

Towards the operation of a dielectric haloscope in cryogenic temperatures

Juan P.A. Maldonado

Physicist – Universidad Nacional de Colombia

MSc student – University of Bonn



What I will talk about

1. Motivation
 - Why the (QCD) axion
 - How to detect it
 - Why MADMAX
 - Why my project
2. Towards a cold calibration and operation of the experiment
3. Summary and next steps: Measurements at CERN 2024

Why (QCD) axion?

Strong CP problem

There is an allowed CP violating term in the QCD lagrangian

$$L_\theta = \theta \frac{g^2}{32\pi^2} F_{\mu\nu}^a \widetilde{F}_a^{\mu\nu}$$

which induces an electric dipole moment of the neutron

$$d_n \approx 10^{-16} \theta \text{ e cm}$$

This, however, is not observed in nature

$$\theta < 10^{-10}$$

Dark Matter

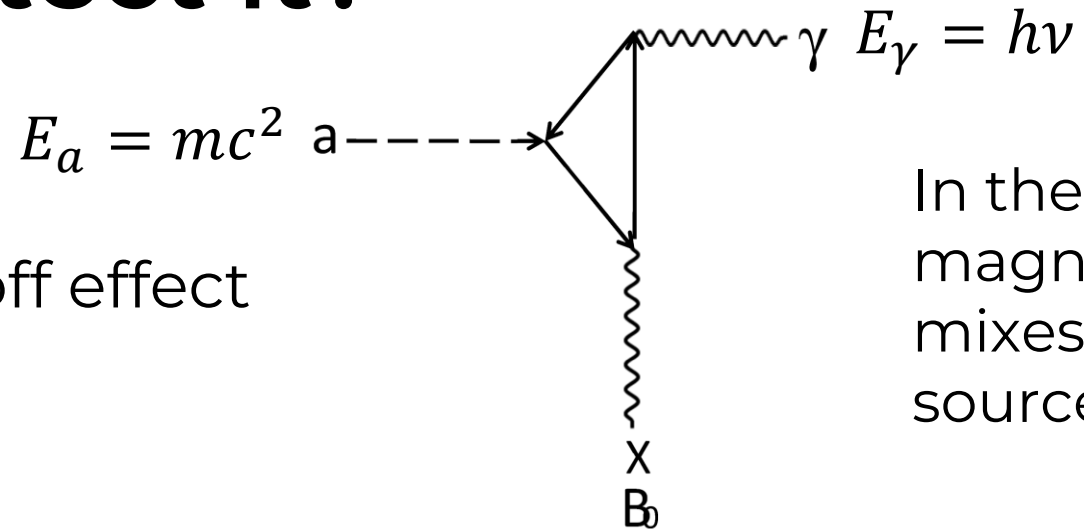
~5 times more dark matter than baryonic matter. Cold, Feebly interacting, not in the Standard Model

Main unknown: DM particle mass → **tuning needed**

How to detect it?

Inverse Primakoff effect

$$E_a = mc^2 \quad a \text{---}$$



In the presence of a magnetic field, an axion mixes with a photon and sources an electric field.

However, there are many other EM sources. For example, thermal radiation!

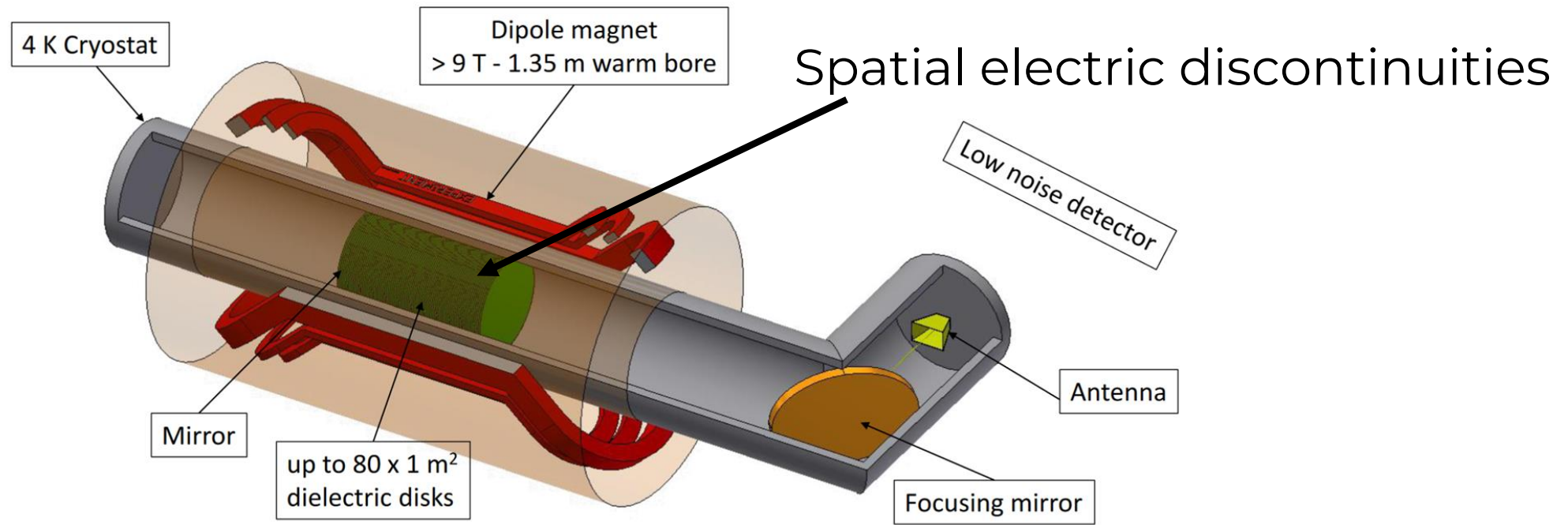


Axion-induced photons ~ 6 per day

Minimum detectable flux ~ 1 photon per second \rightarrow
Signal **amplification needed**

At $\sim 300\text{K}$ there are ~ 100 times more thermally radiated photons per second than at 4K \rightarrow **cooling needed**

Why MADMAX?



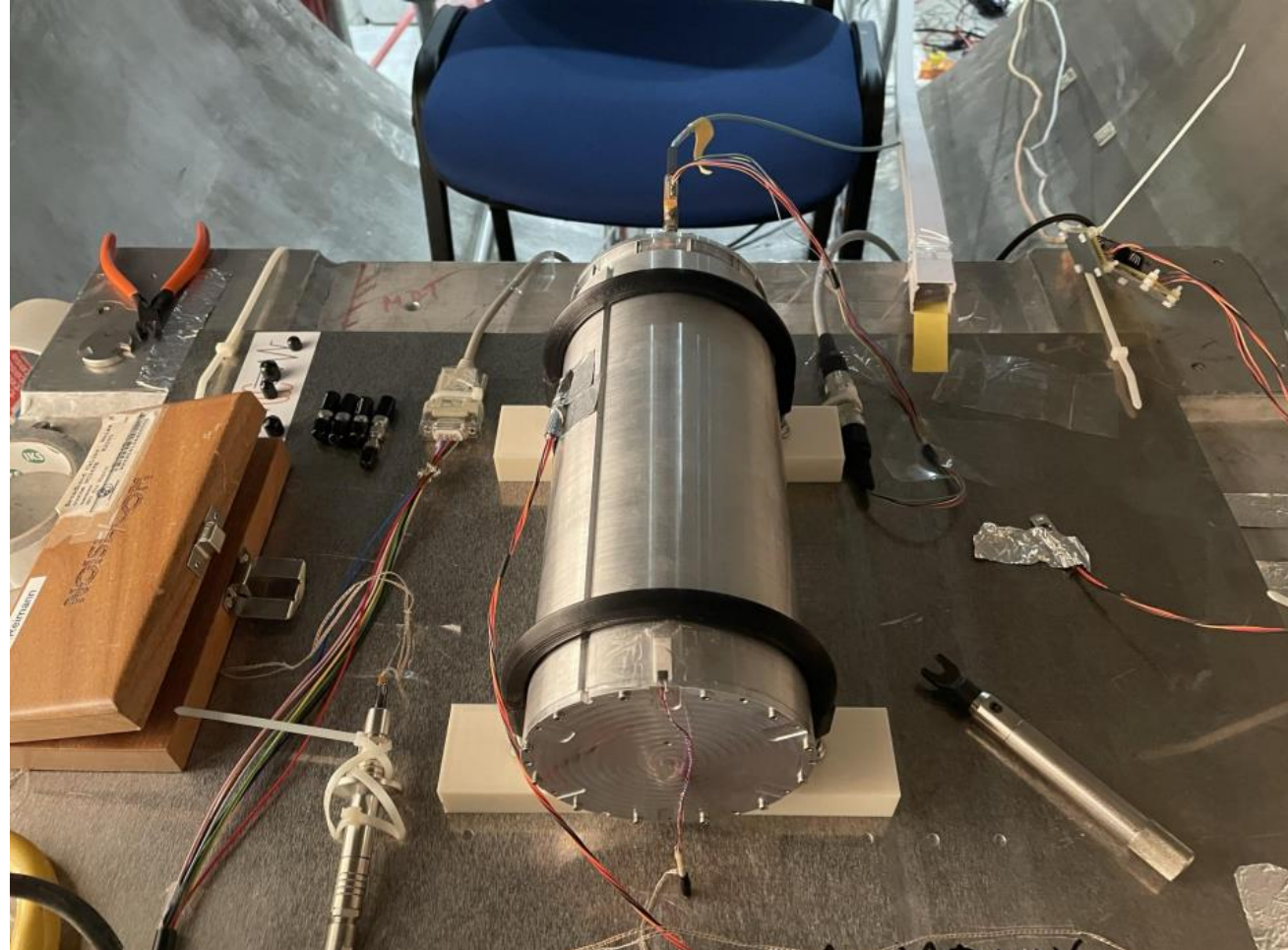
MADMAX addresses the **amplification**, **cooling** and **tuning** problem in a very well motivated range.

Why my project?

The system at 4K is different:
thermal contraction \sim photon
wavelength

The current MADMAX prototype
(CB-100) needs **~ 10 independent
measurements** to be calibrated

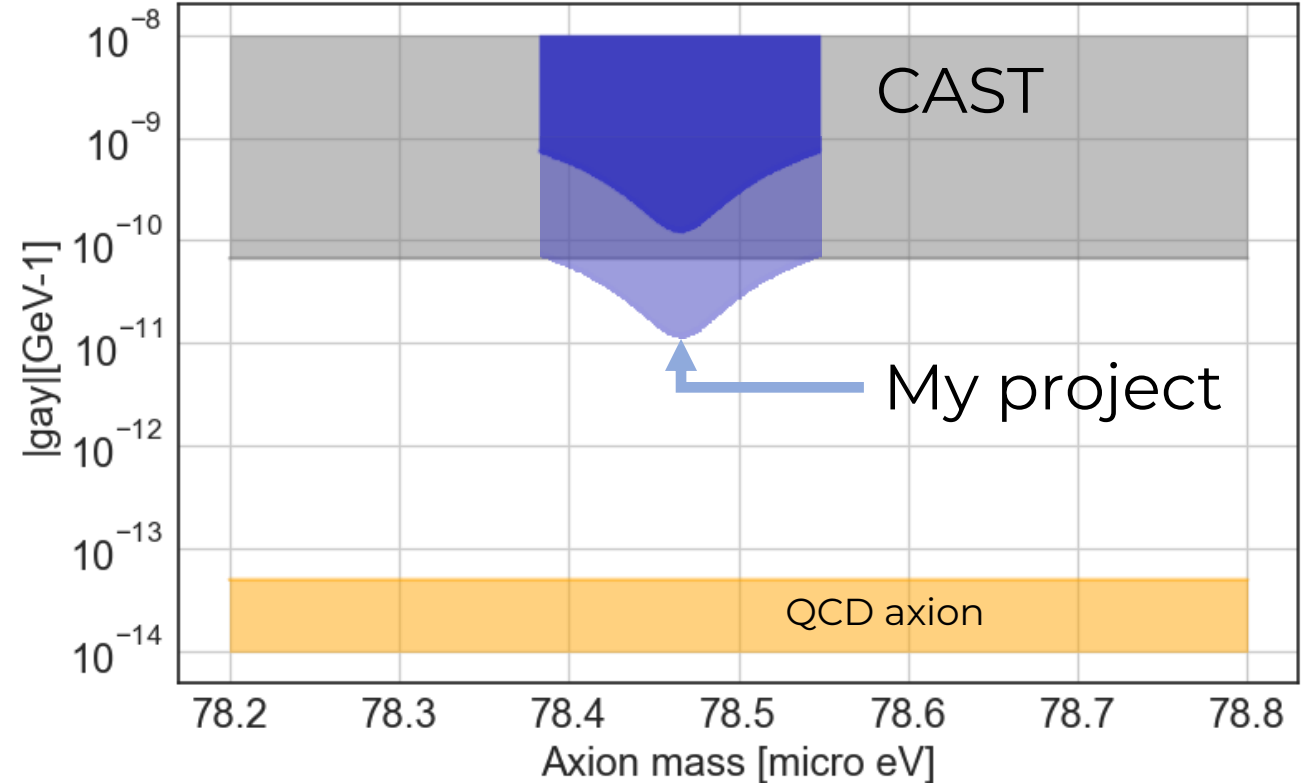
At **300K**, this is done manually in
 ~ 3 hours, but at **4K** this would
require a **~ 1 -month** long
calibration with reduced stability.



Why my project?

Goal: Perform a **semi-automatic calibration** of the experiment **minimizing the number of thermal cycles** and **maximizing reproducibility** and **stability**

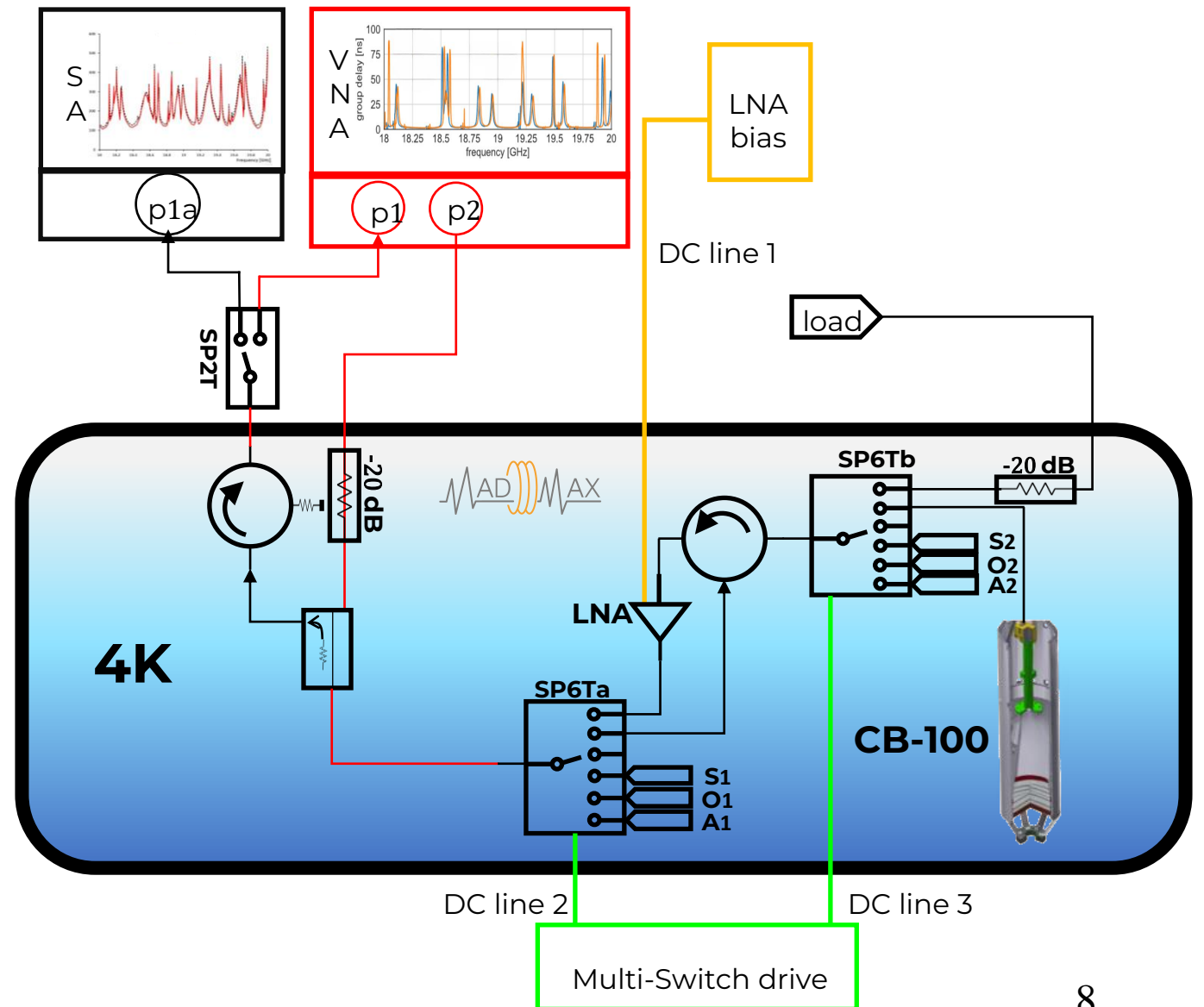
Why? ~1 order of magnitude increase in sensitivity, allowing **DM searches** in so far **unexplored regions**



Towards a cold calibration (+ operation)

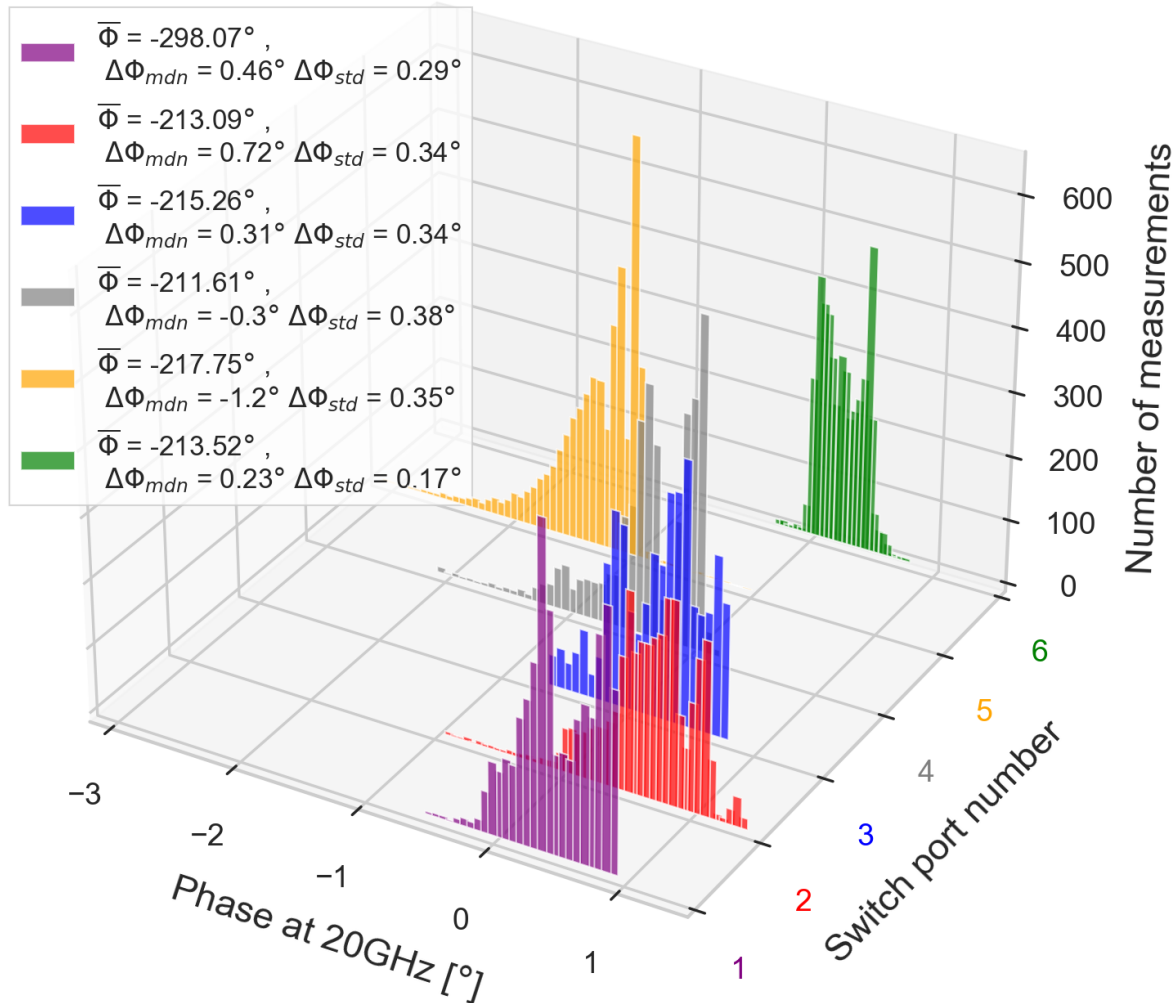
My idea: a **single thermal cycle setup** to perform noise calibration, reflectivity calibration, and data taking

Main challenge: Ensure **stability** and **reproducibility** in the setup



Switch at room temperature

Radial R591722600 S21 48H Phase stability at 297K



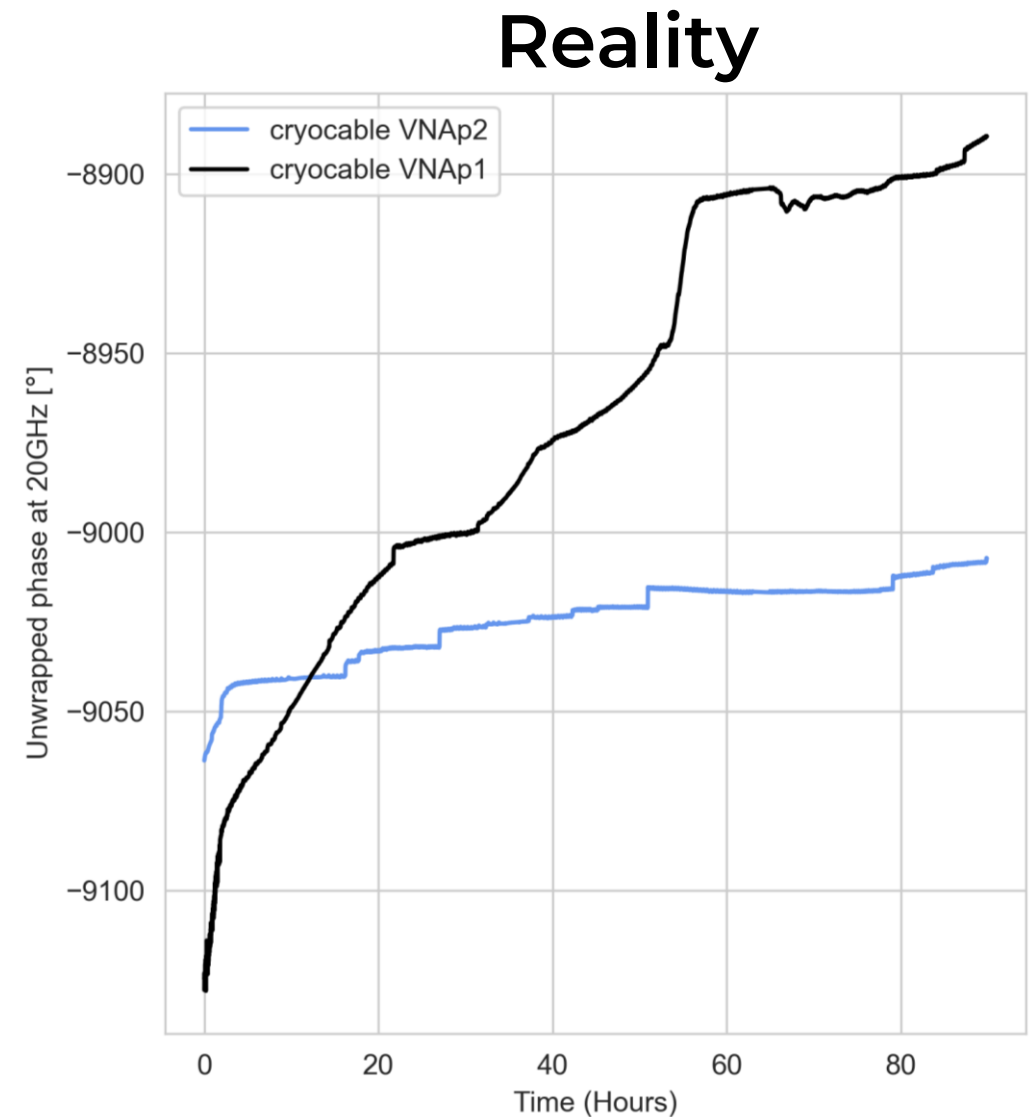
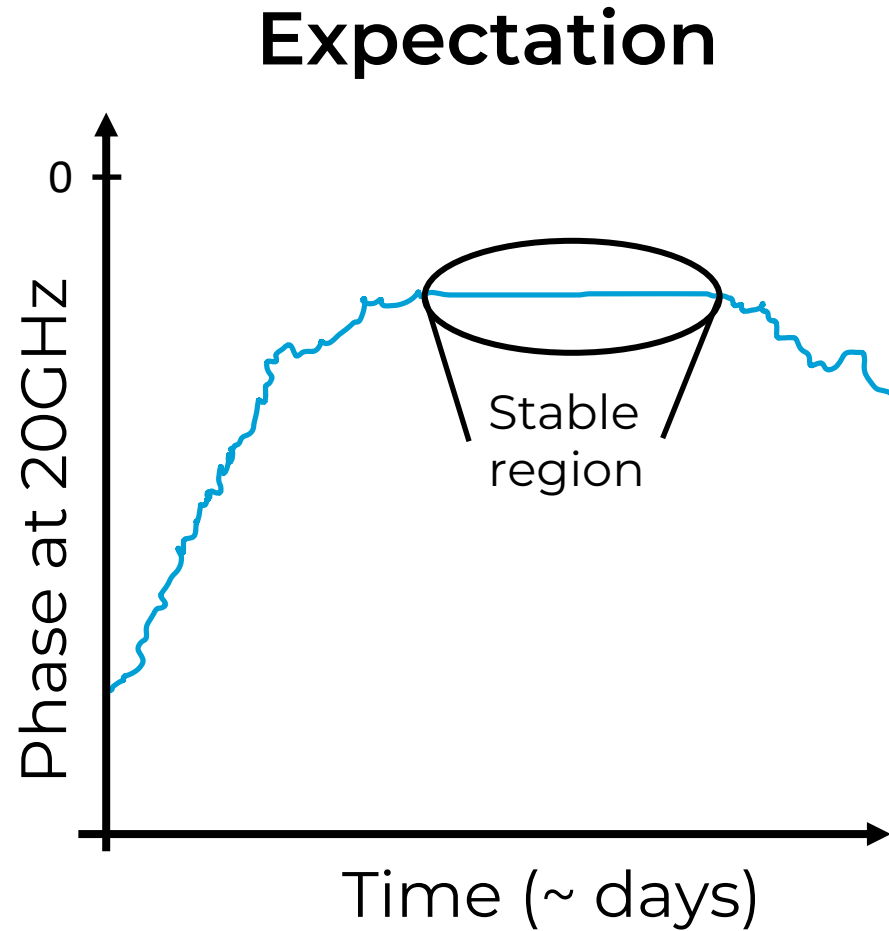
Phase stability study

Important parameter:
standard deviation

De-embedding length using
the statistical median

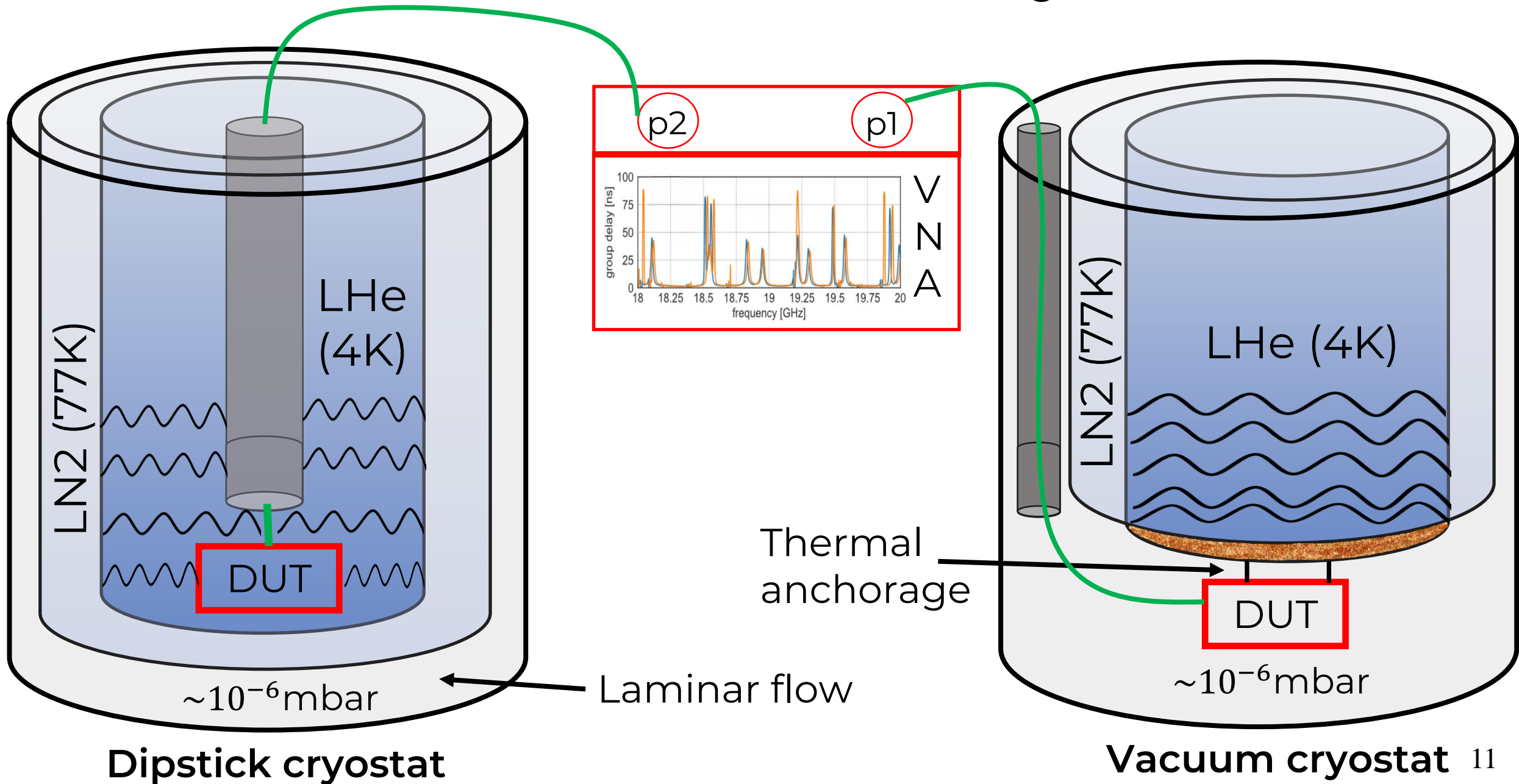
Our requirement: std < 1
degree

Dipstick cryostat



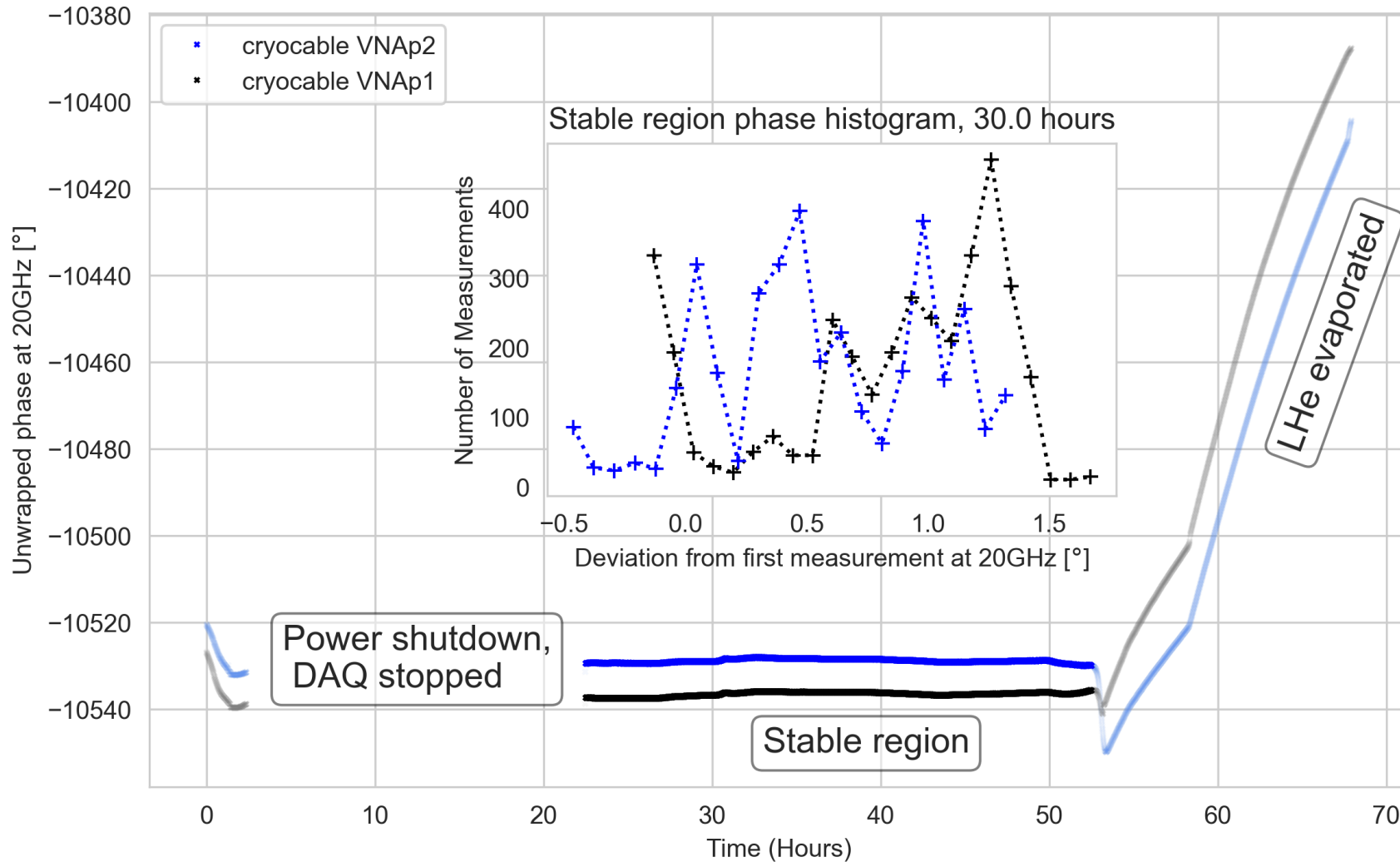
No stable region was found.
Hypothesis: The LN2 fill level affects the stability of the measurement

Difference between cryostats



Vacuum cryostat

Phase evolution in vacuum cryostat using LHe



+30 hours of stability!

Deviations $< 2^\circ$ at 20GHz $\approx 55\mu\text{m}$

To sum up

- Joined MPP 9 months ago with a **central role** in MADMAX: **approach** the **cryogenic calibration** of CB-100 **minimizing the thermal cycles** required
- Devised a **single-cycle calibration and operation** scheme
- Goal: **lead** the **first** MADMAX **cryogenic** dark matter **search** at CERN in 2024
 - What can be accomplished? **~1 order of magnitude increase** in the **sensitivity** of the experiment and **first** MADMAX dark matter **limit** in a so far **unconstrained region**.