



ComPol

A Compton Polarimeter in a Nano Satellite

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TOPICS

- Scientific motivation
- Detector system
- Working principle
- Sensitivity study
- Hardware status

Scientific motivation

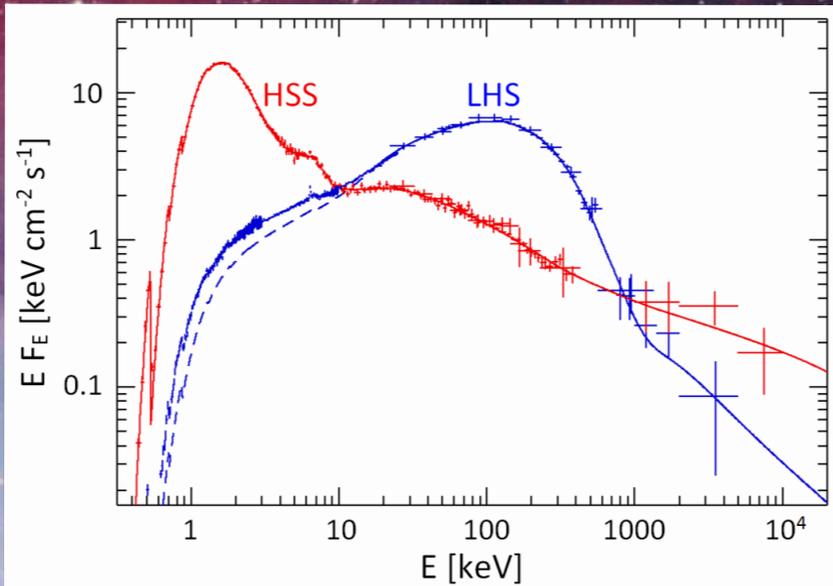
Some facts:

- Black hole binary
 - Black hole ($16 M_{\odot}$)
 - Blue supergiant star ($27 M_{\odot}$)
- Distance: 1.8 kpc
- **Strong X-ray source**

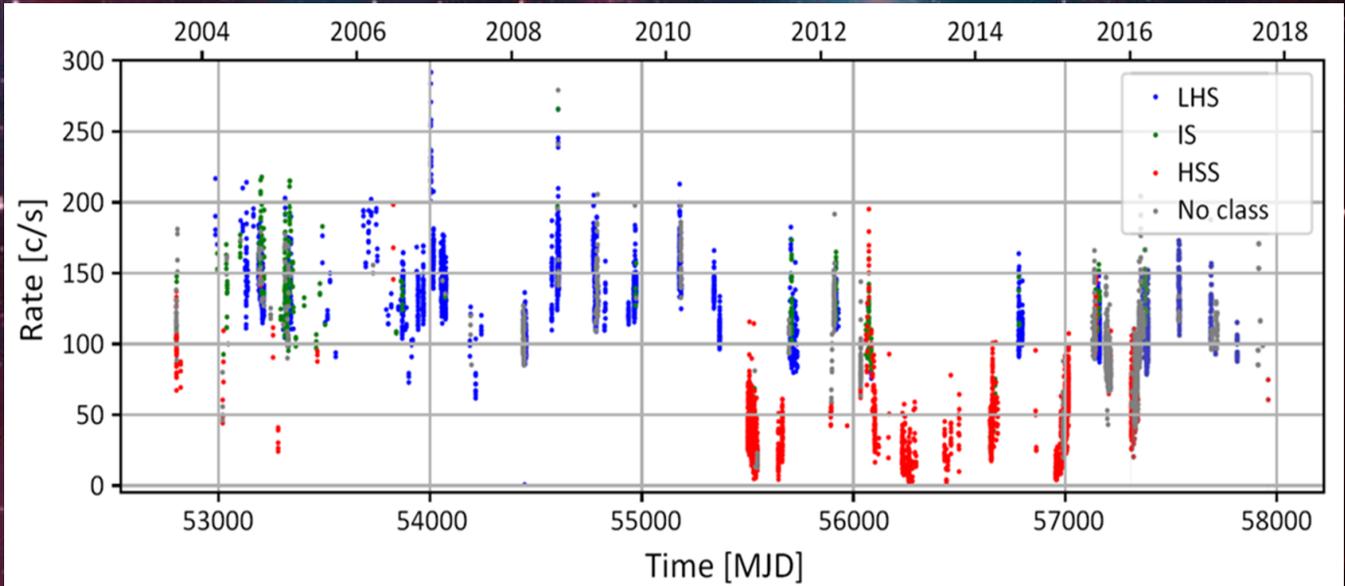


Scientific motivation

Different emission states of Cygnus X-1



Paredes, Josep M. et al. "Gamma Rays from Compact Binary Systems." AIP Conference Proceedings (2008)



F. Cangemi et al., "High energy spectral study of the black hole Cygnus X-1 with INTEGRAL", SF2A, 2018

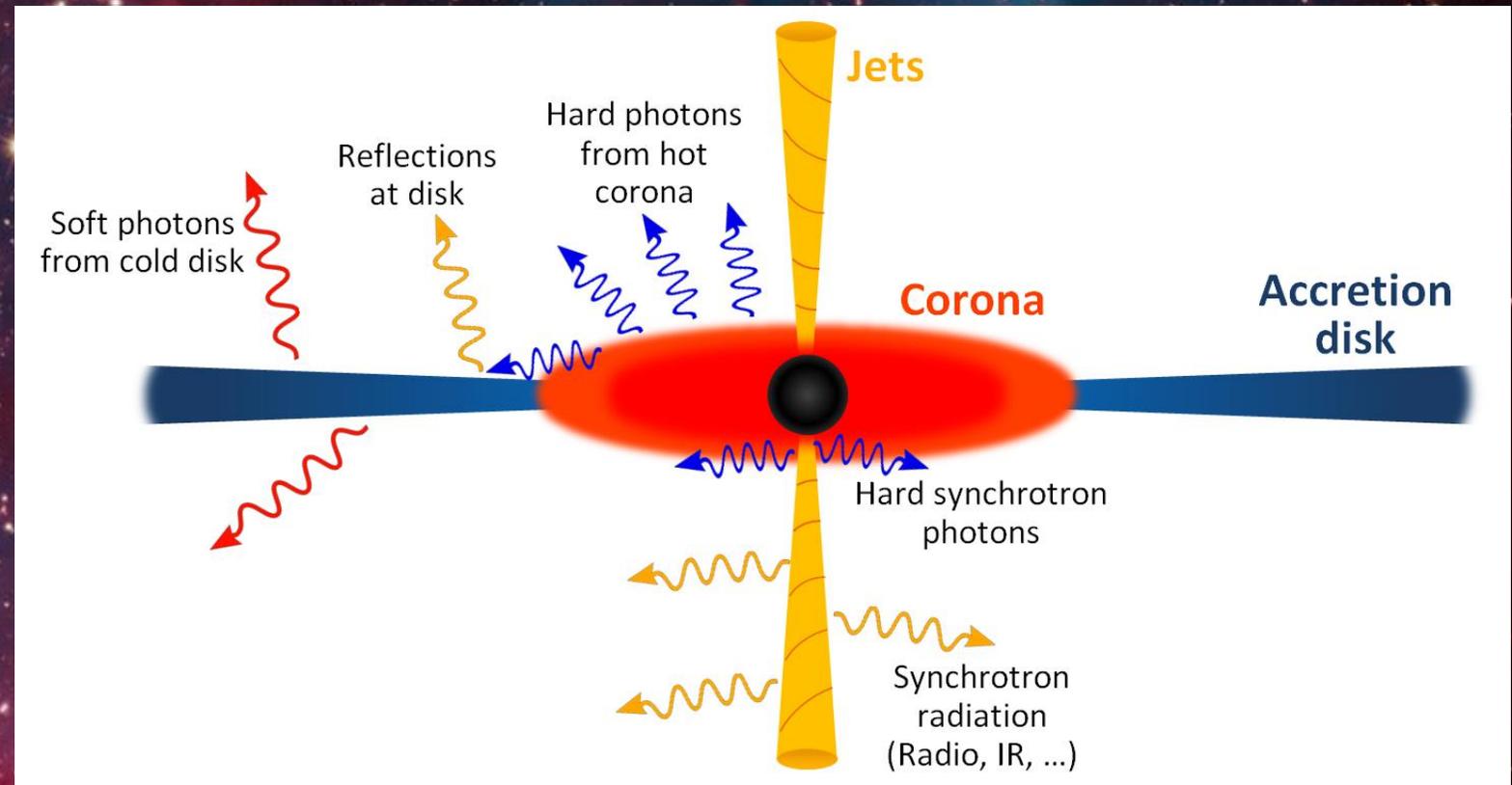
→ Continuous observation necessary

Scientific motivation

Origins of hard X-rays?

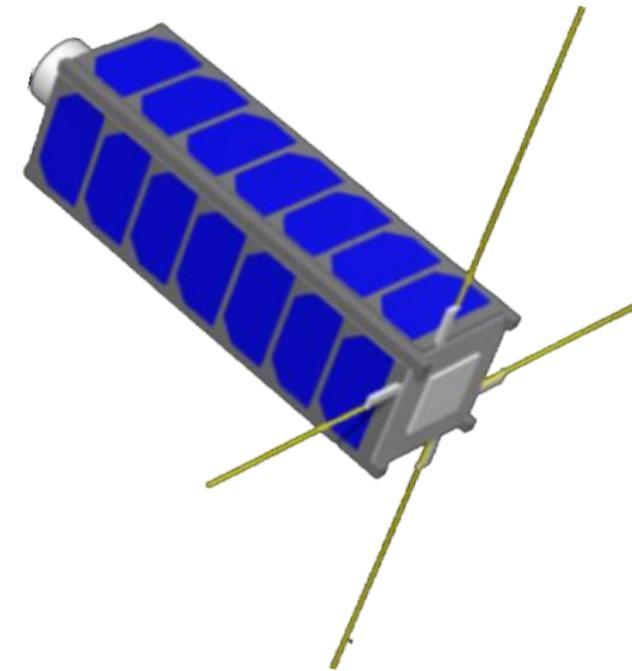
- Inverse Compton in Corona
- Synchrotron radiation (in disk/jets)

→ Polarimetry!

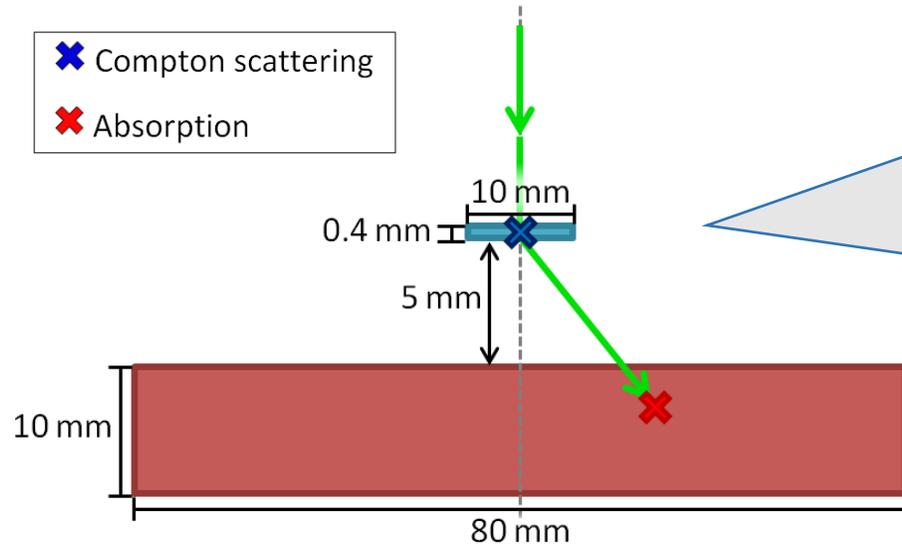


ComPol project (Compton Polarimeter)

- CubeSat mission
3U (10x10x34 cm³)
- Energy range 20 – 200 keV
largely unexplored
- Continuous, long-term pointing at Cygnus X-1
1 year
- Part of the ORIGINS LRSM
Laboratory for Rapid Space Missions



Detector system



- ✕ Compton scattering
- ✕ Absorption

Silicon Drift Detector (SDD)

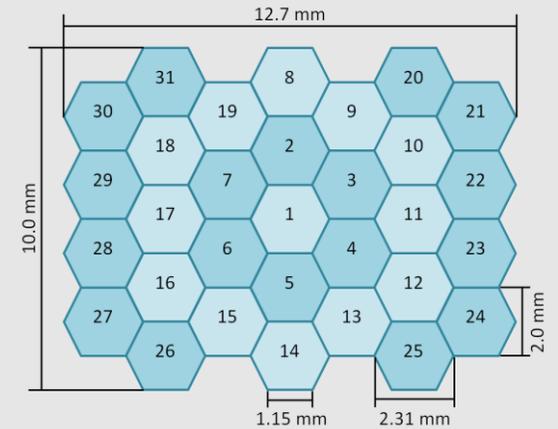
- Used for KATRIN/TRISTAN (Neutrino mass experiment)
- Excellent energy resolution (300 eV @ 20 keV)



MAX PLANCK INSTITUTE FOR PHYSICS

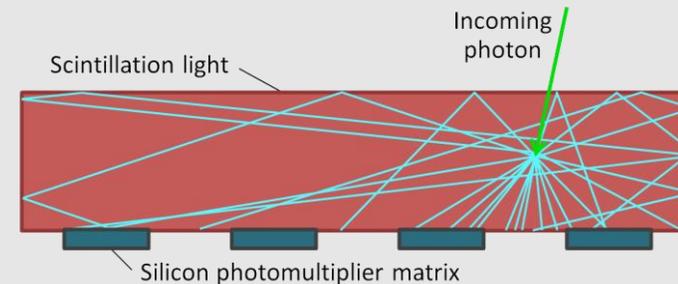


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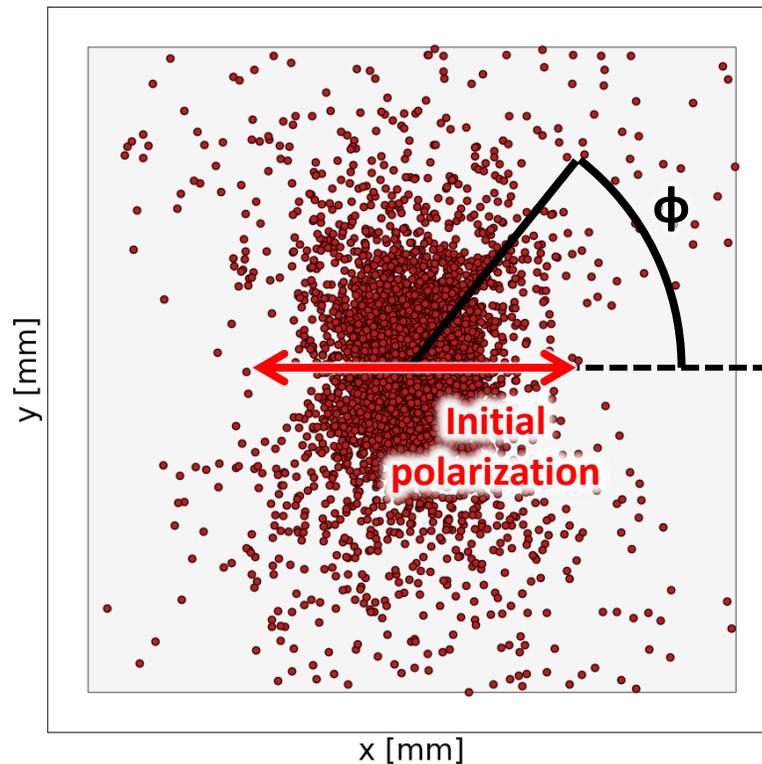
Calorimeter

- Scintillating crystal: CeBr_3
- Silicon photomultiplier matrix on the backside

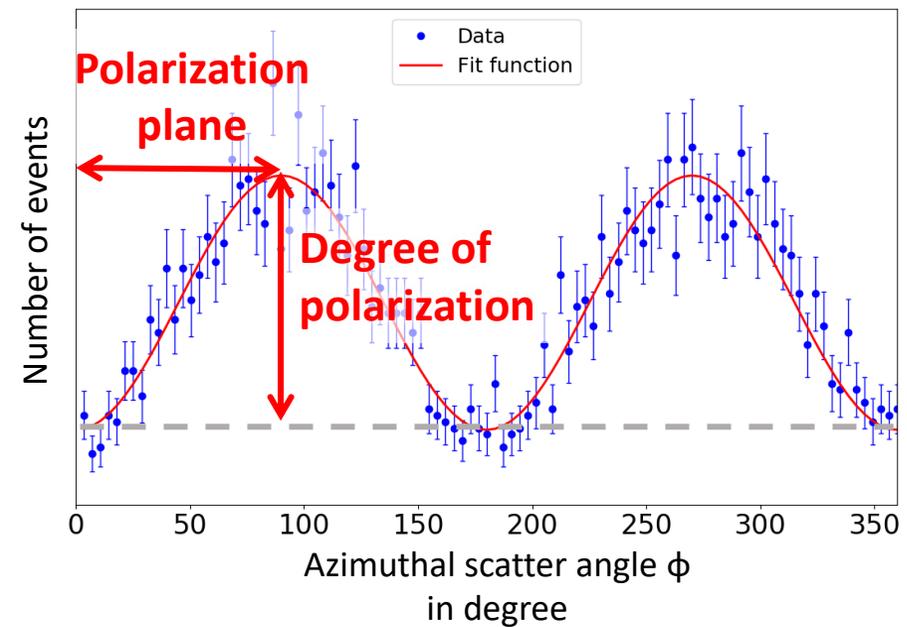


Working principle

Scatter plot in second detector plane



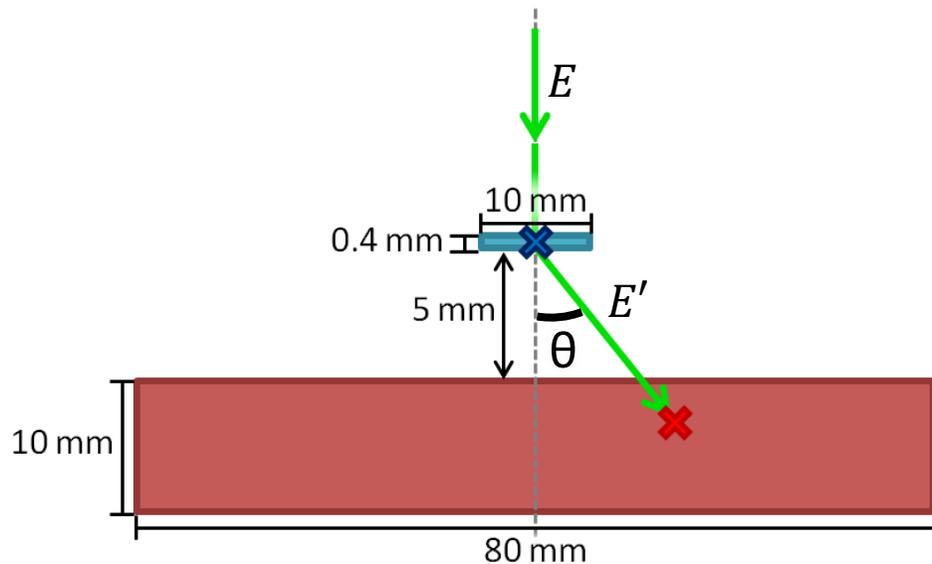
Histogram of the azimuthal scatter angle ϕ



Event reconstruction

Two ways to determine the scatter angle θ

1) via interaction points



2) via energy deposits

$$\Delta\lambda = \lambda' - \lambda = \frac{h}{m_e c} (1 - \cos\theta)$$

$$E = \frac{hc}{\lambda}$$

$$\Rightarrow \cos\theta = 1 - m_e c^2 \left(\frac{1}{E'} - \frac{1}{E} \right)$$

energy after
scattering
 $E' = E_{Cal}$

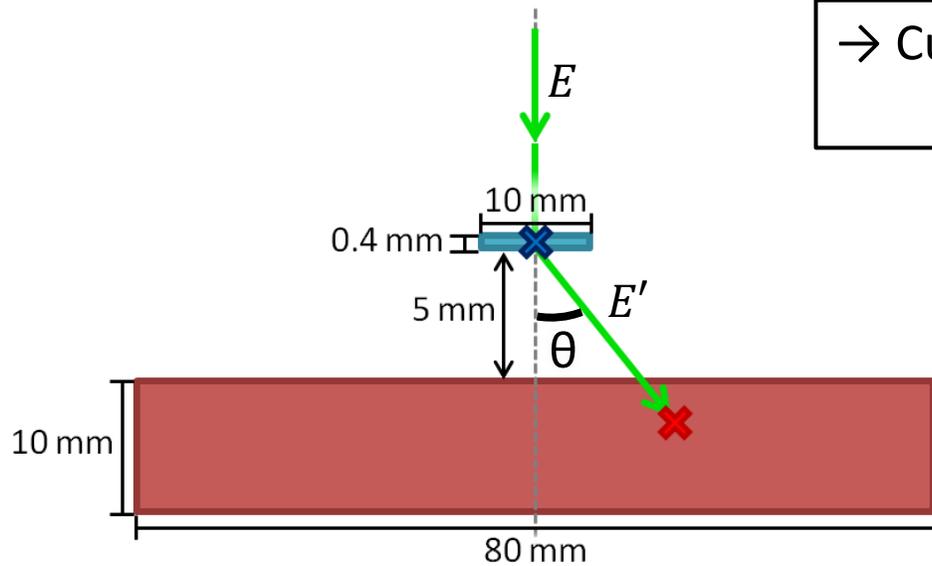
initial energy
 $E = E_{Si} + E_{Cal}$

Event selection

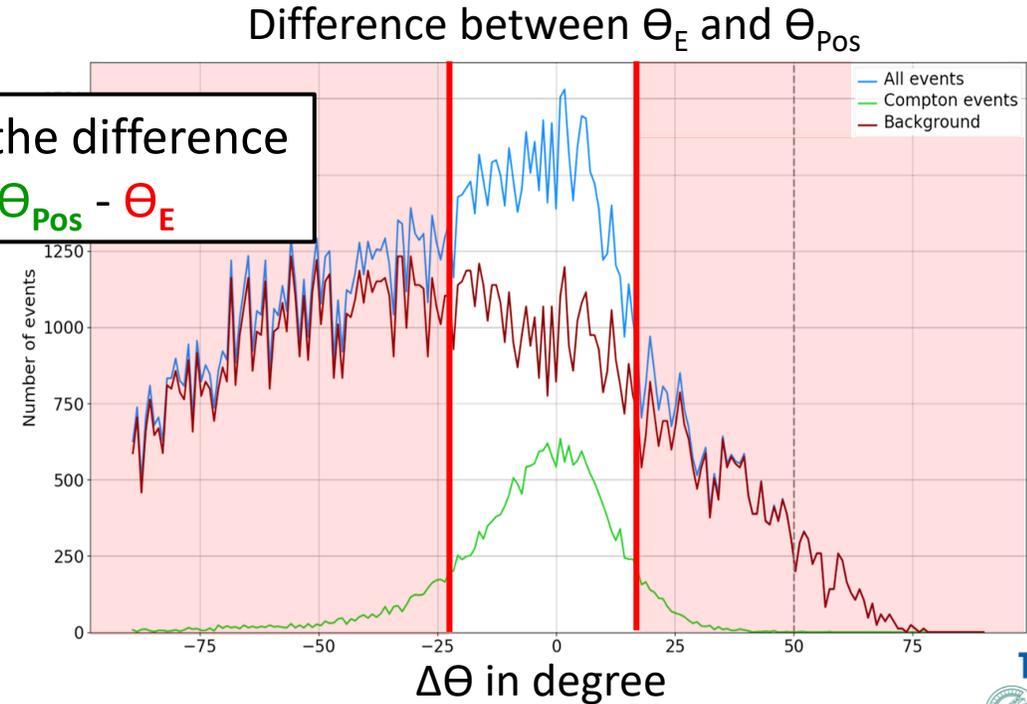
Two ways to determine the scatter angle θ

1) via interaction points

2) via energy deposits



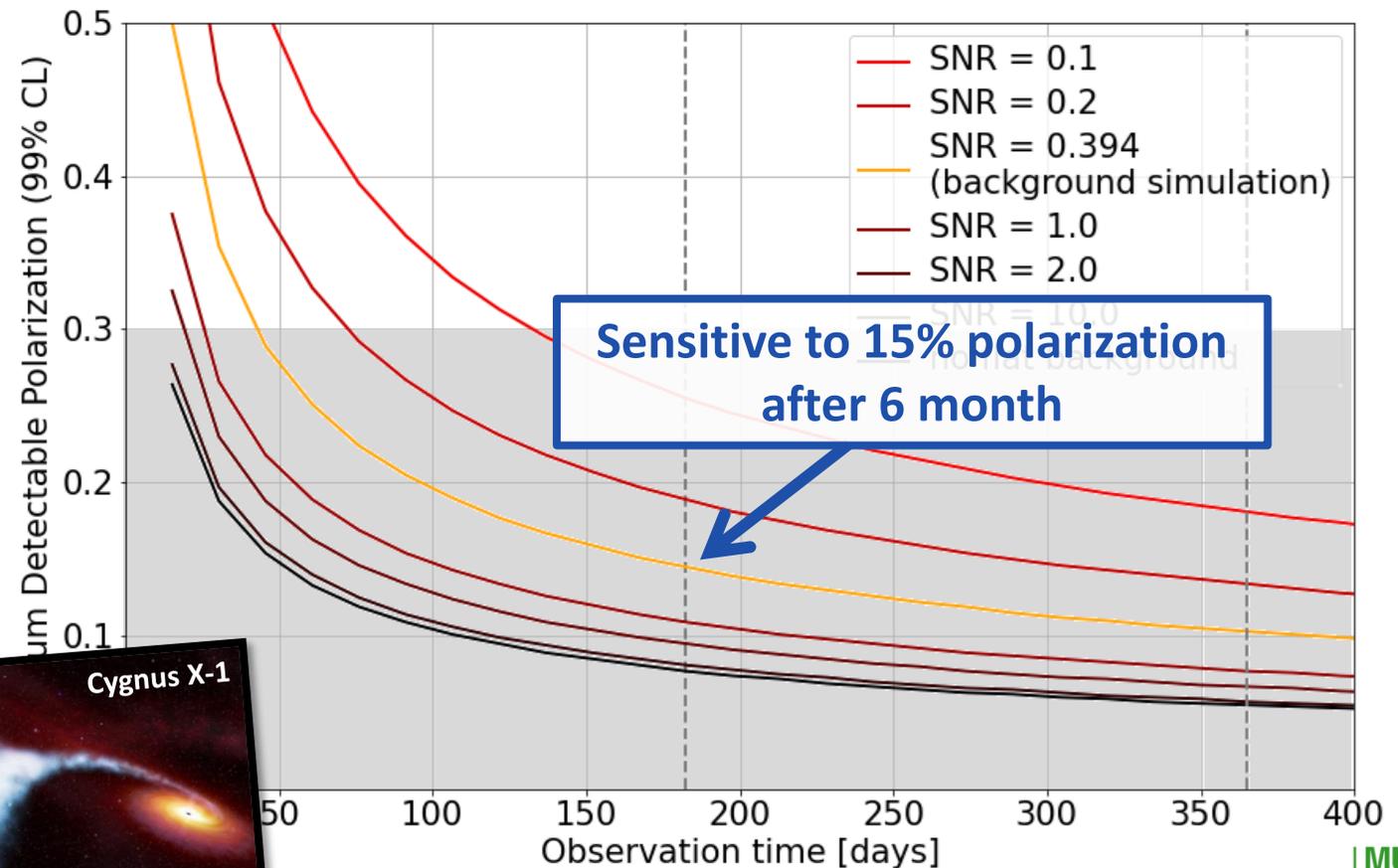
→ Cut on the difference
 $\Delta\theta = \theta_{\text{Pos}} - \theta_E$



Sensitivity studies with Geant4

Simulated effects:

- Real spectra
 - Cygnus X-1
 - Background:
 - γ , e^- , e^+ , n , p^+ , α
 - + Cosmogenic activation
- Whole detector system + Surrounding materials
- Detector effects:
 - Position resolution
 - Energy resolution
 - Energy thresholds



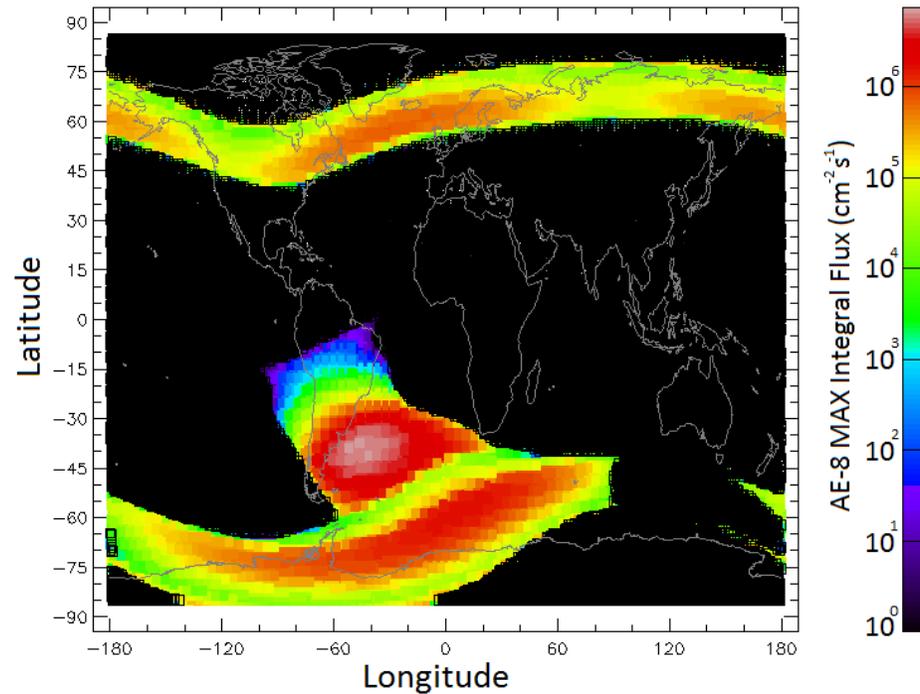
Background study

See talk by
C. Glas

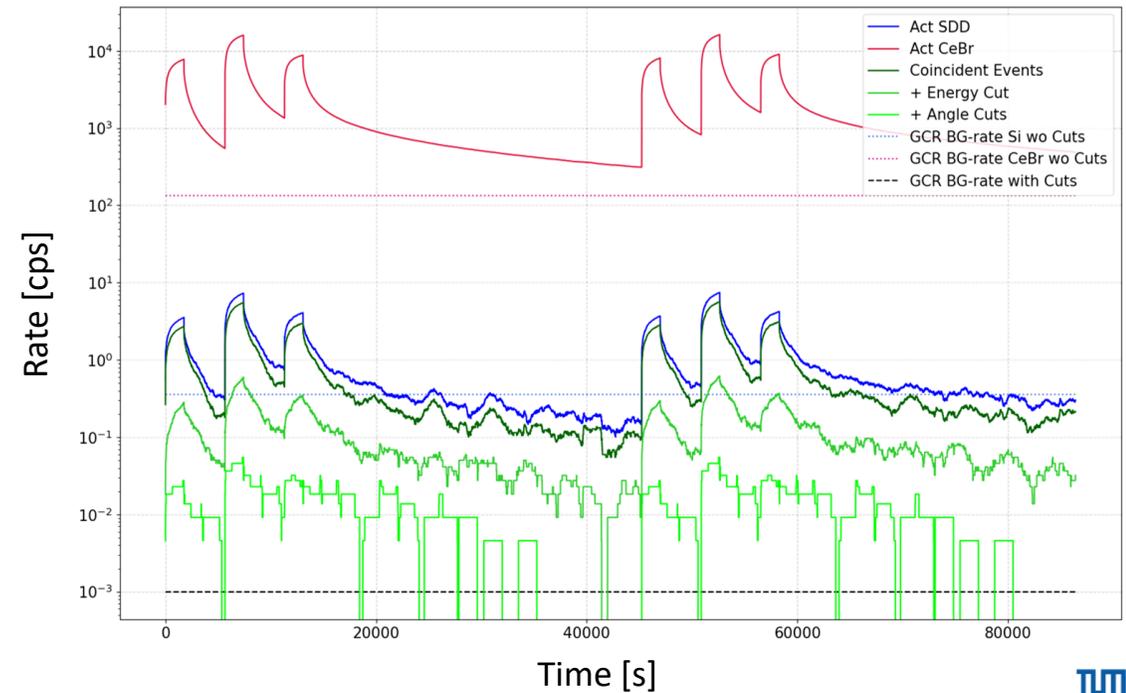


Shielding strategy and Cosmogenic Activation

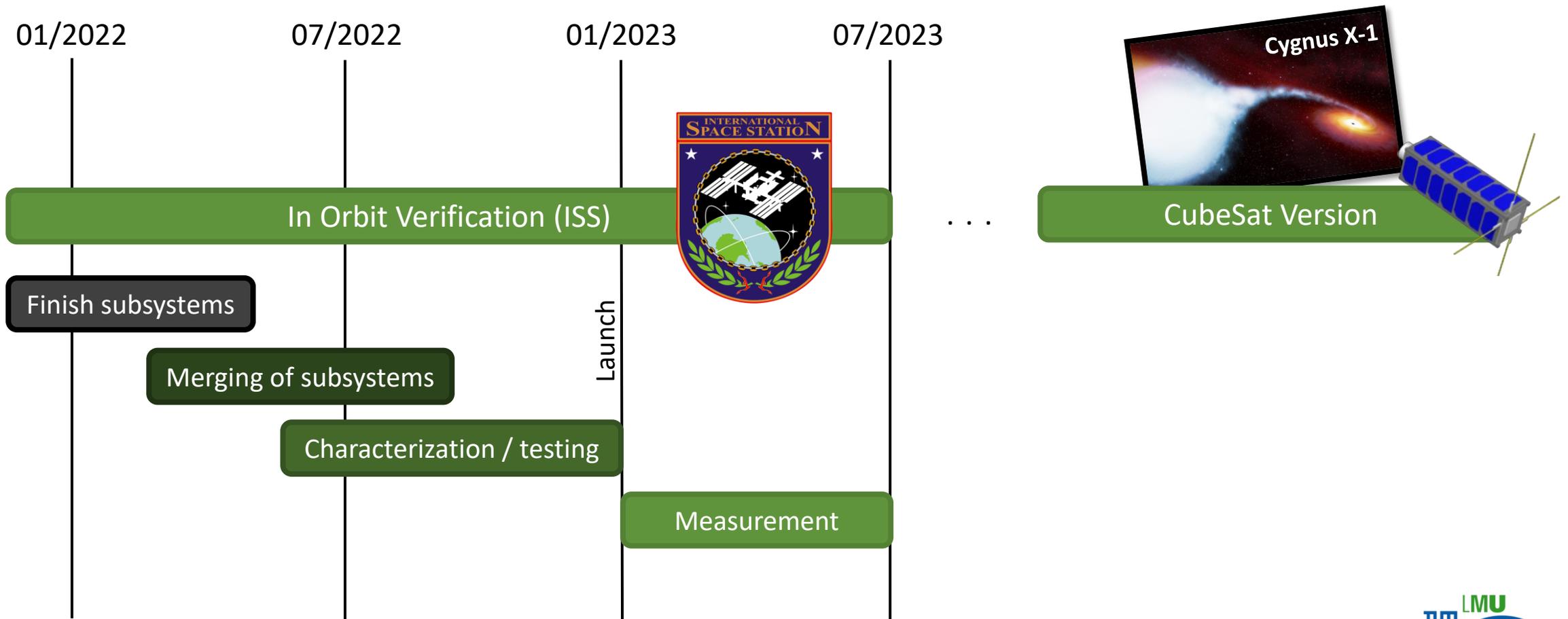
South Atlantic Anomaly



SAA Influence over time

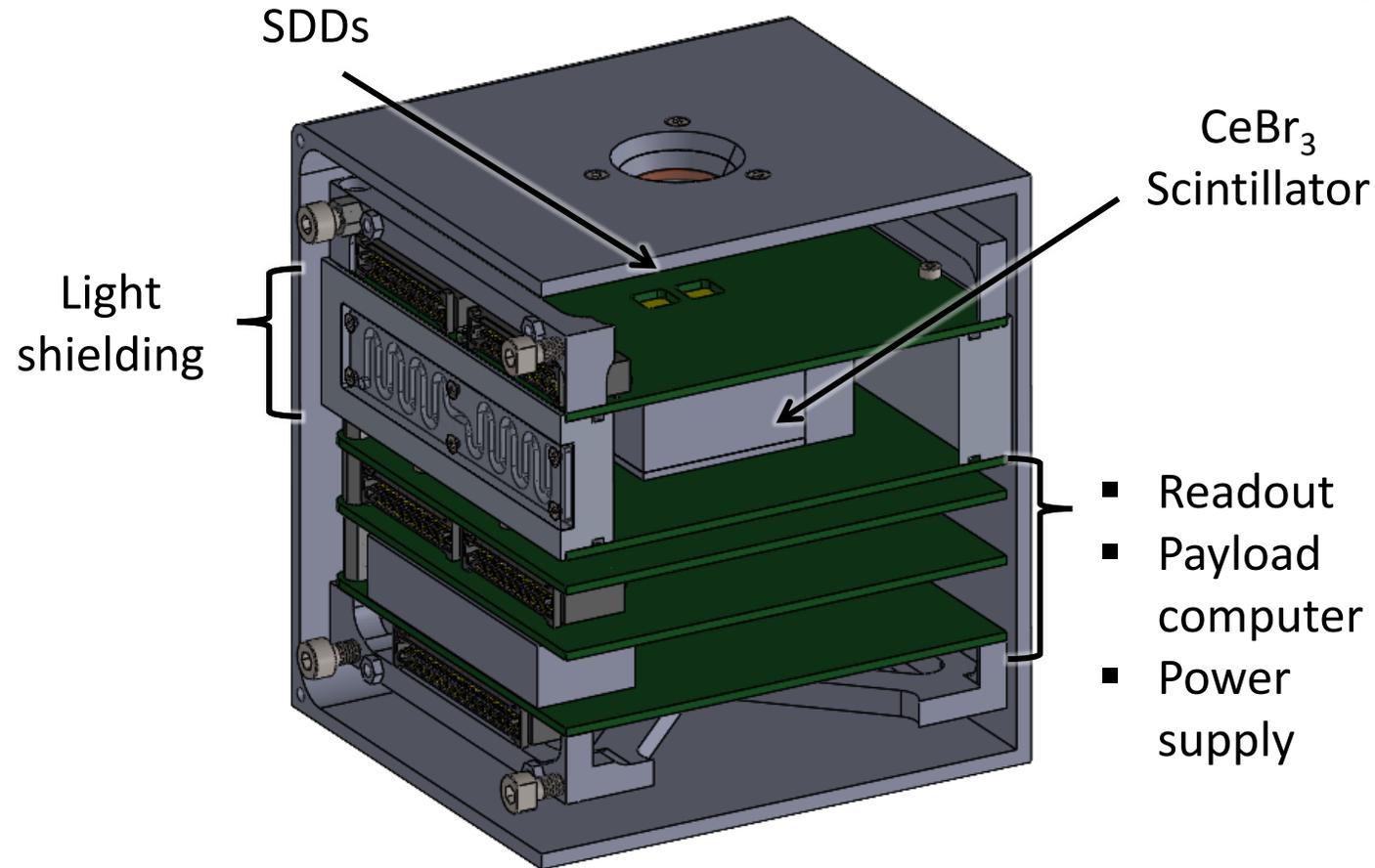


Project schedule



Hardware ISS version

See talk by
K. Geigenberger

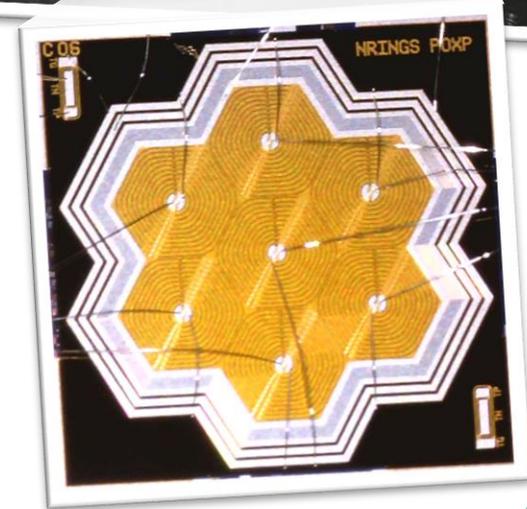
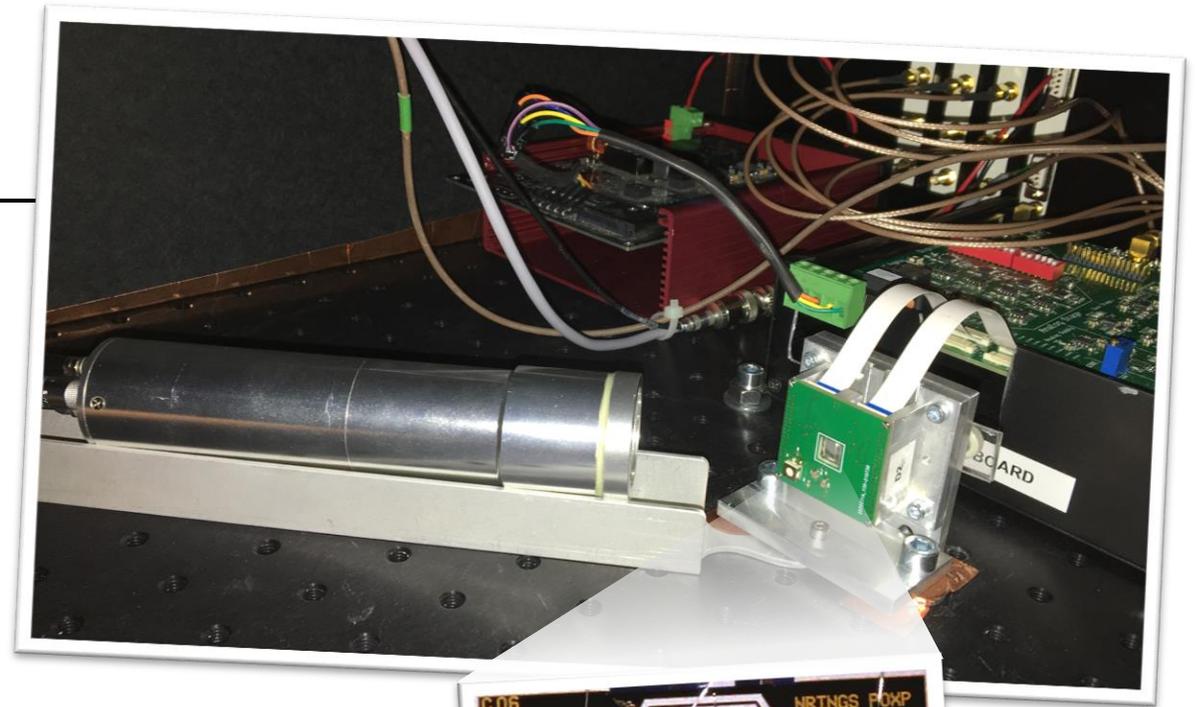
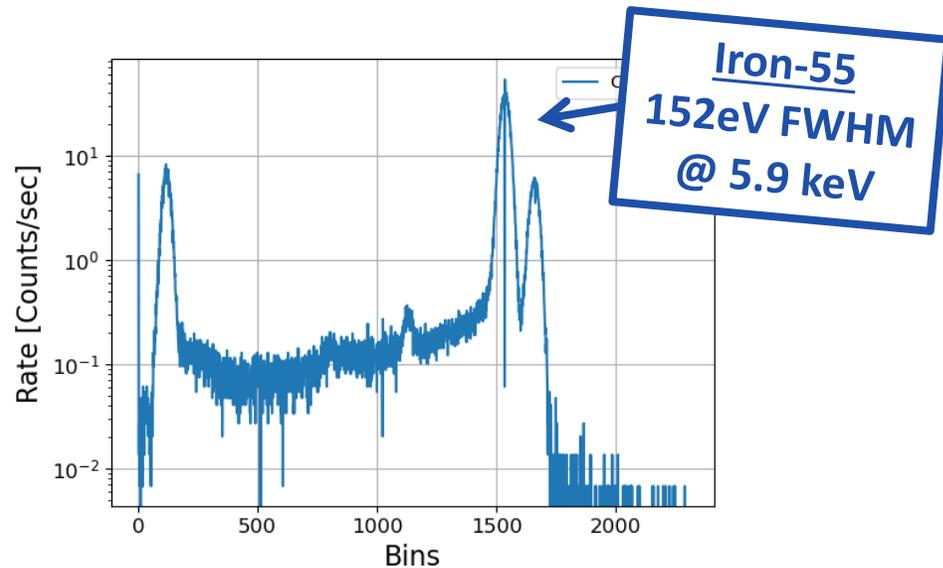


Hardware ISS version

First prototype boards are ready!

Achieved:

- First operation of HLL SDD with SFERA ASIC
- Good performance verified



ComPol project (Compton Polarimeter)

CubeSat mission

- 3U CubeSat \rightarrow (10x10x34 cm³)
- IOV mission: 2023
- CubeSat Launch >2023

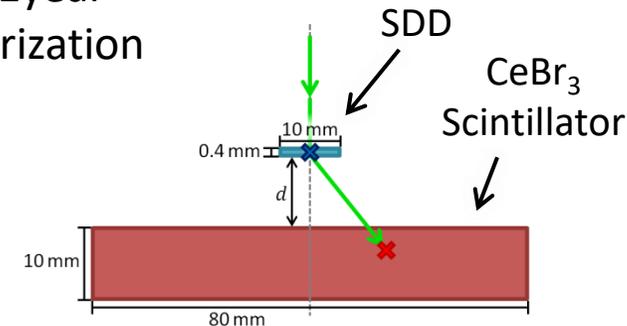
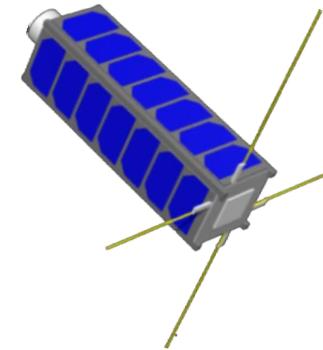


Observe Cygnus X-1

- 20 keV – 2 MeV
- Continuous for >1year
- Spectrum & polarization

Detector system

- Compton telescope
- Stacked detector system



\rightarrow Minimum Detectable Degree of Polarization $\leq 15\%$



Thank you for your attention!

