

# The CONUS experiment and future potential of coherent neutrino scattering

**Manfred Lindner**

*On behalf of the CONUS Collaboration*



**Current Topics in Astroparticle Physics**

Nov. 9-11, 2022

Max-Planck-Institut für Physik, Munich



# Georg and I

## Students at MPP in the early 80's

Georg: neutrinos...     $\leftrightarrow$  I: BSM, colliders, ....

- very good friends
- together teaching assistants for lectures
- training of young group members: supervisors & self-organization



# Back to Munich

Georg: MPP     $\leftrightarrow$  I: TUM

- my interests changing towards neutrinos...  
CP violation in  $\nu$  oscillations  
power of reactor  $\nu$  experiments, ...
- discussions, joint seminars, joint activities:  
→ ISAPP school on astroparticle physics 2006  
→ SFB375, TR27, IMPRS, ...  
→ 2002 Neutrino Conference

## Chairs:

F. v. Feilitzsch

N. Schmitz

## Heavily involved:

Georg

M. Altmann

ML

...others



Max-Planck-  
Institute für  
Physik und  
Astrophysik



# 2006: Moving on...

Georg: MPP     $\leftrightarrow$  I: MPIK, Heidelberg

- **ISAPP summer school on astroparticle physics @MPIK**
  - 2011: lectures by Georg
  - 2019: lectures by Georg
- ...
- **my interests kept expanding:**
  - dark matter: XENON + theory
- **But neutrinos remain a top theme:**
  - theory ...
  - Double Chooz
  - Gerda / LEGEND200
  - STEREO
  - CONUS

# Coherent v Scattering

Z-exchange of  $\nu$  with nucleus

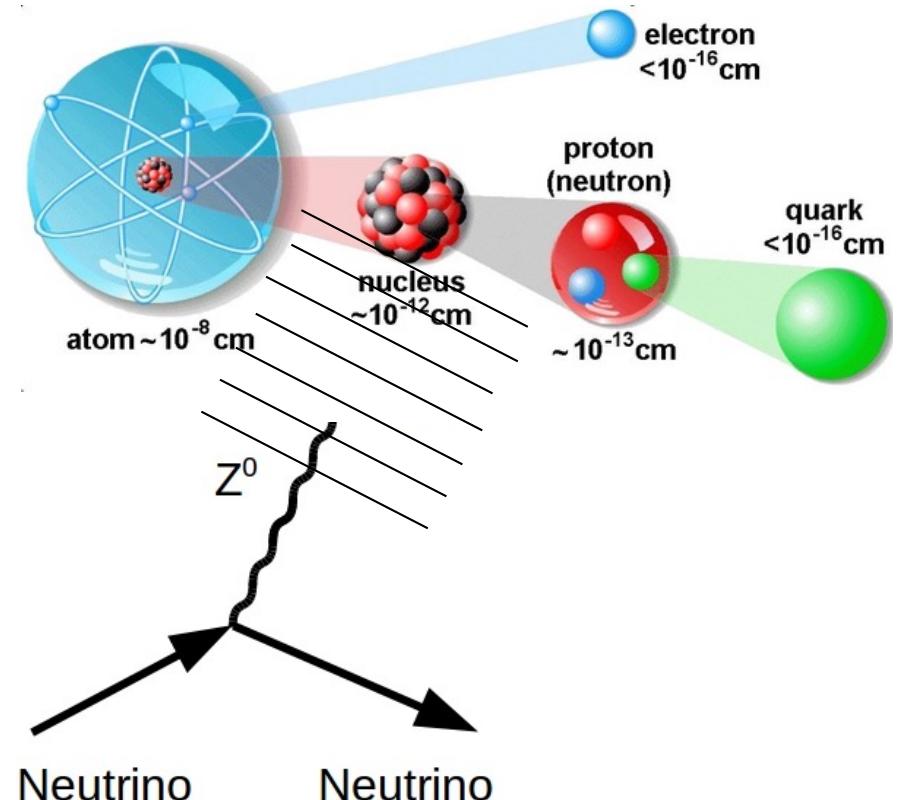
$$Q_w = N - (1 - 4 \sin^2 \theta_w)Z \sim N$$

→ mostly neutrons

momentum ↔ wavelength

Very low momentum

→ nucleus recoils as a whole



Important: Coherence length  $\sim 1/E \rightarrow E_\nu$  below O(50) MeV

→ low energy  $E_\nu \leftrightarrow$  lower cross sections → very high flux!

$$\frac{d\sigma(E_\nu, T)}{dT} = \frac{G_f^2}{4\pi} Q_w^2 M \left(1 - \frac{MT}{2E_\nu^2}\right) F(Q^2) \sim N^2$$

$N \simeq 40 \rightarrow N^2 = 1600 \rightarrow$  detector mass 10t → few kg

# Different experimental Paths

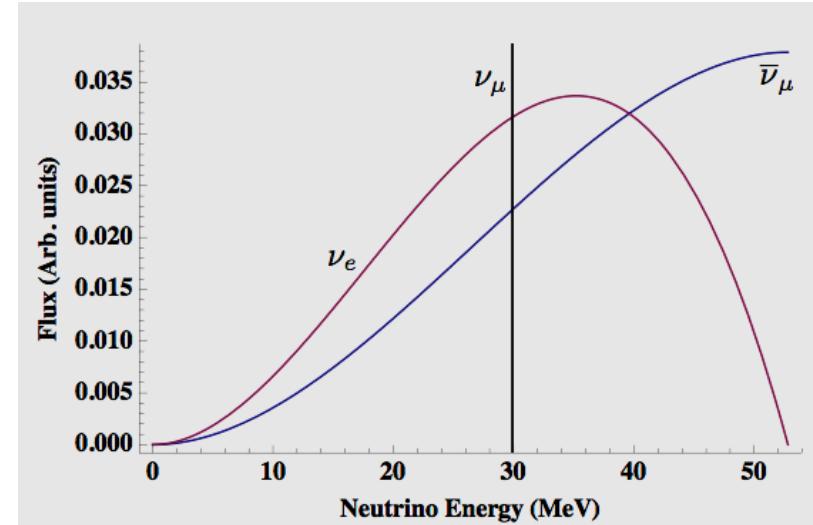
## Low energy $\nu$ 's from accelerators:

- $\pi$ -decay-at-rest (DAR)  $\nu$  source
  - different flavors produced
  - relatively high recoil energies
- form factor & uncertainty
- **1st observation by COHERENT**

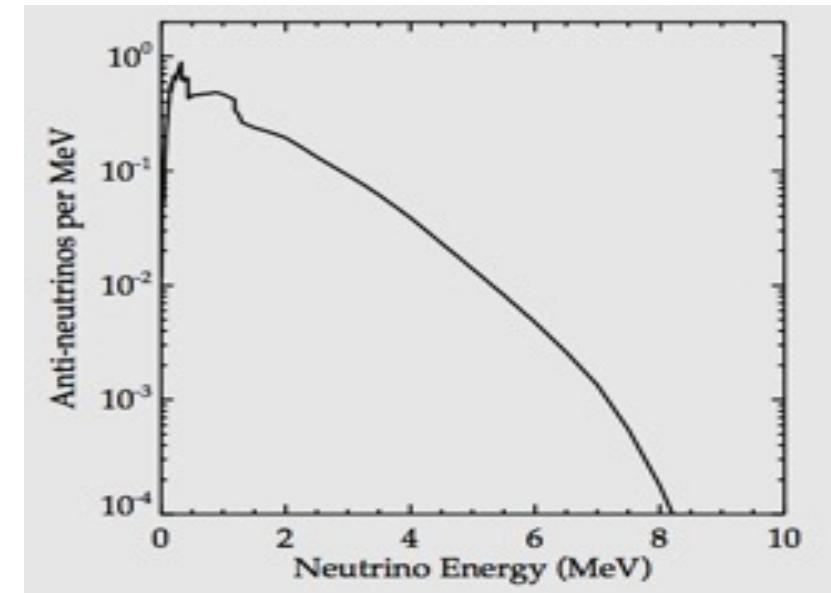
## Reactors:

- lower  $\nu$  energies than accelerators
  - lower cross section → higher flux
  - different flavor content implications for probes of new physics
  - fully coherent ~no form factor
- **CONUS, ...**

## Other: Solar $\nu$ 's @ XENONnT, ...



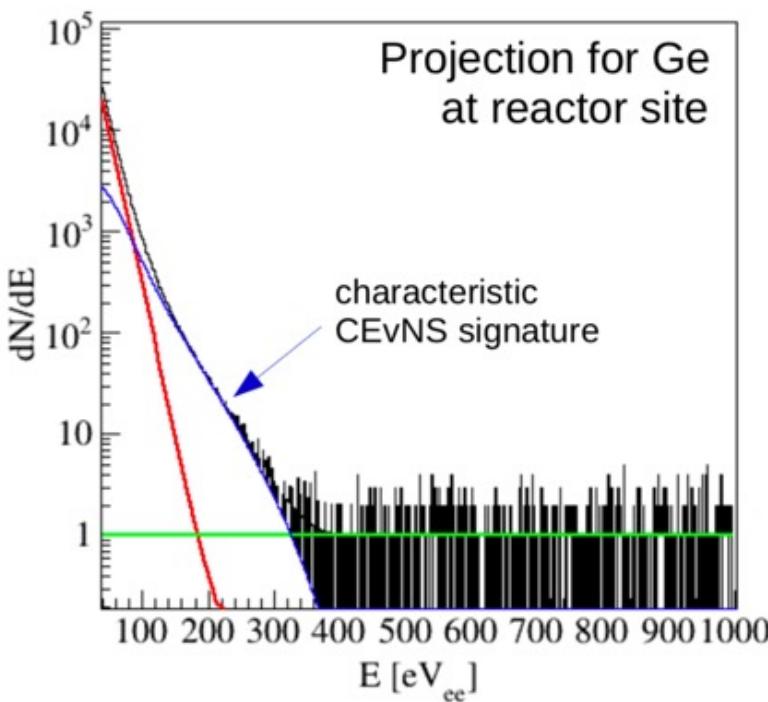
Anderson et al., 1201.3805



# Experimental Requirements

- measure nuclear recoil energy  $T$   
for  $E_\nu = 10 \text{ MeV} \rightarrow T_{\max} \approx 3 \text{ keV}$  (in Ge)
- energy loss due to quenching (Lindhard)  
 $\rightarrow$  Quenching Factor (QF) at low energy  
 $\rightarrow$  include QF uncertainties

$$T_{\max} \approx \frac{2E_\nu^2}{m_n(N+Z)}$$



## requirements of CEvNS detection:

- **very low background**
  - radio-pure materials
  - “virtual depth” shielding
- **low noise threshold (sub keV) + mass**
- **very high  $\nu$  flux**

# The CONUS Experiment

Combine:

- highest neutrino flux → close to power reactor
- lowest detection threshold → on-going R&D
- best background suppression → “virtual depth”



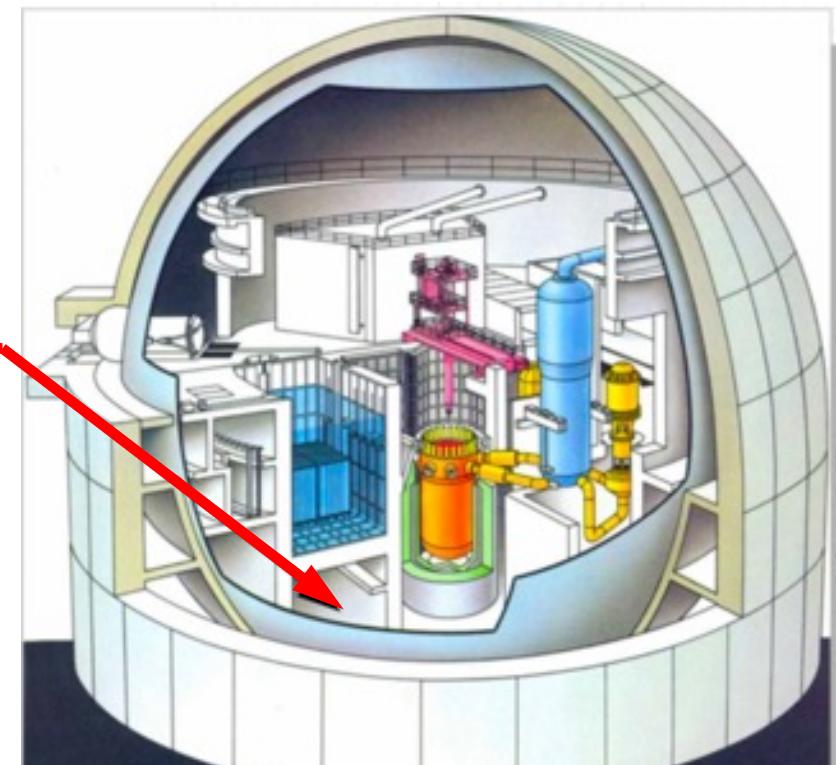
→ COherent NeUtrino Scattering experiment



# The CONUS Reactor Site

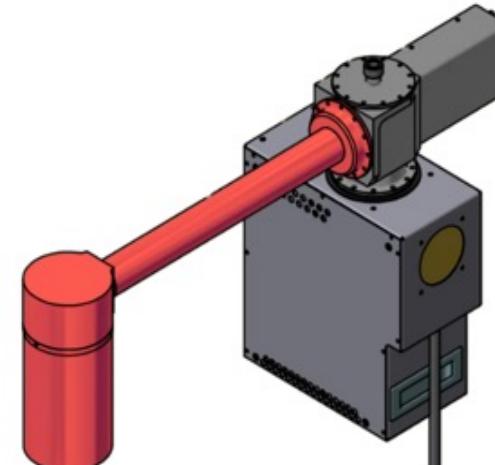
## Brokdorf (Germany) nuclear power plant:

- thermal power  $3.9 \text{ GW}_{\text{th}}$
  - detector @  $d=17\text{m}$
  - $\nu$  flux:  $2.4 \times 10^{13}/\text{cm}^2/\text{s}$
  - very high duty cycle
- very intense integral neutrino flux  
 $E_{\nu}$  up to  $\sim 8 \text{ MeV}$  → fully coherent
- overburden  $10 - 45 \text{ m.w.e}$
  - access during reactor operation
  - measurements of n bckgd with PTB
  - ON/OFF periods
    - bckgd. only measurement
    - seismic detector



# Detectors: CONUS 1-4

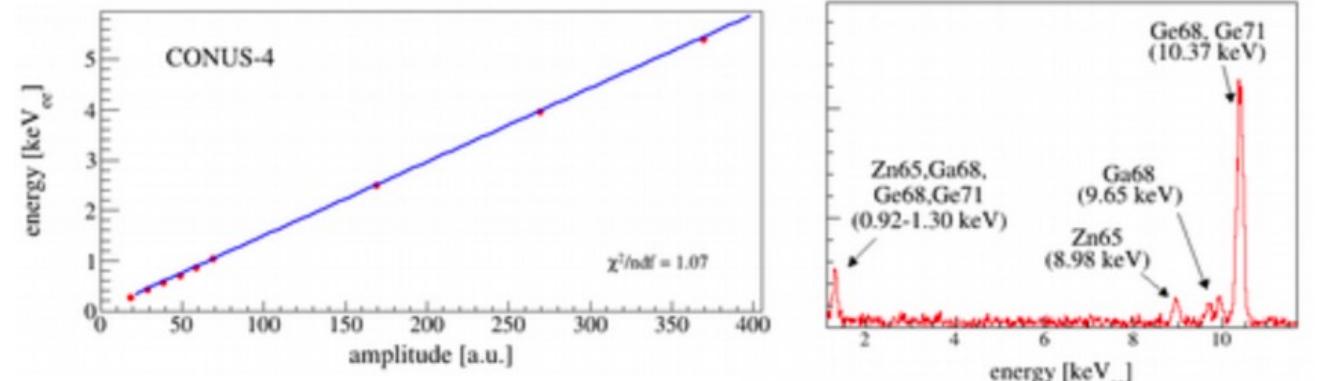
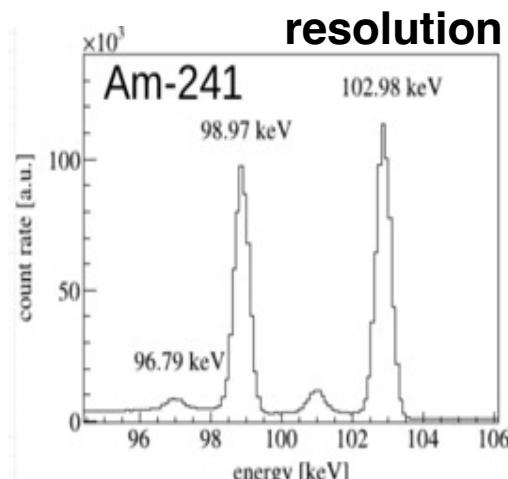
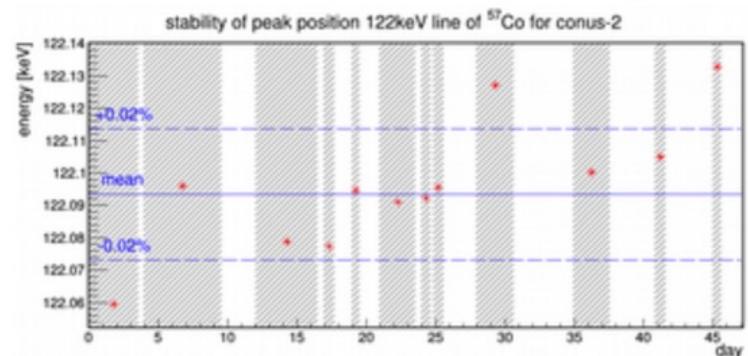
- p-type point contact HPGe
- 4x 1kg – **active mass 3.85kg**
- spec. for pulser res. (FWHM)  $\leq 85\text{eV}$   
→ noise threshold < 300eV
- **electrical PT-cryocoolers**
- ultra low background components
- close R&D collaboration with MIRION



| Detector | Pulser FWHM <sub>P</sub> [eV <sub>ee</sub> ] |
|----------|--|
| CONUS-1  | 69±1   |
| CONUS-2  | 77±1   |
| CONUS-3  | 64±1   |
| CONUS-4  | 68±1   |

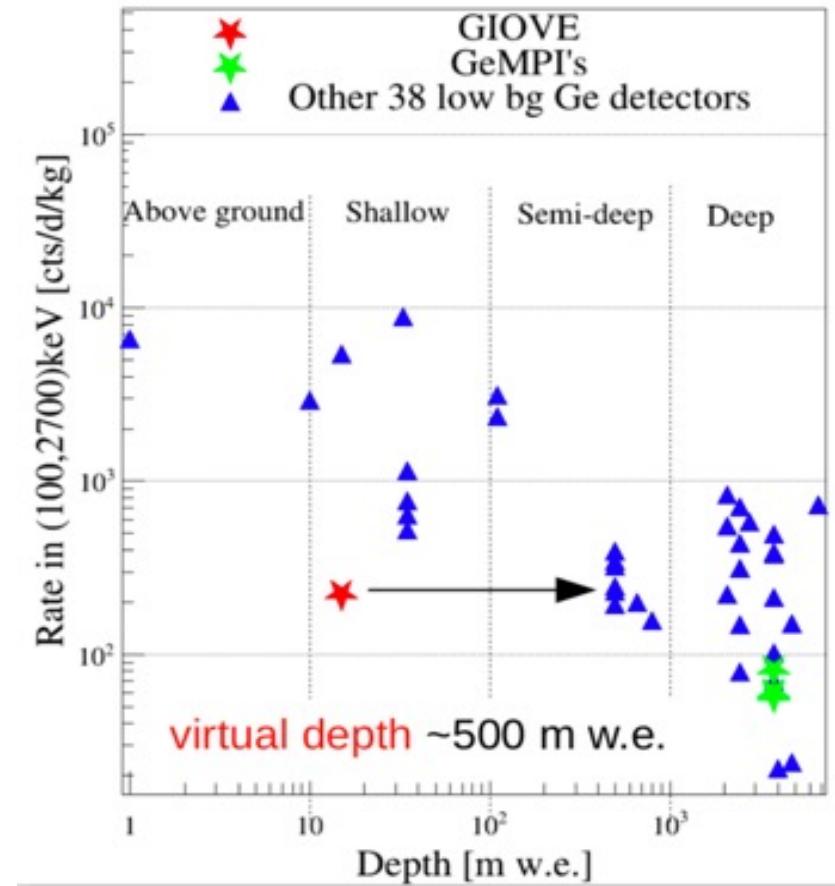
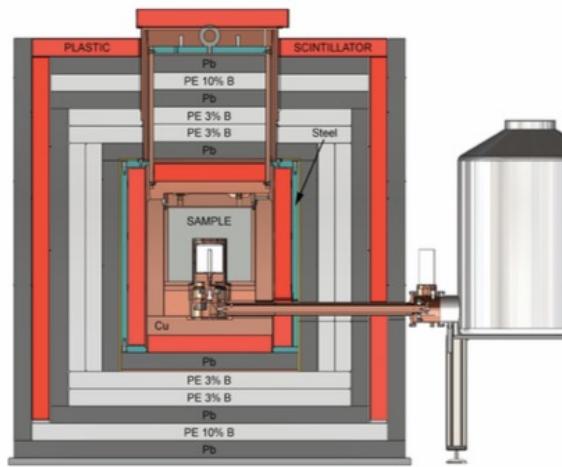
## Long term stability

Under lab. Conditions:  
stan. dev. of peak position:  
**+15eV (+-0.02%)**  
(within 45 days)



# ``Virtual Depth'': The GIOVE Shield

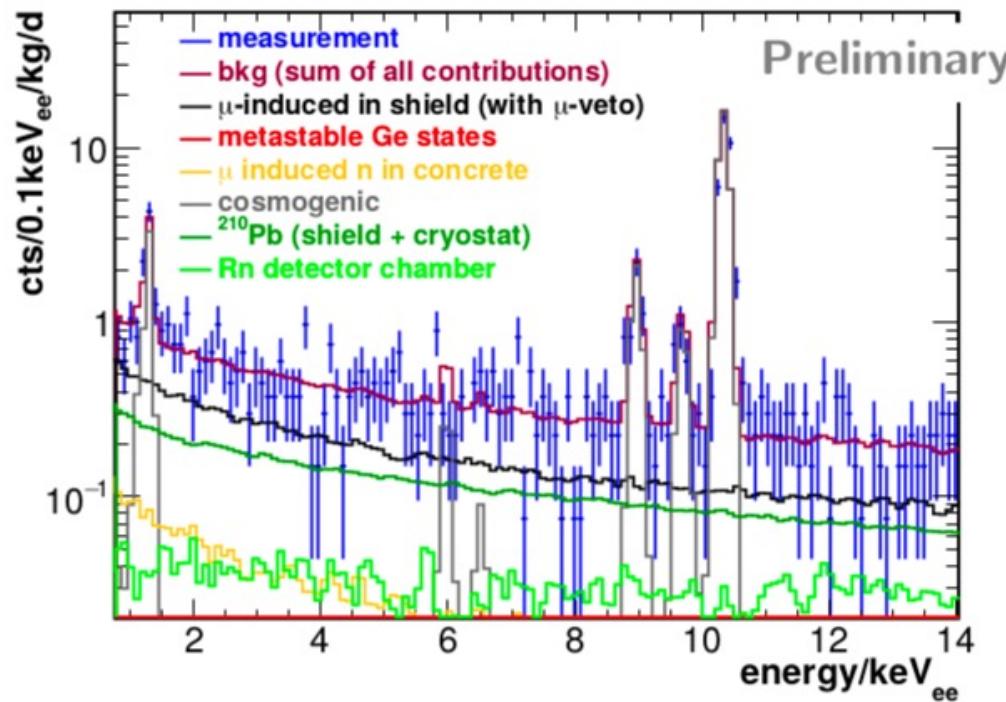
- R&D at MPIK
  - main purpose: material screening @ shallow depth (15 mwe)
  - coaxial HPGe detector ( $m_{act} = 1.8 \text{ kg}$ )
  - radio-pure passive shielding
    - Pb, B-doped PE,  $\mu$ -veto, OFHC Cu
  - active veto: optimized to reduce  $\mu$ 's and  $\mu$ -induced signals
    - plastic scintillators with PMTs
    - 99% muon veto efficiency (dead time  $\sim 2\%$ )
- ( $^{226}\text{Ra}$ :  $70\mu\text{Bq/kg}$ ,  $^{228}\text{Ra}$ :  $110\mu\text{Bq/kg}$ ,  $^{228}\text{Th}$   $50\mu\text{Bq/kg}$ )



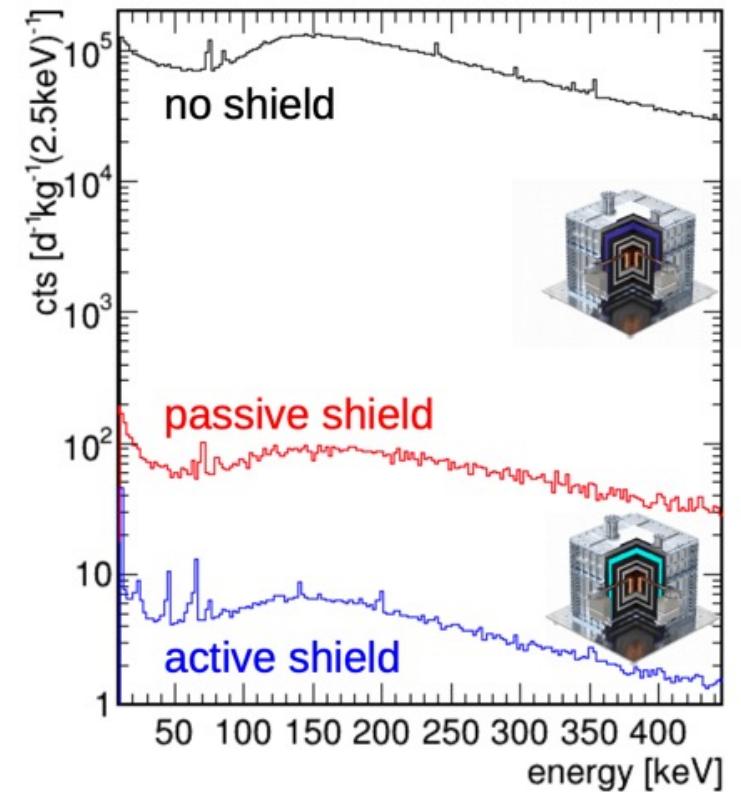
→ ``virtual depth''  
UG projects close to surface  
G.Heusser et al.,  
Eur.Phys.J. C(2015)75:531

# Background Suppression

- Suppression of external natural radioactivity and cosmogenic background by a **factor of  $10^4$**
- Residual background **fully described by MC simulations**  
→ stable background level in [0.5 – 1] keV<sub>ee</sub>:  
**10 counts/kg/d/keV<sub>ee</sub>**



Eur. Phys. J. C 79, 699 (2019)  
arXiv: 2112.09585 (2021)



# Installation @ Reactor

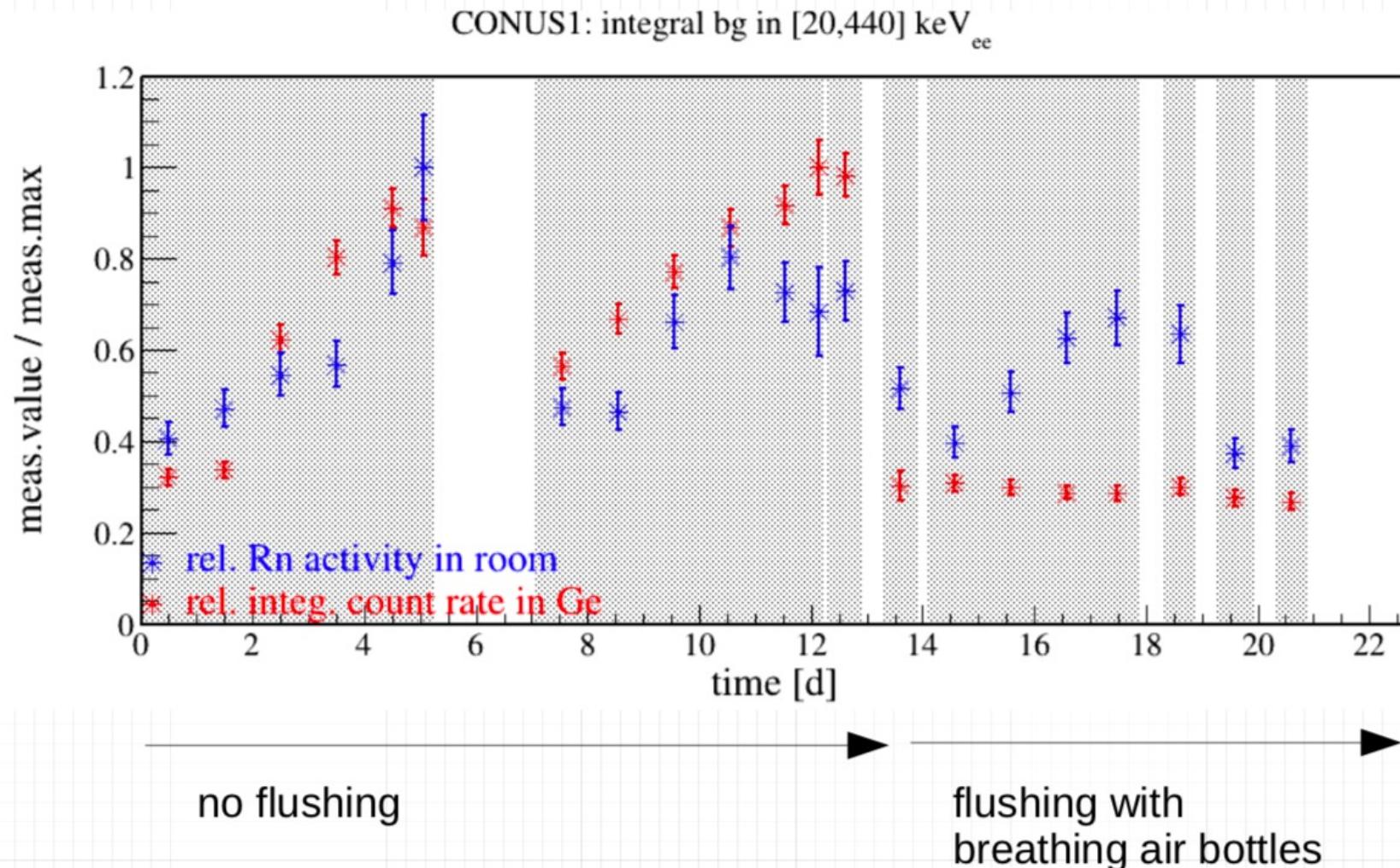
assembly at MPIK UG lab  
→ characterization  
→ commissioning

installation @ Brokdorf  
→ full assembly  
→ commissioning



# Radon Mitigation @ Reactor Site

radon at reactor site: closed room, thick concrete walls → 100-300 Bq/m<sup>3</sup>  
half-life of <sup>222</sup>Rn: 3.8d → counter measure @reactor site: N not allowed  
hermetical sealing + flush with aged breathing air bottles ~1 l/min



# Results

# Ionization Quenching Factor

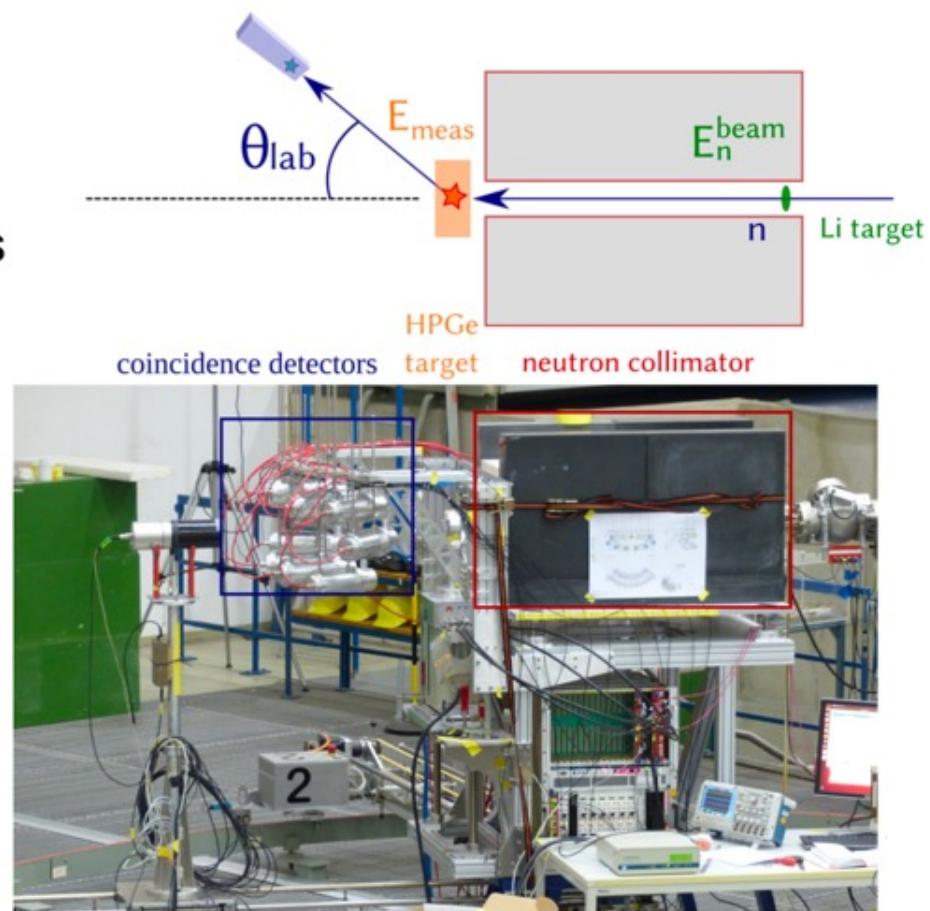
- Ionization quenching factor (IQF)
- Extensively measured from 10-100keV, data lacking in keV range → Conus ROI

$$IQF = \frac{E_{ionization}}{E_{nuclear\ recoil}}$$

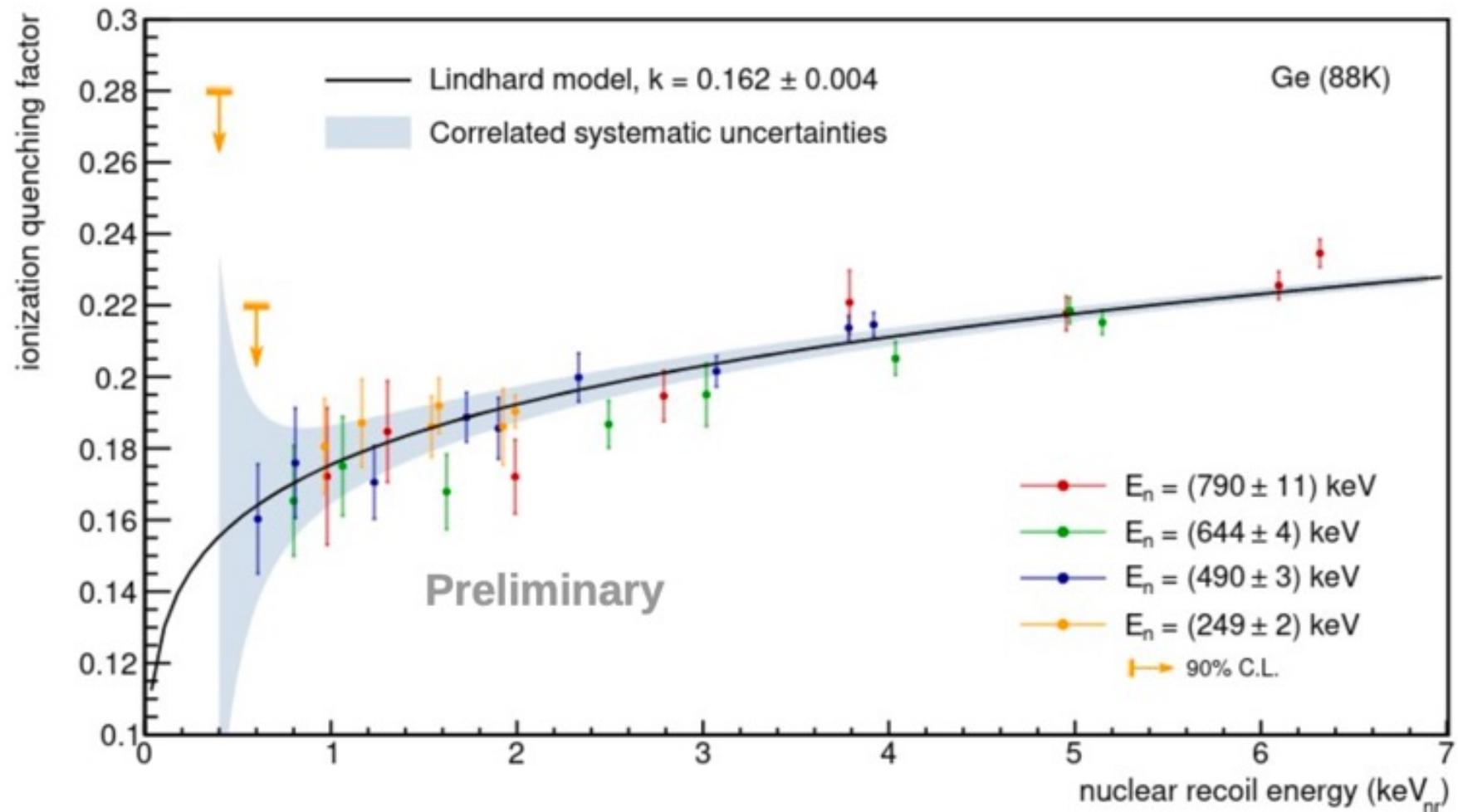
arXiv:2202.03754 (2022)

## Measurement of IQF in Ge:

- direct, model-independent using neutrons
- scientific cooperation with PTB:
  - pulsed proton beam
  - mono-energetic neutrons via  $\text{Li}(p,n)$  reaction
- Experimental setup:
  - neutron collimator
  - thin HPGe target
  - liquid scintillator (LS) array



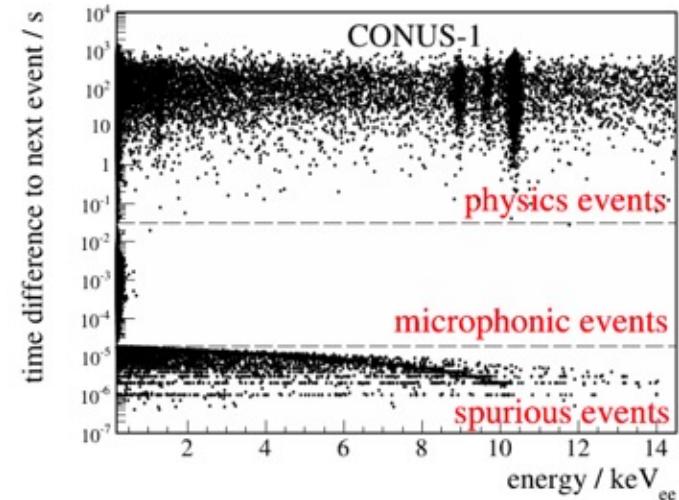
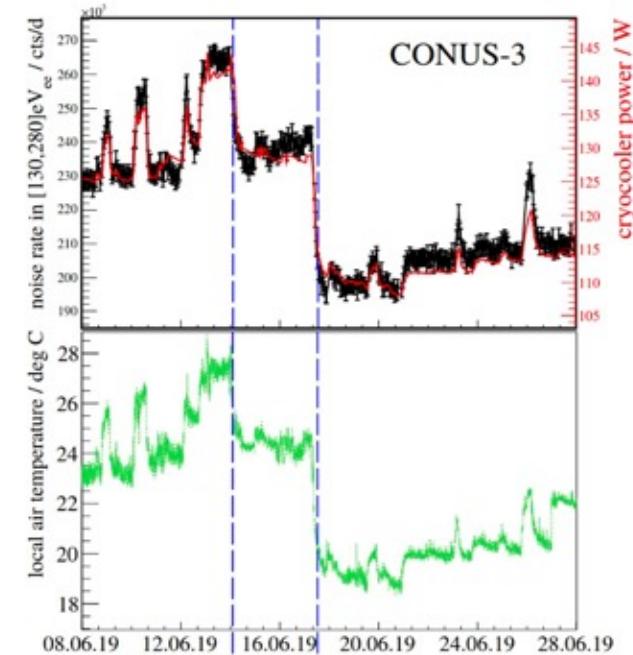
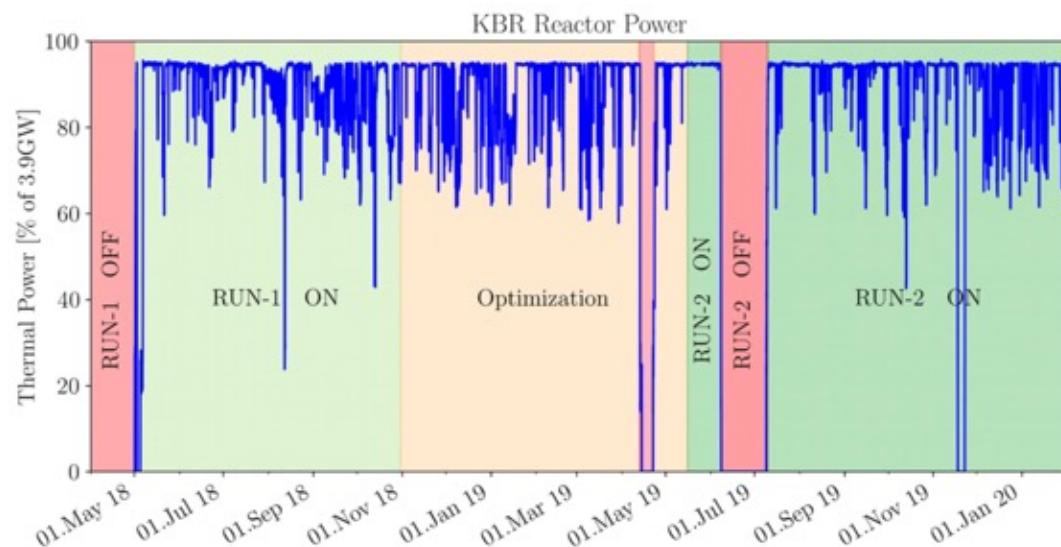
- Model-independent analysis of data ( $\sim 16$ h beam exposure)
  - All major systematic uncertainties included
- ➡ Data compatible with Lindhard model:  $k = 0.162 \pm 0.004$  (stat+syst)



# CEvNS Analysis: Data Selection & Noise Cuts

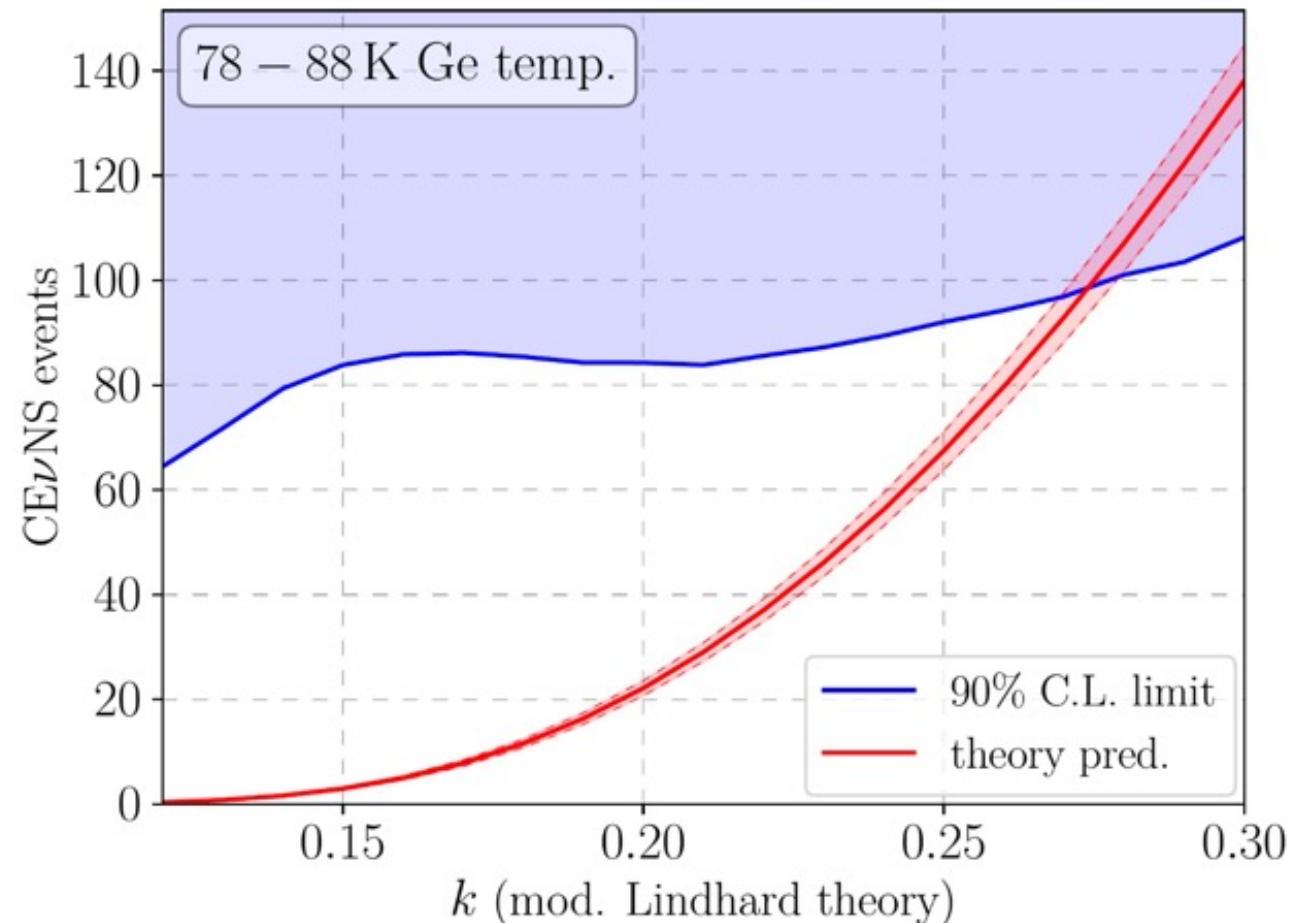
- Noise-temperature correlation cut
  - Reject microphonic and spurious events with time-difference distribution cut
- Run-1+2 exposure after all cuts:  
- 248.7 kg d (reactor-on)  
- 58.8 kg d (reactor-off)

Phys. Rev. Lett. 126, 041804 (2021)



# First Results

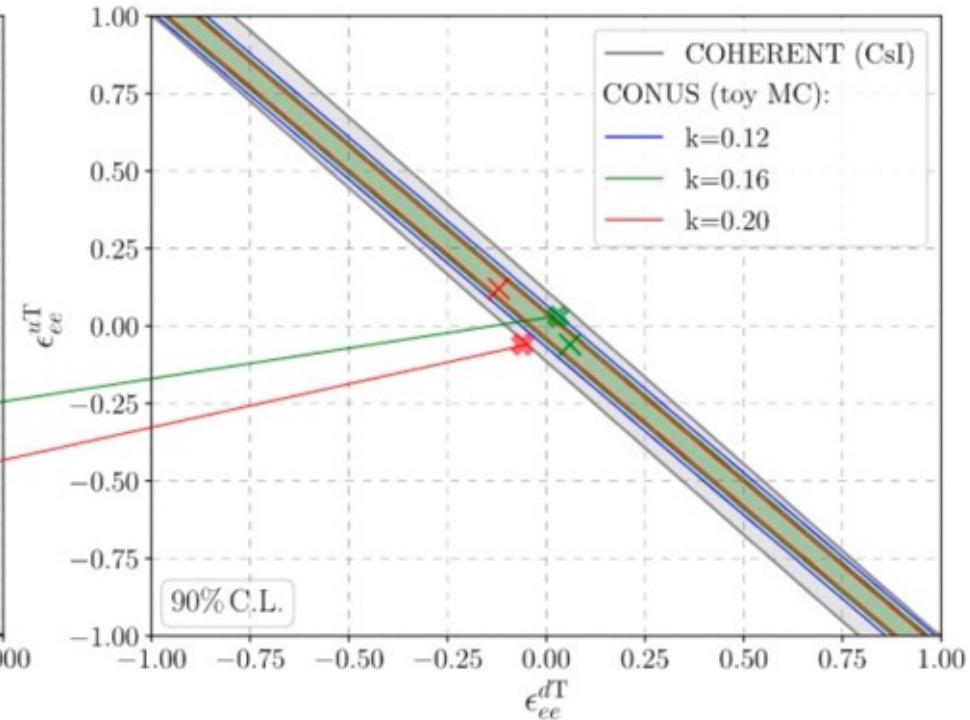
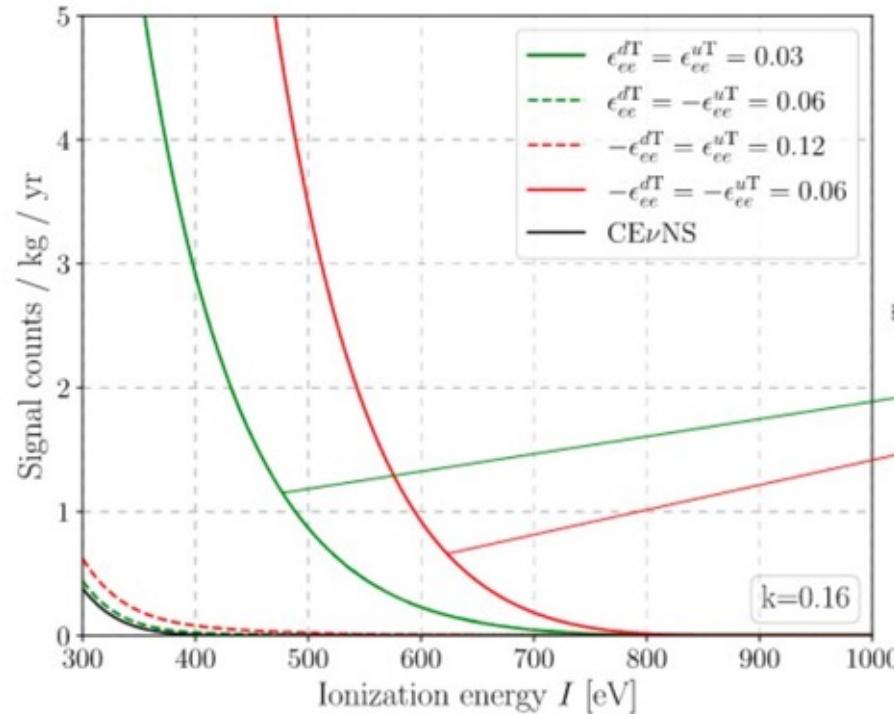
- Best CEvNS limit at reactor:  $< 0.4 \text{ d}^{-1} \text{ kg}^{-1}$  (90 % C.L.)
  - Signal expectation depends on quenching factor
- ➡ For  $k=0.16$ : expected CEvNS signal 17x below CONUS upper limit



# BSM: NSI Limits

- Non Standard Interactions (NSIs):
  - effective vector/tensor operators
  - new couplings  $Q_W \rightarrow Q_{NSI}(\{\epsilon_{\alpha\beta}^q\})$

- Tensor case:  $\left( \frac{d\sigma}{dT_N} \right) = \left( \frac{d\sigma}{dT_N} \right)_{CE\nu NS} + \frac{4G_F^2 M}{\pi} Q_{NSI}^T \left( 1 - \frac{MT}{4E_\nu^2} \right)$



ROI for BSM Analysis: [2 – 8] keV<sub>ee</sub>

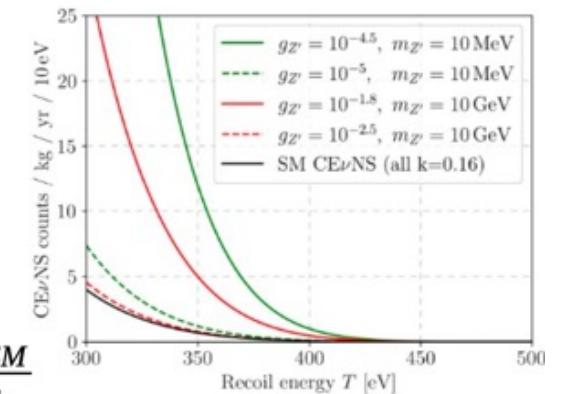
# BSM: Simplified Models

- Simplified models:
  - new light scalar/vector mediators
  - universal couplings

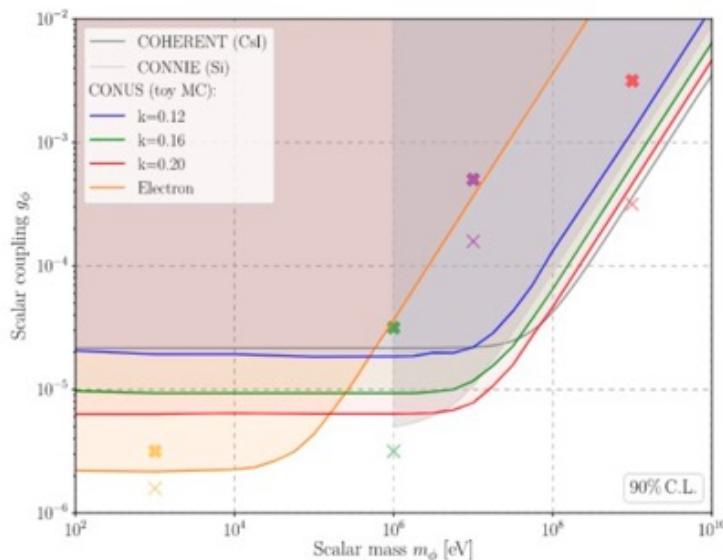
- light scalar boson  $\phi$  : 
$$\frac{d\sigma_\phi}{dT} = \frac{g_\phi^4 (14N + 15.1Z)^2 M^2 T}{4\pi E_\nu^2 (2MT + m_\phi^2)^2}$$

- light vector boson  $Z'$ : 
$$\frac{d\sigma_{Z'}}{dT} = \left(1 - \frac{3g_z^v g_z^q (Z+N)}{\sqrt{2} G_F Q_{SM} (2MT + m_{Z'}^2)}\right)^2 \frac{d\sigma_{SM}}{dT}$$

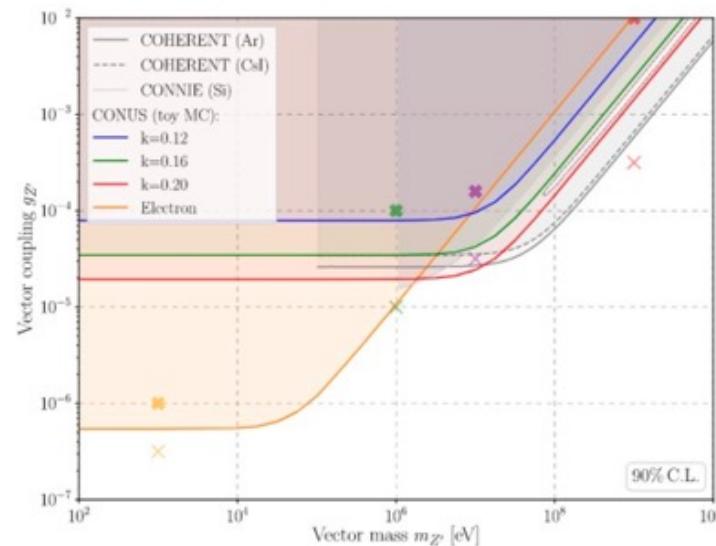
ROI for BSM Analysis: [2 – 8] keV<sub>ee</sub>



Light scalar boson  $\phi$



Light vector boson  $Z'$

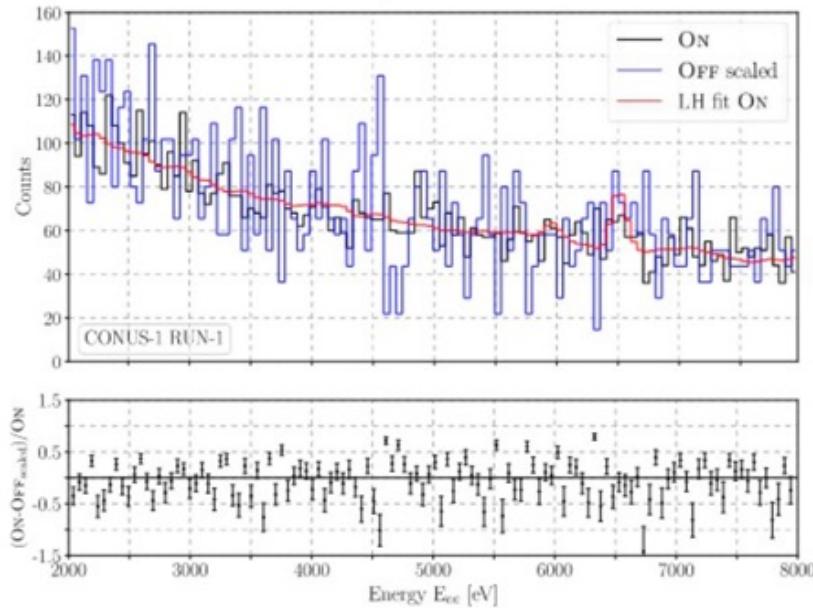


# Electro-magnetic Properties

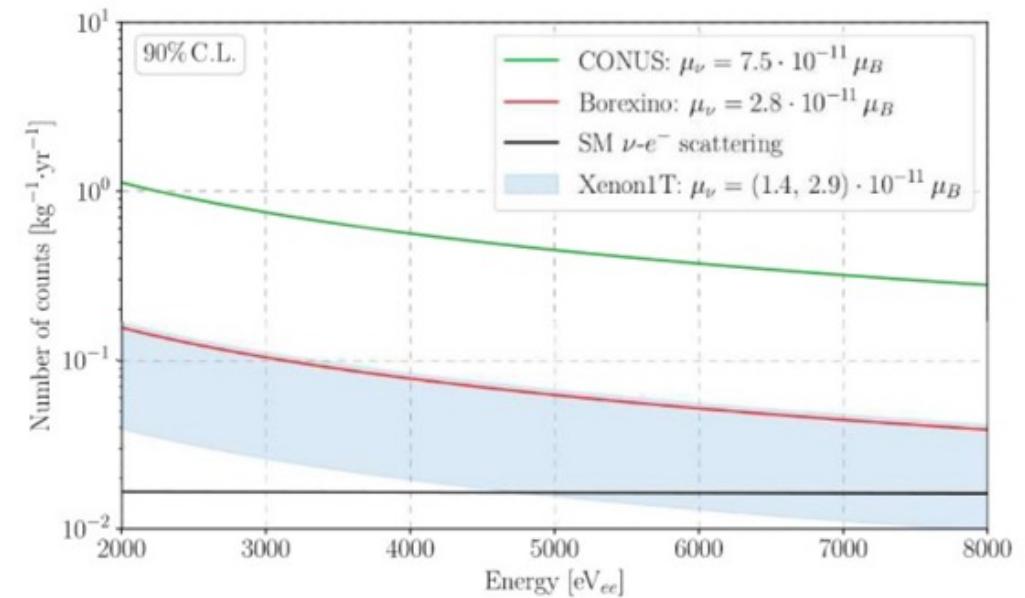
arXiv:2201.12257 (2022)

- Neutrino electron scattering channel:

$$\left(\frac{d\sigma}{dT}\right)_{em} = \pi \frac{\alpha^2}{m_e^2} \left( \frac{1}{T} - \frac{1}{E_\nu} \right) \left( \frac{\mu_\nu}{\mu_B} \right)^2$$



ROI for BSM Analysis: [2 – 8] keV<sub>ee</sub>



magnetic moment  $\mu_\nu < 7.5 \cdot 10^{-11} \mu_B$  (90% C.L.)

millicharges  $q_\nu < 3.3 \cdot 10^{-12} e_0$  (90% C.L.)

19/20

# Soon (plus next Years)

- Brokdorf shut down January 2022 → more off data
- Run-5 analysis (pulse shape analysis) on-going → stay tuned
- Identified a new reactor site: Leibstadt (CH)
  - similar distance & neutrino flux
  - closer position may be possible...
- move upgraded CONUS → CONUS+
  - \* improved detectors (even lower threshold)
  - \* optimized shielding & veto (W,...)
  - \* aim at increased detector mass
  - \* installation planned for next spring
- goals:
  - \* observe CEvNS signal
  - \* improve BSM sensitivity
  - \* demonstrate scalability to O(100)kg

# Longer Term Potential

Upscaling of a working technology to 100kg → very interesting potential  
high statistics → precision → potential for various interesting topics...

**assume:**

**100 kg detector**

**4GW @ 15m**

**flux  $\sim 3 \times 10^{13} / \text{cm}^2 / \text{s}$**

**background 1/kg/day**

**BSMsens=ΔS/S**

**Maneschg, Rink, Salathe, ML**

| Puler/Thresh [eV] | QF=0.15                  | BSMsens            | QF=BF                   | BSMsens            | QF=0.25                              | BSMsens            |
|-------------------|--------------------------|--------------------|-------------------------|--------------------|--------------------------------------|--------------------|
| 40 / 120          | 647 474/<br>8291 / 78.1  | $1 \times 10^{-3}$ | 965 999/<br>10 775/89.7 | $1 \times 10^{-3}$ | $2.9 \times 10^6 /$<br>15 158 / 189  | $6 \times 10^{-4}$ |
| 45 / 135          | 407 092/<br>8 036 / 50.7 | $2 \times 10^{-3}$ | 664 316/<br>10 519/63.2 | $1 \times 10^{-3}$ | $2.1 \times 10^6 /$<br>14 866 / 144  | $7 \times 10^{-4}$ |
| 50 / 150          | 254 745/<br>7780 / 32.7  | $2 \times 10^{-3}$ | 458 072/<br>1 0264/44.6 | $1 \times 10^{-3}$ | $1.6 \times 10^6 /$<br>14 574 / 84.9 | $8 \times 10^{-4}$ |
| 55 / 165          | 158 109/<br>7 524 / 21.0 | $3 \times 10^{-3}$ | 315 843/<br>9 971/31.7  | $2 \times 10^{-3}$ | $1.2 \times 10^6 /$<br>14 318 / 84.9 | $9 \times 10^{-4}$ |
| 60 / 180          | 97 066/<br>7 305 / 13.3  | $3 \times 10^{-3}$ | 217 277/<br>9 716/22.4  | $2 \times 10^{-3}$ | 919 435/<br>13 026 / 65.6            | $1 \times 10^{-3}$ |
| 65 / 195          | 58 827/<br>7 049 / 8.3   | $4 \times 10^{-3}$ | 148 848/<br>9 460/15.7  | $3 \times 10^{-3}$ | 696 196/<br>13 770 / 50.6            | $1 \times 10^{-3}$ |
| 70 / 210          | 35 154/<br>6 830 / 5.1   | $5 \times 10^{-3}$ | 101 386/<br>9 204/11.0  | $3 \times 10^{-3}$ | 527 204/<br>13 514 / 39.0            | $1 \times 10^{-3}$ |
| 75 / 225          | 20 711/<br>6 575 / 3.2   | $7 \times 10^{-3}$ | 68 573/<br>8 949/7.7    | $4 \times 10^{-3}$ | 398 867/<br>13 222 / 30.2            | $2 \times 10^{-3}$ |
| 80 / 240          | 12 042/<br>6 355 / 1.9   | $9 \times 10^{-3}$ | 46 008/<br>8 730/5.27   | $5 \times 10^{-3}$ | 301 231/<br>12 966 / 23.2            | $2 \times 10^{-3}$ |
| 85 / 255          | 6 924/<br>6 136 / 1.1    | $1 \times 10^{-2}$ | 30 598/<br>8 474/3.6    | $6 \times 10^{-3}$ | 226 910/<br>12 711 / 17.9            | $2 \times 10^{-3}$ |

BSMsens=ΔS/S

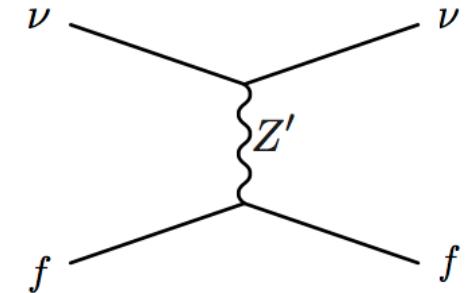
S[1/yr] / B[1/yr] / R=S/B

# Searches for new Physics: NSI's

NSI's  $\leftrightarrow$  new physics at high scales

Which are integrated out

$Z'$ , new scalars, ...  $\rightarrow \epsilon_{ij}$



$$\mathcal{L}_{NSI} \simeq \epsilon_{\alpha\beta} 2\sqrt{2} G_F (\bar{\nu}_L \gamma^\rho \nu_L) (\bar{f}_L \gamma_\rho f_L)$$

$$\frac{d\sigma}{dT}(E_\nu, T) = \frac{G_F^2 M}{\pi} \left(1 - \frac{MT}{2E_\nu^2}\right) \times \left\{ \left[ Z(g_V^p + 2\epsilon_{ee}^{uV} + \epsilon_{ee}^{dV}) + N(g_V^n + \epsilon_{ee}^{uV} + 2\epsilon_{ee}^{dV}) \right]^2 + \sum_{\alpha=\mu,\tau} \left[ Z(2\epsilon_{\alpha e}^{uV} + \epsilon_{\alpha e}^{dV}) + N(\epsilon_{\alpha e}^{uV} + 2\epsilon_{\alpha e}^{dV}) \right]^2 \right\}$$

Barranco et al. 2005

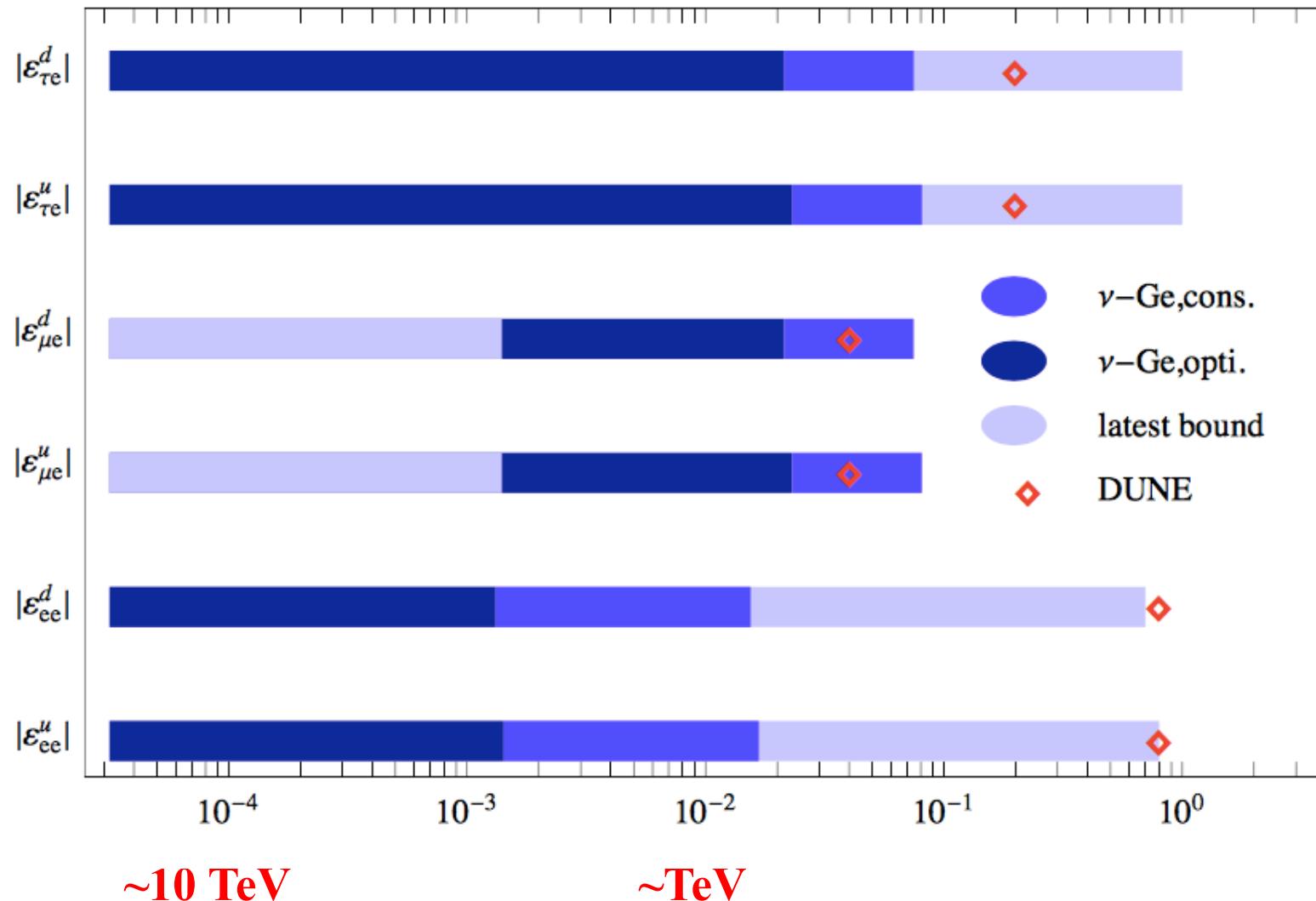
$$|\epsilon| \simeq \frac{M_W^2}{M_{NSI}^2}$$

$\rightarrow$  Competitive method to test TeV scales  
 $\epsilon = 0.01 \leftrightarrow$  TeV scales

# NSI-Potential

100kg detector, 5 years operation @ 4GW

ML, W. Rodejohann, X.Xu



# Precise Measurement of $\sin^2\theta_W$ at low E

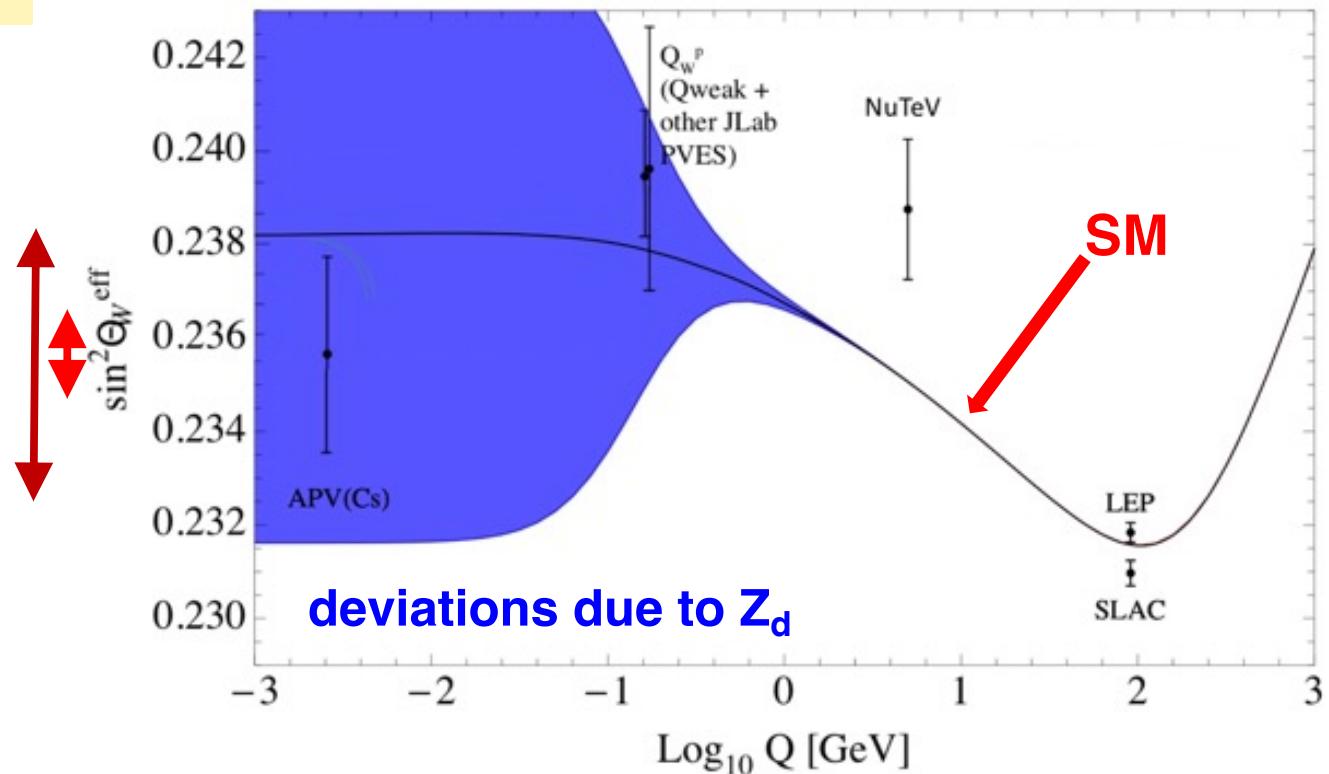
$\sin^2\theta_W$  precisely known in SM  
SM quantum corrections  
→ running  $\sin^2\theta_W^{\text{eff}}$

CEvNS cross-section:  
 $\sigma \sim N - [(1 - 4 * \sin^2\theta_W) Z]^2$

$\underbrace{\quad}_{\simeq 0}$   
→ enhanced sensitivity

BSMsens:  
 $10^{-3} \rightarrow \Delta\sin^2\theta_W = 0.006$   
 $10^{-4} \rightarrow \Delta\sin^2\theta_W = 0.0006$

**potential problem: (g-2) anomaly**  
→ Light dark sector?  
 $Z_d$ ,  $M=150$  MeV; ...other parameters  
See e.g. 1411:4088  
many models lead to similar effects...

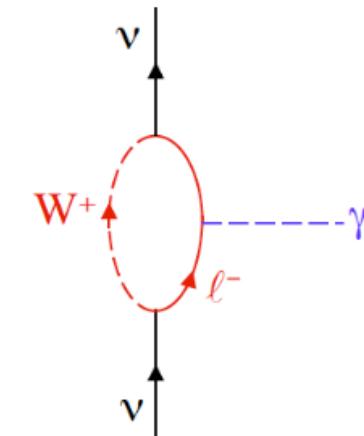


# Searches for new Physics: Magnetic Moments

Magnetic moment for minimal  $\nu$  masses are very tiny:

$$\text{Dirac: } \mu_{kk}^D \simeq 3.2 * 10^{-19} \left( \frac{m_k}{\text{eV}} \right) \mu_B$$

$$\text{Majorana: } \mu_{ll'}^M \lesssim 4 * 10^{-9} \mu_B \left( \frac{M_{ll'}^M}{\text{eV}} \right) \left( \frac{\text{TeV}}{\Lambda} \right)^2 \left| \frac{m_\tau^2}{m_l^2 - m_{l'}^2} \right|$$

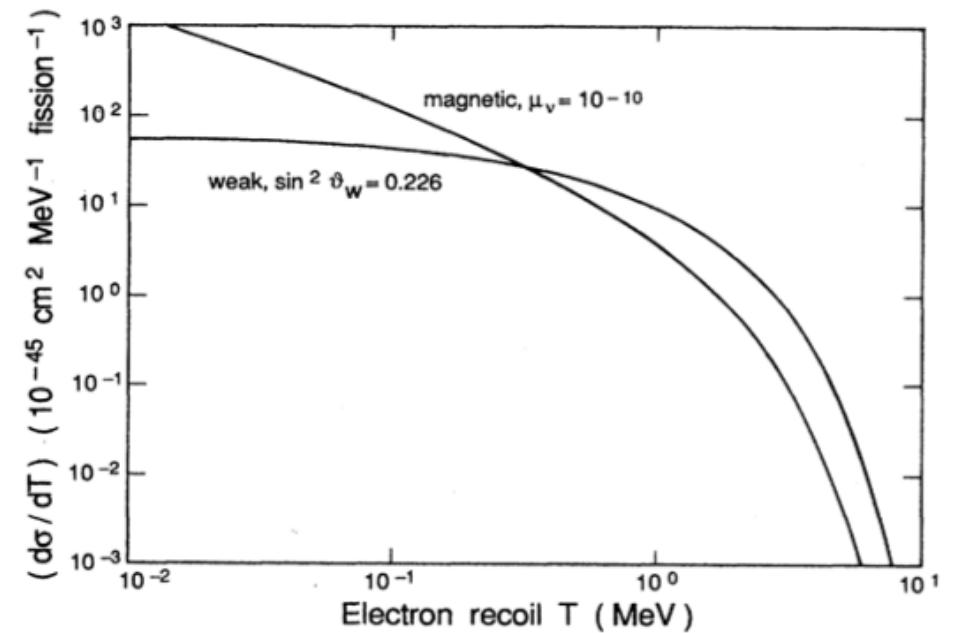


New physics  $\rightarrow$  detectable enhancements due to new physics:  
SUSY, extra dimensions, ...

At least new best limits:  
e-scattering (GEMMA) and astrophysics:

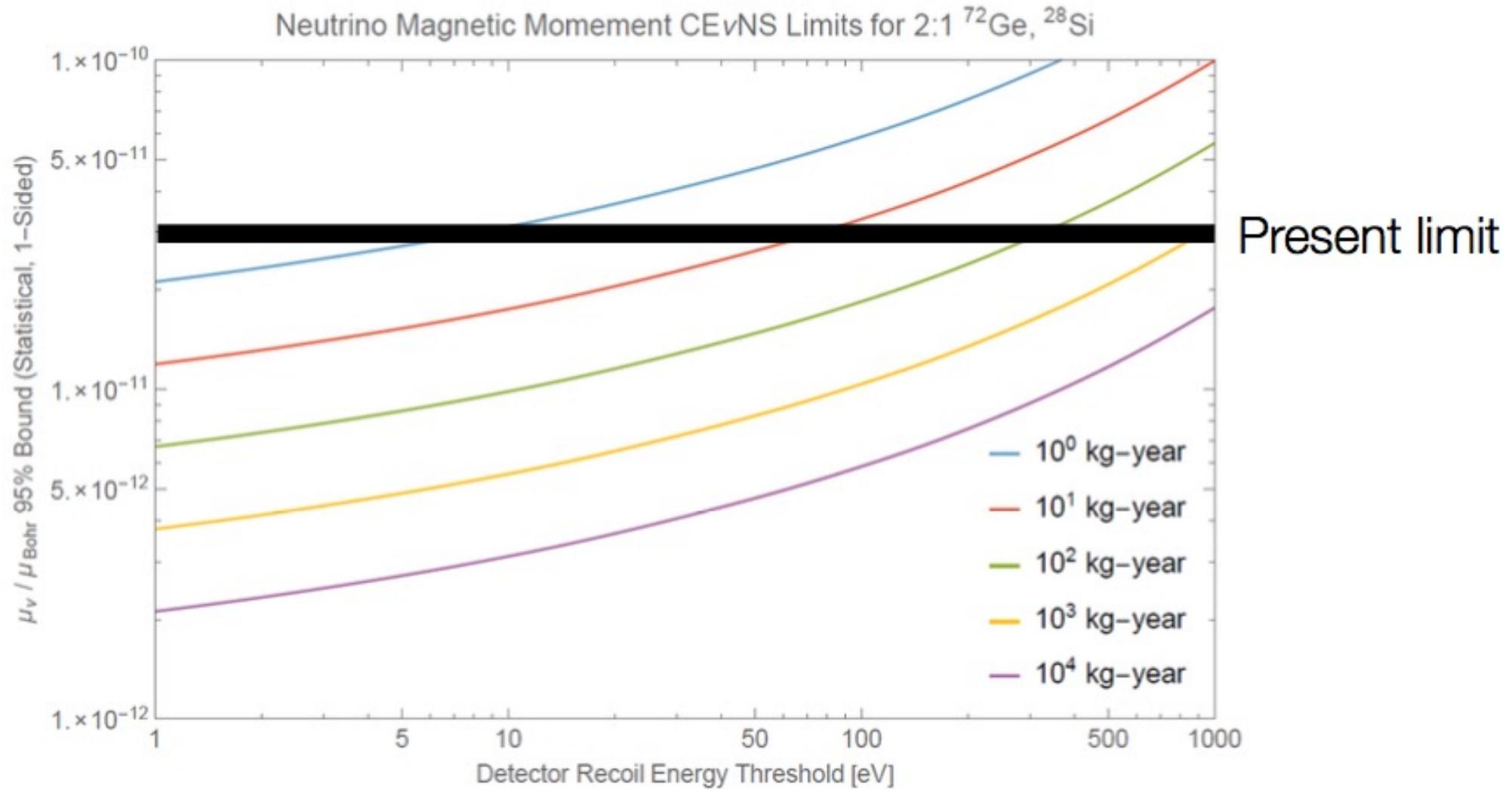
$$\mu_\nu < 3 \times 10^{-11} \mu_b$$

Scattering on protons coherently enhanced:  $\rightarrow$  detectable at low energy (Vogel & Engel 1989)



$$d\sigma \mid \mid \pi \alpha^2 \mu_\nu^2 \left[ 1 - T_R/E_\nu \right] \left[ T_R \right]$$

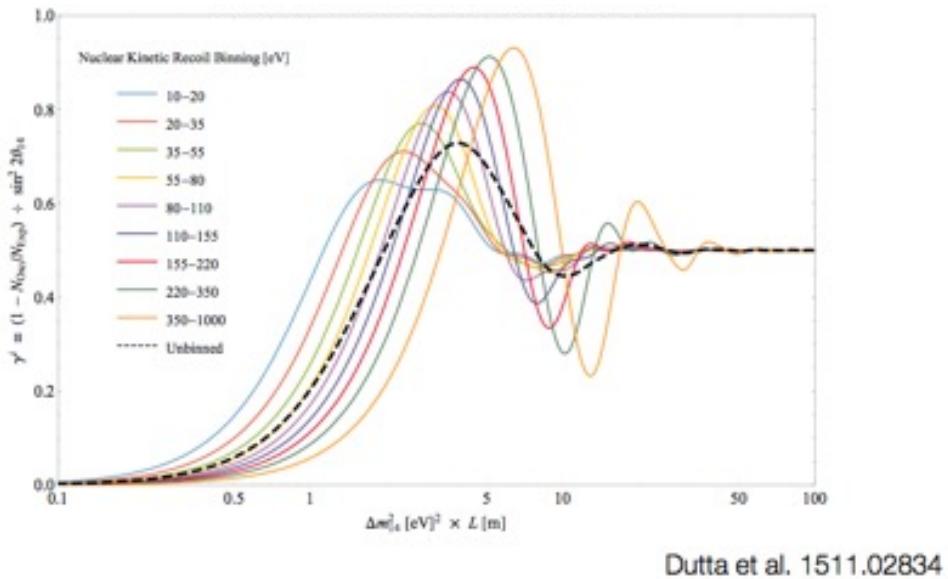
# Potential for Magnetic Moments



**100kg \* 5y = 500 kg-year ; low threshold → one order of magnitude better**

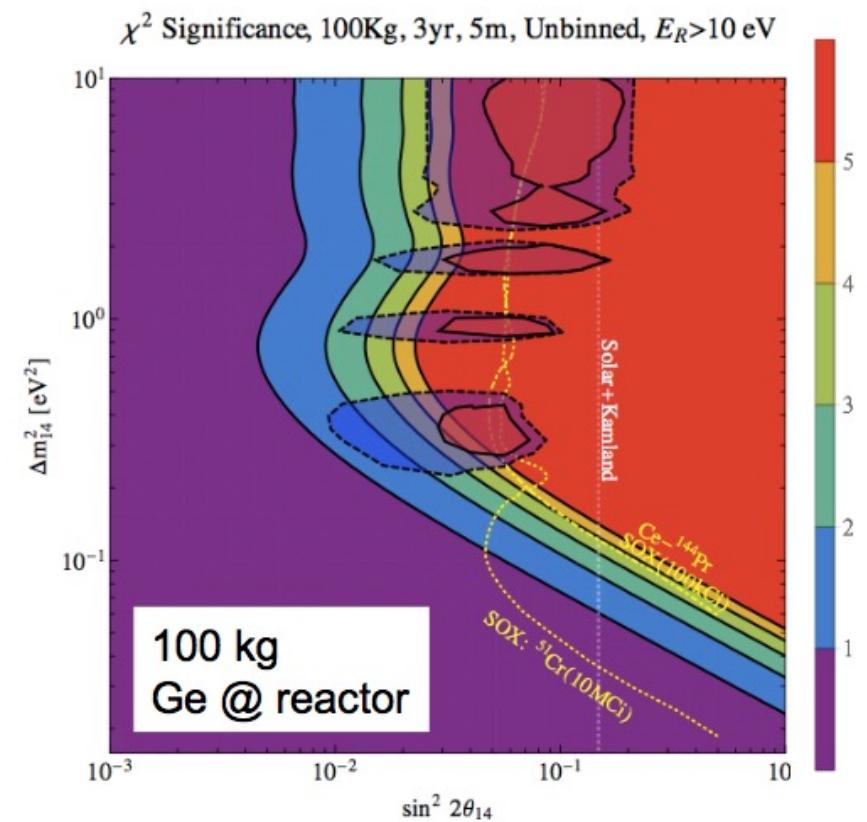
# Searches for new Physics: Sterile $\nu$ 's

- Various indications / hints for sterile neutrinos ?
- eV sterile  $\nu$ 's with small mixing
- keV warm dark matter with tiny mixing
- ...different mass ranges
- any sterile state would motivate more...



$$P(\nu_\alpha \rightarrow \nu_\phi) = 4|U_{\alpha 4}|^2(1 - |U_{\alpha 4}|^2) \sin^2(1.27\Delta m_{41}^2 L/E)$$

- test if / how flux deviates from  $1/R^2$   
 → time scales compared to other projects



B. Dutta et al, arXiv:1511.02834

# Nuclear Structure with coherent Scattering

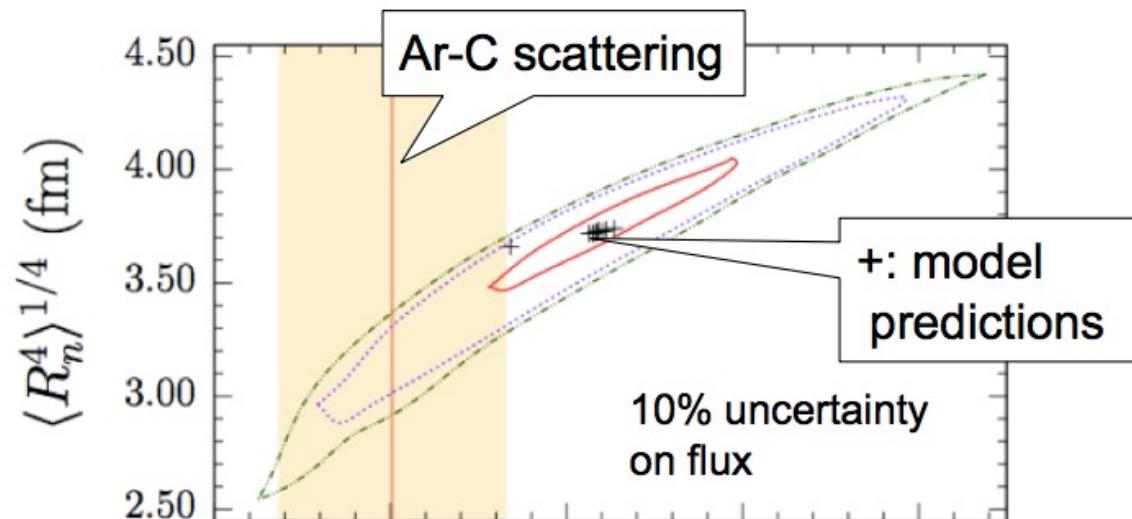
Remember: DAR sources close to de-coherence  $\leftrightarrow$  combine with reactor measurements

$$\frac{d\sigma}{dT} \approx \frac{G_F^2 M}{4\pi} \left(1 - \frac{MT}{2E^2}\right) \left[N F_N(q^2) - Q_W Z F_Z(q^2)\right]^2$$

Nuclear form factors  $F_{N,Z}(q)$   $\sim$  Fourier transforms of N & P densities  
 $\rightarrow$  resolve nuclei (neutrons) in neutrino light

Fit recoil **spectral shape** to determine the  $F(Q^2)$  moments  
(requires very good energy resolution, good systematics control)

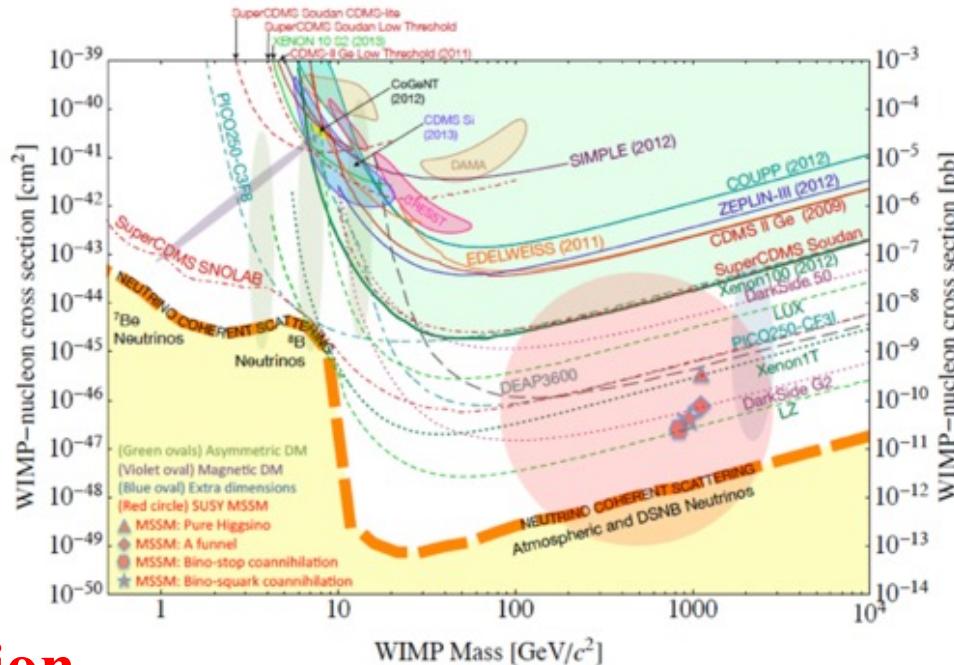
Example:  
tonne-scale  
experiment  
at  $\pi$ DAR source



# CEvNS Connections to more Topics...

## DM connection:

- 1) DM experiments assume coherent DM scattering → test with ν's
- 2) Neutrino floor of direct DM experiments will measure CEvNS  
→ combine different measurements



## CEvNS cross-section

- 3) Important for astrophysical applications: supernovae, ...
- 4) ...

# More Phenomenology / Theory / Applications

- coherent v's → conceptually very interesting questions  
see e.g. [Akhmedov, Arcadi, ML, Vogl, JHEP 1810 \(2018\) 045, arXiv:1806.10962](#)
  - can coherent scattering occur at macroscopic scales?
  - role of the recoil of constituents in quantized picture
  - semi-classical factorization of QFT process into (cross-section) \*  $F(q^2)$  ?
  - ...
- coherence length in QFT approach  
[Egorov, Volobuev: 1902.03602](#)
- many connections to dark matter models
- producing new fermion in CEvNS [Brdar, Rodejohann, Xu: 1810.03626](#)
- effects of CP violating parameters on CEvNS processes  
see e.g. [Sierra, De Romeri, Rojas: arXiv:1906.01156](#)
- Safeguarding...
- ...

# Summary

- **CEvNS was 1<sup>st</sup> observed by COHERENT at  $E_\nu \simeq 30\text{-}50 \text{ MeV}$**
  - **CONUS**
    - close to signal  $\leftrightarrow$  lower QF than expected  $\rightarrow$  run-5 analysis on-going
    - strong BSM limits  $\leftrightarrow$  fully coherent  $\rightarrow$  deviations are BSM
    - improvement: detailed pulse shape analysis – analysis of run 5 ongoing
    - on-going upgrade to CONUS+ at new reactor site
      - \* lower threshold
      - \* increased detector mass
  - **Longer term future: CEvNS will become an interesting tool**
    - upscaling of existing technology to O(100kg)  $\rightarrow$  physics topics:
      - coherent  $\nu$  scattering  $\leftrightarrow$  DM & WIMP scattering, neutrino floor
      - BSM:  $\nu$  magnetic moments, NSIs, steriles,  $\sin^2\theta_W$ , sterile oscillation
      - nuclear form factors with neutrinos  $F(q^2)$
      - reactor  $\nu$  spectrum & anomalies
- $\rightarrow$  very interesting potential of demonstrated technology