

Light dark matter searches with RNDR-DEPFET detectors

HEPHY/FHWN

Classical approach

- Motivated by SUSY, Lee-Weinberg bound
- Particle mass: above 1 GeV/c²
- Main event signature: nuclear recoils
- Experiments: large scale, no discoveries

Light dark matter approach

- Motivated by dark sector, ELDER, SIMP
- Particle mass: above 100 keV/c²
- Main event signature: electron recoils
- Moderate sensitive mass needed







DEPFET detectors

- Technology established in astrophysics (Bepi-Columbo, ATHENA) and particle physics (Belle)
- DEPFET principle:
 - Side wards fully depleted bulk
 - e⁻ are collected in internal gate
 - conductivity of transistor channel is proportional to number of e⁻
 - e⁻ are removed by the clear contact



RNDR-DEPFET detectors

- Instead of removing the e⁻, they are shifted to a second internal gate
- Repetitive independent (n) readout cycles of same signal
- Reduction of noise by $1/\sqrt{n}$



DANAE prototype Setup

- 64x64 RNDR-DEPFET pixels a 50x50 μm²
- Sensor glued on carrier ceramic and wire bonded to flexible PCB
- Sensor mounted in inner and outer Al shielding block with separated front end electronic
- Inner block connected to Stirling cooler, placed in vacuum chamber









Detector Results

Detector performance and calibration

- ⁵⁵Fe spectrum recorded with
 - Calibration on 6 keV lines
- Effects to be considered:
 - Small pixels and thick bulk material → large charge cloud (> four pixels)
 - Specific (slow) r/o in combination with large charge cloud → interframe splits (broken patterns)
- Mitigation by:
 - Long exposure time (in relation to readout time)
 - Small number of repetitions





III. DANAE: Performance – ⁵⁵Fe

1000

2000

3000

4000

Further effects:

- Adapted number of evaluated repetitions at a similar measurement has minor effects
- Calibration of single pixel considers max. four pixel events → small noise results in large patterns → threshold above 10 sigma was applied
- Achievable resolution limited (appr. 170 eV)



8000

7000

Signal (eV)

5000

6000

Calibrated spectra of LED

- Calibration with LED and pixel wise gaussian fit on single electron peaks
- Spectra for different number of averaged repetitions (up to 600)
- Performance affected from generated or lost electrons during readout
- Inspection of single events (slope of signal vs. number of repetition) enables to identify and filter affected events







What to learn from slopes?

- Sufficient large number of repetitions -> distribution with pos. and neg. shoulders evolves
- Deviation from gauss shape -> identify events
- First shoulder one electrons second two, etc.







Conclusions and outlook

Space applications

- Photon number resolved detection in the near IR regime with low dark counts required for various space applications (esp. direct detection of exoplanets)
- Available and proposed technology (<u>https://exoplanets.nasa.gov/internal_resources/1390/</u>):
 - EMCCD (high r/o speed but limited spectral performance multiplicative noise)
 - Skipper CCD (high spectral performance but limited r/o speed)
 - → Radiation damage is a major concern for both technologies due to the ccd principle
- The active pixel detector technology of DEPFETs is expected to be more radiation hard (no lateral charge transfer)

➔ First irradiation tests with conventional DEPFETs at proton accelerator at Wiener Neustadt (MedAustron) with MPE

- A process could be established for irradiation tests (low flux settings, vacuum chamber with beam dump and entrance window)
- Preliminary results → see next talk









Status

- Still non understood high number of generated electrons
- Non reproducible effects on current module observed (affected spectral resolution)

Outlook

- For better understanding and statistics another module shall be assembled this time at HEPHY
- Experience and infrastructure from HEP applications is available (e.g. bonding)
- Set up process for further module assembly





Thank you for the attention