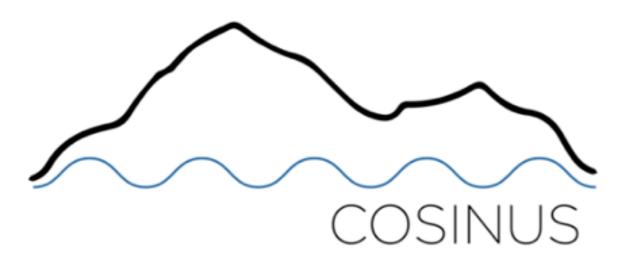
# COSINUS @ Gran Sasso: Construction, Setup and Prototype Detector Results

**Project Review 2021** 

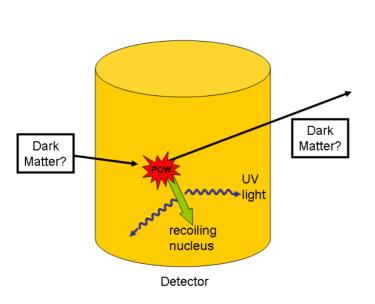
M. Stahlberg

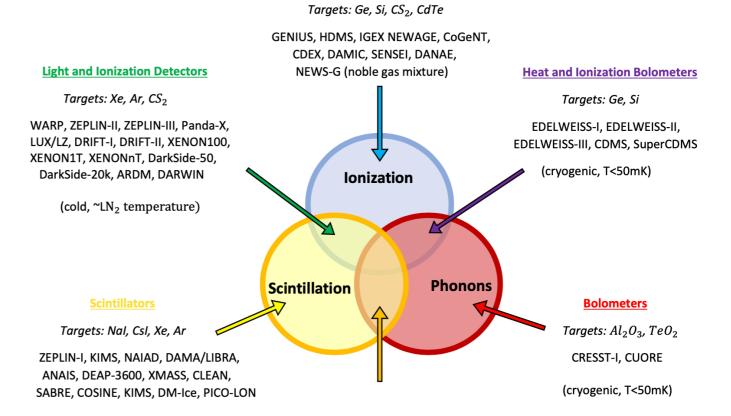




### "Direct" DM Search Landscape

(= Measure small energies deposited in a target material by weakly interacting dark matter particles)





**Ionization Detectors** 

#### **Emulsion**

Targets: NIT (silver halides)
NEWSdm

Targets:  $CaWO_4$ , BGO,  $Al_2O_3$ , Nal, Csl,  $Li_2MoO_4$ ROSEBUD, CRESST-II, CRESST-III,

**Light and Heat Bolometers** 

COSINUS, LUPID
(cryogenic, T<50mK)

#### **Bubbles and Droplets**

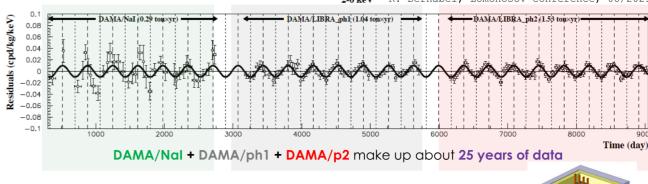
Targets:  $CF_3Br$ ,  $CF_3I$ ,  $C_3F_8$ ,  $C_4F_{10}$ COUPP, PICASSO, SIMPLE, PICO-2L, PICO-60, PICO-250



### Why COSINUS?

- Long-standing disagreement between "DAMA-modulation" and null-results from other experiments
- So far, few experiments with same target material; constraints from COSINE & ANAIS
- COSINUS is the only experiment with eventby-event separation (!)

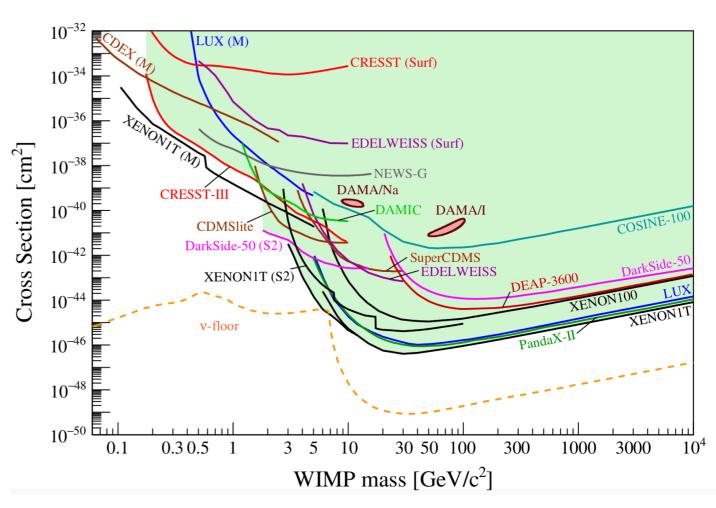
Scenario: Spin-independent elastic DM-nucleus scattering, "standard" DM halo



Total exposure: 2.86 tonne years

Statistical significance: 13.7  $\sigma$ 

positive evidence for the presence of DM particles in the galactic halo

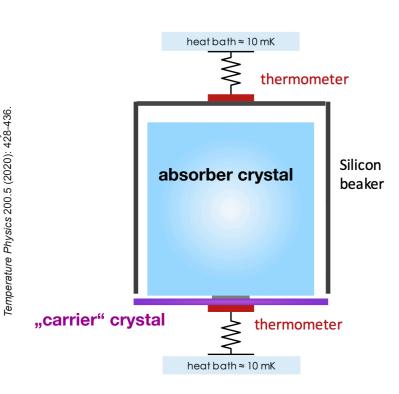


Billard, Julien, et al. arXiv: Direct Detection of Dark Matter--APPEC Committee Report. No. arXiv: 2104.07634. 2021.

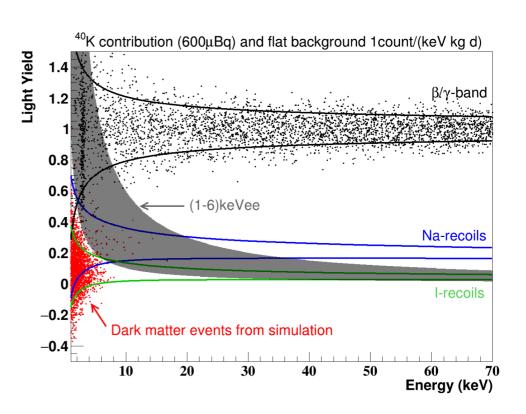


"Cryogenic Observatory for Signatures seen in Next-generation Underground Searches"

#### "Baseline" design

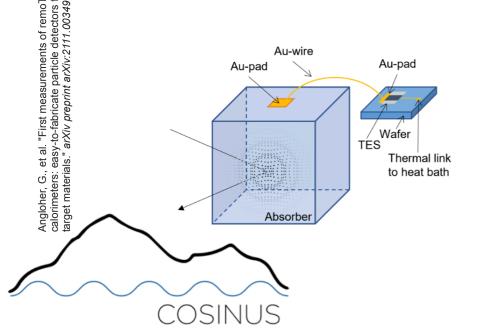






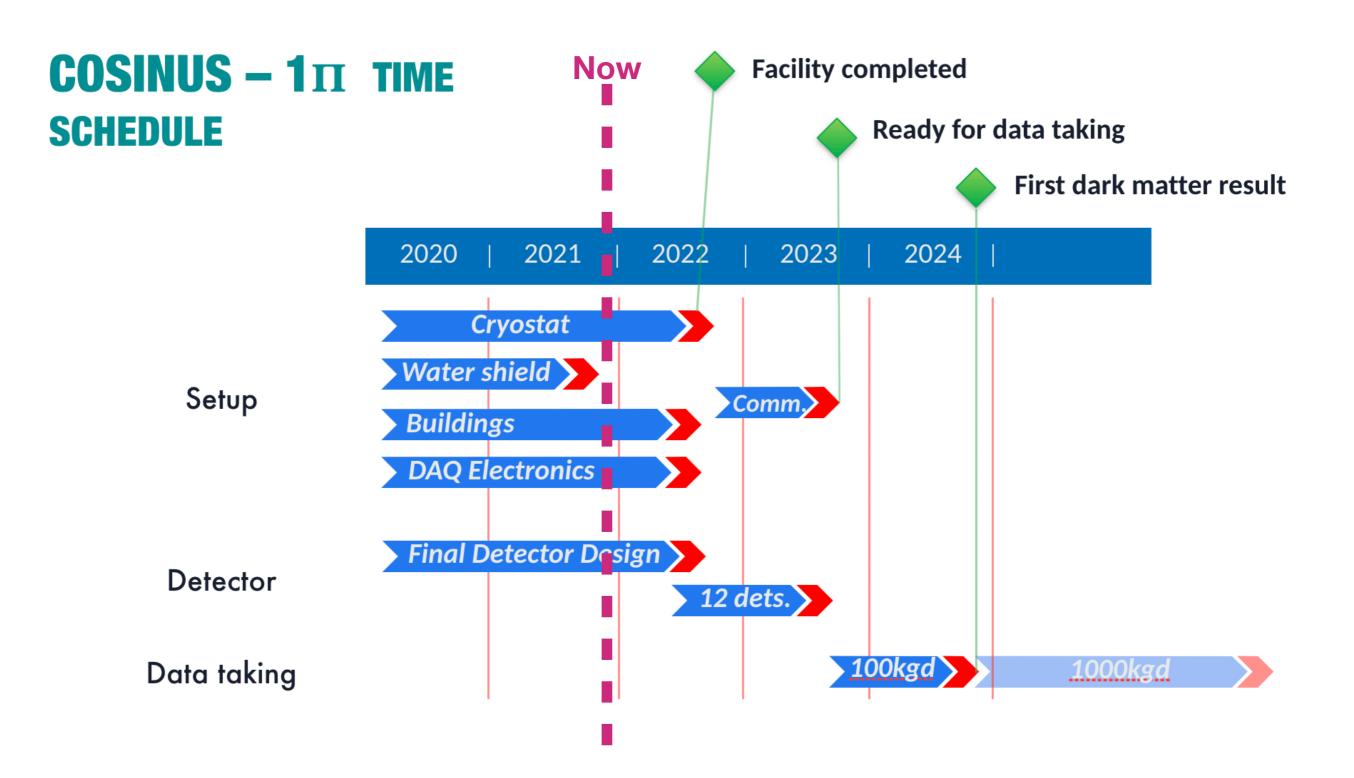
Simulation with 100 kg day exposure and 1 keV threshold

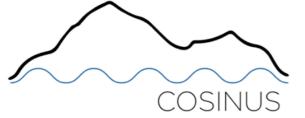
#### "remoTES" design

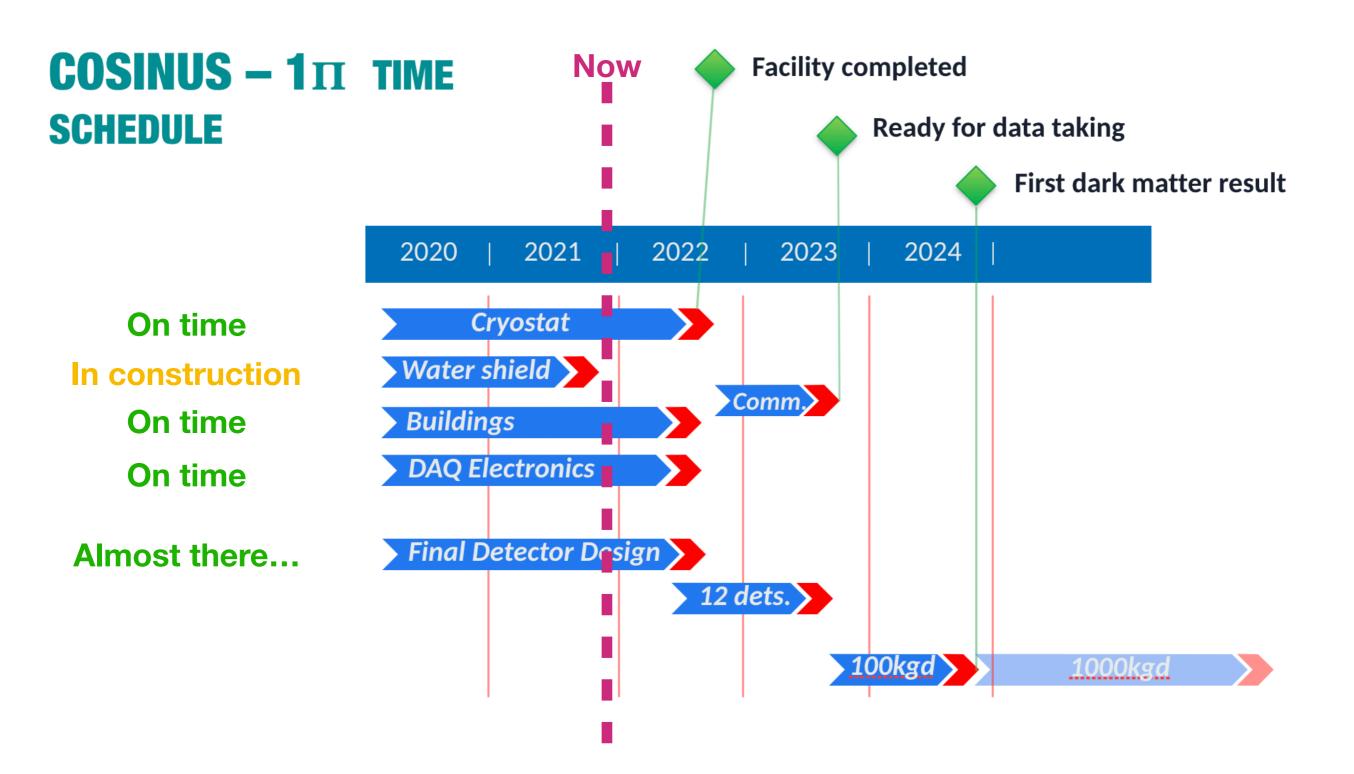


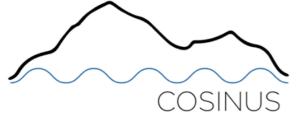
Phonon (~90%) + Scintillation (~10%) signal

- -> Factor 10 improvement over light only (e.g. DAMA)
- -> Nuclear recoil energy threshold instead of electron equivalent
- -> Use cryogenic sensor technology and know-how from CRESST









### COSINUS - Who we are





















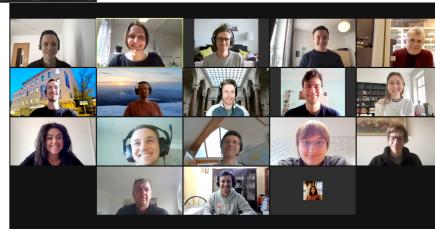


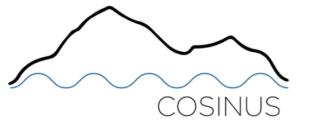












### Collaboration Meeting 10/2021 @LNGS





### COSINUS group at MPP, 12/2020









Karoline Schaeffner (Group Leader)





Franz Proebst (Senior scientist)

Michele Mancuso (Staff scientist)



Godehard Angloher (Senior scientist, part time)



Robert Stadler (Engineer)



Christopher Jablonski (Engineer)



Martin Stahlberg (Postdoc)

Group started in 2019 as Max Planck Research Group



### COSINUS group at MPP, 12/2021









Karoline Schaeffner (Group Leader)

#### **Support from CRESST**



Federica Petricca (Senior scientist)



Franz Proebst (Senior scientist)



Michele Mancuso (Staff scientist)



Godehard Angloher (Senior scientist, part time)



Robert Stadler (Engineer)



Christopher Jablonski (Engineer)



Torsten Frank (Senior scientist, part time)



Karl-Heinz Ackermann (Engineer, part time)



Peter Mühlbauer (Technician, part time)



Moritz Kellermann (PhD student)



Mukund R. Bharadwaj (PhD student)



Martin Stahlberg (Postdoc)



Vanessa Zema (Postdoc)

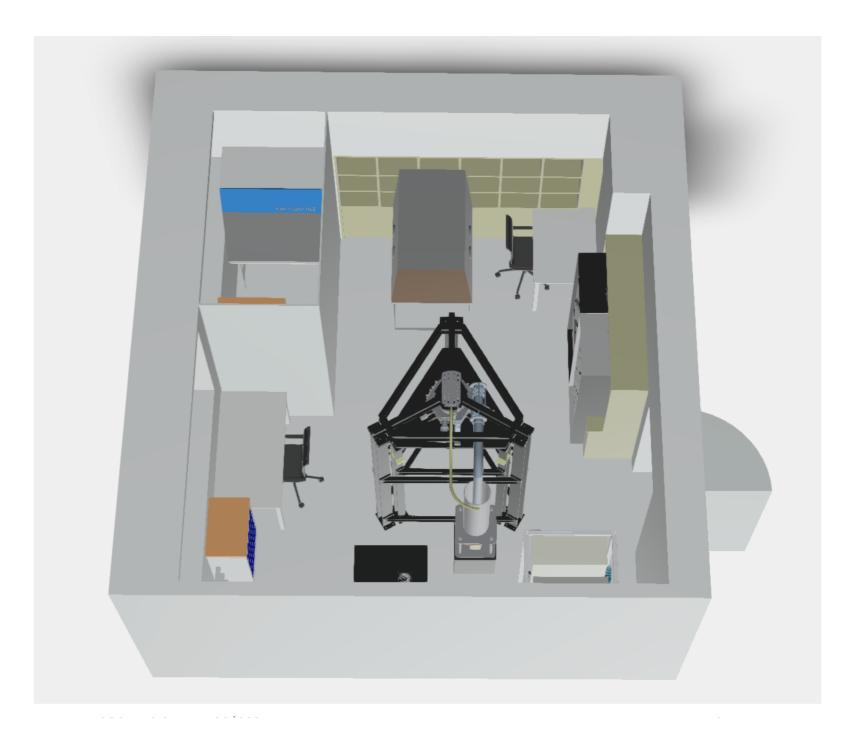


Henrik Ansorge (Working Student)



### COSINUS lab

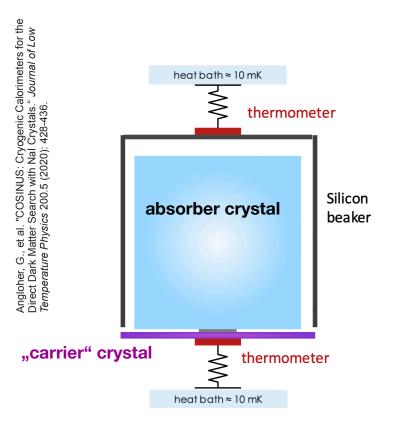
- Glove box for detector mounting
- Nitrogen cabinet for storage of detector components
- Sealing machine for Nal crystals
- Working tables
- Delivery in March:
  - → Dry cryostat

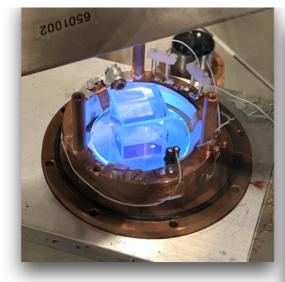




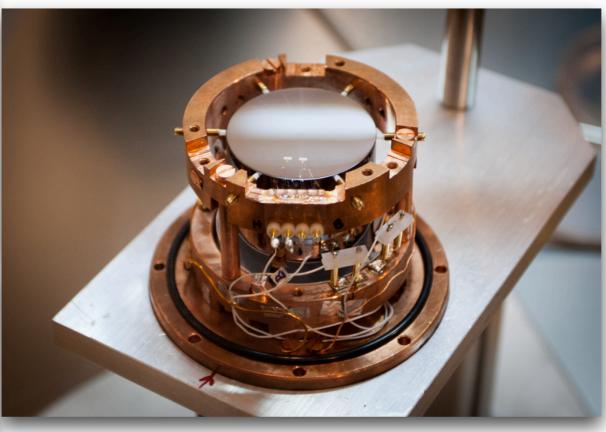
### "Baseline" detector design

- Si beaker encloses Nal absorber, which is interfaced to "carrier" crystal with Transition Edge Sensor (TES)
- Nal is hygroscopic -> handling difficult, interface problematic
- Carrier can collect light (!)







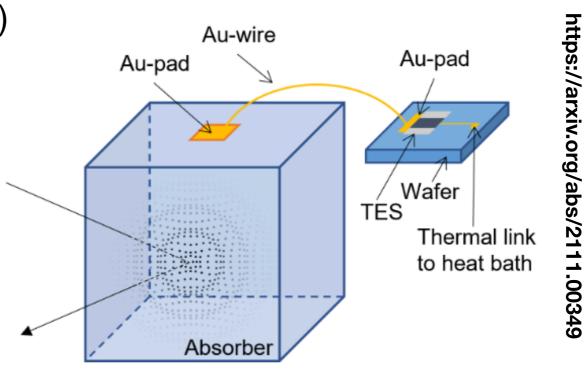




### "remoTES"

- Detector concept by Matt Pyle
- First measurements by COSINUS group @MPP
- Idea: avoid problems of attaching TES with carrier to absorber crystal
- Influence of Au link and pads on the signal shape
- Wafer is reusable, easier to fabricate (!)







**Project Review 2021** 

### DAQ & Software

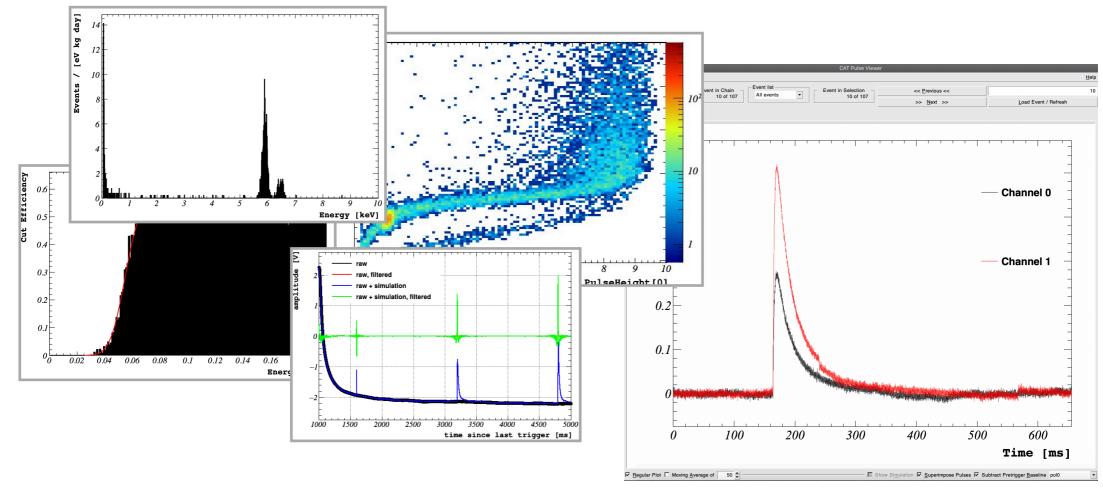
DAQ electronics & readout software developed by



Analysis Software developed at MPP in collaboration with HEPHY

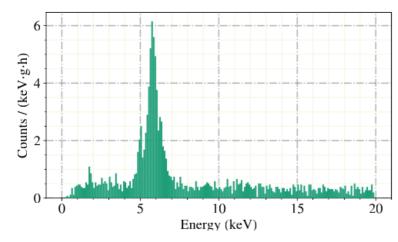


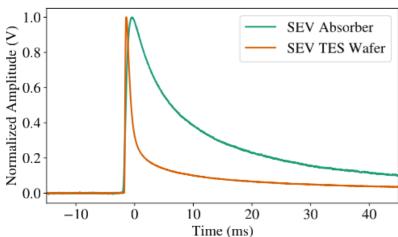
All of this used also in CRESST





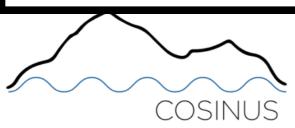
#### remoTES - first results





https://arxiv.org/abs/2111.00349

# Many thanks to the CRESST group for their support and access to their facility!



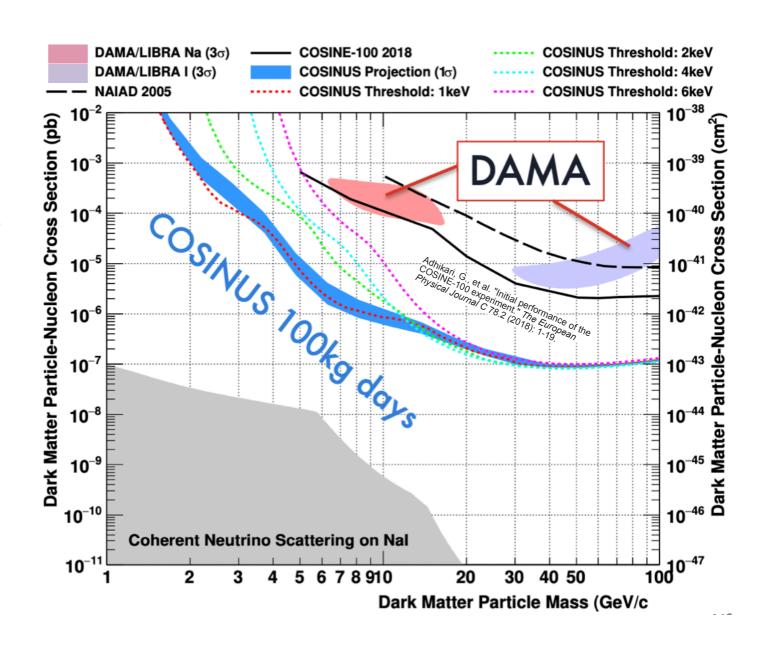
Run	Absorber material	Absorber size [mm]	Au-pad properties	Au wire properties	TES-type	Energy threshold [eV]
557	$CdWO_4$	20x10x5	200nm	17 μm	LD type	600*
			sputtered	wedge bond	on $Al_2O_3$	
561	Si	20x10x5	200nm	17 μm	LD type	900*
			sputtered	wedge bond	on $Al_2O_3$	
561	Si	20x10x5	200nm	17 μm	istick-TES	900*
			sputtered	wedge bond	on Si	
562	Si	20x10x5	200nm	10 μm	istick-TES	770
			sputtered	wedge bond	on Si	
564	CsI	Ø <b>=</b> 20	300nm	17 μm	LD type	18000*
		h=2	sputtered	wedge bond	on $Al_2O_3$	
564	CsI	Ø=20	600	17	istick-TES	100000*
		h=2	sputtered	wedge bond	on Si	
566	Si	20x10x5	200nm	17 μm	LD type	323
			sputtered	glued on pad	on $Al_2O_3$	
567	Si	2021025	200	17, bi	LD type	800*
			sputtered	glue spot on pad	on $Al_2O_3$	
569	Si	20x10x5	no pad	17 μm, glued	LD type	40000*
			sputtered	directly on Si	on $Al_2O_3$	
569	Si	20x10x5	400nm foil	17 μm	istick-TES	665
			RRR=15	glued on pad	on Si	
			glued on Si			
571	Si	20x10x5	400nm foil	17 μm	istick-TES	735
			RRR=15	wedge bond	on Si	
			glued on Si			
571	${ m TeO_2}$	20x10x5	400nm foil	17 μm	LD type	<4000*
			DDD 15	wedge bond	$\alpha$ n Al $\alpha$ O $\alpha$	
			glued on TeO <sub>2</sub>			
573	${ m TeO}_2$	20x10x5	400nm foil	17 μm	LD type	1100
			RRR=15	2 wedge bonds	on $Al_2O_3$	
			gland on ToO2			
574	${ m TeO}_2$	20x10x5	400nm foil	17 μm	LD type	3760
			RRR=15	6 wedge bonds	on $Al_2O_3$	
			glued on $TeO_2$			
577	${ m TeO}_2$	20x10x5	400nm foil	17 μm	LD type	2989
			RRR=15	4 wedge bonds	on $Al_2O_3$	
			glued on TeO <sub>2</sub>			

### Physics Reach

Nal absorber mass: ~100g

-> 10 detectors x 100days = 100kg days

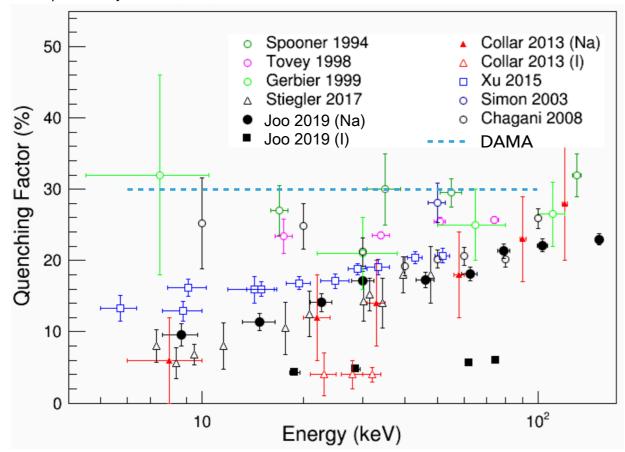
No rate -> no modulation (!)



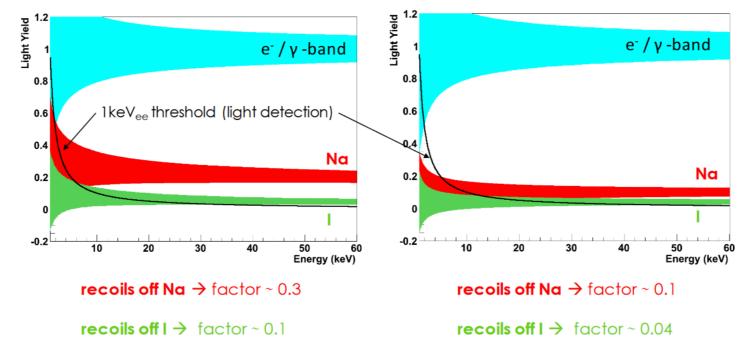


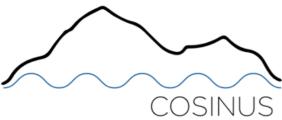
#### "QUENCHING FACTOR MYSTERY"

Modified from: Joo, H. W., et al. "Quenching factor measurement for NaI (TI) scintillation crystal." Astroparticle Physics 108 (2019): 50-56.



- Measurements of quenching factors (QF) at room temperature do not agree
- In particular, role of TI is unclear (usually crystals are doped)
- Strong influence of QF on NR energy scale
   -> comparability of modulation studies (!)
- COSINUS will provide the first cryogenic QF measurement for Nal



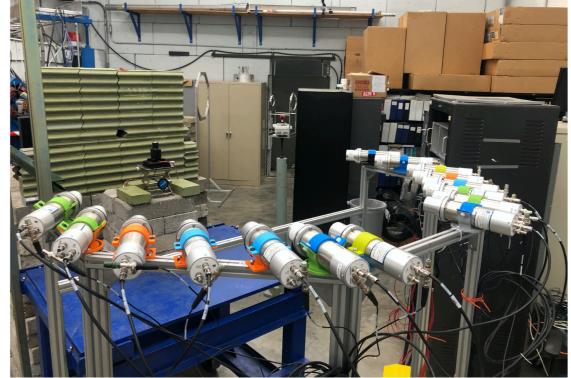


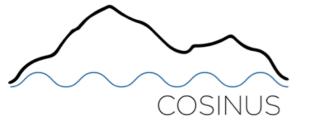
#### "QUENCHING FACTOR MYSTERY"

- Initiated QF Measurement at TUNL (Triangle Universities Nuclear Laboratory),
   performed in September 2021 by the Barbeau group
- ▶ Monoenergetic neutron beam (~3.3MeV), detectors at different scattering angles
- Five ultrapure crystal samples from SICCAS with different TI dopant levels tested
- Analysis ongoing!

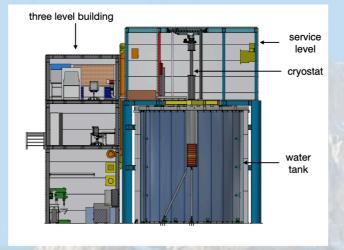




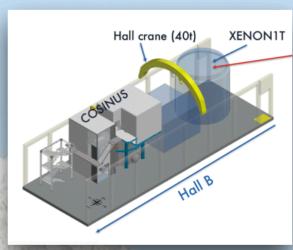










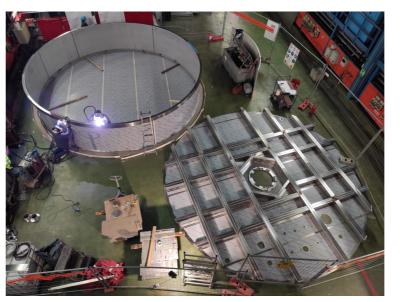




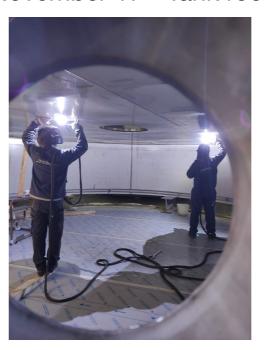
### Construction progressing...

October 10 - Screed installation ("Estrich")





November 17 - Tank roof

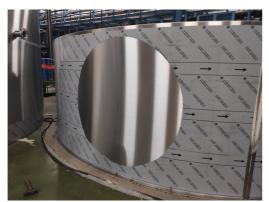


December 07 - Third ring

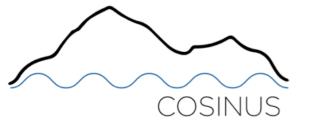


**Today:** Fourth ring & manhole

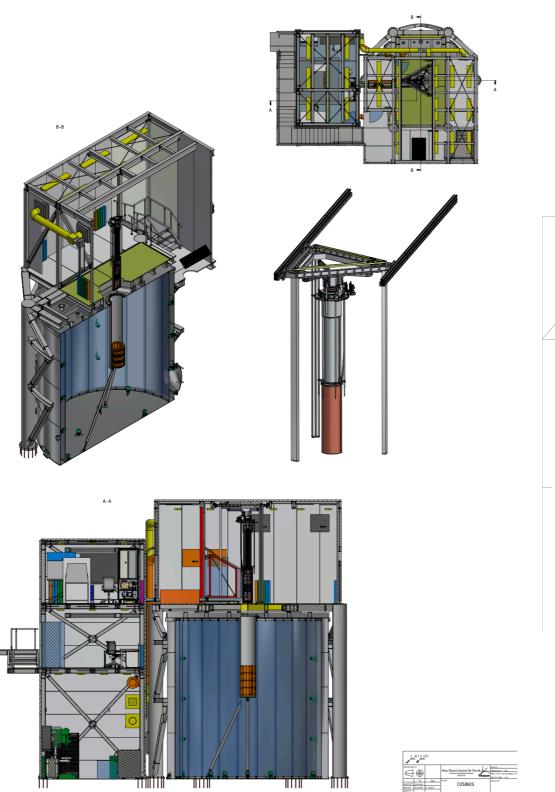


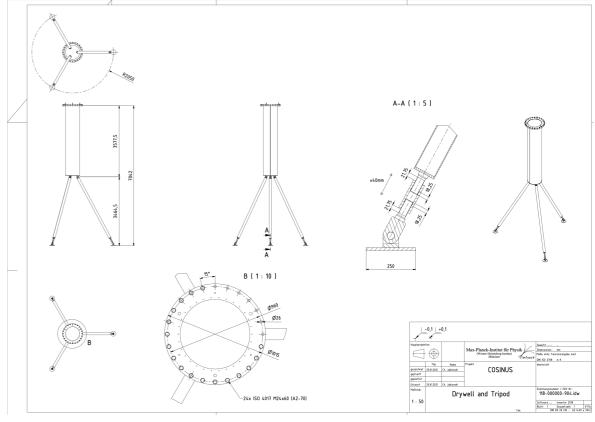


Tank will be completed by Christmas!



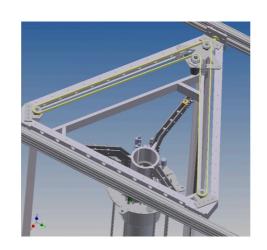
### Lifting system, Drywell & Tripod







### Lifting System



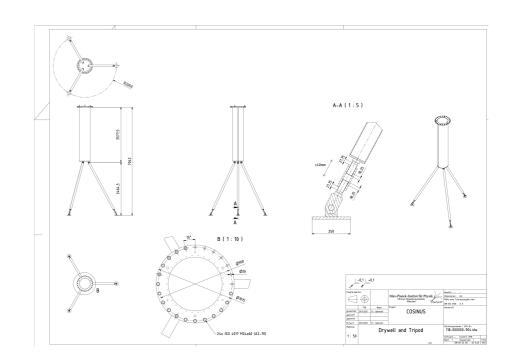
- Designed at MPP
- Motor-drive delivered
- Control unit in preparation
- System will be tested and commissioned at MPP

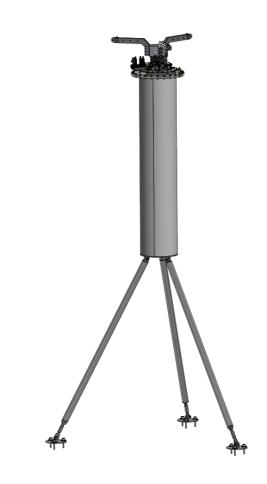




### Tripod & Drywell

- Designed at MPP
- Screening and careful selection of material
- Manufacturing started
- Delivery to LNGS planned for 2021

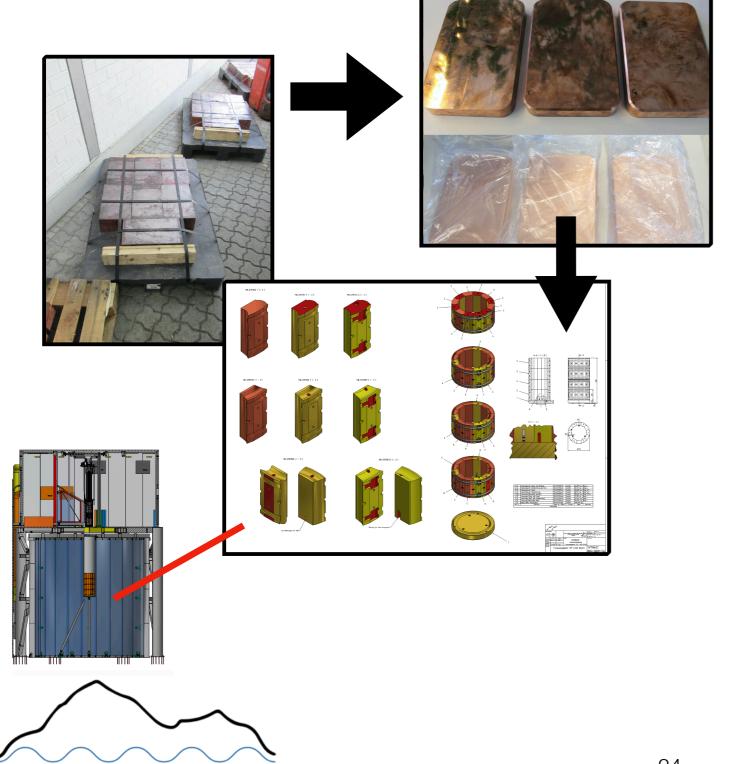






### Copper (external) shield

#### **Design & materials screening**



COSINUS

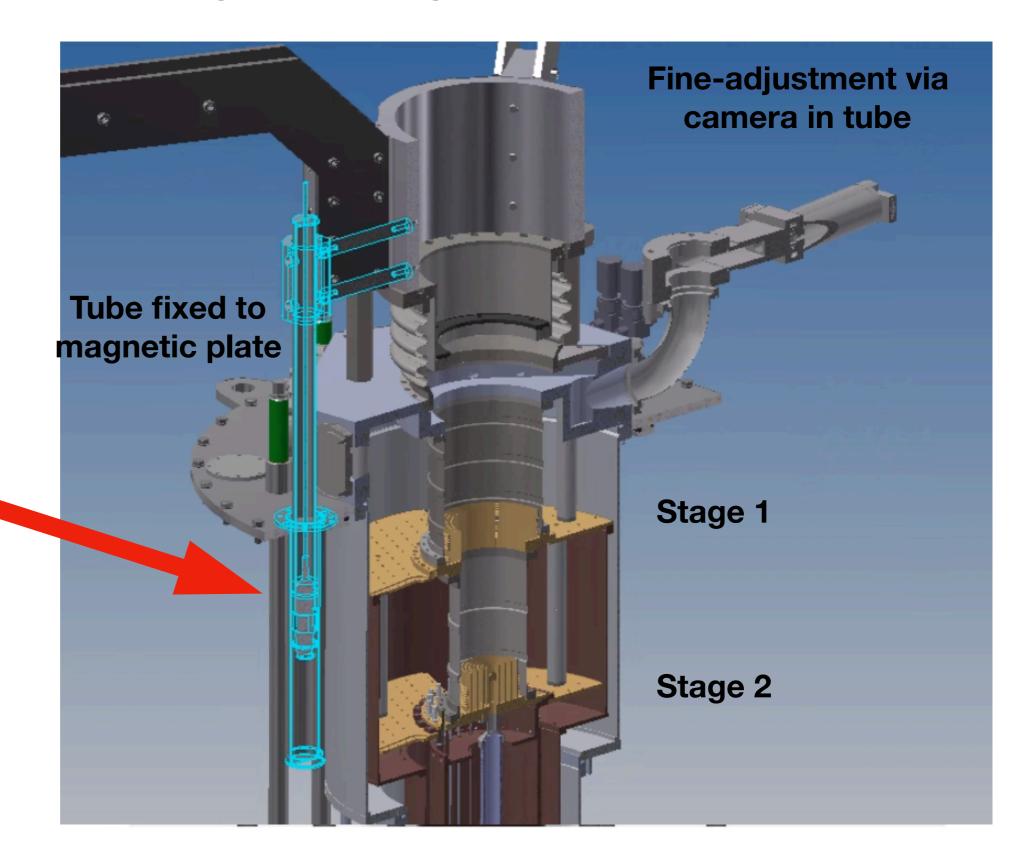
#### **Aluminum prototype**







### Interior: Finalized design for alignment

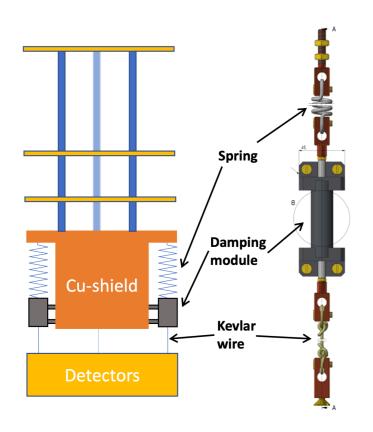




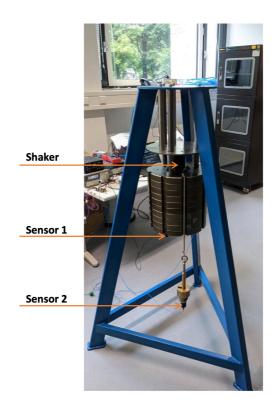
**USB Camera** 



### Decoupling at the detector stage



- Optimize spring & resonance frequency of damping modules
- Magnetic eddy current damping



Mockup system



### Cryostats: Visit to Cryoconcept in 11/2021





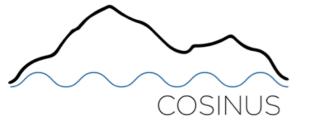


Dilution units



## **Summary**

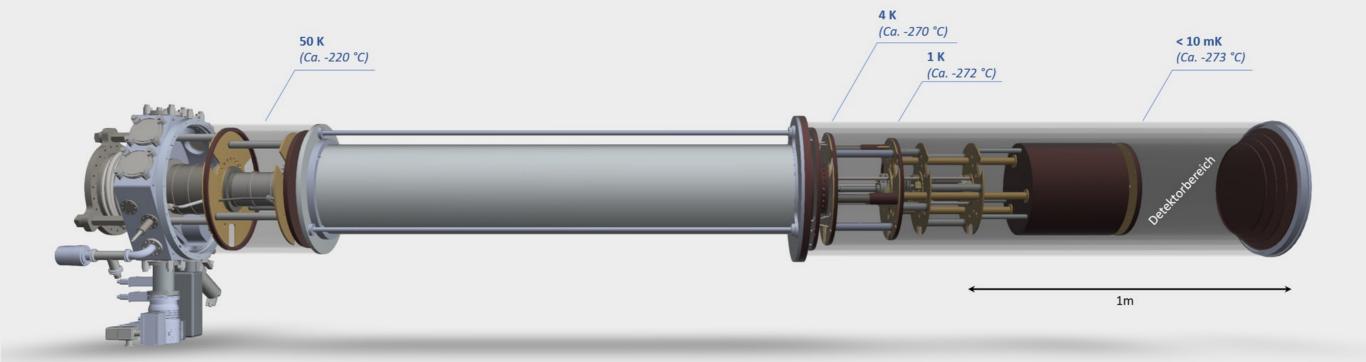
- COSINUS Facility construction has started at LNGS
- "remoTES" solves several problems
   -> approaching final detector design
- Cryostats will be delivered in March & June 2022
- Designs for water & air systems
- Prototype for shielding ready; production starts in January
- Construction of Drywell, Tripod and Lifting System almost done
- Group at MPP growing!

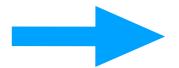


- Building a new experiment of this size would not be possible without the help and support of a lot of people
- Many thanks especially to the CRESST group, the mechanical and electronics workshop, the construction team, the procurement department, D.Werner and the directors!

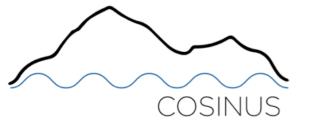








www.cosinus.it



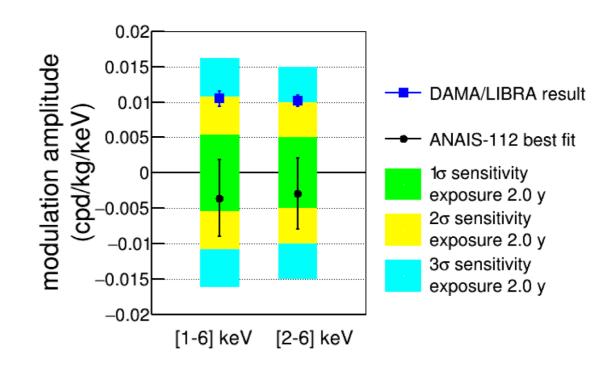
### Thank you!

### Backup: COSINE-100 and ANAIS modulation results

Govinda, Adhikari, and COSINE-100 collaboration. "Dark Matter Searches with the COSINE-100 Experiment." *Journal of Physics: Conference Series*. Vol. 1468. No. 1. IOP Publishing, 2020.

COSINE-100, 2.51 yrs 99.7% C.L. **Preliminary** Amplitude (cpd/keV/kg) 95.5% C.L. 68.3% C.L. 3σ DAMA region **Expected Standard Halo** 0.02 50 150 200 250 100 300 350 Phase (Days)

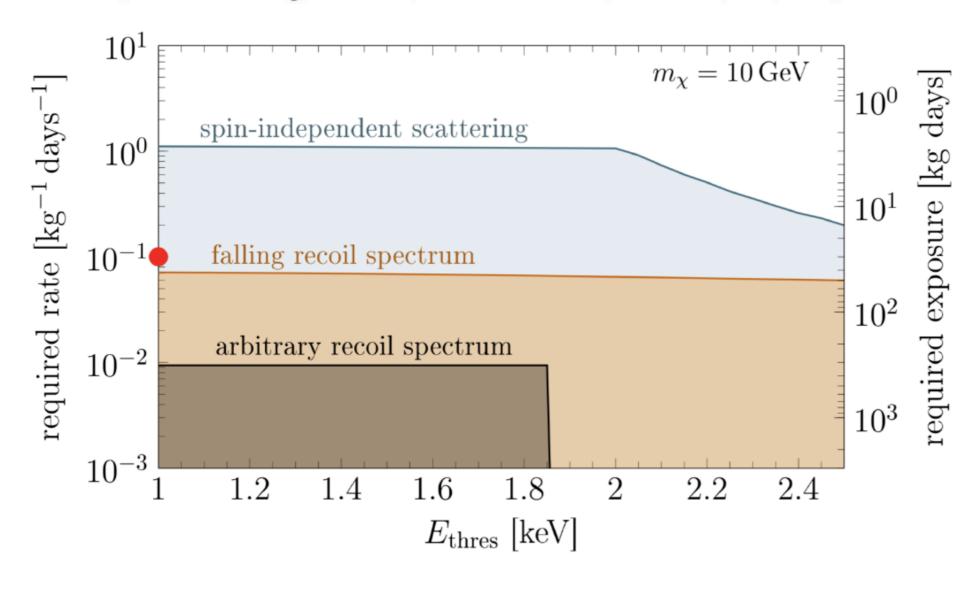
Amaré, J., et al. "ANAIS-112 status: two years results on annual modulation." *Journal of Physics: Conference Series*. Vol. 1468. No. 1. IOP Publishing, 2020.



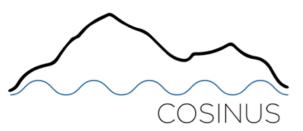


### Backup: Model-Independent Tests

F. Kahlhöfer, K. Schmidt-Hoberg, K. Schäffner, F. Reindl and S. Wild , JCAP 1805 (2018) no.05, 074



- Modulation amplitude cannot exceed total rate!
- No assumption on dark matter halo
- For model-independent cross-check would need ~1.8keV nuclear recoil threshold, ~300kgd exposure

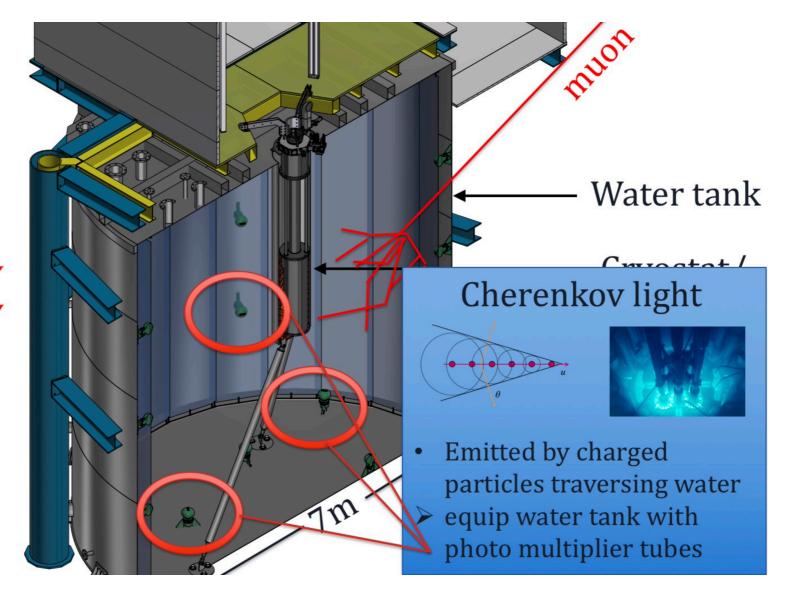


### Backup: Muon Veto

Muons create secondary particles: most dangerous are neutrons  $R \simeq 4 \text{ cts} / (\text{kg year})$ 

→ Too high: Veto mandatory

Neutrons from radioactive processes:  $R \simeq 0.05 \text{ cts} / (\text{kg year})$ 



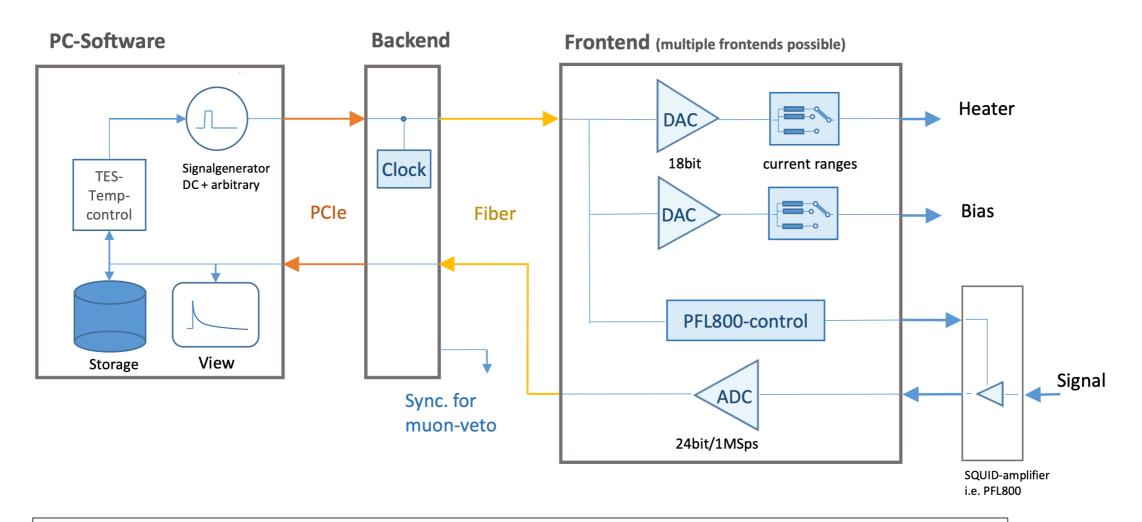
- From the start or not?
- Technical specifications of veto clear; design mostly finished



### **Backup: DAQ Electronics**



- "Versatile Data AcQuisition" combining several electronic parts
- Developed at HEPHY in Vienna, two prototypes so far



#### **VDAQ-Devices:**

- Backend: Ready available PCIe-FPGA-card
- Frontend: 19"- crate. Plug-in cards for analog out (heater, bias) and analog in (signal).

