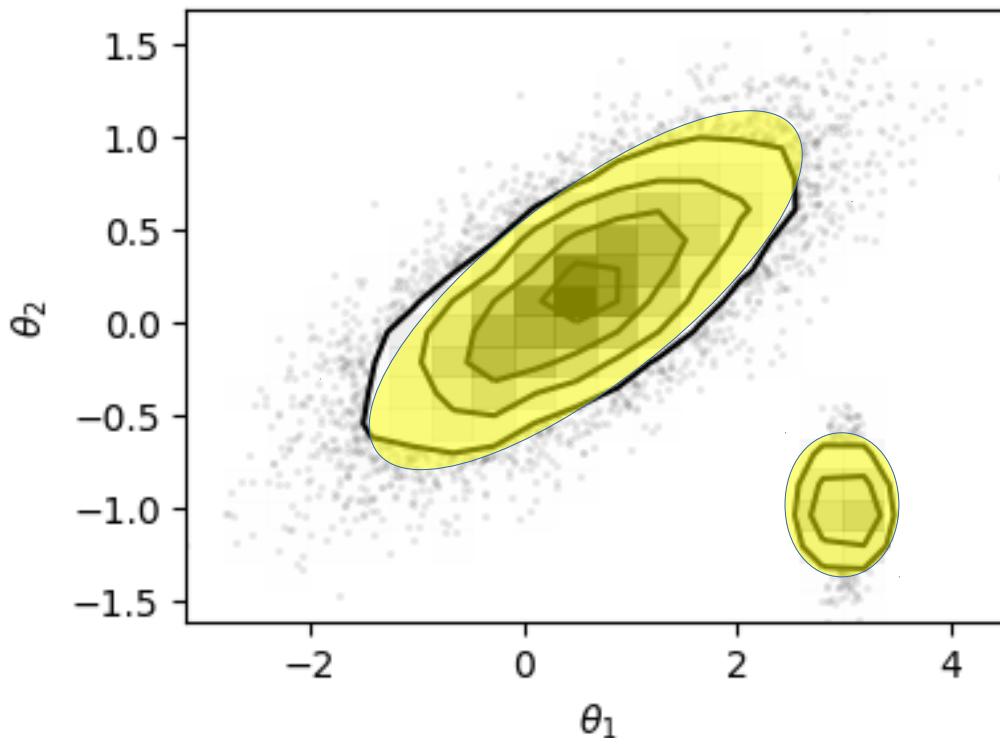
Billiard walks with gradients:  
Betancourt (2011), Skilling (2012)

# Multi-dimensional case



- Ordering of samples well-defined by  $L$
- “Indep. of dim”  
depends primarily on constrained sampler!
- Not limited to continuous spaces of fixed dimensions

# Region sampling



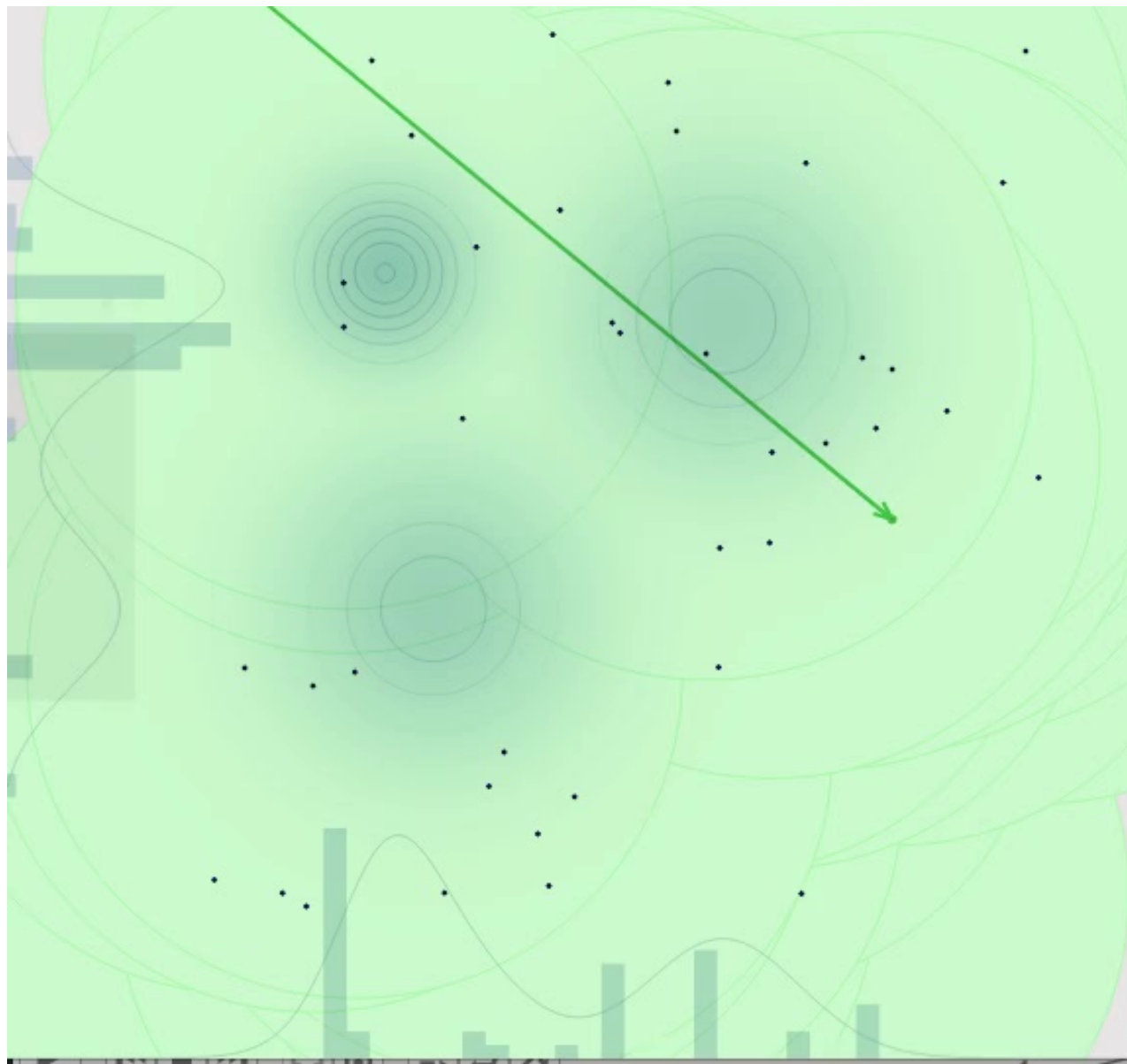
- Ellipsoidal sampling  
Mukherjee+06
- Multi-ellipsoidal sampling  
Shaw+07,  
Feroz&Hobson08,  
Feroz+09
  - MultiNest (pymultinest)
- Very popular!

Clustering live points:  
x-means (Feroz&Hobson 2008)  
Iterative Jarvis-Patrick (Handley+2015)

Wilks' theorem  
Elliptical distributions

uniformly sampled, pure, slowly changing

# Animation of MLFriends NS



For efficiency:

Reconstruct a region

Sample uniformly

RadFriends: ellipsoid  
around each point  
Circle size cross-  
validated to recover  
current points

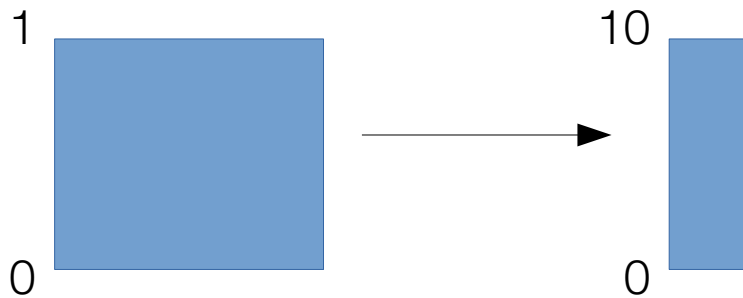
Alternative: MCMC

<https://chi-feng.github.io/mcmc-demo/app.html>



# Specifying priors

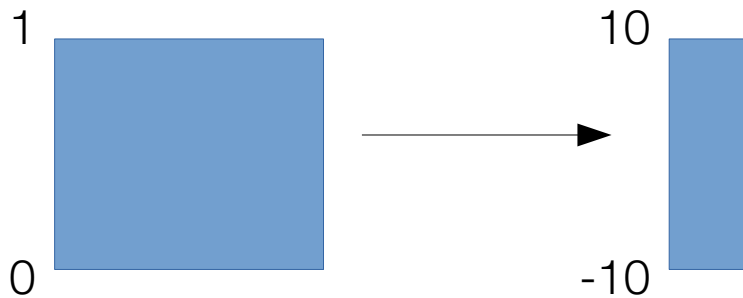
Transform uniform cube to physical parameters



```
def my_prior_transform(cube):  
    # cube is a d-dimensional array  
    params = cube.copy()  
    # from 0 to 10  
    params[0] = cube[0] * 10  
    return params
```

# Specifying priors

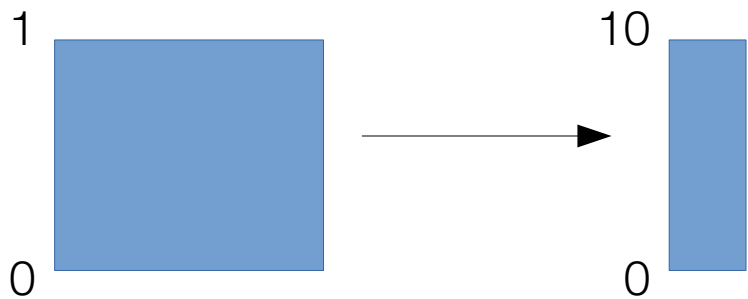
Transform uniform cube to physical parameters



```
def my_prior_transform(cube):  
    # cube is a d-dimensional array  
    params = cube.copy()  
    # from -10 to +10  
    params[0] = cube[0] * 20 - 10  
    return params
```

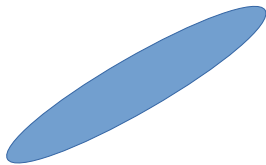
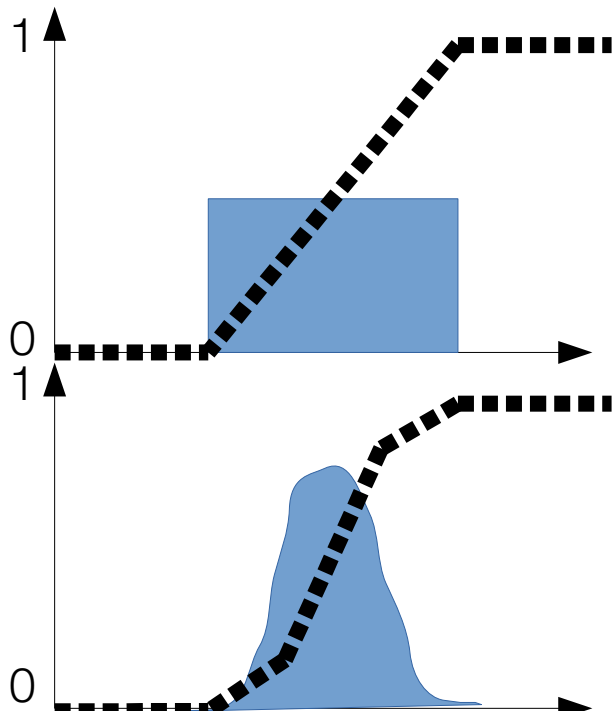
# Specifying priors

Transform uniform cube to physical parameters



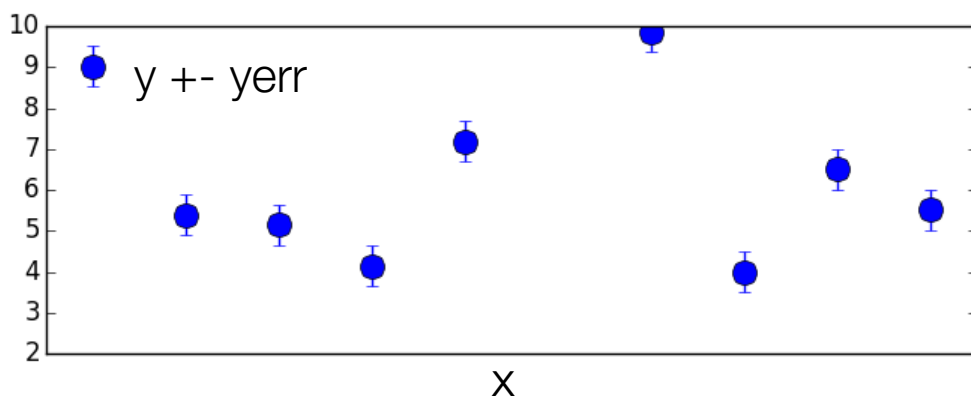
```
def my_prior_transform(cube):  
    # cube is a d-dimensional array  
    params = cube.copy()  
    # from -10 to +10  
    params[0] = cube[0] * 20 - 10  
    # from 1 to 1010  
    params[1] = 10**(cube[1] * 10)  
    # via inverse CDF  
    params[2] = rv.ppf(cube[2])  
    return params
```

```
# Gaussian prior 5+/-1  
rv = scipy.stats.norm(5, 1)
```



# Specifying likelihoods

## Gaussian example

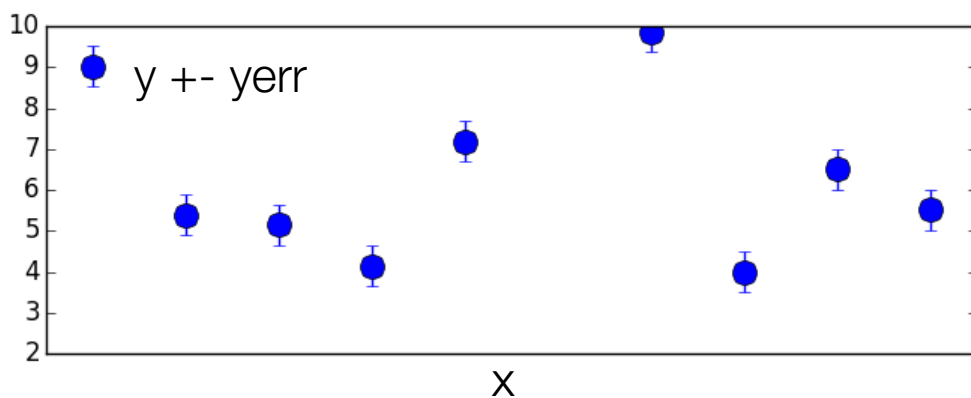


```
mydata = numpy.loadtxt("mydata.txt")  
x, y, yerr = mydata.transpose()
```

```
def my_likelihood(params):  
    # params is a d-dimensional array  
    # already transformed  
    a, b, c, d = params  
    # compute model prediction:  
    m = (a + numpy.sin(x * c)) * d  
    # compute gaussian likelihood  
    return -0.5*(((m - y)/yerr)**2).sum()
```

# Specifying likelihoods

## Gaussian example



Arbitrarily complex model calculation →

Arbitrarily complex data uncertainties →

```
mydata = numpy.loadtxt("mydata.txt")
x, y, yerr = mydata.transpose()
```

```
def my_likelihood(params):
    # params is a d-dimensional array
    # already transformed
    a, b, c, d = params
    # compute model prediction:
    m = (a + numpy.sin(x * c)) * d
    # compute gaussian likelihood
    return -0.5*(((m - y)/yerr)**2).sum()
```

```
result = solve(
    LogLikelihood=my_likelihood,
    Prior=my_prior_transform,
    n_dims=4,
    outputfiles_basename='mysine_')
```

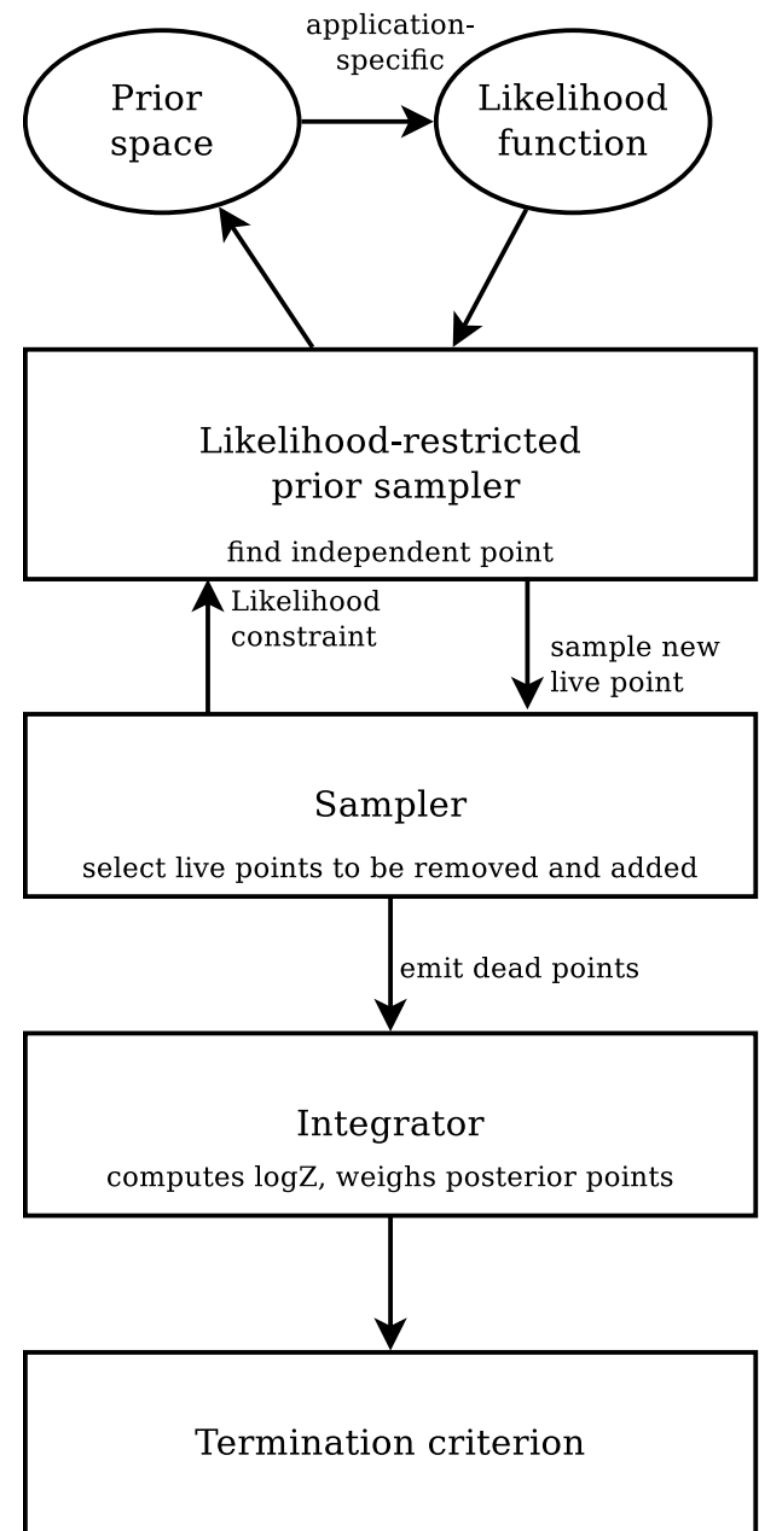
# --- Tutorial session ---

<https://johannesbuchner.github.io/UltraNest/example-sine-line.html>

- Fit a time series with a sine
- Parameter estimation and Z computation
- Visualisations

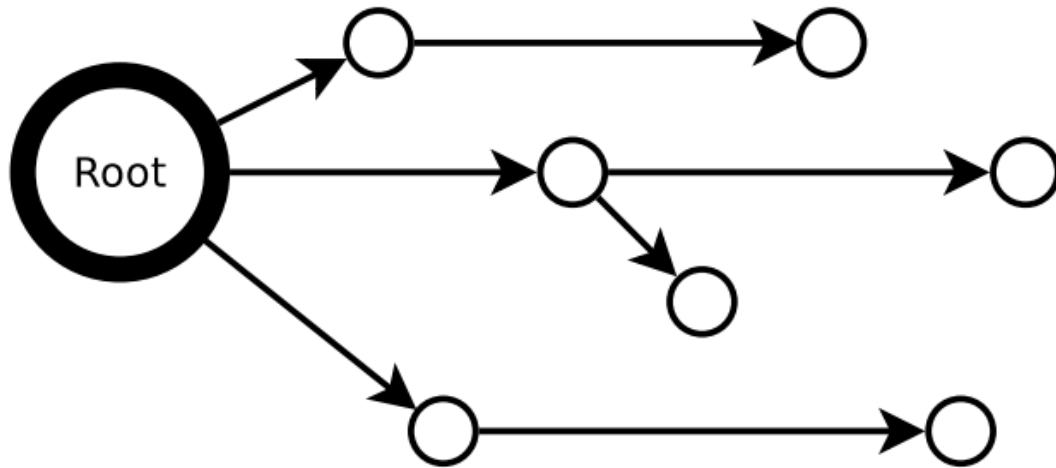
# Components

- Advanced NS topics
  - Integrators
  - Error estimates
  - Diagnostics: Tests & Visualisations
  - LRPS
    - Local steps
    - Region based
    - Hybrid
  - Termination criteria
  - Parallelisation
  - Extensions





# Tree search view



- Root = prior Volume
- Outgoing edges: split volume
- # parallel edges = # of live points

Assume queue, sorted by likelihood value  
add all children of root node to queue

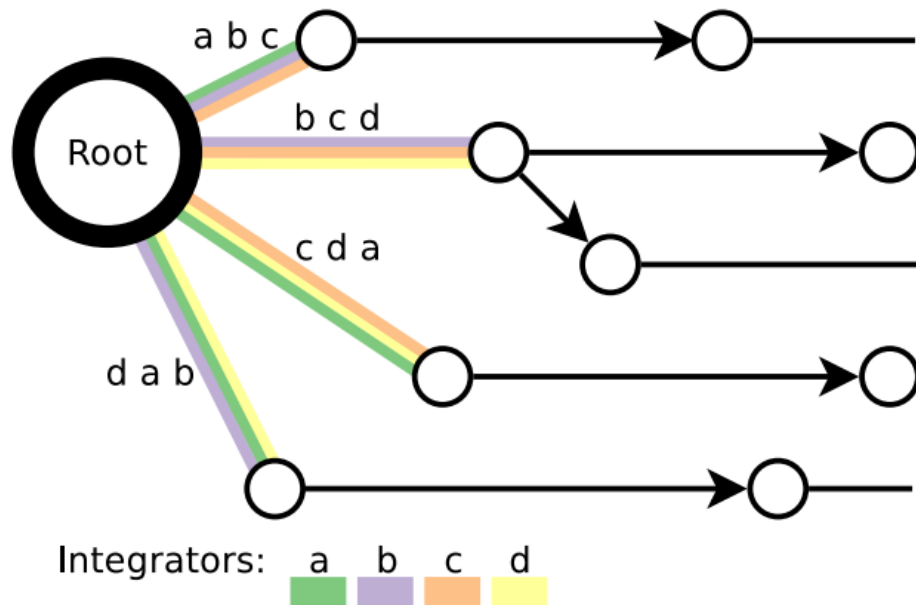
```
while queue is not empty:  
    Nlive = length of queue  
    obtain and remove next node from queue
```

```
    store dead node with weight  $L * V$   
    shrink volume by  $1 / Nlive$ 
```

```
    if not terminating:  
        expand node
```

```
    add all children of node to Q
```

# Error estimation



Estimates L noise.

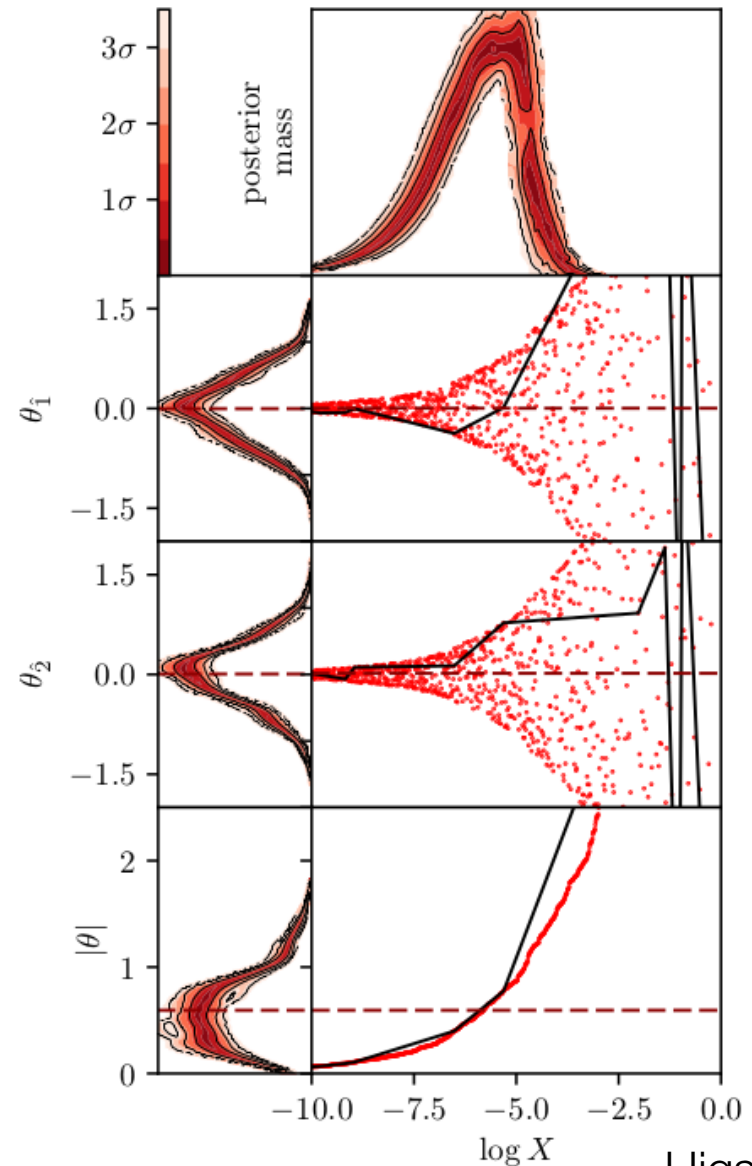
To estimate V noise, change volume shrinkage

$dV=1/N \rightarrow \text{beta}(N, 1)$  randomly

- Multiple integrators
  - Blinded to some threads
- 
- LRPS validation
  - Integration validation

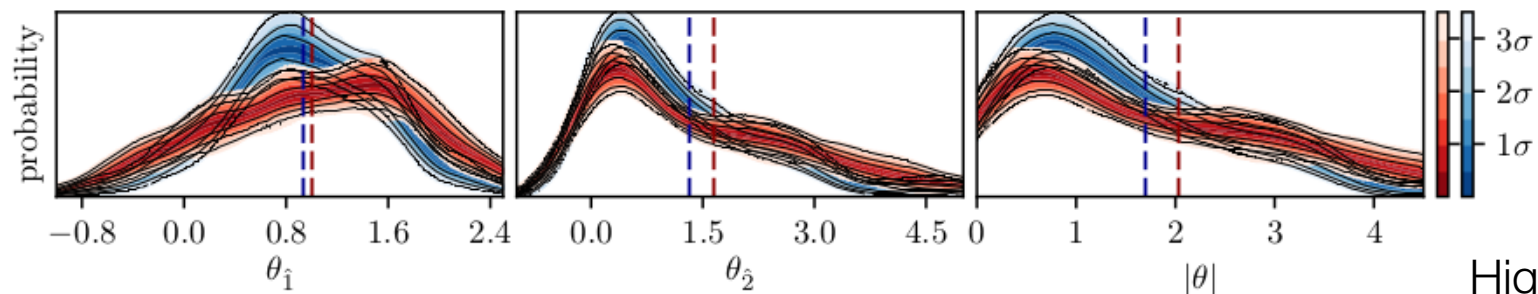
# Diagnostics: Visualisations

- Each leave-k-out integrator gives results  
→ uncertainty in posterior  
  
(+uncertainty in  $Z$ )

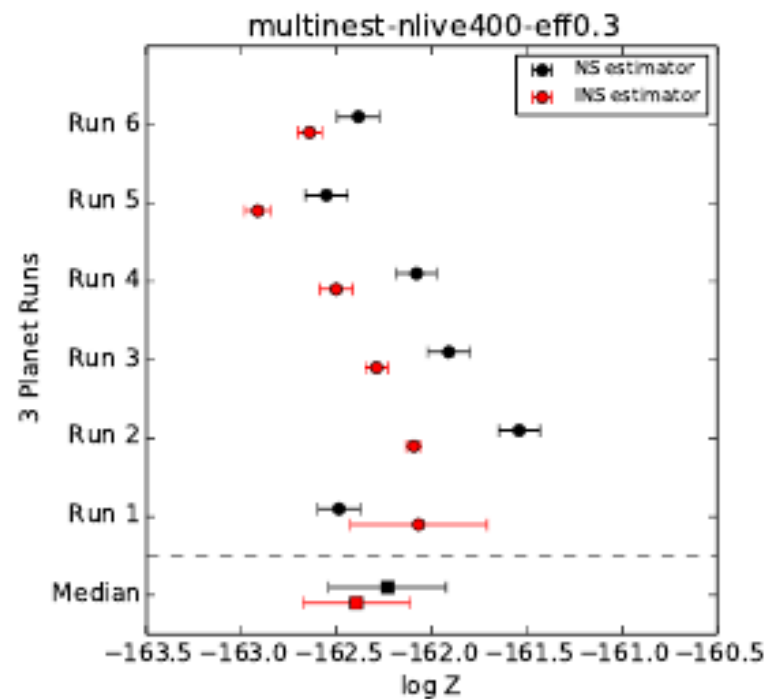


# Diagnostics: Visualisations

- Comparison of multiple runs



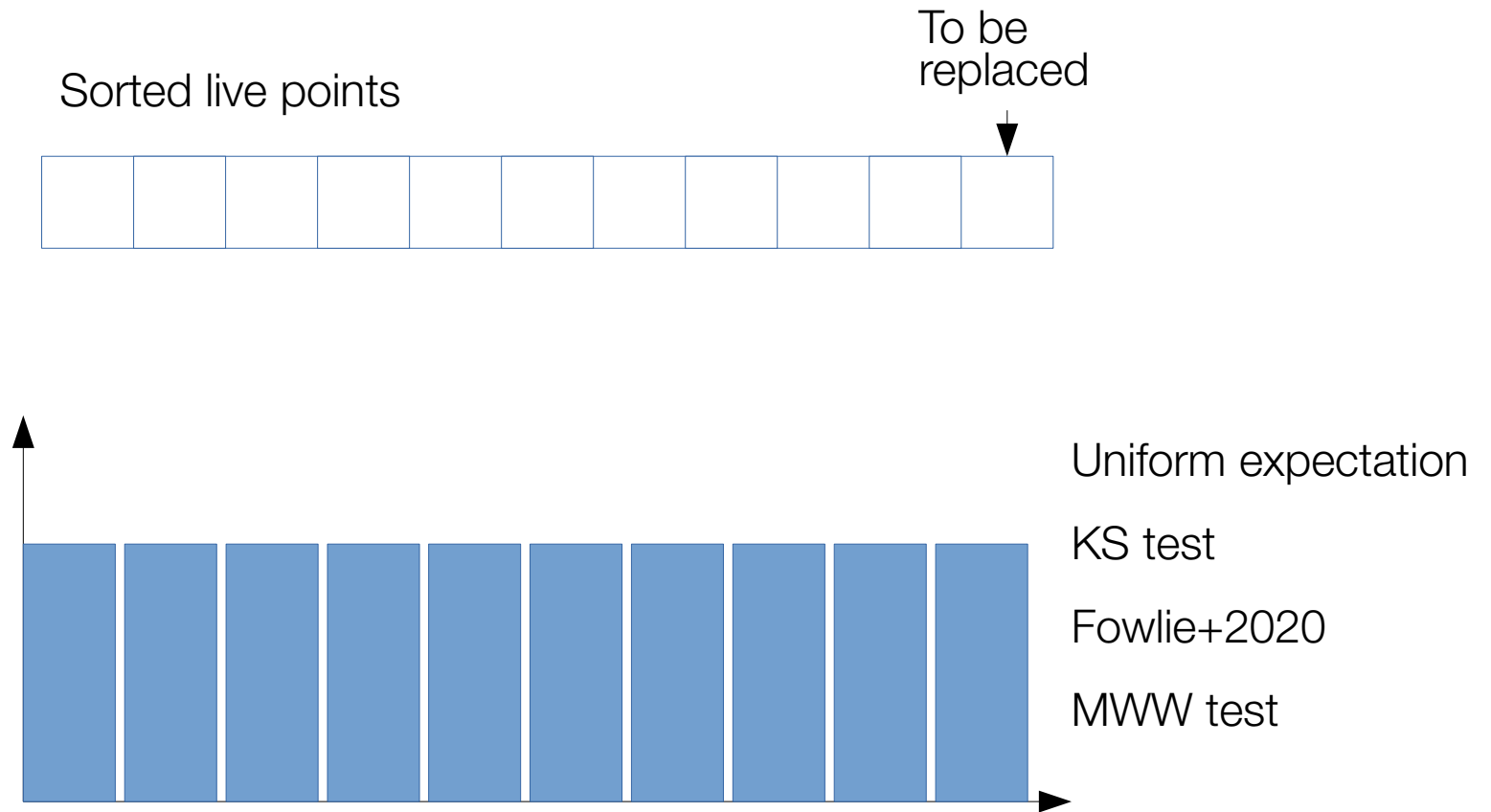
Higson+19



Nelson+20

# Diagnostics: tests

## Insertion order



# Parallelisation

- Within the likelihood
- LRPS farmed out
- Removing multiple points
- Merging independent runs
- Analyzing multiple data sets simultaneously

# Extensions

- Dynamic NS
- Using thrown-away points

Higson+17

Importance NS

Chopin & Roberts 07, 08,10

- No hard borders:

- Diffusive NS
- Daemonic NS

Brewer+11

- HMC integrations

Habeck15

- 

- Daemonic NS for smooth border
- Thermometer as diagnostics



# Related algorithms

- Simulated annealing
  - Special cooling schedule
- Sequential Monte Carlo <https://arxiv.org/abs/1805.03924>
  - Population is reused, special proposal
- Bridge, path sampling
- Importance sampling

# Future

- Theory: SMC–NS
- Deeper HMC-NS blends
- More implementations in other languages

Parallelisation  
Low latency  
Resuming  
Easy to install  
High-d  
Reliable

Review on Nested Sampling methods in prep.